

Recent oscillation results from NOvA



NNN18

November 2018

Gregory Pawloski

University of Minnesota



What is NOvA

Long baseline neutrino experiment

$E \approx 2$ GeV (off-axis narrow band beam)

$L = 810$ km

Oscillations governed by Δm_{32}^2 (Δm_{31}^2)

NuMI beam produced at Fermilab

ν_μ and $\bar{\nu}_\mu$ beam modes

Analyzed 8.85×10^{20} POT with ν_μ mode

Analyzed 6.91×10^{20} POT with $\bar{\nu}_\mu$ mode

$(\bar{\nu}_\mu) \rightarrow (\bar{\nu}_x)$ oscillations

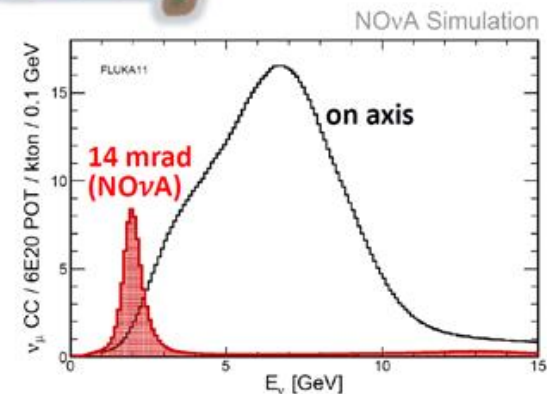
Two detector experiment

Near detector (Fermilab, IL)

Measure beam before oscillation

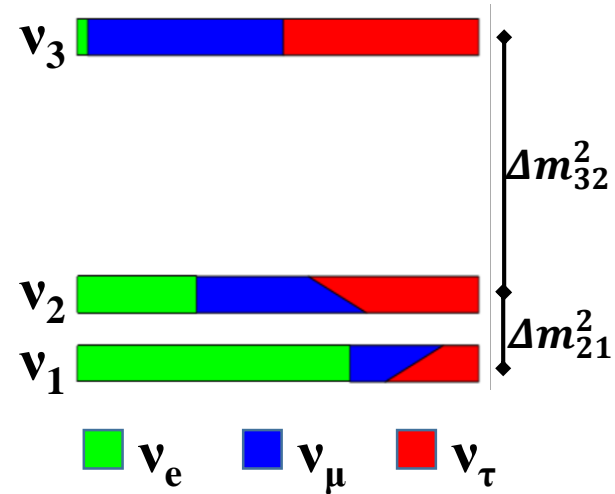
Far Detector (Ash River, MN)

Measure oscillated beam





Physics Goals



Long baseline neutrino oscillation measurements

ν_μ CC disappearance

ν_e CC appearance (long baseline $\pm 30\%$ matter effect)

Sensitive to: θ_{23} , δ_{CP} , Δm_{32}^2 (*Mass Hierarchy*)

NC disappearance

Sensitive to Sterile Neutrinos: θ_{24} , θ_{34} , Δm_{41}^2

Non-oscillation Measurements

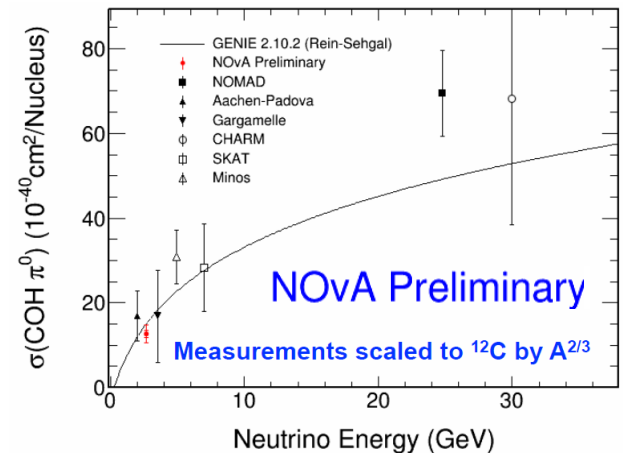
Near Detector cross sections measurements

Supernova detection

Exotic phenomena

Monopoles, neutrino magnetic moment, etc

NC Coherent Pion Production Measurement

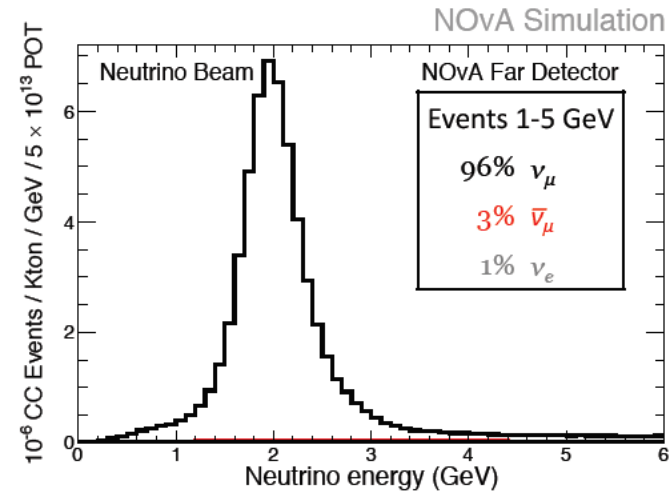
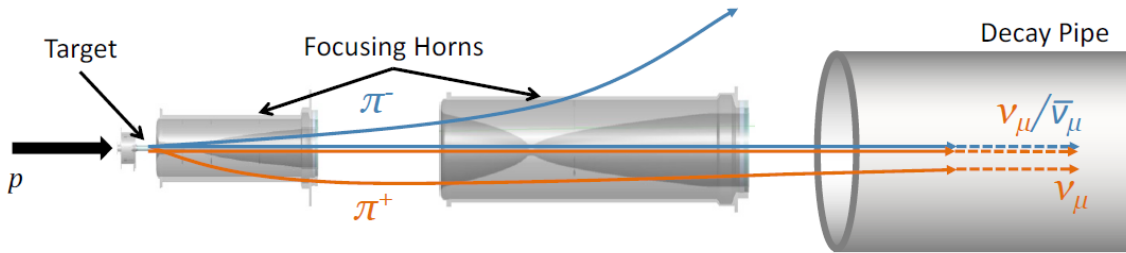


$$\sigma = 14.0 \pm 0.9(\text{stat.}) \pm 2.1(\text{syst.}) \times 10^{-40} \text{cm}^2/\text{nucleus}$$



NuMI Beam

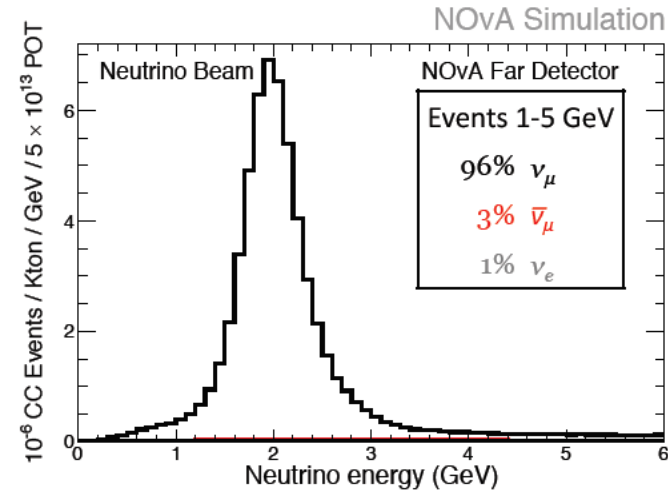
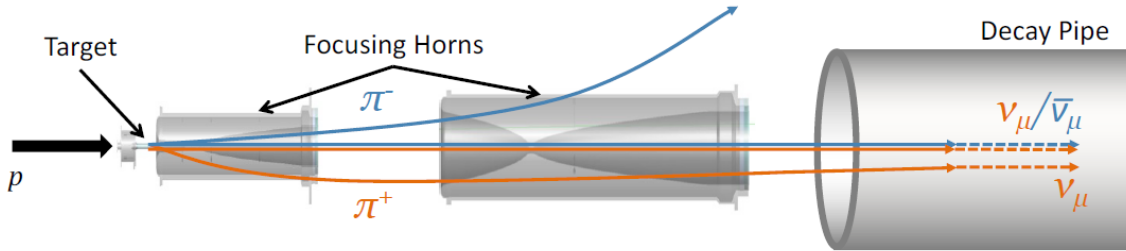
Neutrino Mode



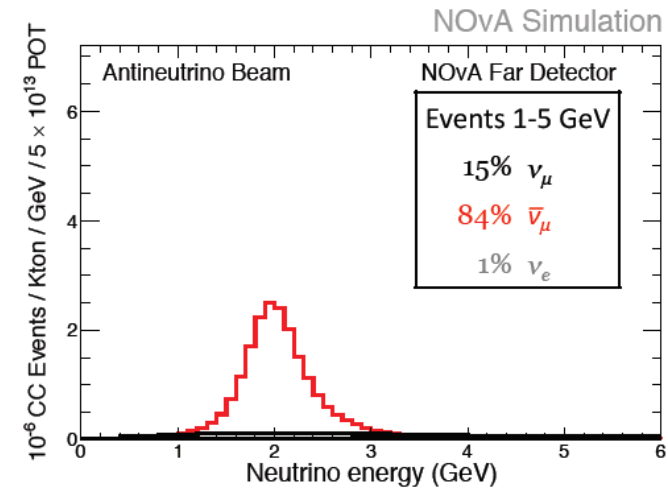
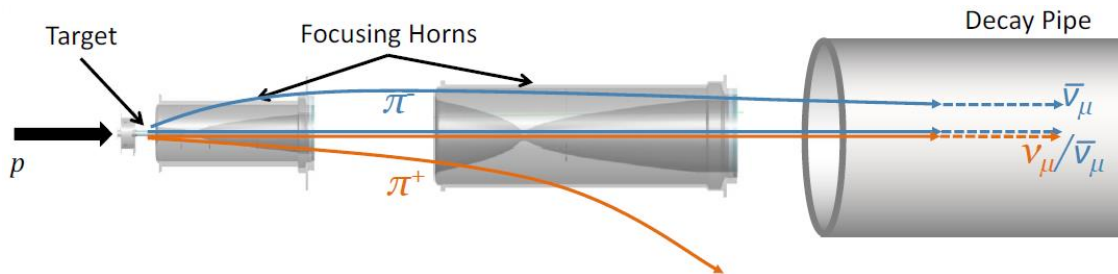


