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## Electrodisintegration of $^{16}\text{O}$ and determination of astrophysical S-factors of the inverse reaction

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After more than five decades of experimental effort the rate of  $\alpha$  on  $^{12}\text{C}$  radiative capture at astrophysical energies ( $\sim 0.3$  MeV above threshold) is not determined with desired precision and it is a cause of the largest uncertainty contribution in modeling of evolution of massive stars and underlying nucleosynthesis. By using the windowless gas jet target and modern energy-recovery linear accelerators (ERLs, CBETA at Cornell, NY, USA and MESA in Mainz, Germany) to reach high luminosity, a high precision measurement of the electron scattering on  $^{16}\text{O}$  nucleus would provide a method to determine the rate of the  $\alpha$  on  $^{12}\text{C}$  radiative capture for energy range  $< 2$  MeV with a superb precision compared to previous experiments [1]. The feasibility of this method still needs to be studied. This could be done in a moderate luminosity experiment at existing electron accelerator sites by measuring the rate at  $> 2$  MeV where the cross section is much larger.

[1] I. Frišćić, T. W. Donnelly, and R. G. Milner, Phys. Rev. C 100, (2019) 025804

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### Attendance

### Contact Email

### Scheduling Constraints

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