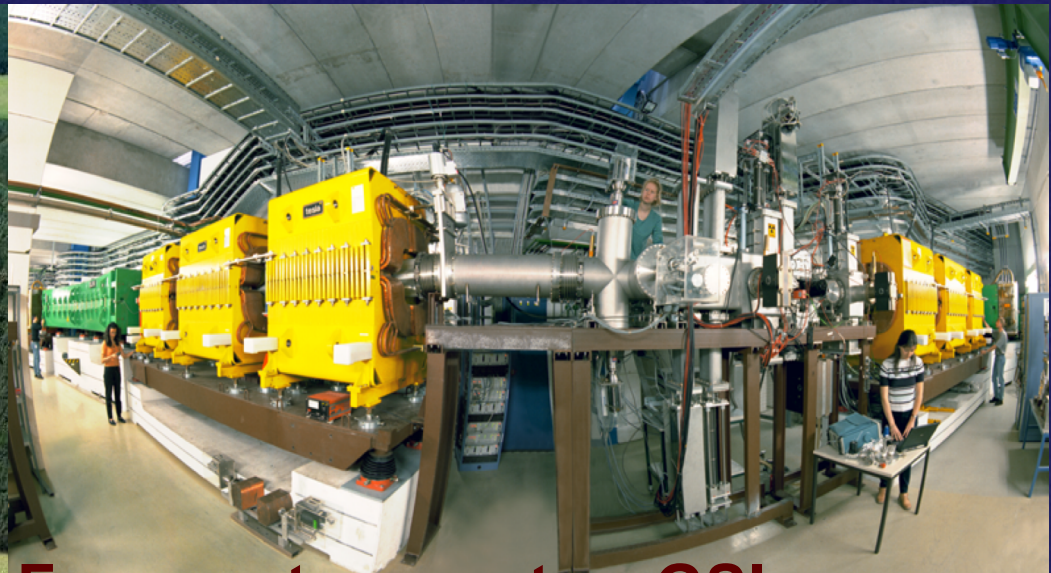


Determination of Proton Radii of Neutron-rich Oxygen Isotopes

Satbir Kaur



**GSI,
Germany**



Fragment separator, GSI

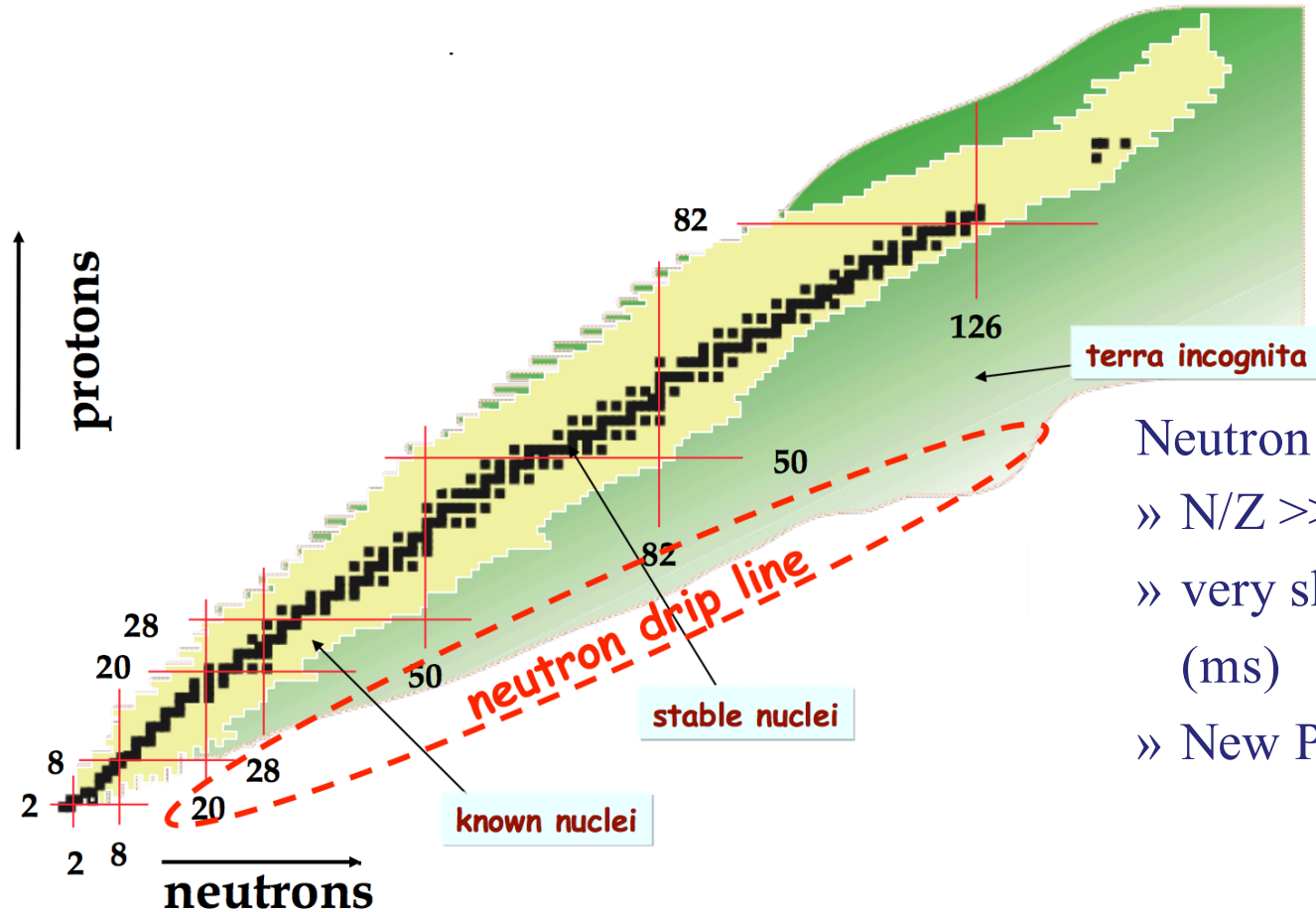


Outline

- Introduction
 - Nuclear landscape and neutron-rich nuclei.
- Scientific Motivation
 - Motivation to study oxygen isotopes.
 - Importance of Proton Radii (R_p)
- Methods to measure R_p and Charge changing cross sections (σ_{cc})
- Experimental Setup
- Results
- Summary



Neutron-rich nuclei



- Neutron Rich Nuclei
- » $N/Z \gg 1$
 - » very short half life (ms)
 - » New Physics.



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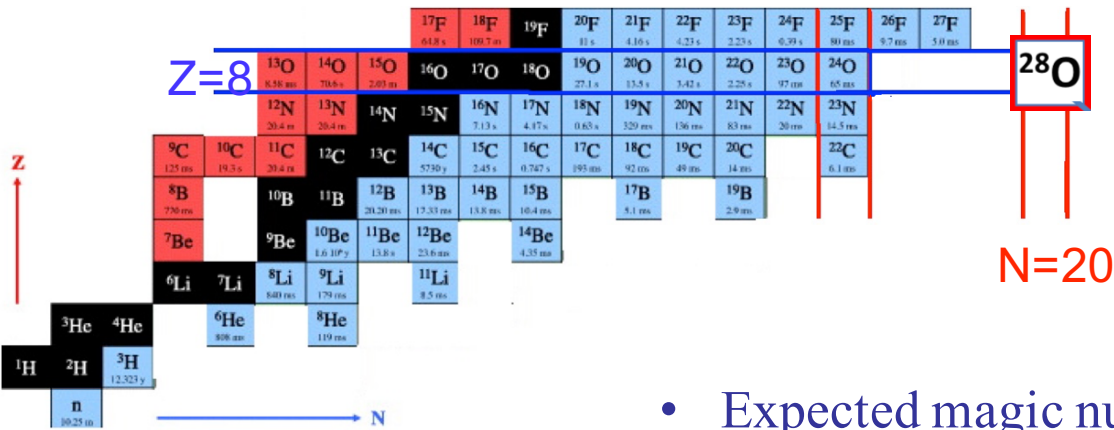
➤ Description of Experimental Setup

➤ Results

➤ Summary



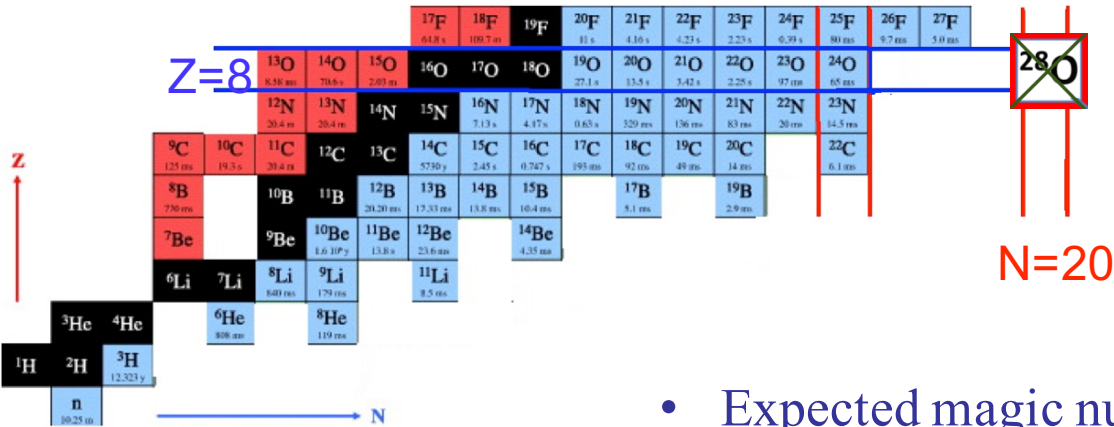
Interesting neutron-rich oxygen isotopes.



- Expected magic number nucleus



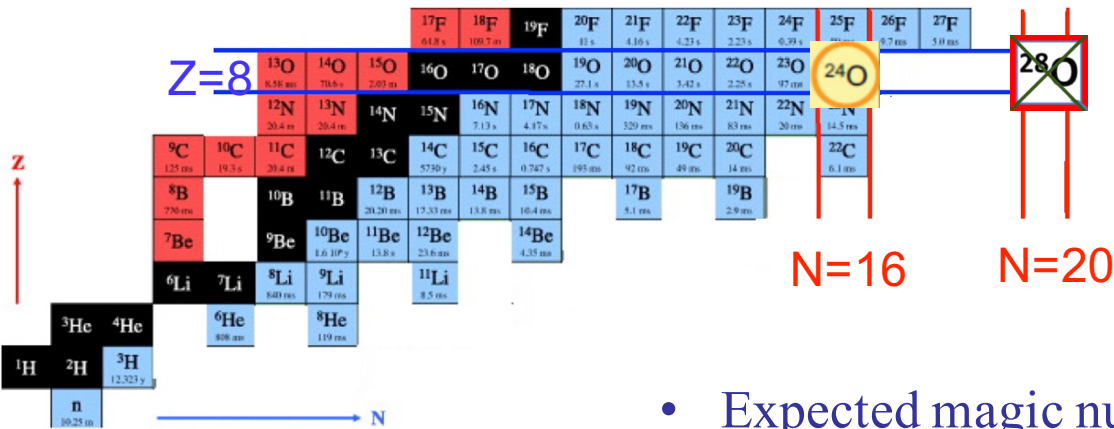
Interesting neutron-rich oxygen isotopes.



