# 3rd Order Nonlinear Susceptibility Calculations of Kr for Lyman-α Laser

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# The ALPHA-g Collaboration at CERN

Antihydrogen Laser PHysics Apparatus

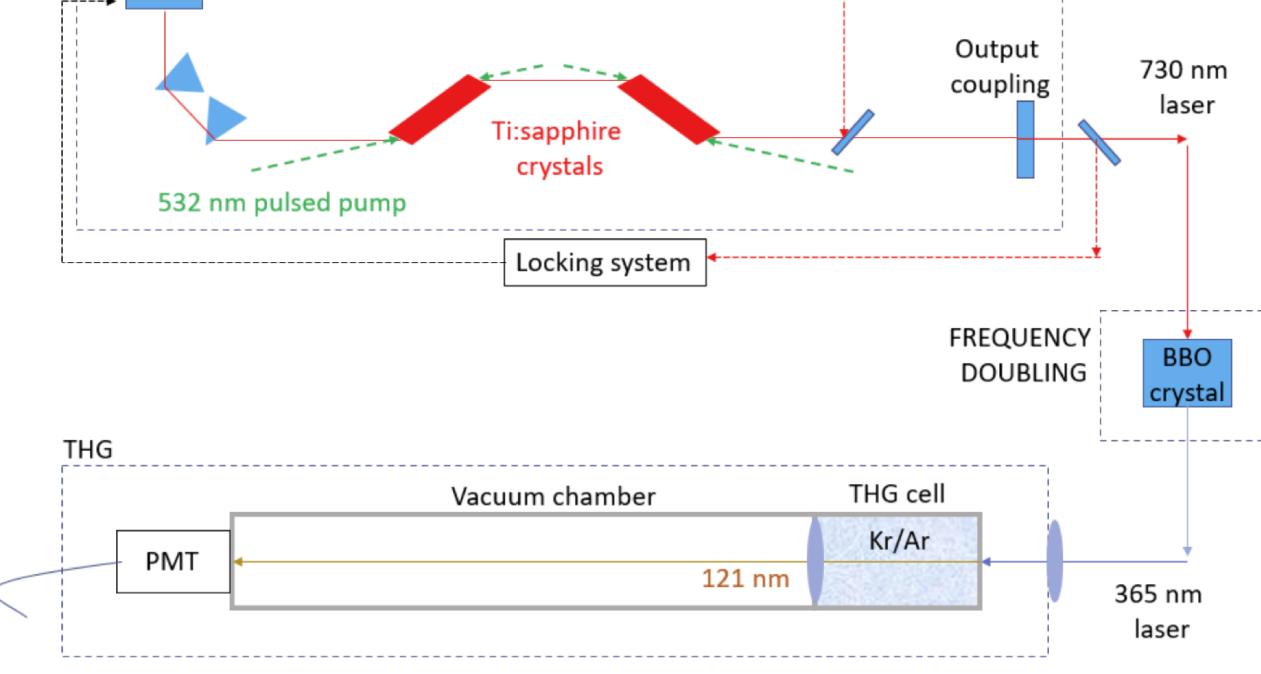
At ALPHA-g, the trajectories of antihydrogen atoms as they leave a magnetic trap in order to investigate the gravitational behaviour of the anti-atom. A Lyman- $\alpha$  (121 nm) laser is needed to cool the atoms to approximately 50 mK. My work involved calculating 3rd-order  $\chi$  for the laser set-up.

### Laser Set-Up

	Seed laser	
LASER OSCILLATOR	730 nm	

#### Nonlinear Optical Processes

When light hits certain materials, the response of the system is not linear in relation to the magnitude of the applied electric field. Each  $\chi$  term represents the nth-order nonlinear susceptibility.



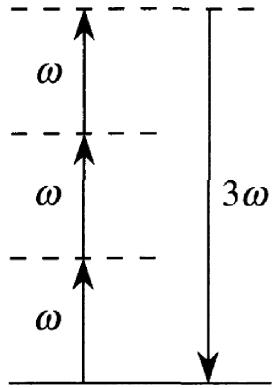
## Calculations

$$\tilde{P}(t) = \epsilon_0 [\chi^{(1)} \tilde{E}(t) + \chi^{(2)} \tilde{E}^2(t) + \chi^{(3)} \tilde{E}^3(t) + \dots]$$

We focus on the third-order nonlinear processes  $\chi^{(3)}$  which involve 3 different incident waves to interact.