NuMI Beam

Neutrino Mode



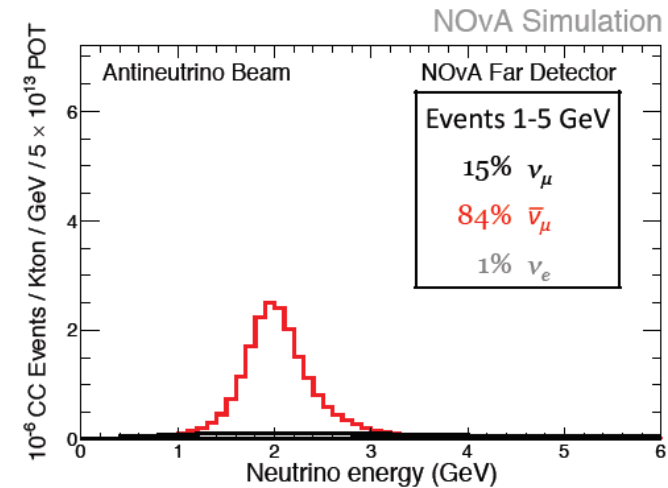
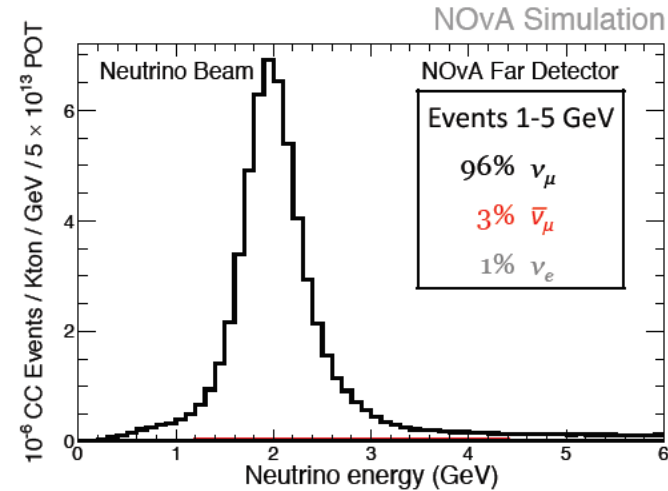
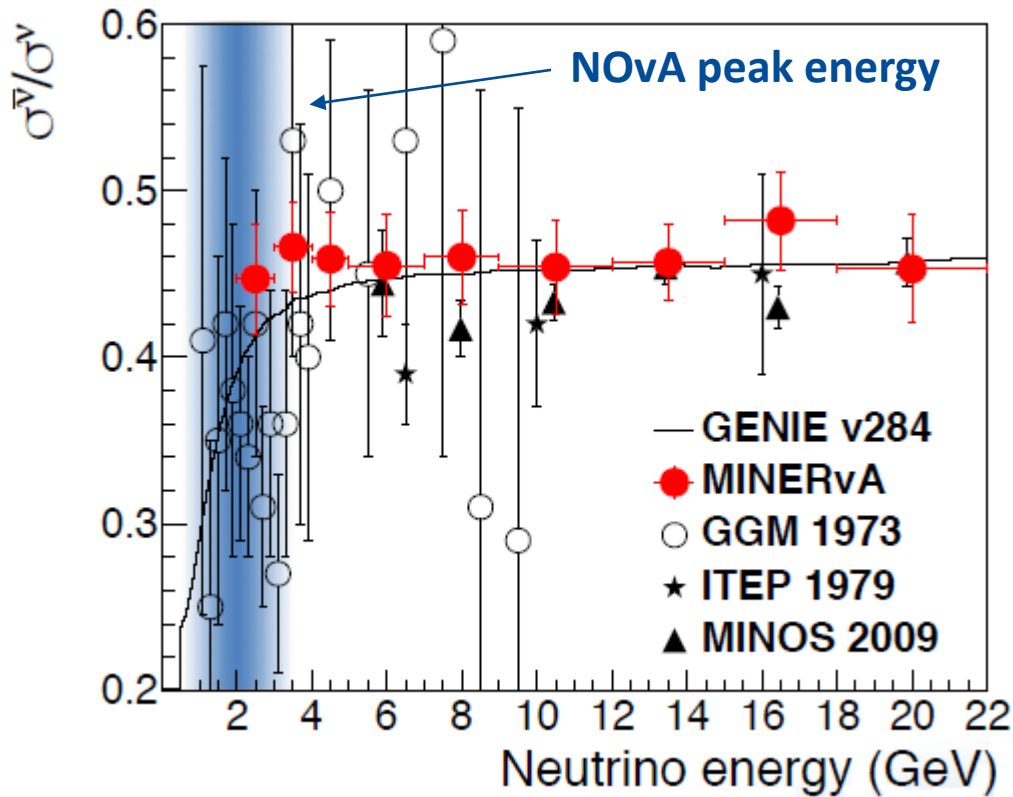
Antineutrino Mode





NuMI Beam

For NOvA energies, the antineutrino cross-section is ~ 2.8 times lower than the neutrino

