ANALYSIS TECHNIQUES AND SYSTEMATIC UNCERTAINTIES FOR HYPER-K

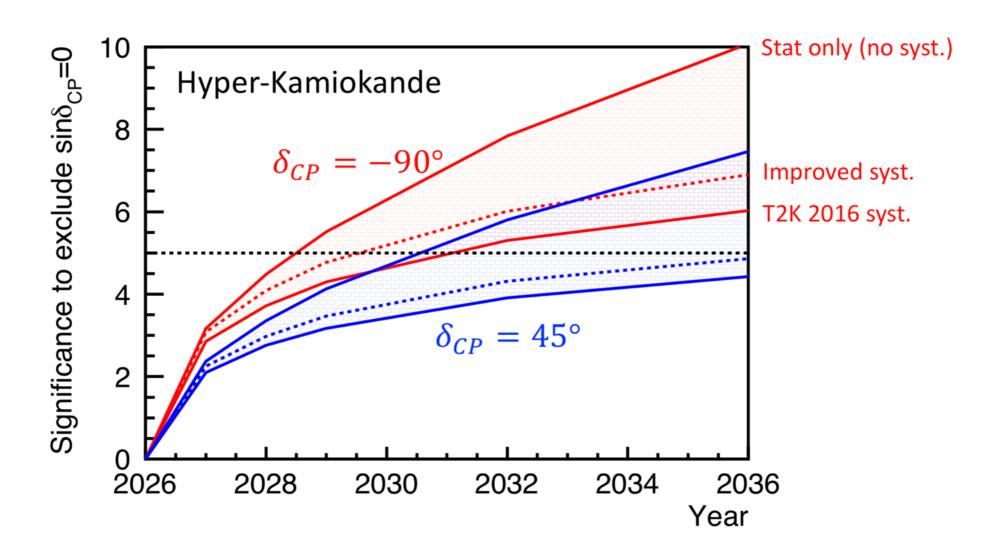
Mark Hartz TRIUMF & Kavli IPMU

NNN 2018 Hyper-K Satellite Meeting Oct. 21, 2018

MAXIMIZING THE SENSITIVITY OF HYPER-K

- ➤ Hyper-K: unprecedented sensitivity for a range of physics measurements
- ➤ To take full advantage of the Hyper-K detector:
 - ➤ Control systematic errors with near/intermediate detectors, external measurements, calibration
 - > Optimize analysis tools and methods to maximize signal and control backgrounds
- ➤ In this talk I will cover:
 - ➤ Physics topics that motivate systematic error reduction and analysis improvements
 - Controlling systematic errors with near detector measurements
 - Overview of analysis tools and areas for improvement

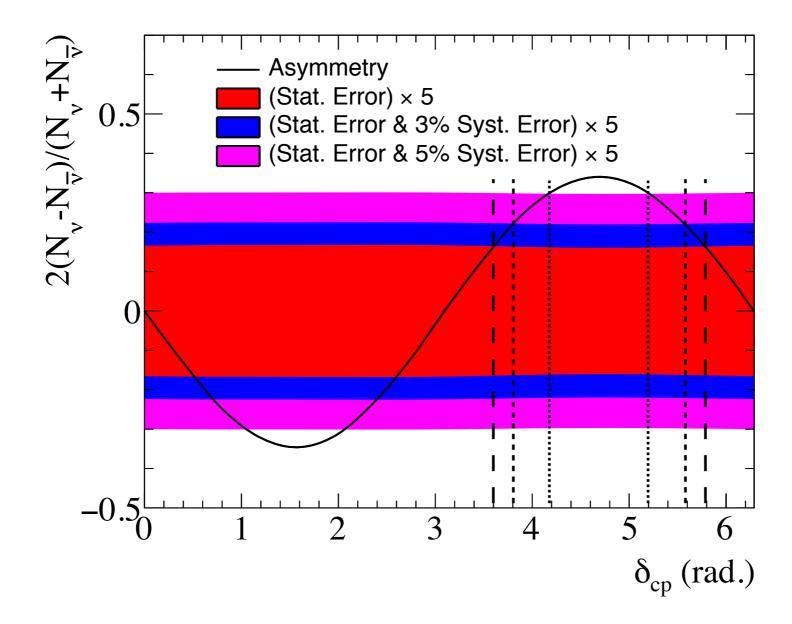
IMPACT OF SYSTEMATIC ERRORS ON CPV MEASUREMENT



- ➤ The sensitivity for the CP violation discovery strongly depends on the systematic error level that can be achieved
- ightharpoonup T2K has controlled errors to $\sim 6\%$
- ➤ We need reduction of the systematic errors for Hyper-K