- Expected magic number nucleus ^{28}O unbound (Tarasov et al., 1997)



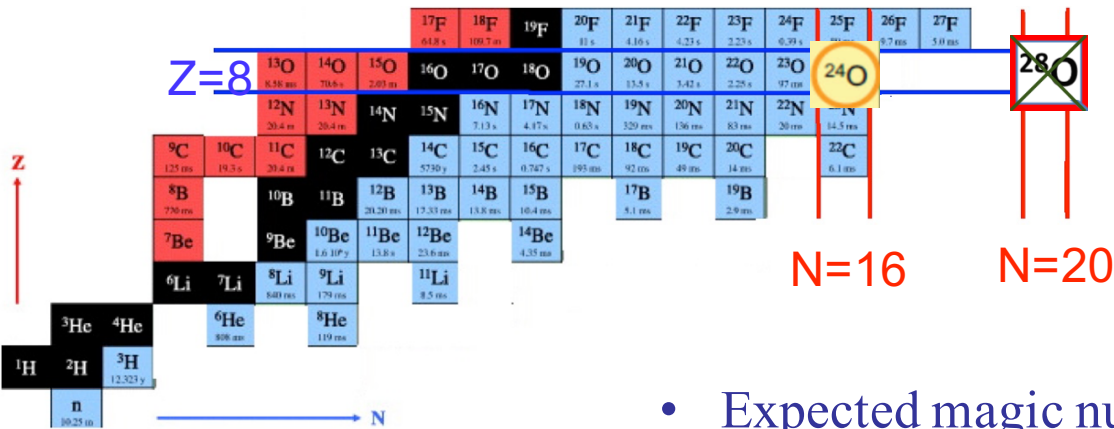
Interesting neutron-rich oxygen isotopes.



- Expected magic number nucleus ^{28}O unbound (Tarasov et al., 1997)
- Neutron drip line of O at ^{24}O ($N=16$). (Hoffman et al 2008, Lunderberg et al 2012).



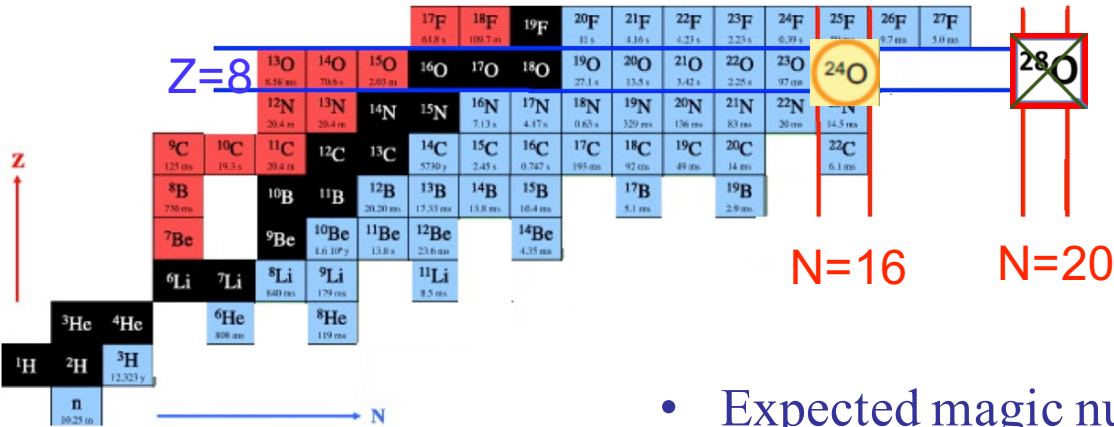
Interesting neutron-rich oxygen isotopes.



- Expected magic number nucleus ²⁸O unbound (Tarasov et al., 1997)
- Neutron drip line of O at ²⁴O. (Hoffman et al 2008, Lunderberg et al 2012).
- ²⁴O doubly magic with new magic number (N =16). (A. Ozawa et al., (2000), R. Kanungo et al., (2002).) R. Kanungo,(2009))



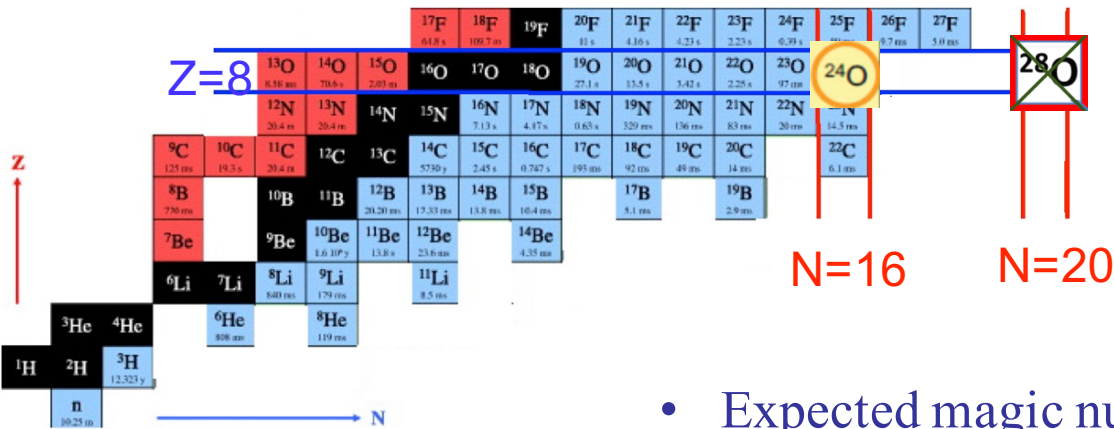
Interesting neutron-rich oxygen isotopes.



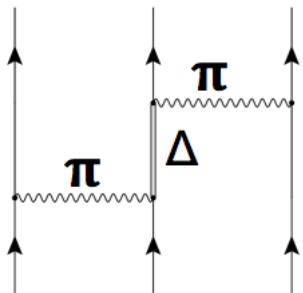
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- To yield the drip line at ^{24}O , 3NF required .



Interesting neutron-rich oxygen isotopes.



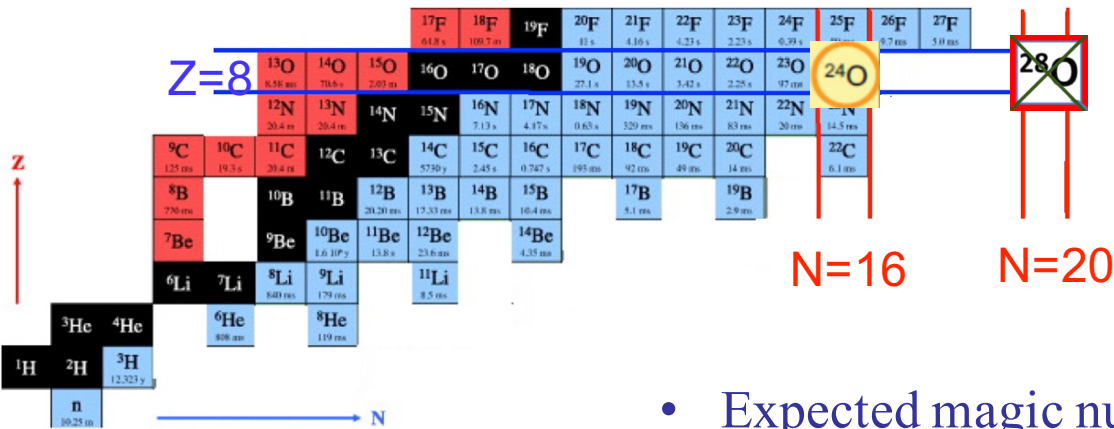
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- To yield the drip line at ^{24}O , 3NF required .



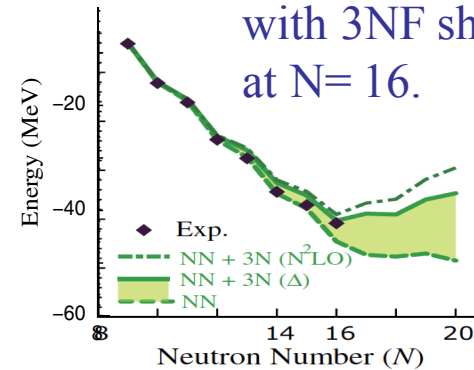
An example of 3N force



Interesting neutron-rich oxygen isotopes.

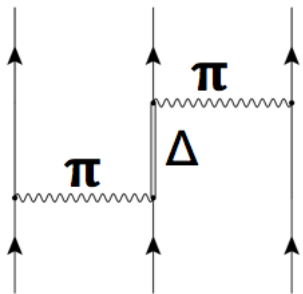


Ground state energies with 3NF show a kink at N=16.



Otsuka et al., (2010)

- Expected magic number nucleus ^{28}O unbound (Tarasov et al., 1997)
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An example of 3N force



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➤ Methods to measure R_p and Charge changing cross sections (σ_{cc})

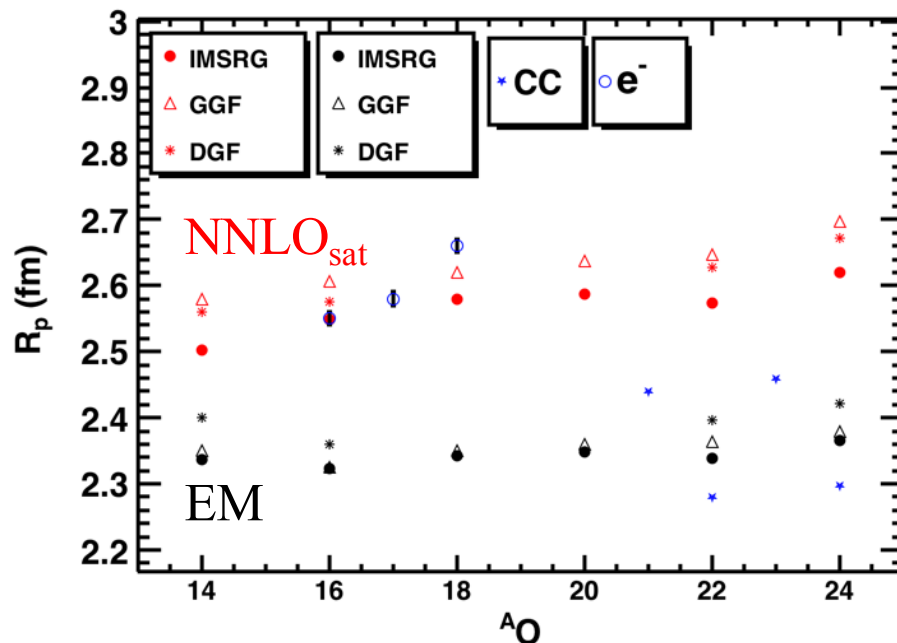
➤ Description of Experimental Setup

➤ Results

➤ Summary



Test of ab-initio theories



Calculated R_p with EM and NNLO_{sat} interaction

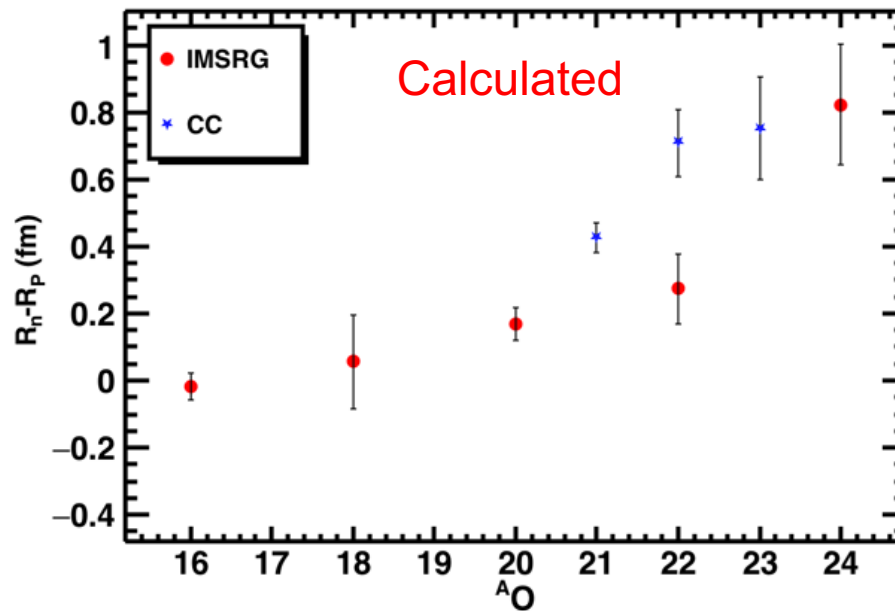
IMSRG, SCGF calculations (Lapoux et al., (2016)), Couple cluster (CC) (Hagen et. al, (2012))
Rp e⁻ scattering experiment (Atomic Data and Nuclear Data Tables, 2013)

Proton radii of neutron rich oxygen isotopes not measured till date.



