Supernova and solar neutrinos: the present and the future

Erin O'Sullivan Stockholm University, Oskar Klein Centre NNN 2018 – Vancouver With X-rays, which penetrate much more than ordinary light, you can see inside your hand. With neutrinos, which penetrate much more even than x-rays, you can look inside the Sun.

- Ray Davis Jr, Nobel Ceremony 2002



Solar neutrinos

Solar neutrinos are produced in thermonuclear reactions









Complementary approaches to measuring the solar upturn/transition



9

Looking ahead: Transition region



SNO+ has the potential to measure pep neutrinos to 10% uncertainty in 1 year (depending on backgrounds)

Hyper-K will measure the upturn with 3-5 **σ** significance

Measuring the last unknowns of the solar neutrinos



Supernova neutrinos

Thermodynamics imprinted on neutrinos



Detecting supernova neutrinos in IceCube

Abbasi+ 2011



Mainly inverse beta decay ($\overline{\mathbf{v}}_{e} + \mathbf{p} \rightarrow \mathbf{n} + \mathbf{e}^{+}$)

IceCube has the best neutrino rate measurement from nearby

supernovae



15

IceCube can measure features imprinted in the timing structure

Tamborra+ 14



Detecting supernova neutrinos in water



Measuring the supernova direction



Super-K can measure the direction of a SN with half angle of 3-5° @ 10 kpc (x2 times better with Gd) Hyper-K can measure the direction of a SN with half angle of 1-2°@ 10 kpc 18

Liquid scintillator detectors have unique NC channels



Liquid scintillator can provide an early-early warning for a supernova



Time [day]

Lead measures non- $\overline{\mathbf{v}}_{e}$ channels

Astronomically patient

```
\mathbf{v}_{e} + <sup>208</sup>Pb → <sup>207</sup>Bi + n + e<sup>-</sup>
\mathbf{v}_{e} + <sup>208</sup>Pb → <sup>206</sup>Bi + 2n + e<sup>-</sup>
\mathbf{v}_{x} + <sup>208</sup>Pb → <sup>207</sup>Pb + n
\mathbf{v}_{x} + <sup>208</sup>Pb → <sup>206</sup>Pb + 2n
```



Liquid argon: a \mathbf{v}_{e} measurement from nearby supernovae

 $v_{e} + {}^{40}\text{Ar} \rightarrow e^{-} + {}^{40}\text{K}^{*}$



22

The technologies: complementarity

	Water	lce	Argon	Lead	Scint
v _e	v	v			~
Ve	(🖌)		~	~	
V _x				~	~
Low energy					~
Pointing	~				
Energy info	v		~		~
How many events?	Super-K	IceCube	DUNE	HALO	SNO+
	10,000	790,000	3,000	~tens	7,000

Neutrino detectors work together to sound the alarm



Neutrino detectors connected to the SNEWS system



Summary

- New solar neutrino measurements with next-gen detectors will look for new physics, particularly in the vacuum-matter transition region, and will look for the remaining unmeasured solar neutrinos
- Using the entire suite of technologies available to measure supernova neutrinos will allow us to capitalize on this once-in-a-lifetime event

Early supernova explosion produces mostly \mathbf{v}_{e}



Janka+ 2007

 $p + e^{-} \rightarrow n + v_e$

All flavours at late times



Janka+ 2007

ν_{e,μ,τ},ν_{e,μ,τ}

 $e^++e^-\rightarrow v + v$ $n+e^+ \rightarrow p+\overline{v}_e$ $p+e^{-} \rightarrow n+v_{e}$

hep discovery potential