**Triple Harmonic Generation** 

3 photons of frequency  $\omega$  are destroyed, and 1 photon of frequency  $3\omega$  is created.



97945*cm*<sup>-</sup>

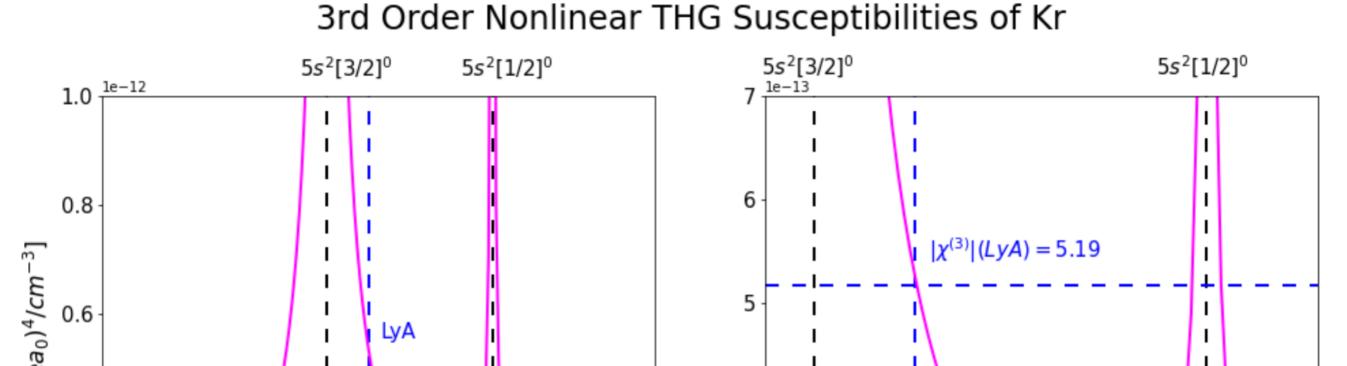
 $\chi^{(3)}$  is a tensor where the indices k,m, and n represent three transitions that correspond to the new energy level when an  $\omega$  is absorbed ( $\omega$ , 2 $\omega$ , and 3 $\omega$ ).  $\chi^{(3)}$  is a summation of all allowed paths.

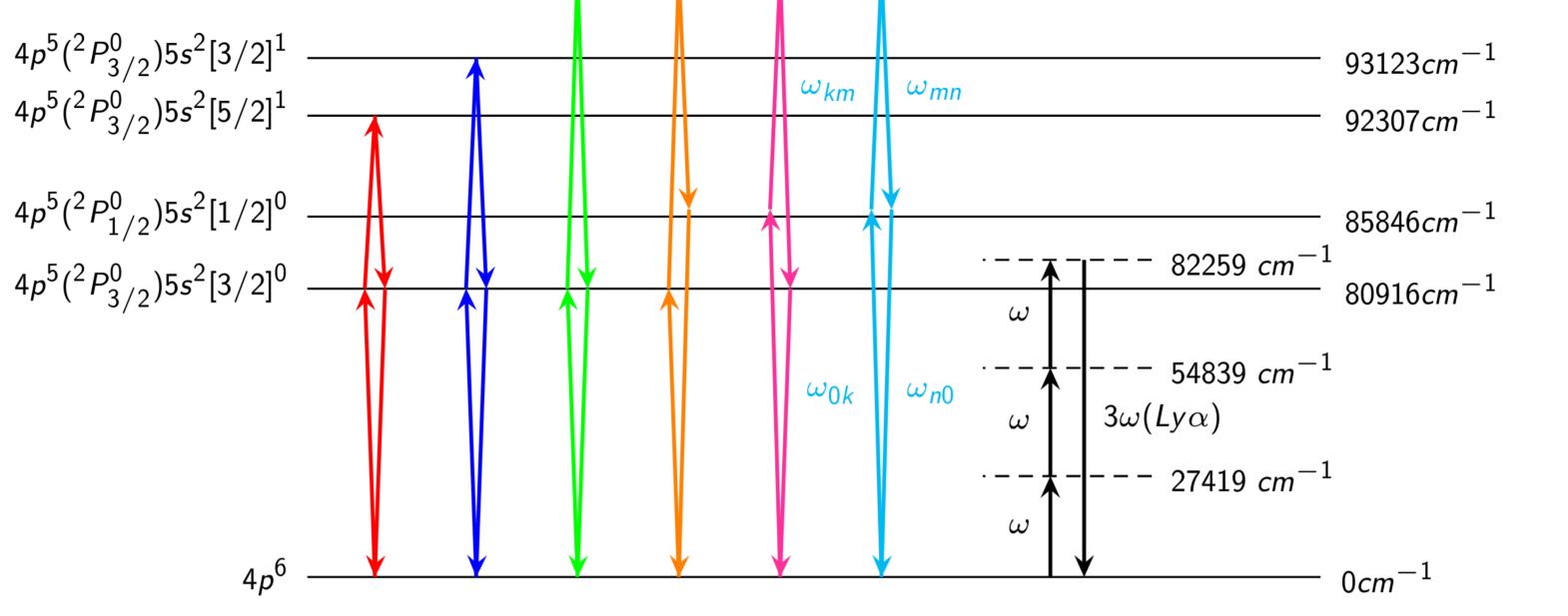
 $4p^{5}(^{2}P^{0}_{1/2})5s^{2}[3/2]^{1}$ 