Determination of neutron skin

Neutron skin Thickness $\rightarrow \delta R = \langle r_n^2 \rangle^{1/2} - \langle r_p^2 \rangle^{1/2}$



Neutron skin calculated from measured R_m and calculated R_p .
 R_n determined using matter radii (R_m) from A. Ozawa et al., (2001)

R_p data required to determine neutron skin thickness.



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Techniques to measure Charge radii

Electron Scattering :

$$F(q) = \frac{4\pi}{qZ} \int_0^{\infty} \rho_{\text{ch}}(r) \sin(qr) r \, dr$$

Limited to long lived nuclei only

$F(q)$ carries information about charge distributions.

Isotope shift: change in energy of atomic levels of different isotopes.

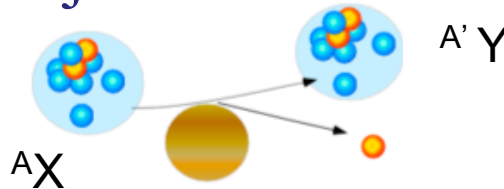
Limitations

- high intensity beams with low energy difficult to produce for very short lived nuclei.
- Not applicable to all neutron rich nuclei.

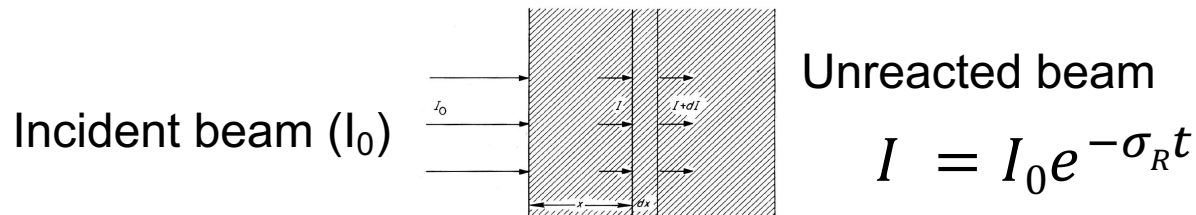


Proton radii from Charge Changing Cross Sections (σ_{cc})

It is the cross-section for reactions leading to any change of the atomic number of the projectile nucleus.



Principle of Measurement : Transmission type measurement



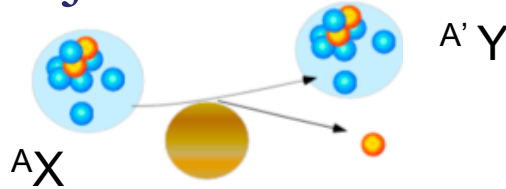
$$N_{sameZ} = N_0 e^{-\sigma_{cc} t}$$

$$\sigma_{cc} = \frac{1}{t} \ln \frac{N_0}{N_{sameZ}}$$

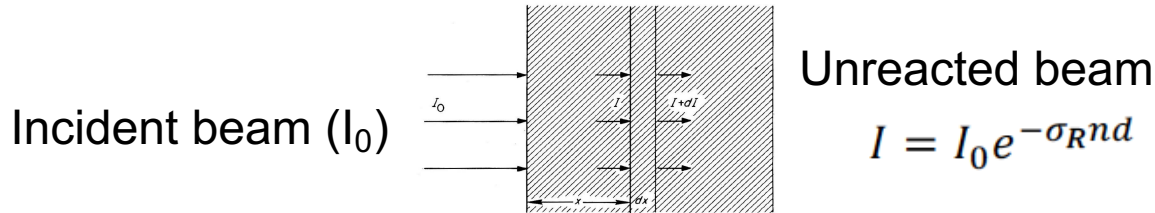


Proton radii from Charge Changing Cross Sections (σ_{cc})

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$$\sigma_{cc} = \frac{1}{t} \ln \frac{\left(\frac{N_{sameZ}}{N_{in}} \right)_{R_{out}}}{\left(\frac{N_{sameZ}}{N_{in}} \right)_{R_{in}}}$$

Target out

Target in



Outline

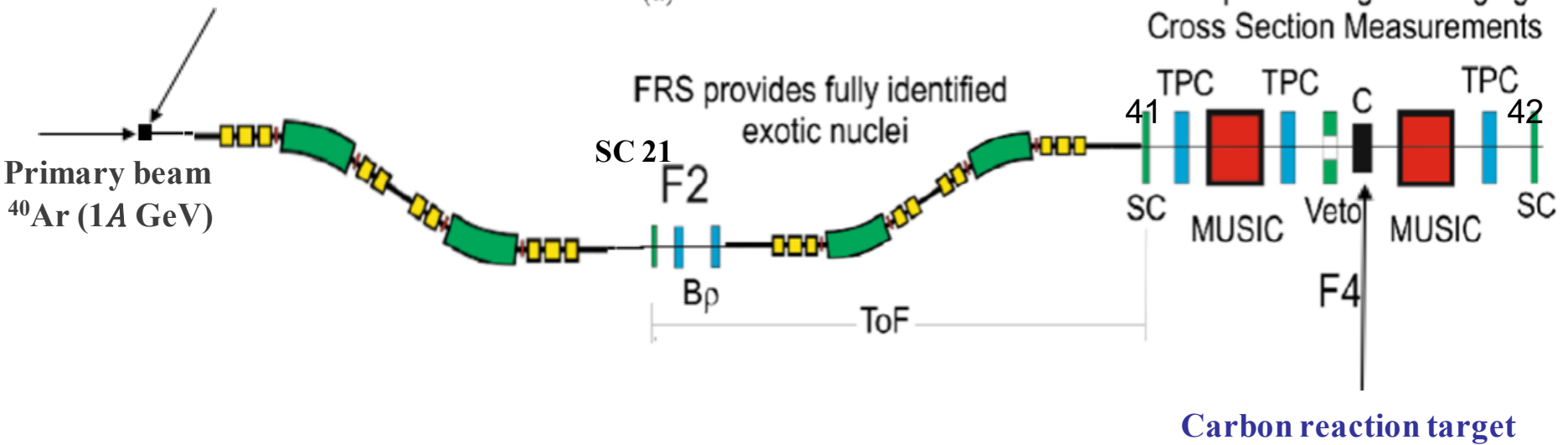
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Detector Setup

$^{16-24}\text{O}$ produced from fragmentation of $1\text{A GeV } ^{40}\text{Ar}$ beam at Fragment Separator, GSI, Germany.