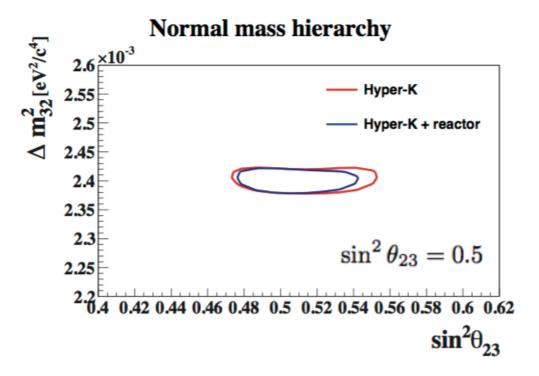
HYPER-K PHYSICS - CP VIOLATION

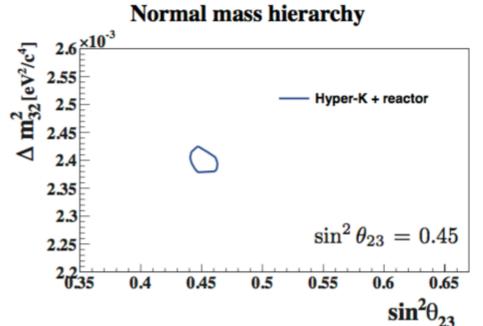
- ➤ Hyper-K collects ~2000 neutrino and antineutrino candidates
- \sim 3% statistical error on the relative number of neutrino and antineutrino candidates



➤ Controlling systematic errors to the 3% level or better is necessary or sensitivity is significantly degraded

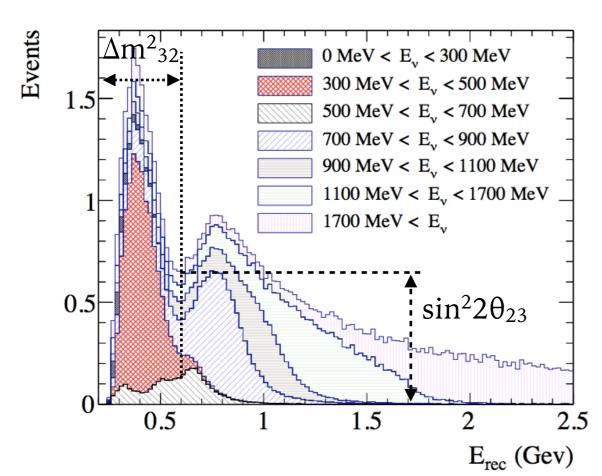
HYPER-K PHYSICS - LBL DISAPPEARANCE CHANNEL





Hyper-K Design Report: arXiv:1805.04163

► We aim for a 0.6% error on Δm^2_{32} and as low as a 1.3% error on $\sin^2\theta_{23}$ parameters



- ► Determine the position of the oscillation dip with 0.5% accuracy for Δm^2_{32} measurement
- Accurately predict the feed-down contributions into the dip region for sin²θ₂₃ measurement

NEUTRON DETECTION

➤ Neutron detection is used in a broad range of physics in Hyper-K, including high and low energy signals

Proton decay:

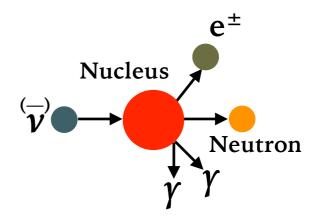
- ➤ Reject atmospheric background neutrino events
- ► 70% tagging efficiency = 70% rejection rate in $p\rightarrow e^+\pi^0$ channel

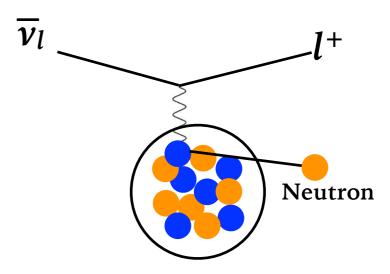
Atmospheric neutrinos:

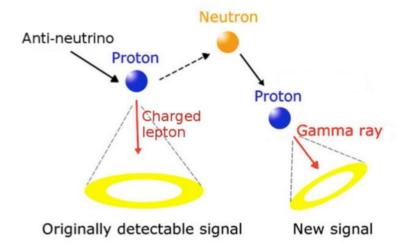
➤ Statistical separation of neutrinos and antineutrinos is improved by neutron tagging of antineutrino candidates

Supernova relic neutrinos:

➤ Tag neutron to select inverse beta decay events and reject solar neutrino backgrounds