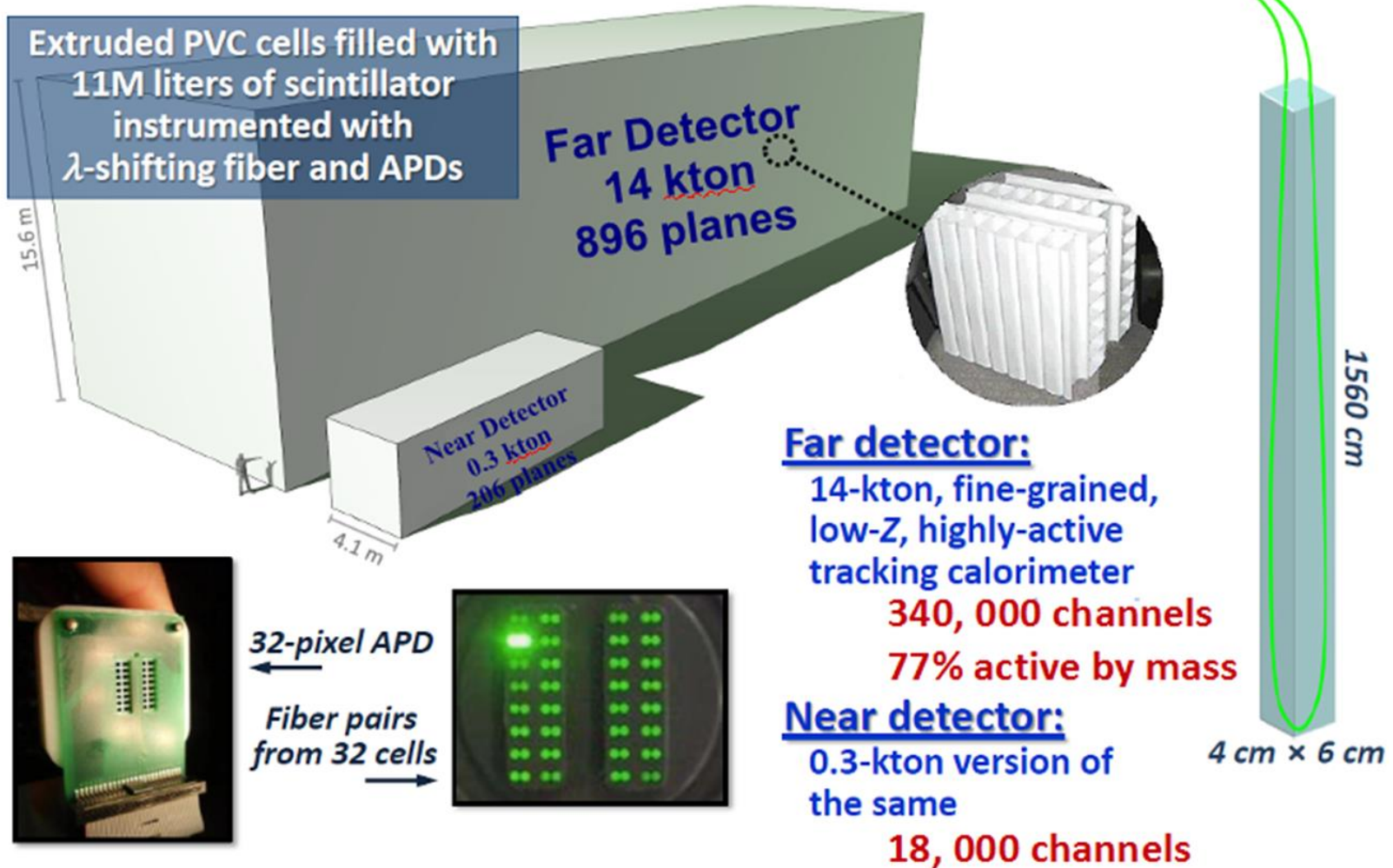




Two Detector Experiment

A NOvA cell

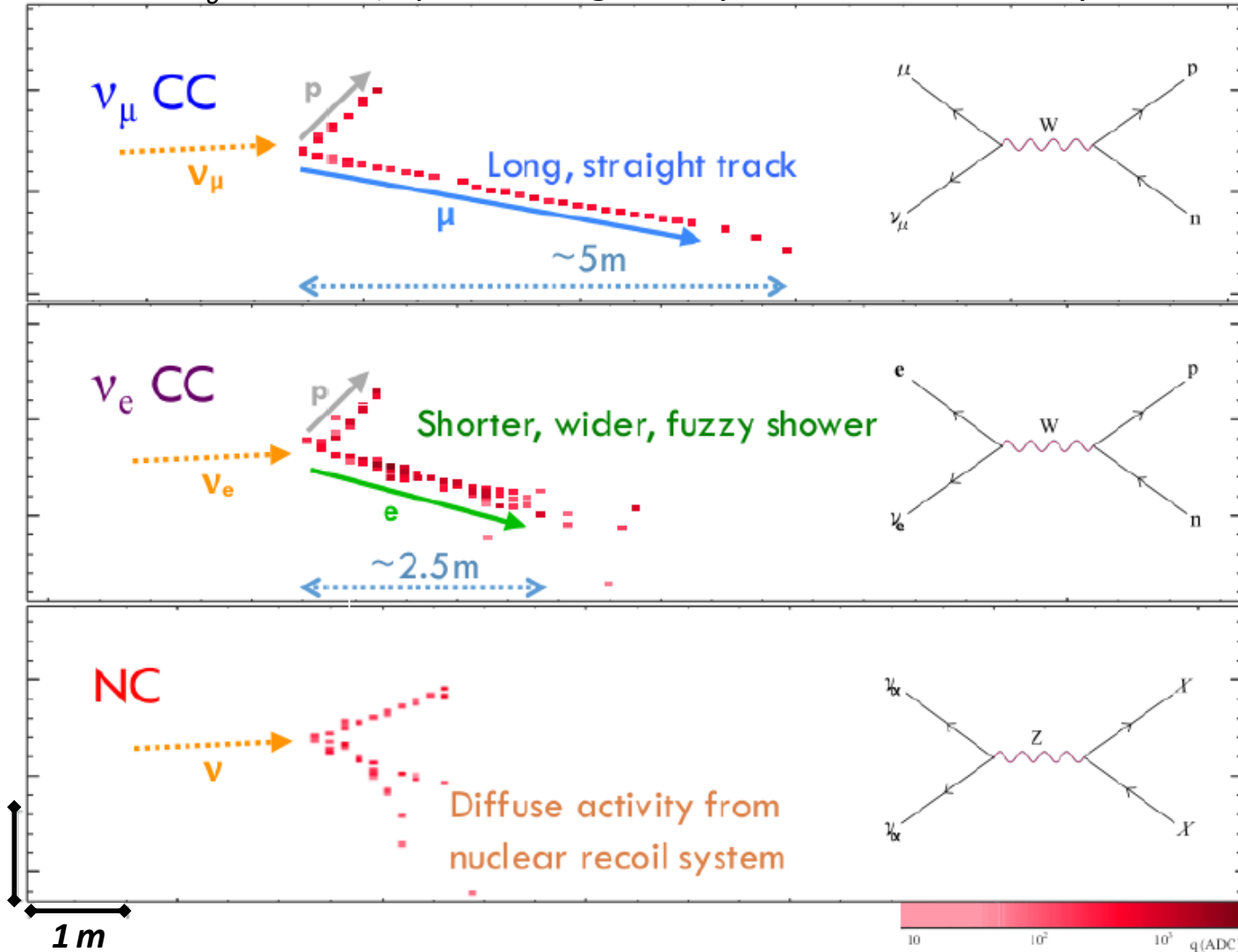
To APD





Event Topologies

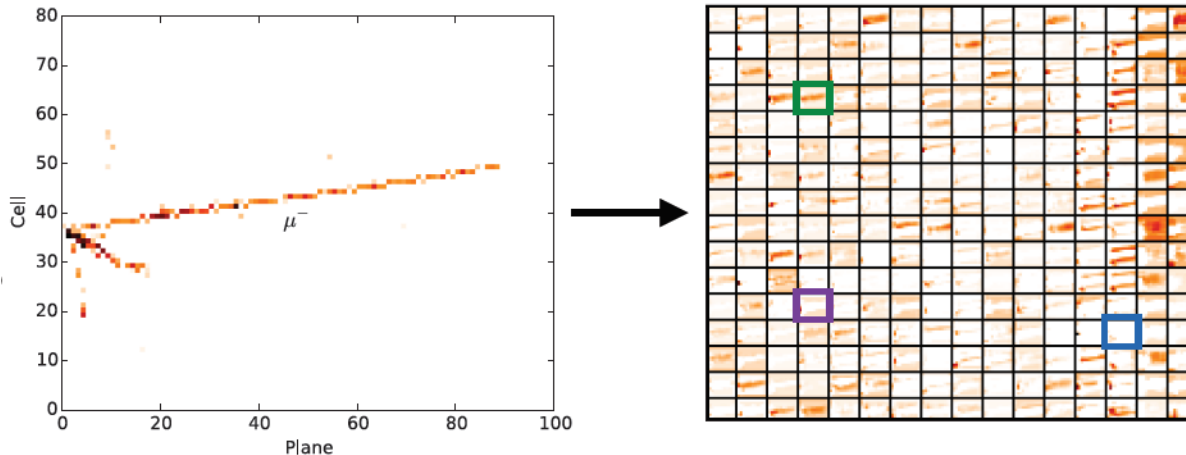
$X_0 = 38$ cm (6 planes longitudinally, 10 cells transversely)



Events classified with Convolutional Visual Network (CVN)

A deep learning algorithm similar to image recognition software
 Cells are like pixels and the energy depositions are like colors
 Filters pick out event features

JINST 11 (2016) P09001



One algorithm classifies:

1) ν_{μ} CC

2) ν_e CC

3) NC

4) Cosmic



More selection criteria

Data Quality, preselection, containment, and cosmic rejection

e.g. min number hits, directionality cuts, distance from detector edges, etc

Reduces cosmic background from $\sim 10^6$ events to less than 10 events (CVN does a lot too)

Analysis specific

ν_μ disappearance analysis

Additional muon track ID based on kNN (track length, dE/dx, scattering, quality)

Events separated into 4 subsamples by hadronic energy fraction

Lower had E fraction \rightarrow fraction has better energy resolution and purity

ν_e appearance analysis

Reclaim events that fail primary (core) cuts by reexamining with alternate cosmic cuts

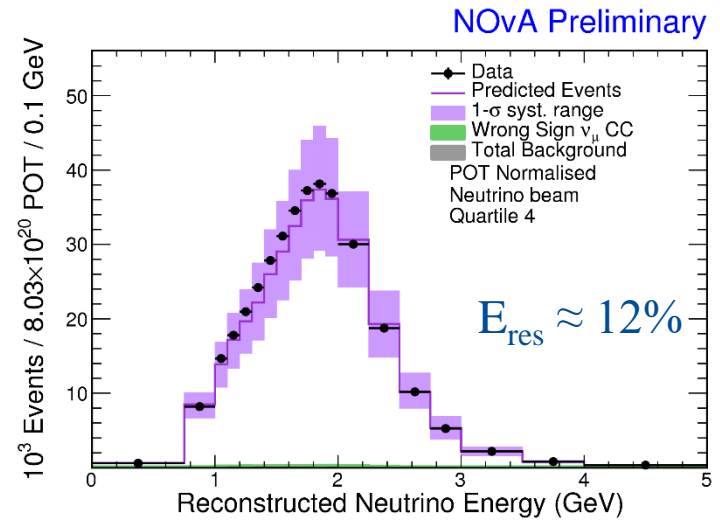
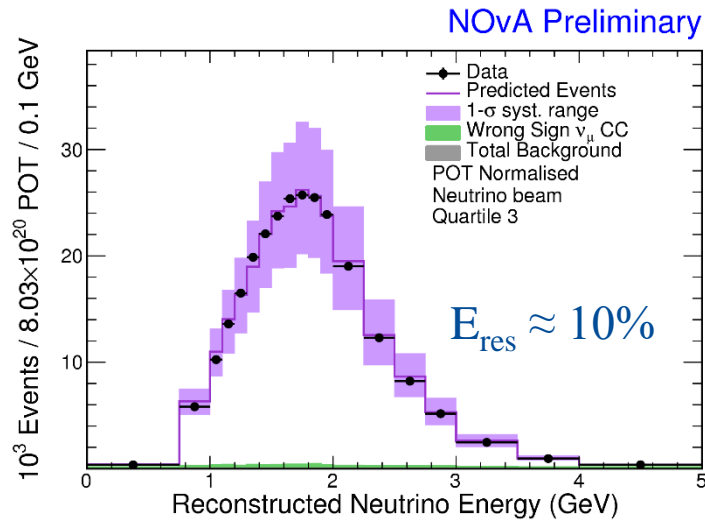
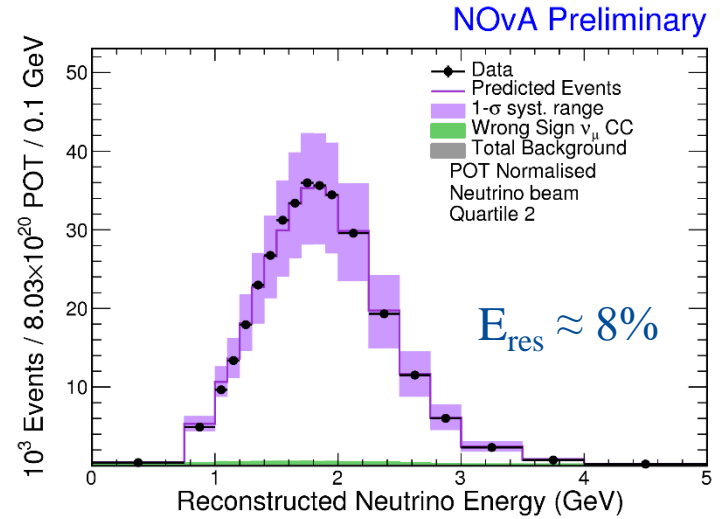
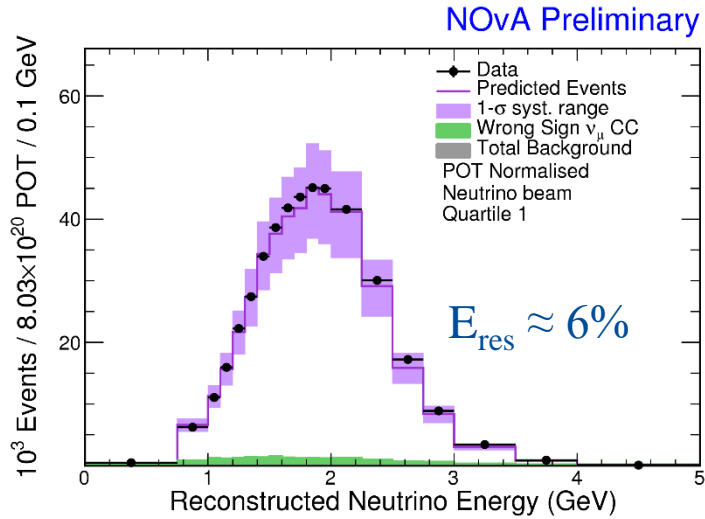
Forms a peripheral sample