6.3 g/cm^2 thick Be production target

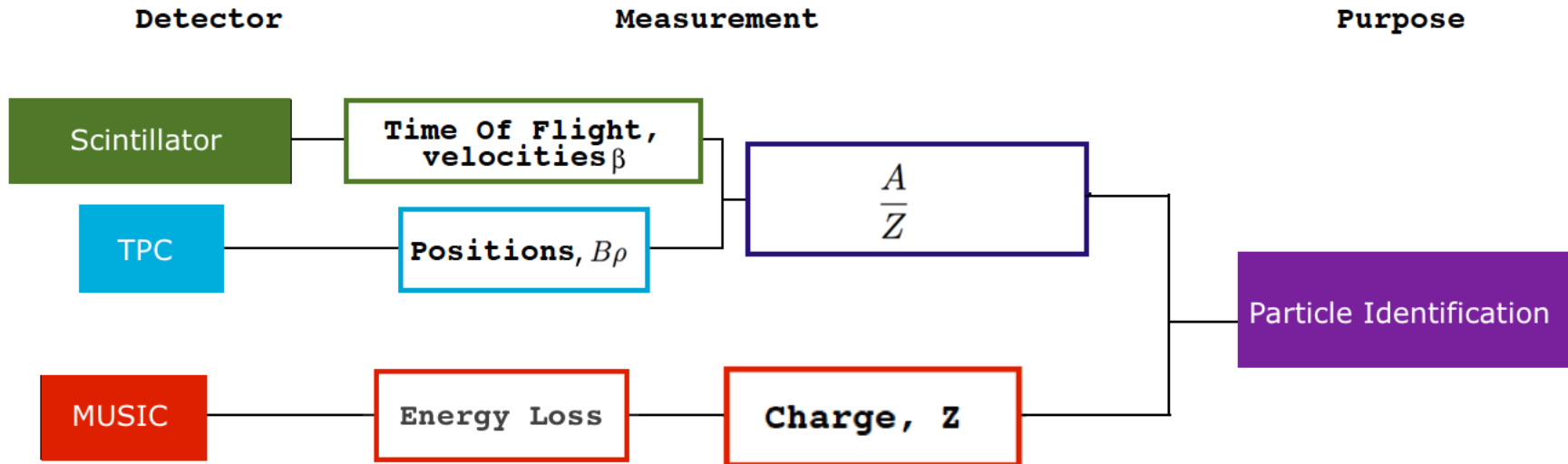




Particle Identification Spectrum

$$\frac{A}{Z} = \frac{e B \rho}{u \gamma \beta c}$$

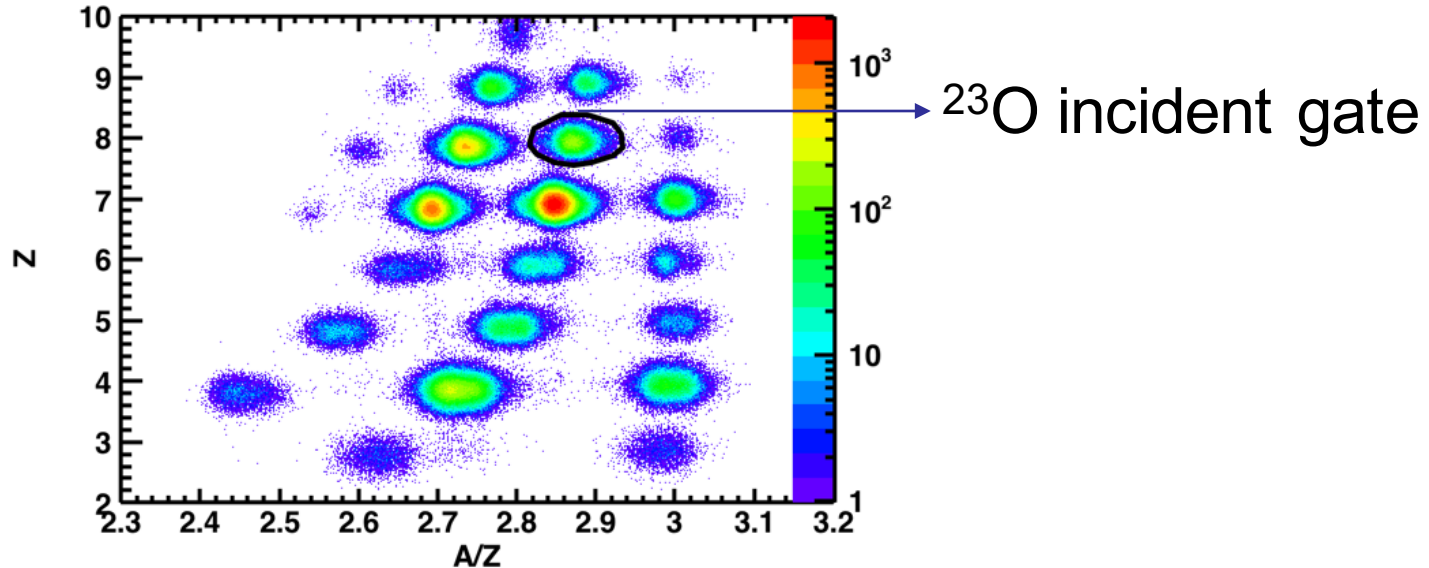
Magnetic rigidity, $B\rho = B\rho_{central} \left(1 - \frac{M_B x_{F2} - x_{F4}}{D_B} \right)$



Multisampling Ionization Chamber

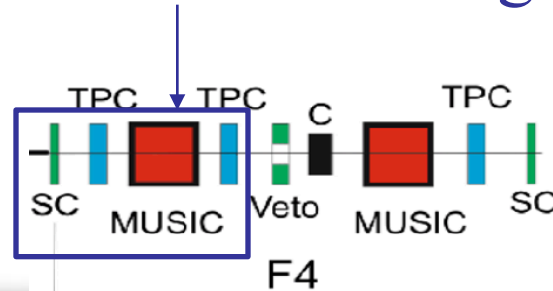


Particle Identification Spectrum



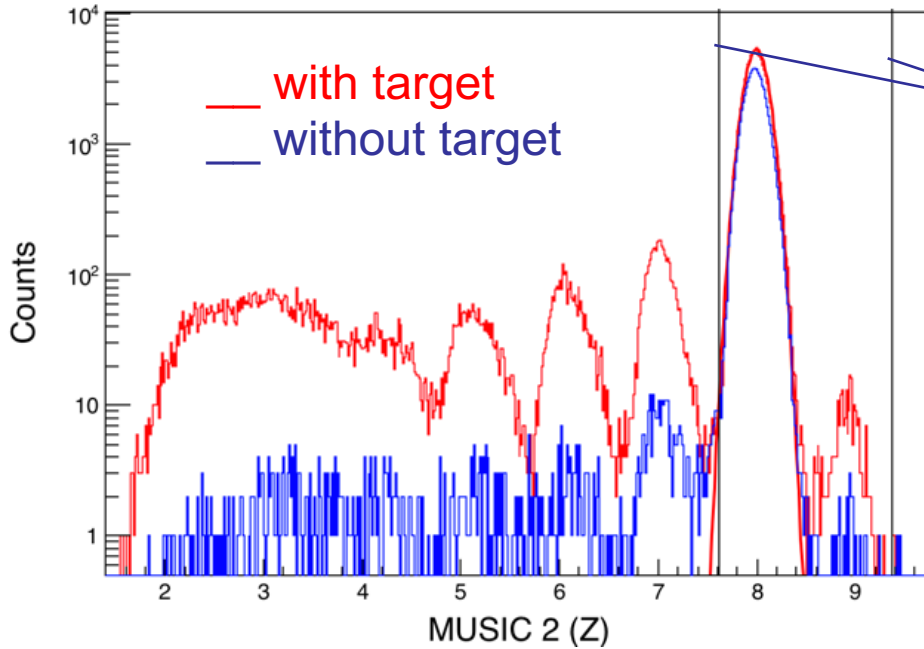
PID after removing spurious events

PID spectrum for ^{23}O before the target





Z identification after the target

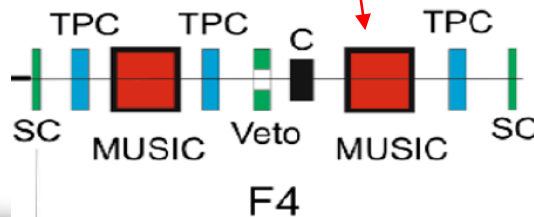


N_{sameZ} gate for ^{23}O
 3.5σ of Z=8 and Z=9 Peak

Transmission ratio $(R) = \frac{N_{sameZ}}{N_{in}}$

$$\sigma_{cc} = \frac{1}{t} \ln \frac{R_{out}}{R_{in}}$$

Energy loss spectrum in MUSIC detector after the target with ^{23}O incident beam selected.





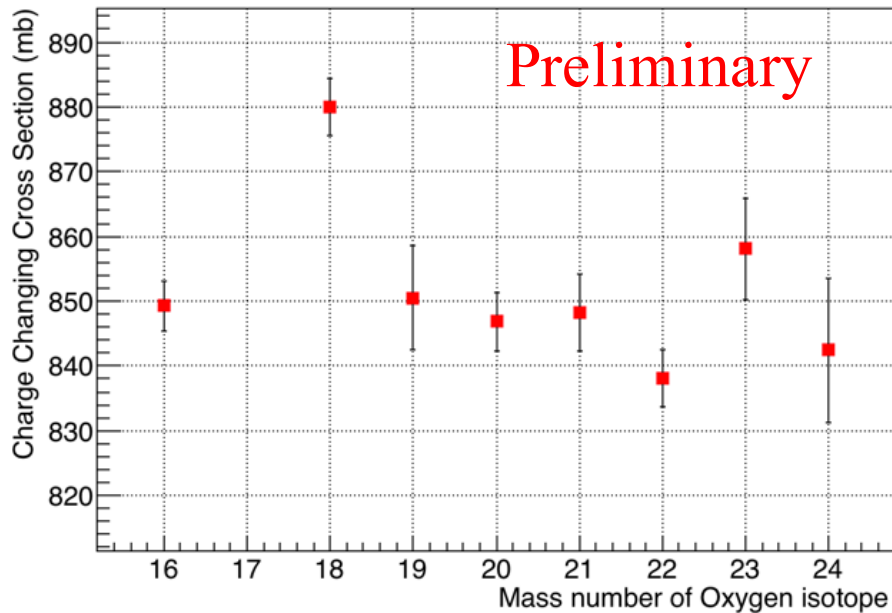
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Results

σ_{cc} of O isotopes



- An increase in σ_{cc} of ^{18}O .
- The σ_{cc} of $^{19-21}\text{O}$ shows flat trend.
- The σ_{cc} of ^{22}O decreases followed by an increase for ^{23}O .



R_p of ^{16}O and ^{18}O from the σ_{cc}

Glauber Model formalism

$$\sigma_{cc} = \int d\mathbf{b} P_{cc}(\mathbf{b}).$$

$$P_{cc}^{\text{dir}}(\mathbf{b}) = 1 - \exp \left(-2 \sum_{N=p,n} \iint ds dt T_{\text{Proj}}^{(p)}(\mathbf{s}) T_{\text{Target}}^{(N)}(\mathbf{t}) \right. \\ \left. \times \text{Re} \Gamma_{pN}(\mathbf{b} + \mathbf{s} - \mathbf{t}) \right),$$

where $T_{\text{Proj}}^{(p)}(\mathbf{s}) = \int_{-\infty}^{\infty} dz \rho_P^{(p)}(\mathbf{r})$ Proton density

and $\Gamma_{NN}(\mathbf{b}) = \frac{1 - i\alpha}{4\pi\beta} \sigma_{NN}^{\text{tot}} \exp\left(-\frac{\mathbf{b}^2}{2\beta}\right)$

(NN cross section) (finite range parameter)



R_p of ^{16}O and ^{18}O from the σ_{cc}

Glauber Model formalism

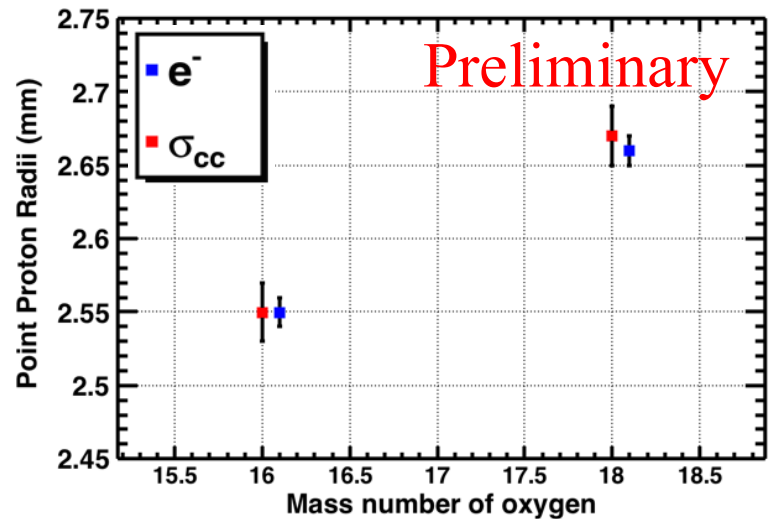
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where $T_P^{(p)}(\mathbf{s}) = \int_{-\infty}^{\infty} dz \rho_P^{(p)}(\mathbf{r})$ Proton density

and $\Gamma_{NN}(\mathbf{b}) = \frac{1 - i\alpha}{4\pi\beta} \sigma_{NN}^{\text{tot}} \exp\left(-\frac{b^2}{2\beta}\right)$ (finite range parameter)

(NN cross section)



R_p for ^{16}O and ^{18}O Agree with e^- scattering experiments.



Summary

- » σ_{cc} measurement is a new method to determine R_p of neutron-rich isotopes.
- » R_p determined from σ_{cc} for ^{16}O and ^{18}O are consistent with electron scattering experiment.
- » The first measurement of R_p for $^{19-24}\text{O}$ is underway.
- » The measured R_p will be used for first determination of neutron skin of O isotopes.
- » The measured R_p will also verify various newly developed models.