$$\chi^{(3)} \alpha \frac{\text{dipole moments}}{\text{resonances}}$$

- Wigner-Eckart Theorem separates dipole moment into corresponding Clebsch-Gordan coefficient and reduced matrix element
- Reduced matrix element calculated using oscillator strength and transition probability
- 3ω close to resonance to yield a higher nonlinear susceptibility without high chance of absorption

#### **Results and Conclusion**





**Figure:** Energy Level Diagram to show allowed Transitions in Kr, and THG

Results: 9.2 \* 10<sup>-37</sup> esu/cm<sup>-3</sup>

Mahon et al. (1979): 11 \* 10<sup>-37</sup> esu/cm<sup>-3</sup>

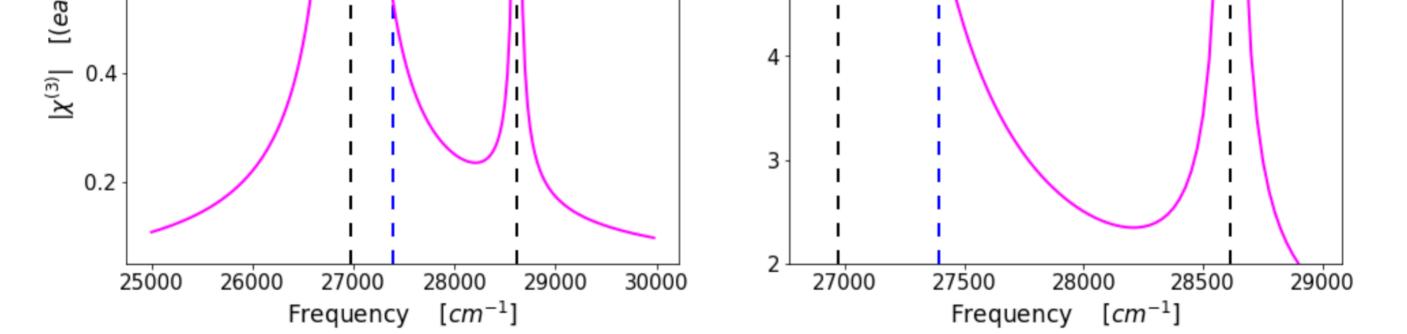


Figure:  $\chi^{(3)}$  values for Kr where  $\omega$  ranges from 25000 to 30000  $cm^{-1}$  or approximately 330-400 nm.  $\chi^{(3)}$  at LyA is approximately 9.2  $* 10^{-37} esu/cm^{-3}$  with 20 percent uncertainty.



Since our values are on the same magnitude as values found in literature, our method is verified

Discrepancy comes from more significant digits to transition probabilities and oscillator strengths

Discovery, accelerated