Separate core events into 2 subsamples by CVN value

Higher CVN \rightarrow better purity

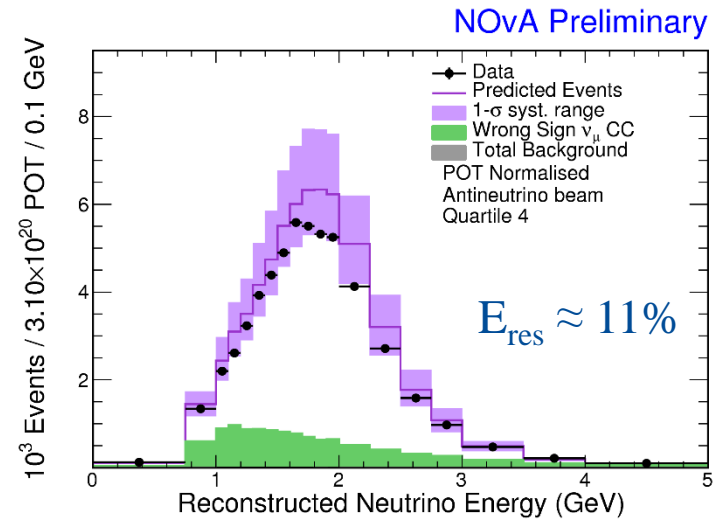
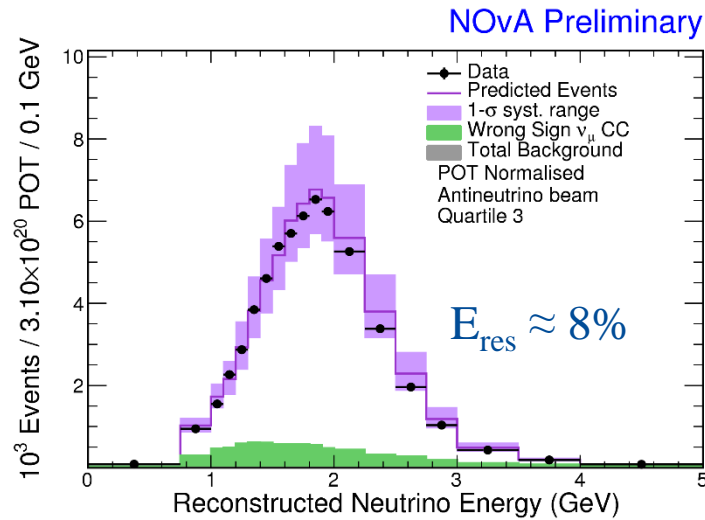
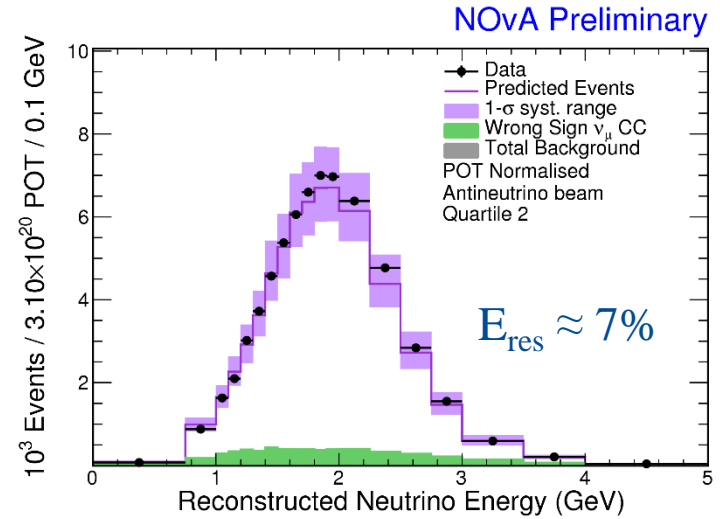
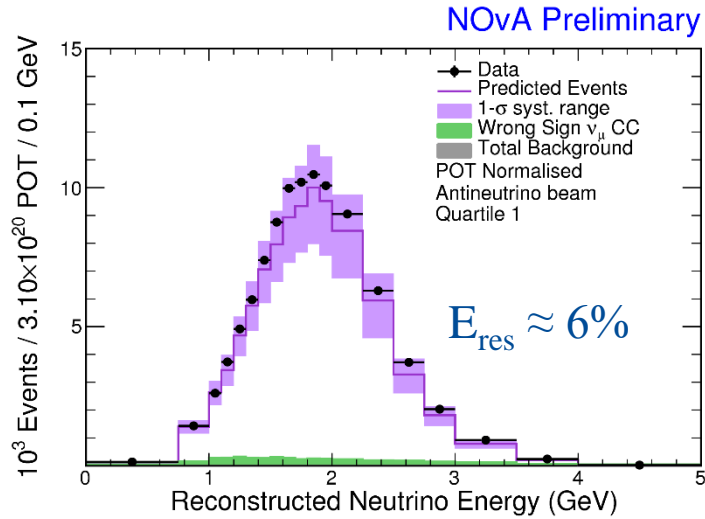