Acknowledgements

Supervisor : Dr. Rituparna Kanungo

W.Horiuchi, F. Ameil, J. Atkinson, Y. Ayyad, S. Bagchi, D. Cortina-Gil, I. Dillmann, A. Estrade, A. Evdokimov, F. Farinon, H.Geissel, G. Guastalla, R. Janik, R. Knobel, J. Kurcewicz, Yu. A. Litvinov, M. Marta, M.Mostazo, I. Mukha, C. Nociforo, H.J. Ong, S. Pietri, A. Prochazka, C. Scheidenberger, B. Sitar, P.Strmen, M. Takechi, J. Tanaka, I. Tanihata, S.Terashima, J. Vargas, H.Weick, and J. S.Winfield



GSI Helmholtzzentrum für Schwerionenforschung GmbH

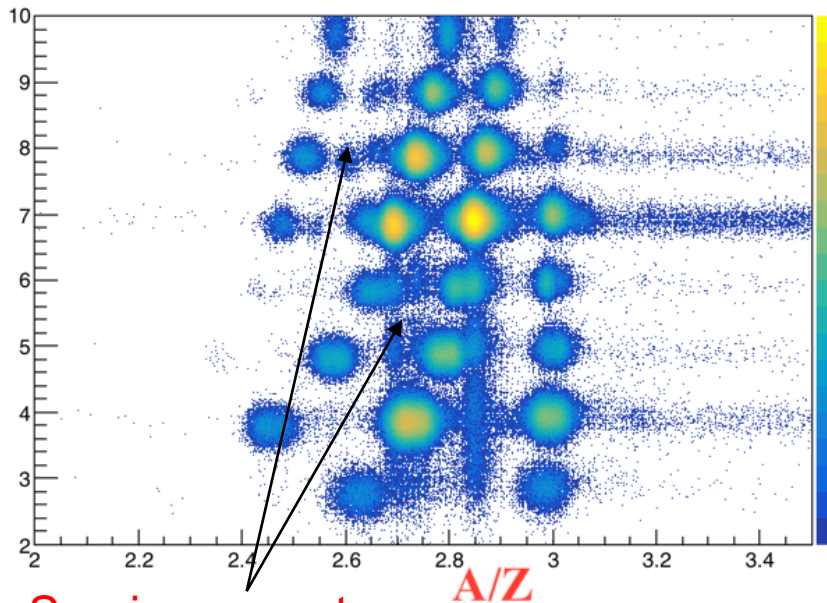


Thank You

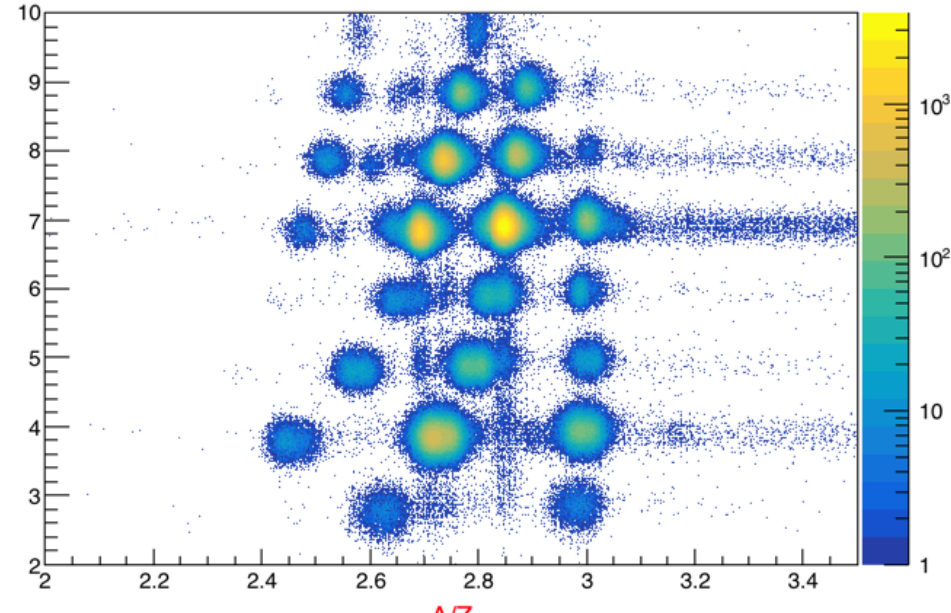


Particle Identification(PID) for ^{23}O

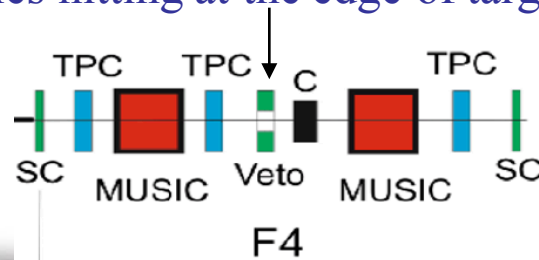
PID after veto rejection



Spurious events



Veto_Detector \rightarrow Plastic scintillator to reject particles hitting at the edge of target





Glauber Model and σ_{cc}

Glauber Model applied successfully to Boron isotopes

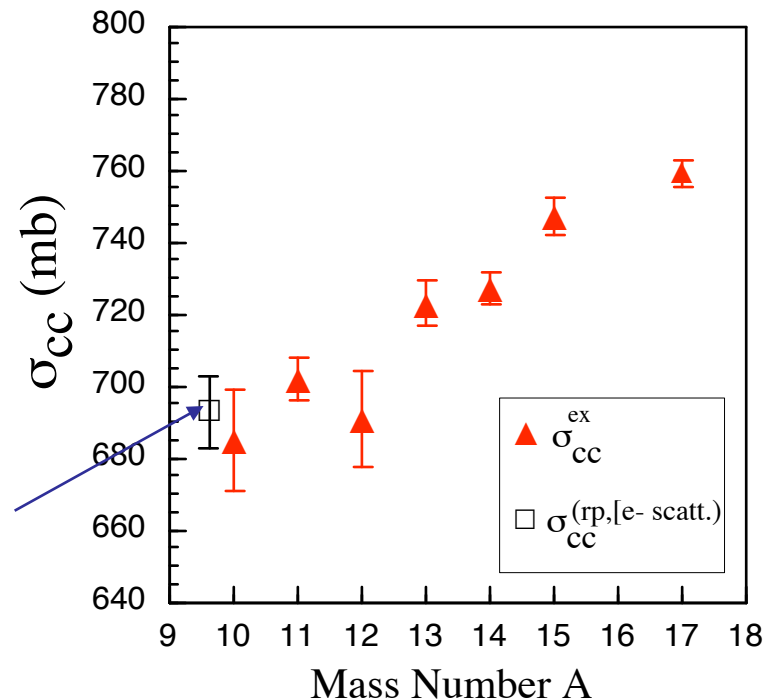
$$\Gamma_{NN}(\mathbf{b}) = \frac{1 - i\alpha}{4\pi\beta} \sigma_{NN}^{\text{tot}} \exp\left(-\frac{b^2}{2\beta}\right)$$

where α is the ratio of the real to the imaginary part of (σ_{NN}) scattering amplitude in the forward direction,

(σ_{pn}^{tot}) is the pp (pn) total cross sections,

β is the slope parameter of the elastic scattering differential cross section.

$$\beta_{pN} = \frac{1 + \alpha_{pN}^2}{16\pi} \sigma_{pN}^{\text{tot}}$$



(A. Estradé et al., Phys. Rev. Let.113, 132501,(2014))



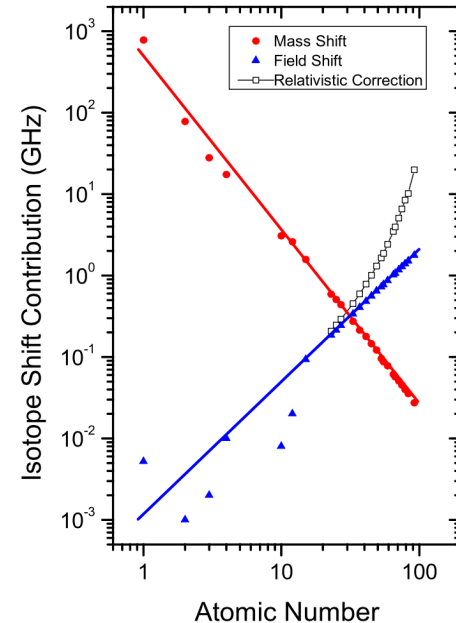
Limitations of isotope shift

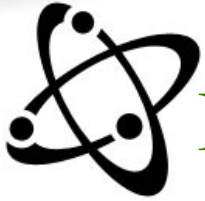
Isotope shift: change in energy of atomic levels of different isotopes.

$$\delta\nu_{A,A'} = \delta\nu_{A,A'}^{\text{MS}} + K_{\text{FS}} \delta\langle r^2 \rangle_{A,A'}$$

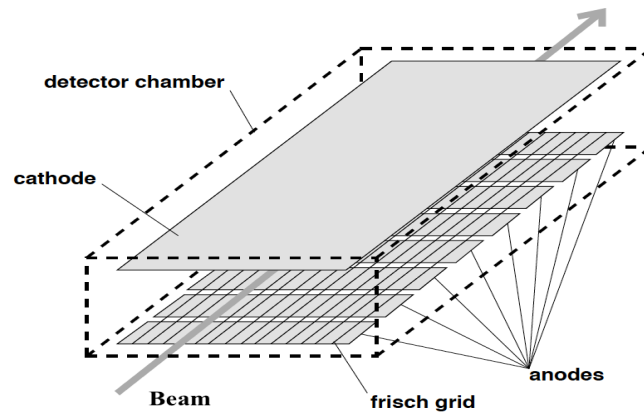
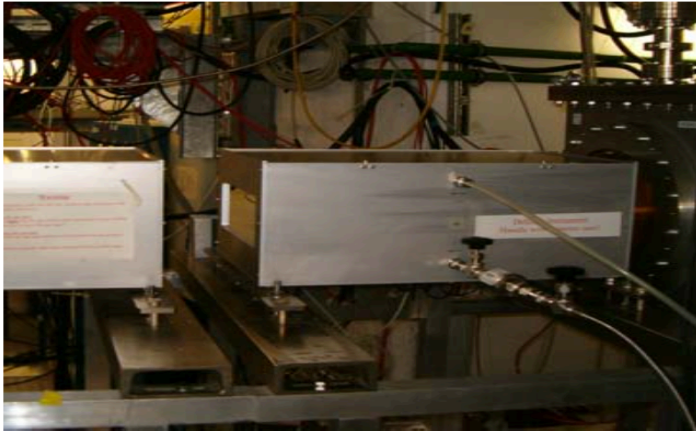
Limitations

- low energy and good intensity beams difficult to produce for all neutron rich isotopes.
- Mass shift term dominates for O.
- Many body calculations complicated





Multisampling Ionization Chamber(MUSIC)



Counting Gas -CF₄
Pressure- 1 bar
Dimensions-200 x 80 x 100mm

Bethe formula for energy loss

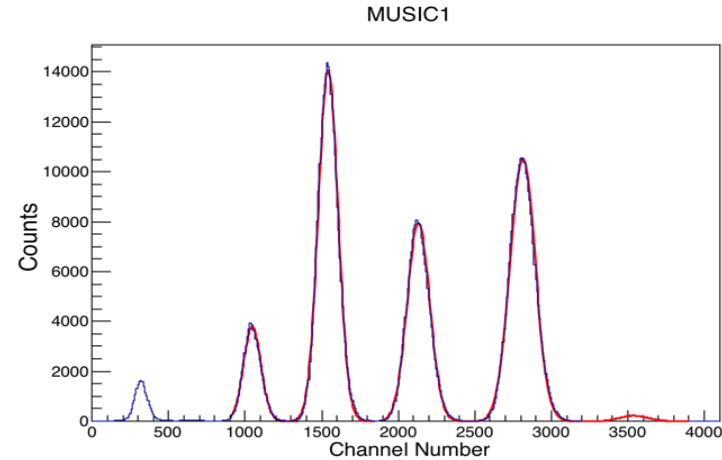
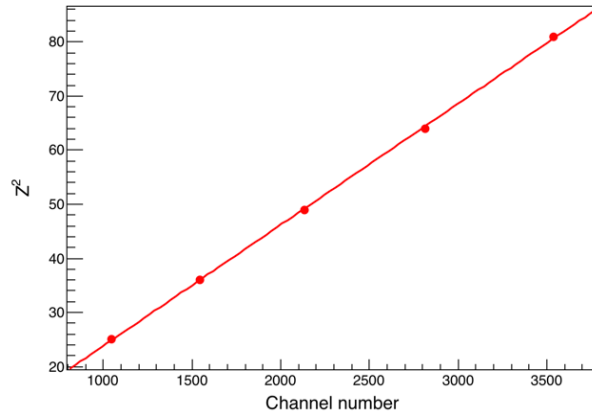
$$\frac{-dE}{ds} = \frac{4\pi Z_p^2}{m_e c^2 \beta^2} \left(\frac{e^2}{4\pi\epsilon_0} \right)^2 Z_t N_t \left(\ln \frac{2m_e v^2}{I} - \ln(1 - \beta^2) - \beta^2 \right)$$

Geometric average of signals from each anode gives us the energy loss

$$dE_{raw} = (e_1 \cdot e_2 \cdot e_3 \cdot e_4 \cdot e_5 \cdot e_6 \cdot e_7 \cdot e_8)^{1/8}$$

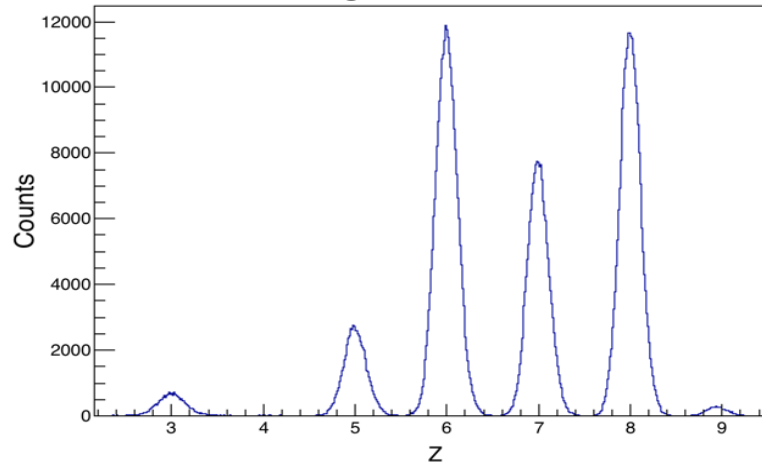


Calibration of MUSIC Detector



Points correspond to mean of the peaks obtained from gaussian fit of each peak in MUSIC energy spectrum

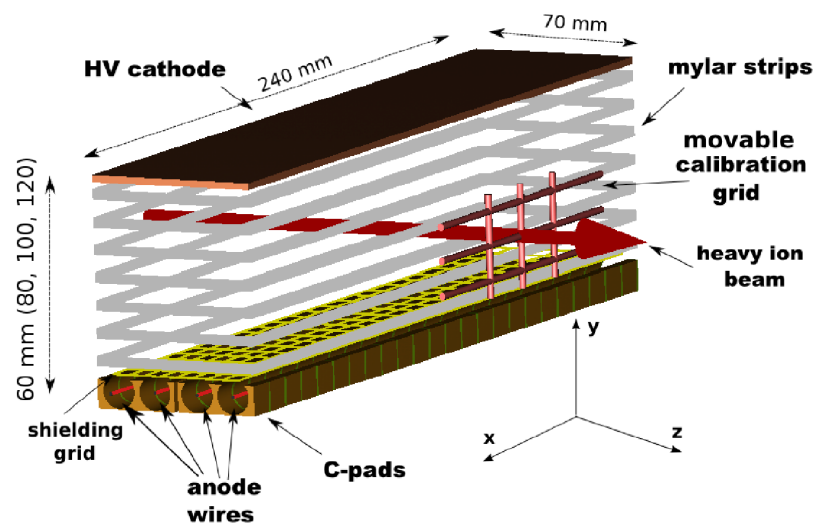
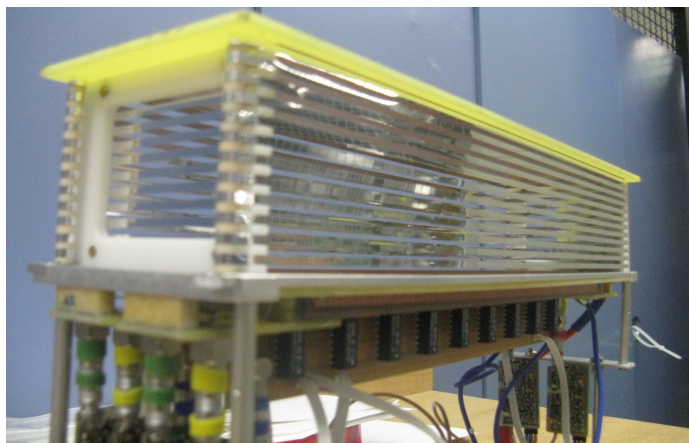
$$Z^2 = g * c + o$$



Calibrated MUSIC spectrum



Time Projection Chamber

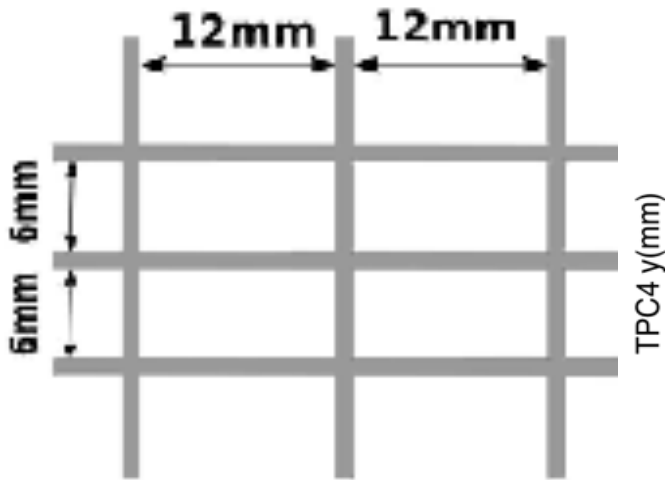


$$y = w_d t_a + y_{\text{off}}$$

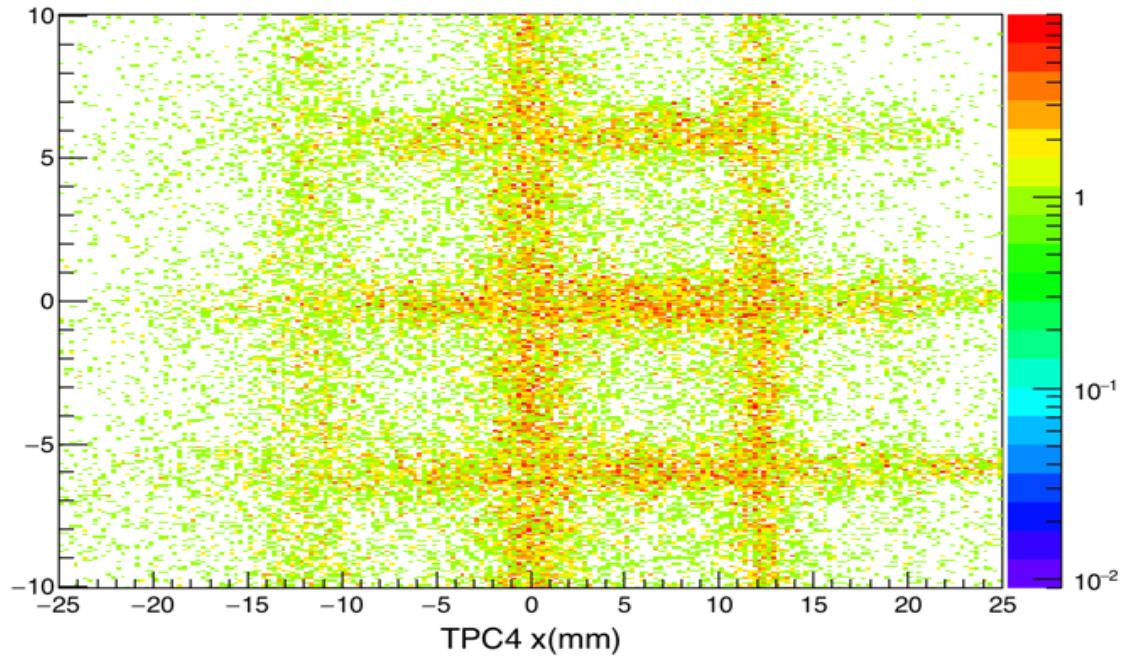
$$x = w(t_1 - t_r) + x_{\text{off}}$$



TPC Detector Calibration



Schematic of scintillator fibres used for calibration of TPC



2 dimensional position spectrum showing structure of grid using stable beam.



Plastic scintillator

Scintillator: Energy deposited \longrightarrow light.
PMT: Light \longrightarrow electrical pulse

Each scintillator (at F2 and F4) had photomultiplier modules on both sides which gives two measurements for each detector.

$$TOF_{RR} = |T_{41R} - T_{21R}|$$

$$TOF_{LL} = |T_{41L} - T_{21L}|$$

$$TOF = \frac{(TOF_{RR} + TOF_{LL})}{2}$$



Calibration of TOF from Scintillator Detector

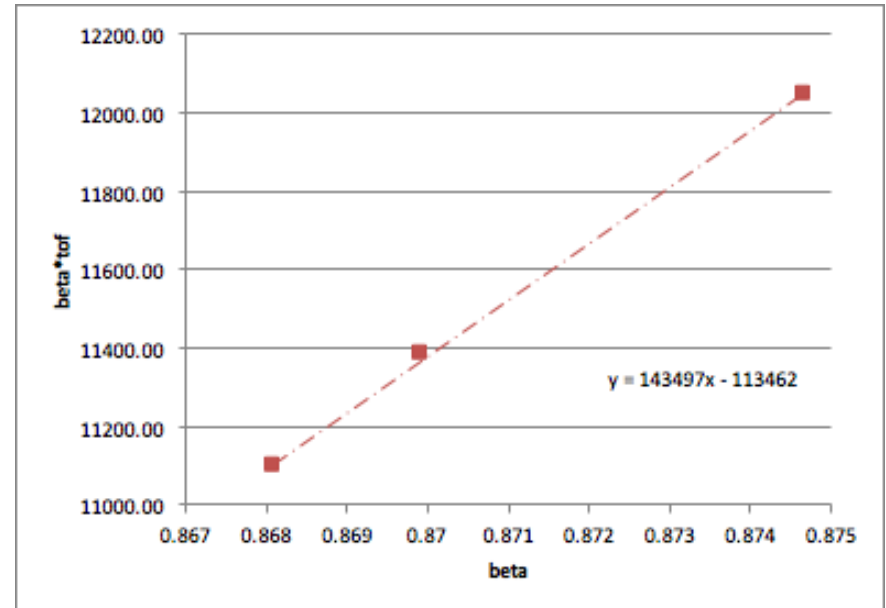
Time of flight is calibrated using stable primary beam with three different velocities.