ND Selection - ν_μ CC



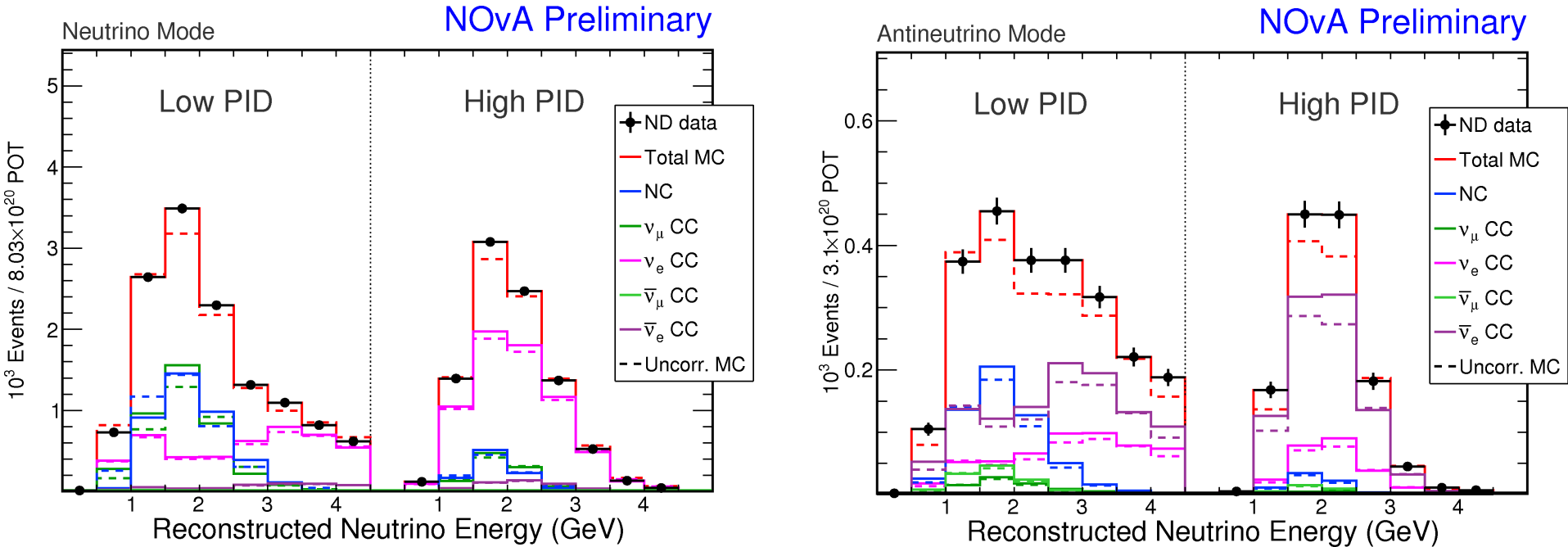


ND Selection - $\bar{\nu}_\mu$ CC





ND Selection - $\nu_e + \bar{\nu}_e$ CC

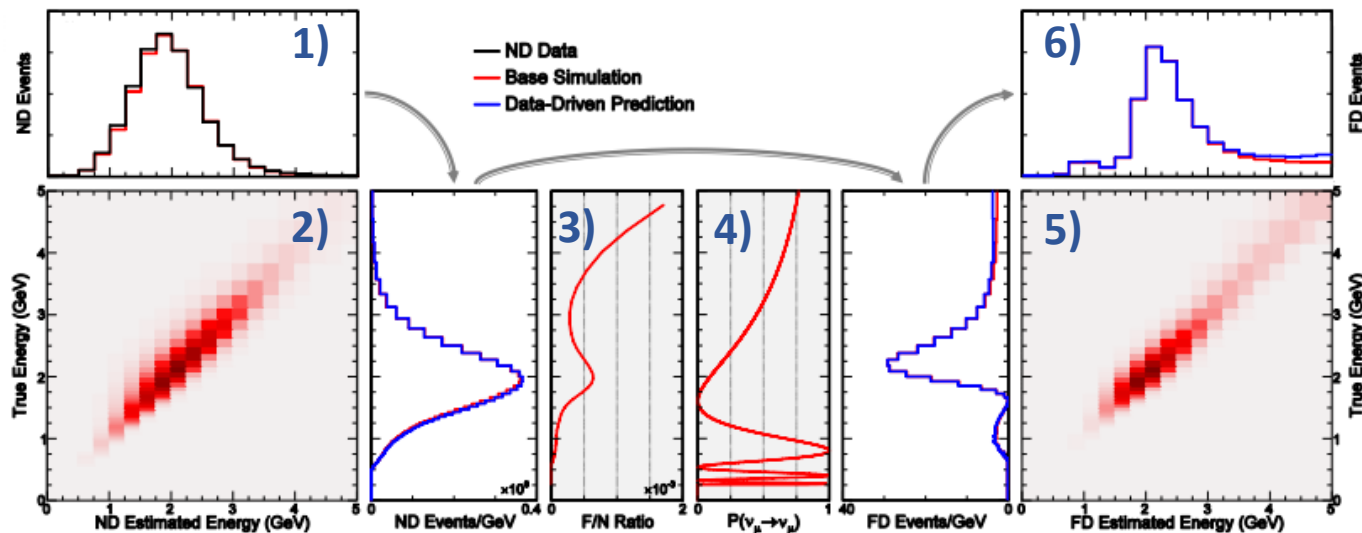


Near detector sample is Far Detector appearance background sample

ν_μ -CC, NC, and intrinsic beam ν_e -CC extrapolate differently to FD

Extrapolation

- 1) Select events in ND (use data)
- 2) Map ND reco E to true E (use simulation)
- 3) Apply ratio of FD events to ND events in bins of true E (use simulation)
Takes into account differences between two detectors
- 4) Apply oscillation probability on FD true E events (use simulation)
- 5) Map FD true E to reco E (use simulation)
- 6) Oscillated FD prediction



Don't need to separately measure flux, cross-section, efficiencies, etc in ND

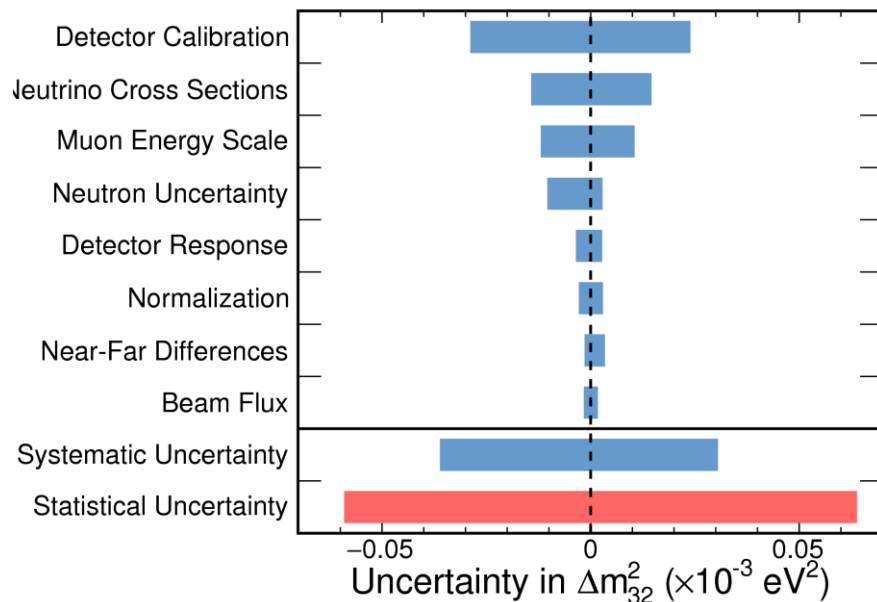
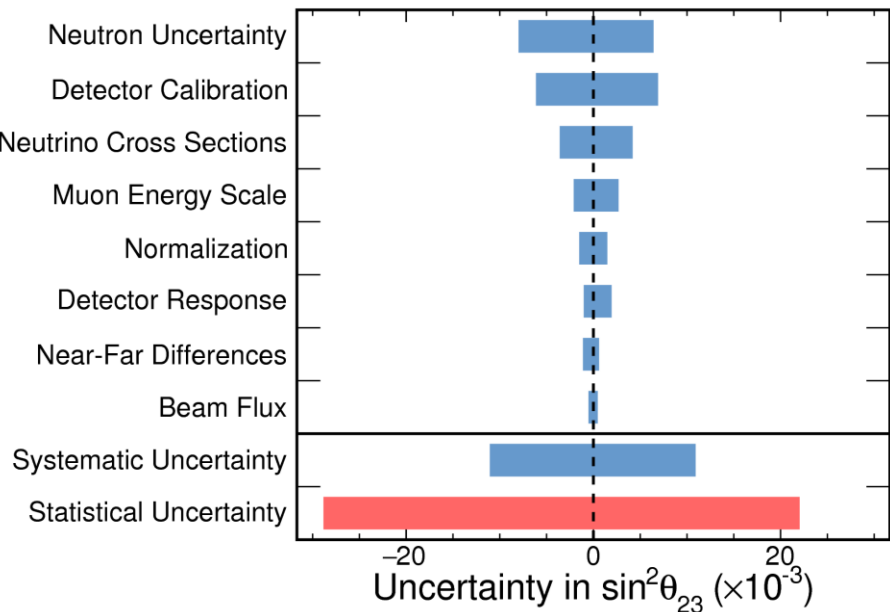
Systematics accounted for by altering simulation at steps 2, 3, 4, and 5



Systematics - $\nu_{\mu} + \bar{\nu}_{\mu}$ CC

NOvA Preliminary

NOvA Preliminary

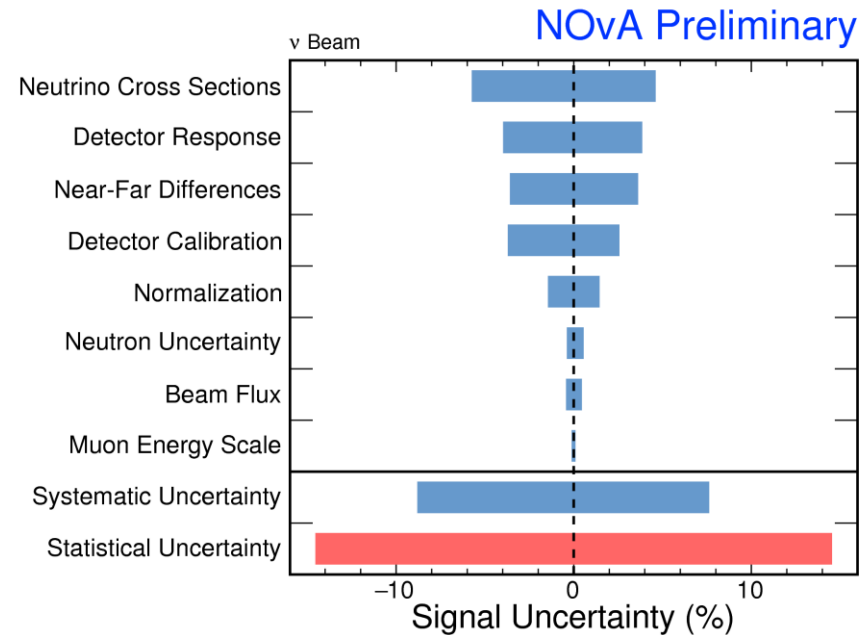
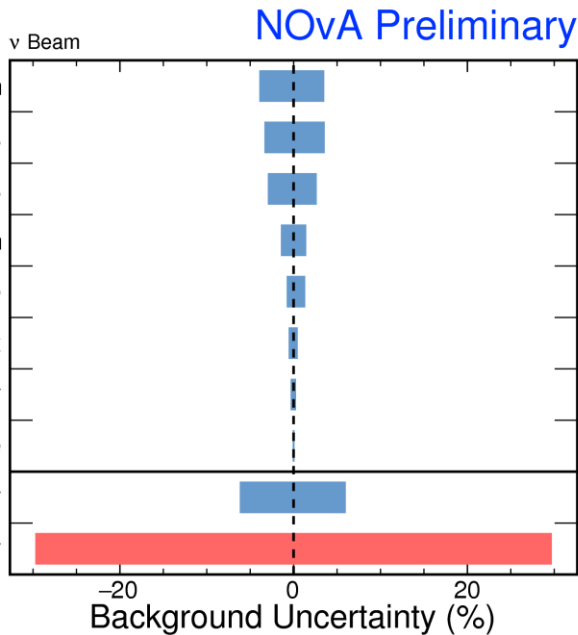


Statistics Limited

Dominant systematics related to neutron modeling, calorimetric energy calibration, and cross sections