$$\text{beta} * \text{tof} = \text{Flight path} + \text{beta} * \text{tof offset}$$

$$\text{tof offset} = 143497$$

$$\text{flight path} = 113462$$

$$\text{beta} = \frac{\text{flight path}}{\text{TOF} - \text{TOFoffset}}$$



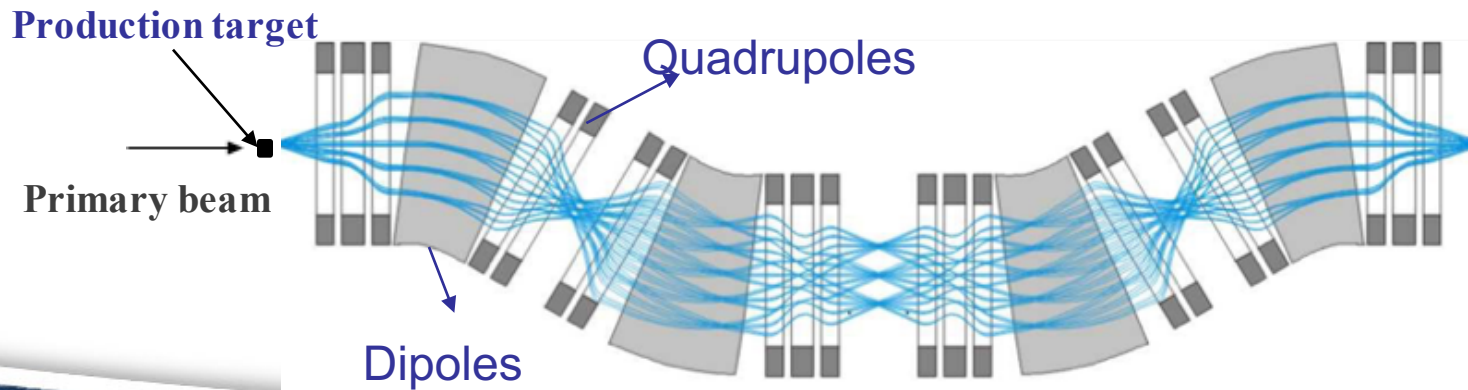
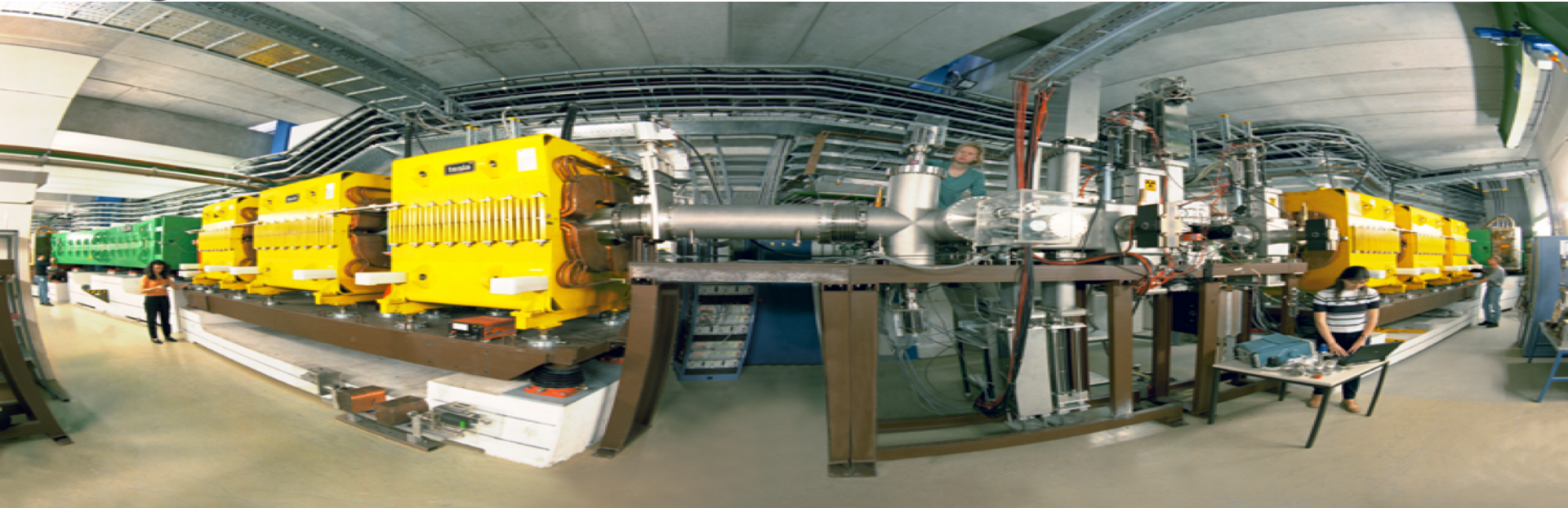


Chiral effective field theory

	2N Force	3N Force	4N Force
LO $(Q/\Lambda_\chi)^0$		—	—
NLO $(Q/\Lambda_\chi)^2$		—	—
N ² LO $(Q/\Lambda_\chi)^3$			—
N ³ LO $(Q/\Lambda_\chi)^4$			



Fragment Separator GSI





Charge Radii Relation to Point Proton Radii

Root mean square charge radius r_c is given by

$$\langle r_c^2 \rangle = \langle r_p^2 \rangle + \langle R_p^2 \rangle + \frac{N}{Z} \langle R_n^2 \rangle + \frac{3\hbar^2}{4mp^2c^2}$$

r_p is the radius of point proton distribution of a nucleus

R_p and R_n are the charge radii of free proton and free neutron

last term is so called Darwin–Foldy term

Matter Radii Related to Point Proton Radii and point neutron radii

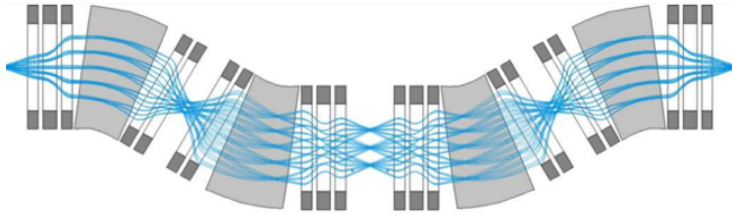
$$\langle r_m^2 \rangle = \frac{Z}{A} \langle r_p^2 \rangle + \frac{N}{A} \langle r_n^2 \rangle$$

I. Tanihata et al. / Progress in Particle and Nuclear Physics 68 (2013) 215–313



Ion optics of Fragment separator

FRS working as an achromatic system
with a dispersive mid-plane



$$\begin{pmatrix} x \\ a \\ y \\ b \\ s \\ \delta_p \end{pmatrix} = M \begin{pmatrix} x_0 \\ a_0 \\ y_0 \\ b_0 \\ s_0 \\ \delta_{p0} \end{pmatrix}.$$

$$a = \frac{dx}{ds} \quad b = \frac{dy}{ds}$$

$$\delta_p = \frac{p - p_0}{p_0}$$

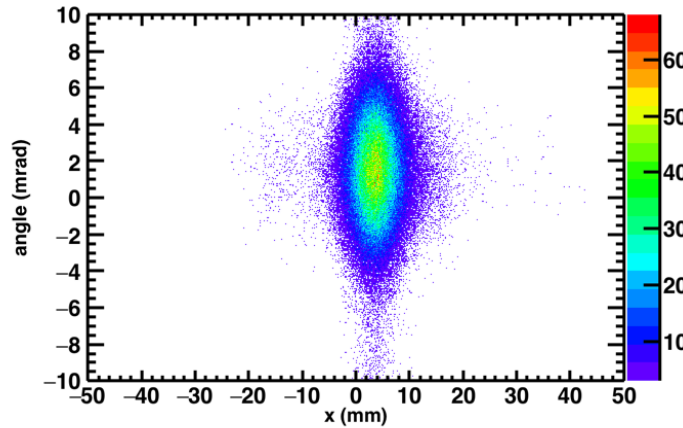
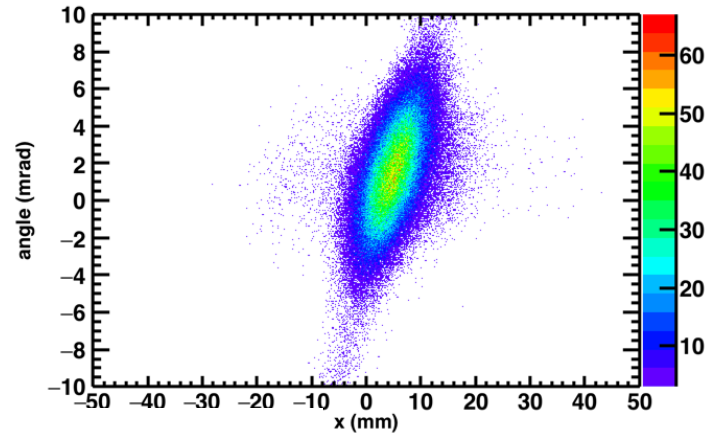
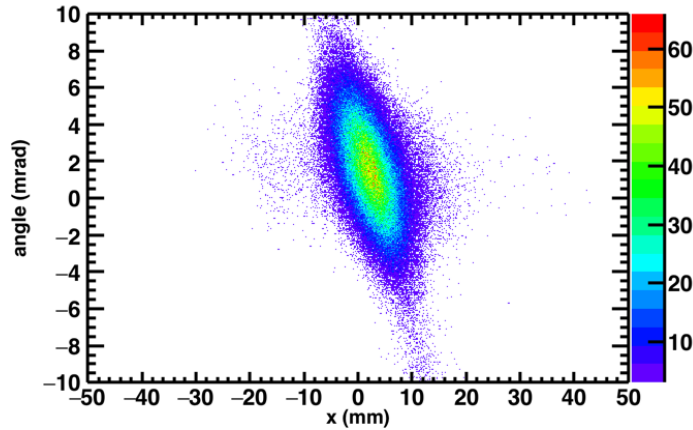
Magnification ← $M = \begin{pmatrix} (x|x) & (x|a) & 0 & 0 & 0 & (x|\delta) \\ (a|x) & (a|a) & 0 & 0 & 0 & (a|\delta) \\ 0 & 0 & (y|y) & (y|b) & 0 & 0 \\ 0 & 0 & (b|y) & (b|b) & 0 & 0 \\ (s|x) & (s|a) & 0 & 0 & 1 & (s|\delta) \\ 0 & 0 & 0 & 0 & 0 & 1 \end{pmatrix} \rightarrow D$

$$\delta_{F2} = \frac{1}{D_B} [x_{F4} - (x|x)_B x_{F2}]$$

$$\delta_{F2} = \frac{p - p_B}{p_B} = \frac{\chi - \chi_B}{\chi_B} \longrightarrow \chi = (1 + \delta_{F2})\chi_B$$



Focal plane location using position from TPC



The x-position at the image plane is independent from the incident angle of the beam.