Systematics - ν_e CC



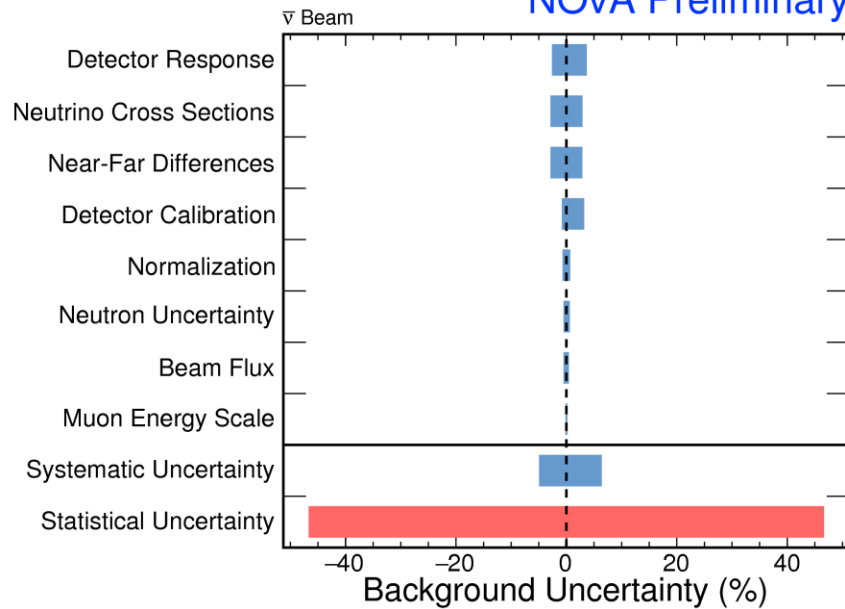
Statistics Limited

Dominant systematics related to cross sections and calorimetric energy calibration

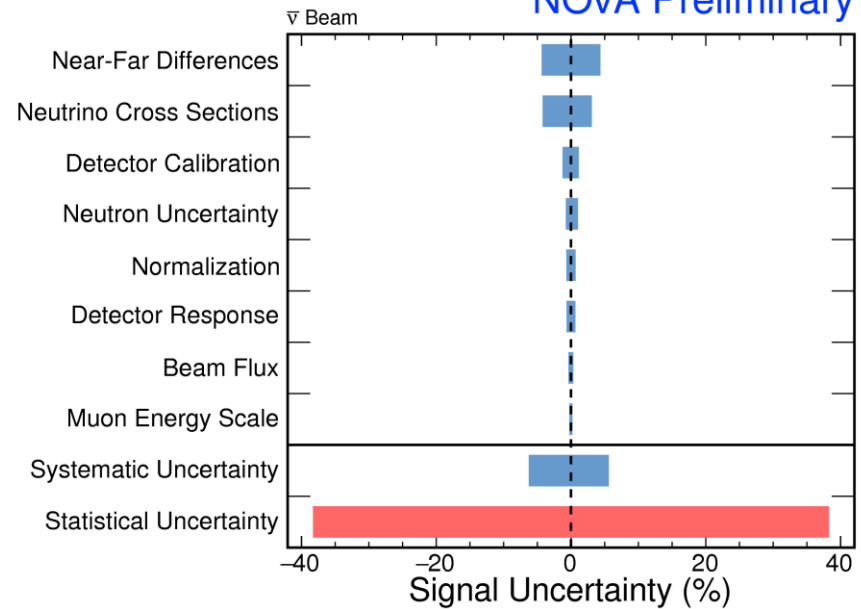


Systematics - $\bar{\nu}_e$ CC

NOvA Preliminary



NOvA Preliminary



Statistics Limited

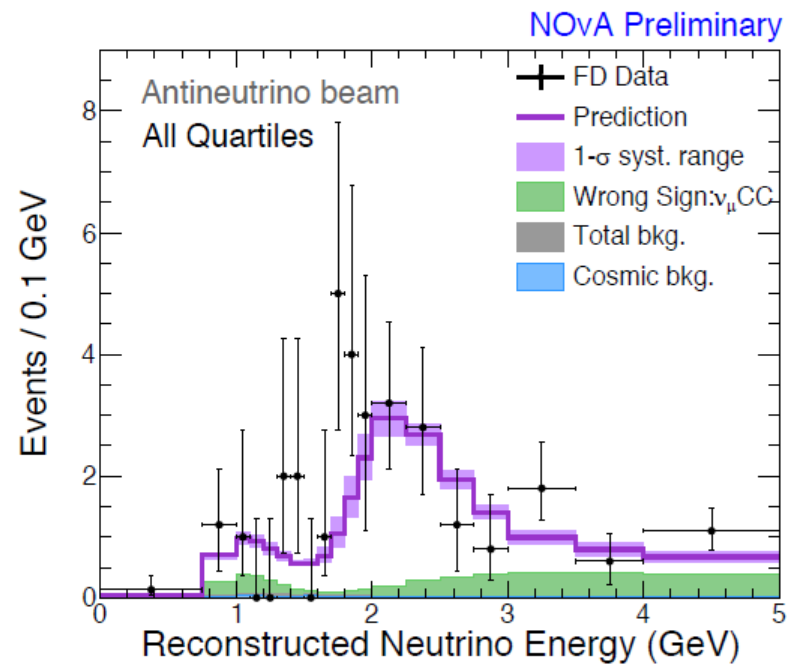
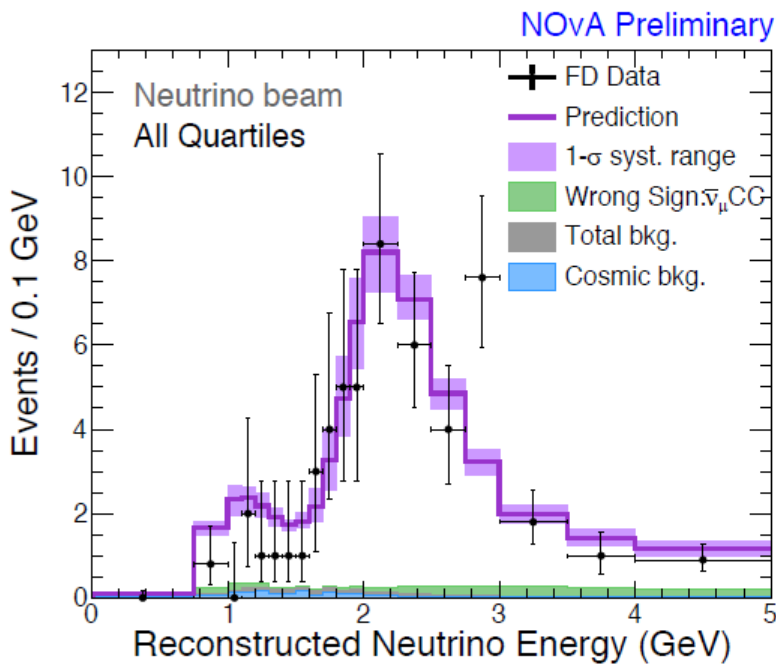


Results



ν_{μ} CC Disappearance Results

	Total Observed	Best Fit Prediction	Total Background	Cosmic Background	Beam Background
Neutrino Mode	113	121	3.3	2.1	1.2
Antineutrino Mode	65	50	1.1	0.5	0.6