Nuclear Density distributions

Fermi or woods saxon form

$$\rho(r) = \rho_0 / (1 + \exp((r - c)/z))$$

c is the radius of distribution to a point where density falls to half and z is diffuseness related to thickness of surface region.

Harmonic oscillator density

$$\rho(r) = \rho_0 (1 + \alpha(r/a)^2) \exp(-(r/a)^2)$$

$$\alpha = \alpha_0 a_0^2 / (a^2 + \frac{3}{2} \alpha_0 (a^2 - a_0^2))$$

$$a_0^2 = (a^2 - a_p^2) A / (A - 1)$$

$$\alpha_0 = (Z - 2) / 3; \quad a_p^2 = \frac{2}{3} \langle r^2 \rangle_{\text{proton}}$$

where a is the size parameter.



Proton Radii from σ_{CC}

Glauber Model

(R. J. Glauber:, 1959)

$$\sigma_R = \iint [1 - T(\mathbf{b})] db$$

At high energies

- » Nucleons follow straight lines trajectories .
- » Interaction of projectile and the target governed by individual nucleon nucleon cross section.

Probability of interaction $[p(b)] \propto \sigma_{nn}, \rho_P(r,z)$ and $\rho_T(r,z)dz$



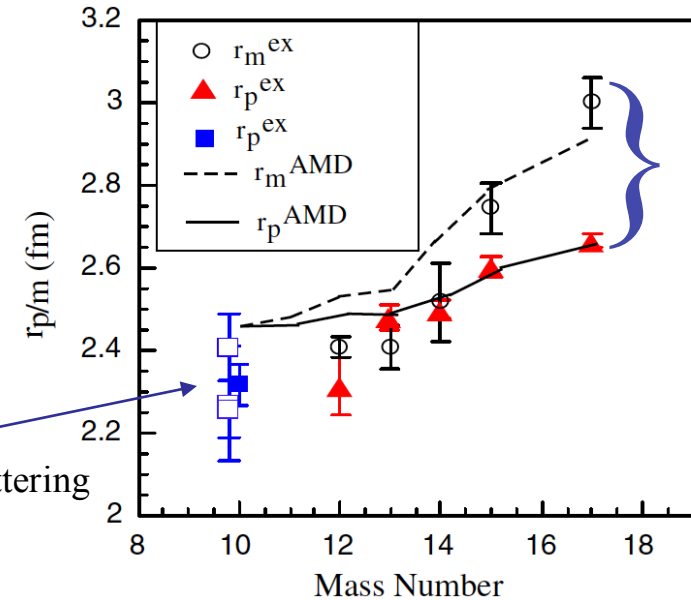
Proton radii measured from σ_{cc}

Glauber Model applied successfully to B and C isotopes

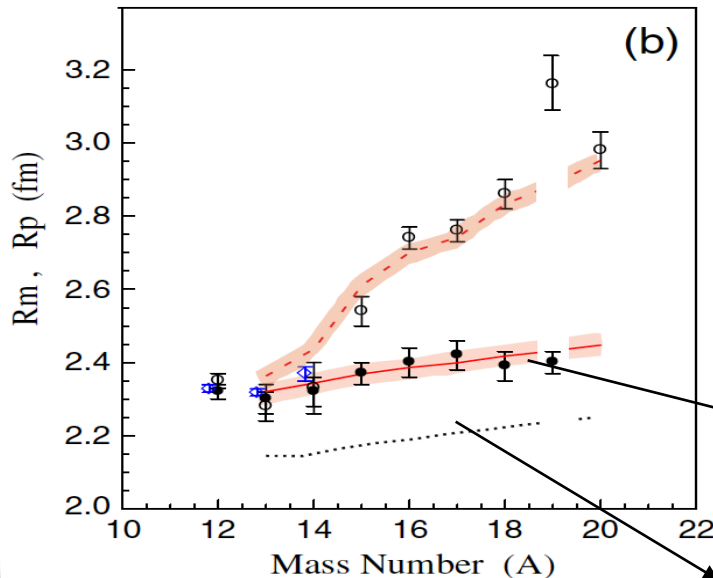
Boron isotopes

(A. Estradé et al., Phys. Rev. Lett. 113, 132501, (2014))

- A thick neutron skin of 0.51 ± 0.11 fm was observed in ^{17}B .
- ($\delta R = 0.15 \pm 0.04$ fm for ^{208}Pb)



Proton radii from electron scattering



Carbon isotopes

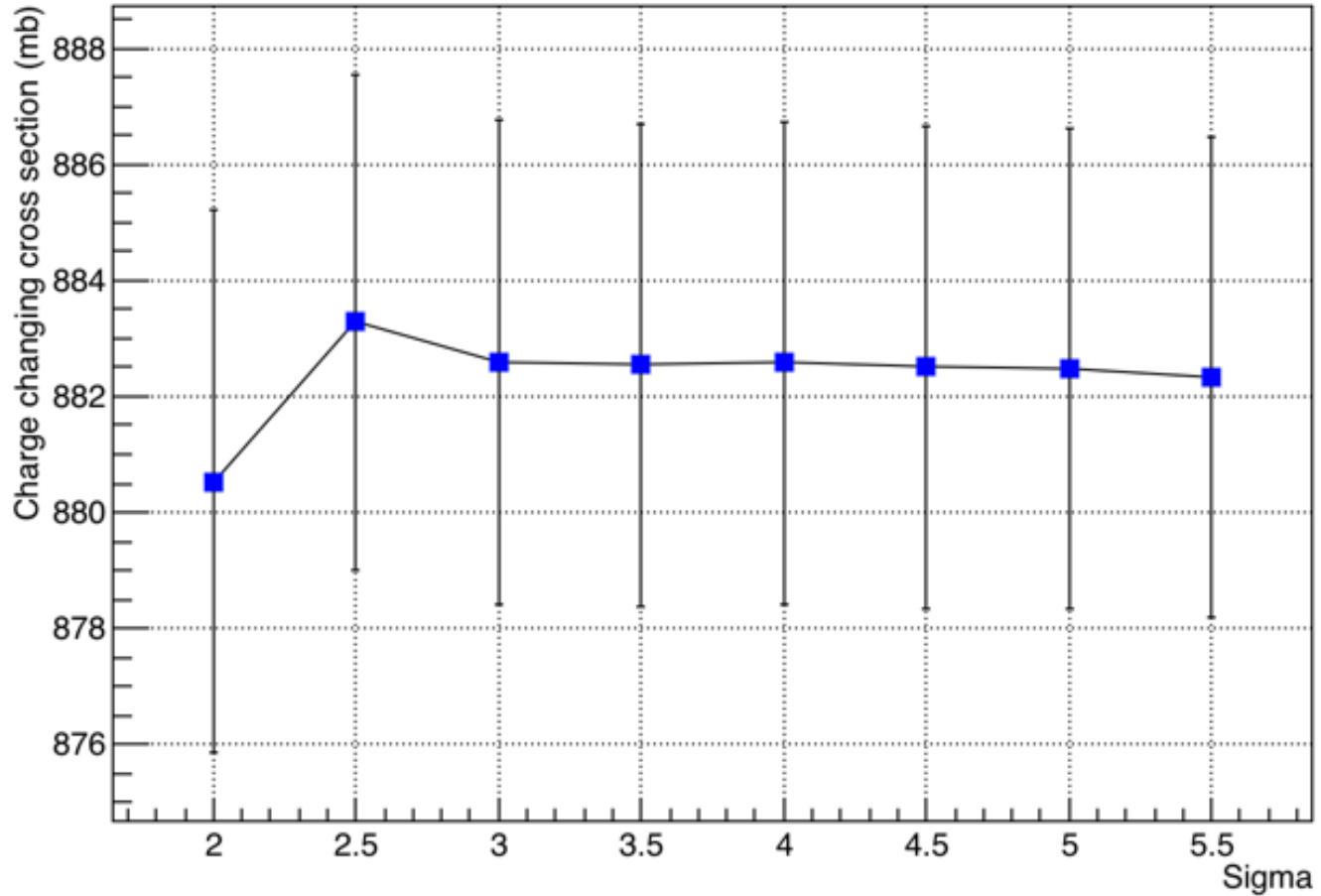
R. Kanungo et al. Phys. Rev. Lett. 117, 102501 (2016).

Proton radii consistent with 3N forces

Proton radii with NN forces



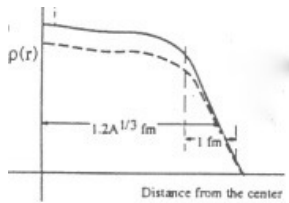
180 Sigma CC variation with different NsameZ gate





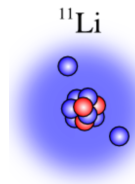
Nuclear Structure from Nuclear Radii

For Stable Nucleus

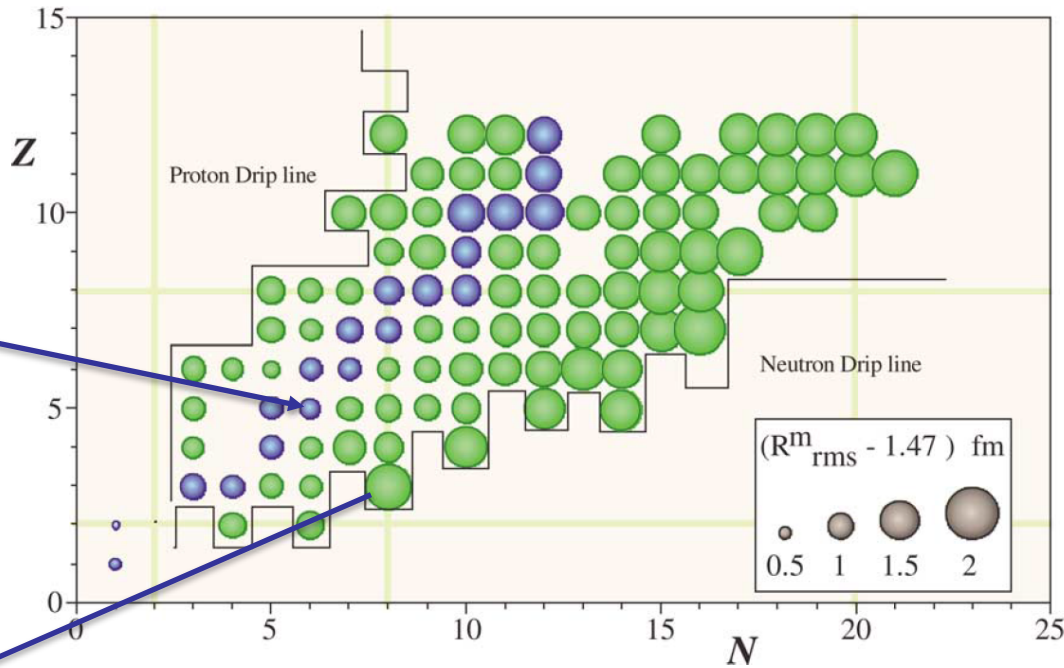


^{11}B ($R \approx 2$ fm)

Neutron halo



^{11}Li
($R \approx 3.5$ fm)



I. Tanihata, R. Kanungo / C. R. Physique 4 (2003) 437–449

$$R \neq 1.2 A^{1/3}$$



PID for 240

lb_veto && music41_tpc4 && music41_asc41 && music41_tpc5 && sc41lrt_music41 && Toflivstofrr