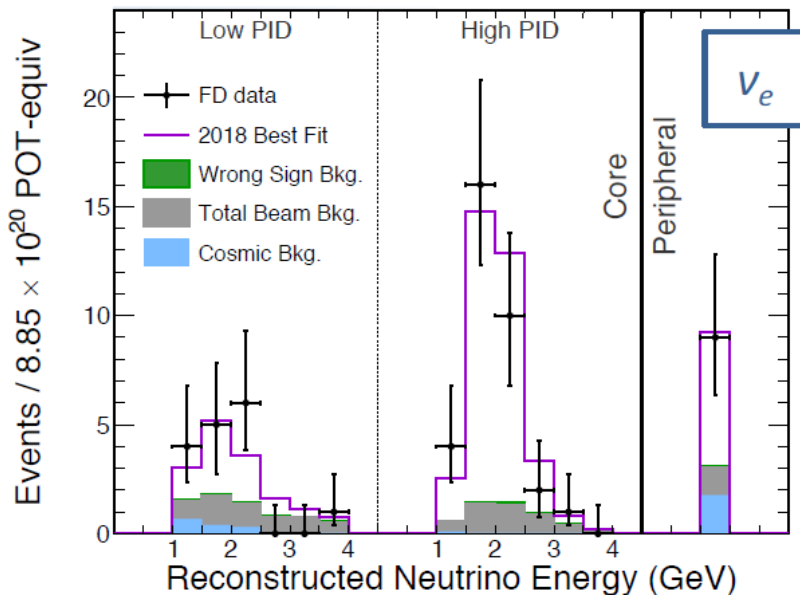




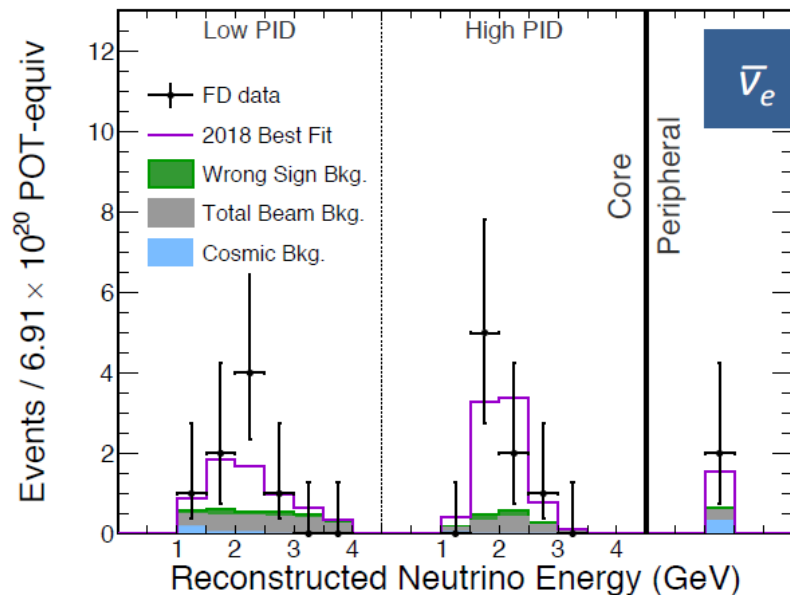
ν_e CC Appearance Results

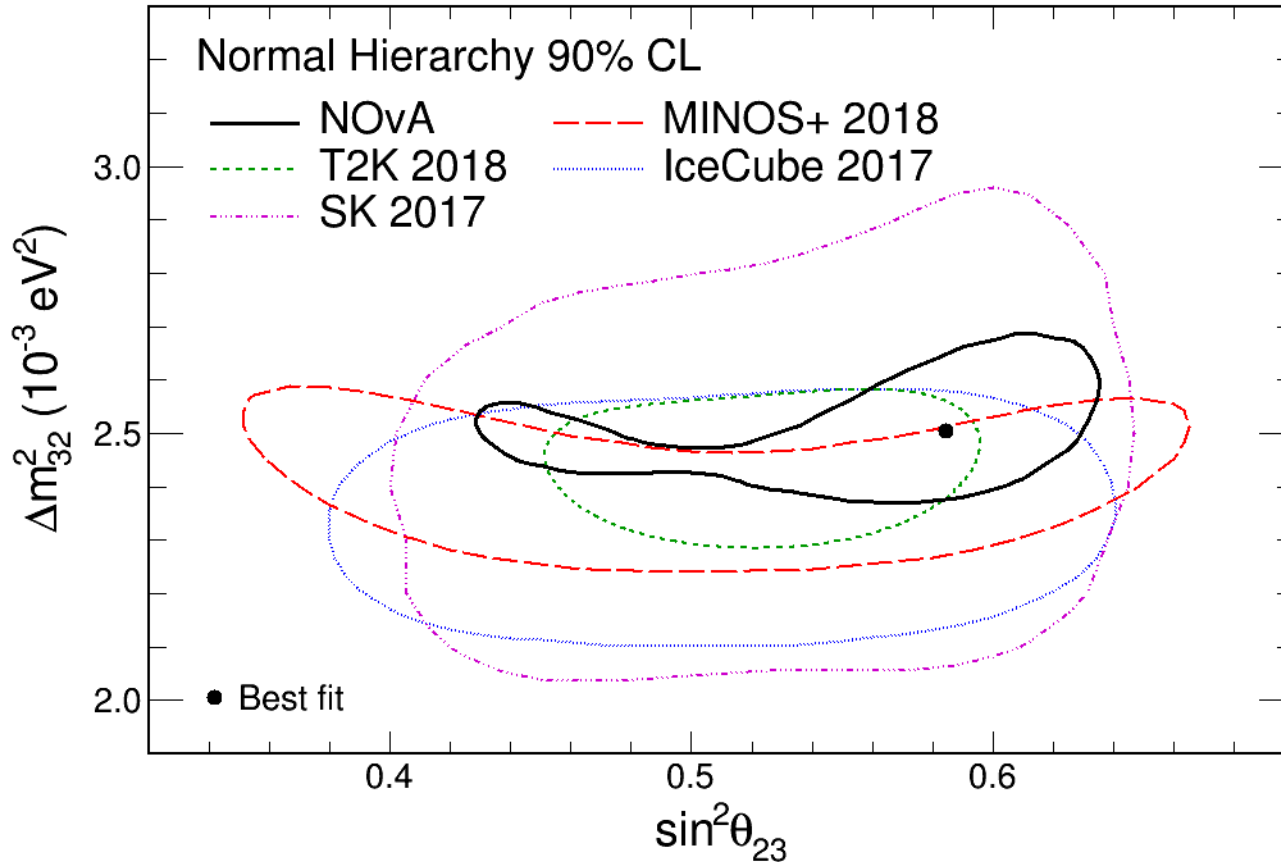
	Total Observed	Best Fit Prediction	Wrong Sign	Cosmic Background	Beam Background
Neutrino Mode	58	59	0.7	3.3	11.1
Antineutrino Mode	18	15.9	1.1	0.7	3.5

NOvA Preliminary



NOvA Preliminary





Best

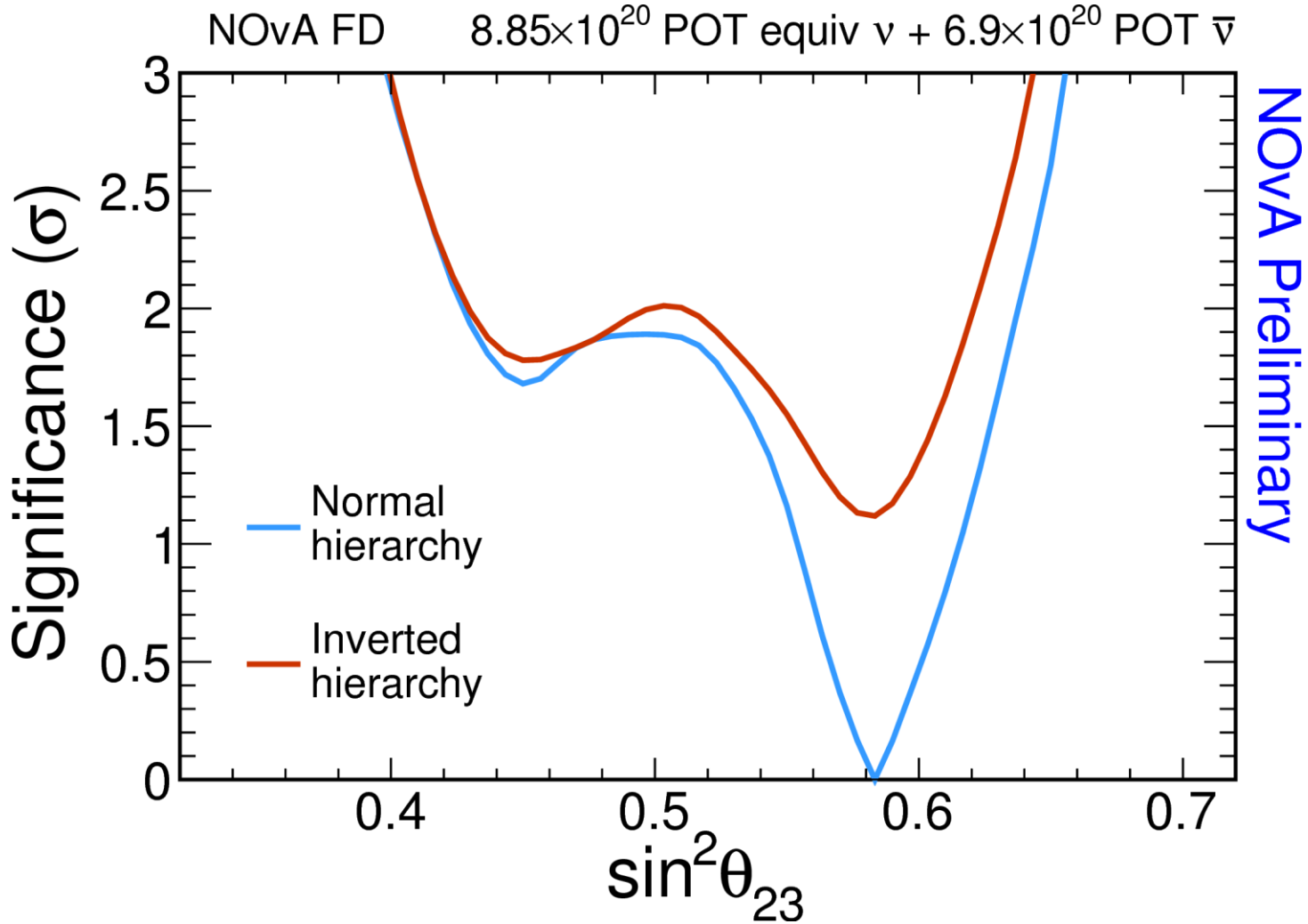
$$\Delta m_{32}^2 = 2.51^{+0.12}_{-0.08} \times 10^{-3} \text{ eV}^2$$

Fit

$$\sin^2(\theta_{23}) = 0.58 \pm 0.03$$



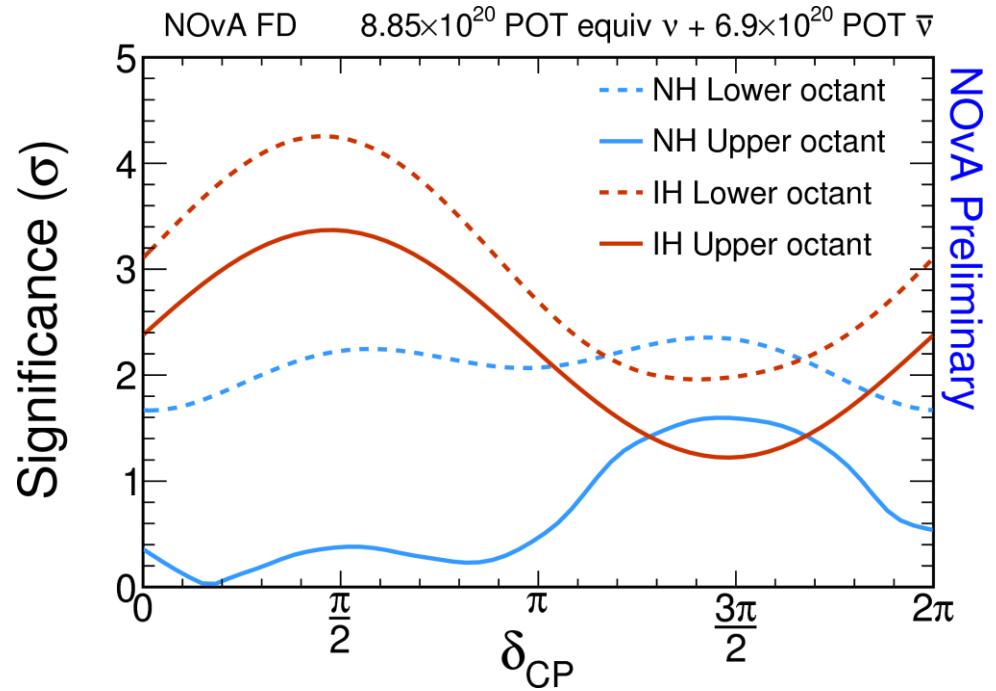
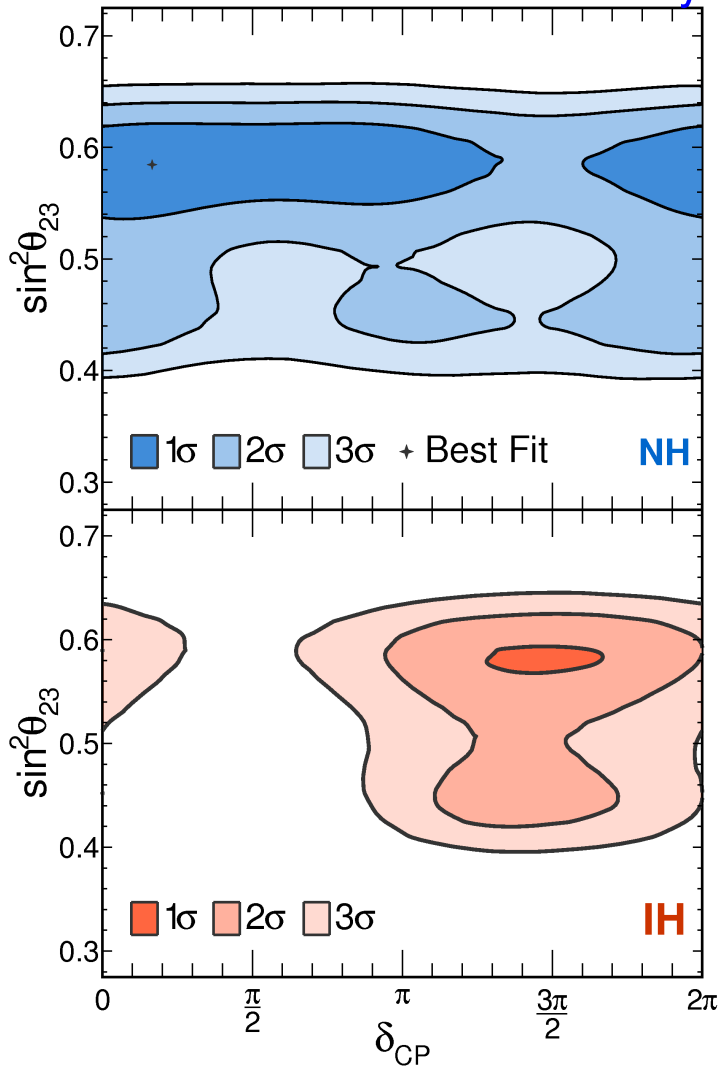
Joint Results



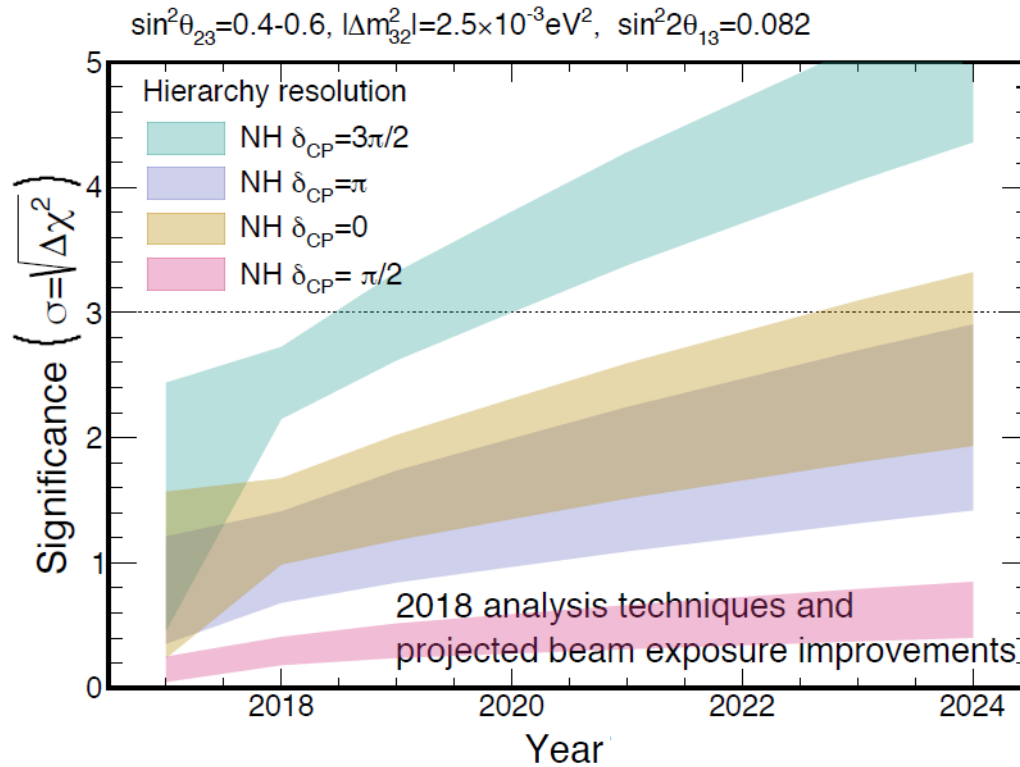


Joint Results

NOvA Preliminary



Prefer Normal Hierarchy at 1.8σ



3σ sensitivity to the hierarchy possible in 2020 with favorable parameters
 3σ sensitivity for 30-50% of δ_{CP} range by 2024

2σ sensitivity to CP violation in 2024 for favorable parameters

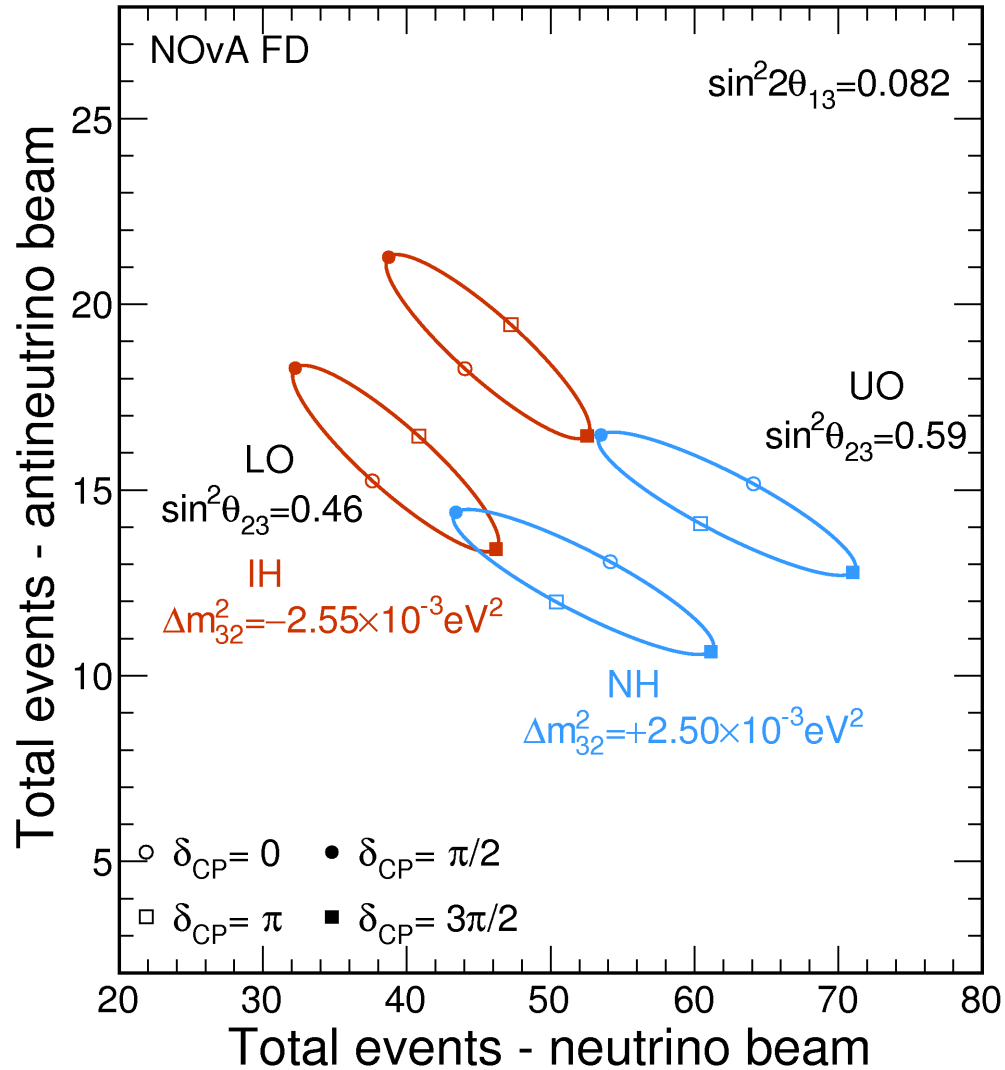
Thank You





Appearance Probability

NOvA Simulation





Appearance Probability

NOvA Preliminary