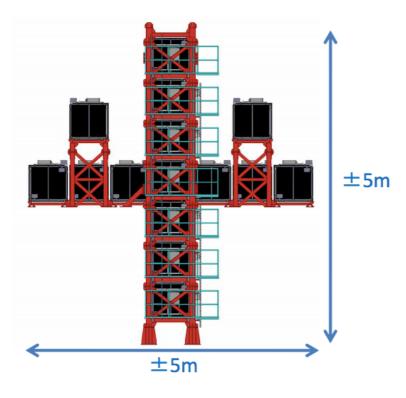




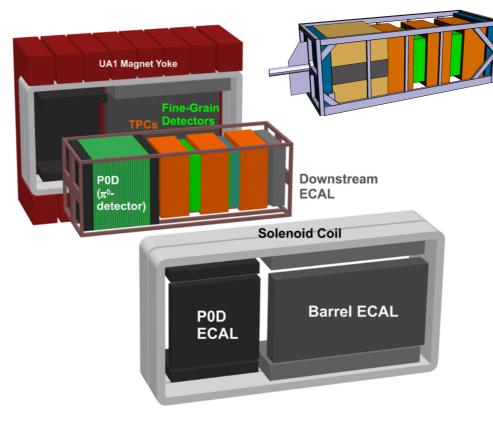
SYSTEMATIC ERRORS: NEAR AND INTERMEDIATE DETECTORS

BASELINE DESIGN FOR NEAR/INTERMEDIATE DETECTORS

On-axis Detector (INGRID)



Off-axis Magnetized Tracker (ND280→ND280 Upgrade→??)



Off-axis spanning intermediate water Cherenkov detector (IWCD)



- ➤ On-axis detector: measure beam direction, monitor event rate
- ➤ Off-axis magnetized tracker: charge separation (measurement of wrong-sign background), study of recoil system
 - ➤ Expect upgrades of detector inherited from T2K will be necessary
- ➤ Off-axis spanning water Cherenkov detector: intrinsic backgrounds, electron (anti)neutrino cross-sections, neutrino energy vs. observables, H₂O target, neutron multiplicity measurement

CONTROLLING BACKGROUND FOR CPV MEASUREMENT

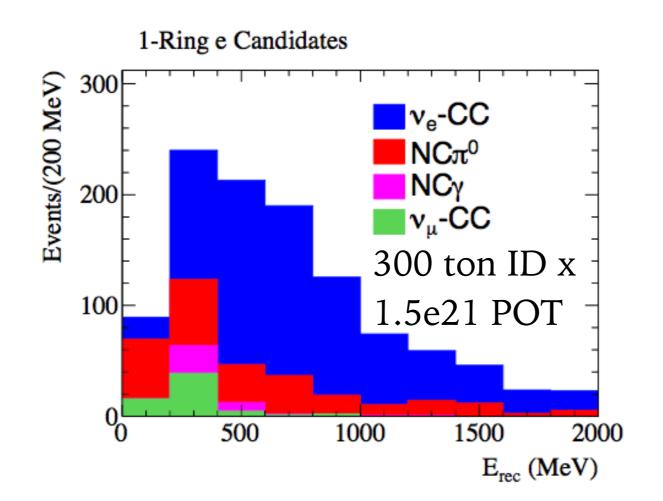
Sample Composition:

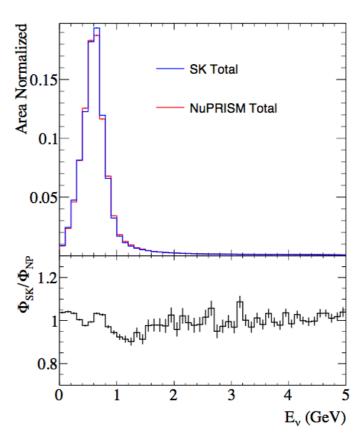
	Neutrino Candidates	Antineutrino Candidates
Signal	80%	62%
Wrong-sign Background	1%	11%
Intrinsic electron (anti)neutrino & NC	19%	27%

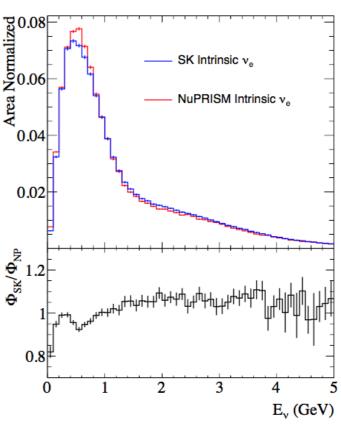
- ➤ We aim for a 1% systematic error contribution from the wrong-sign background and intrinsic electron (anti)neutrino & NC background respectively
- ➤ Wrong-sign background must be measured with 9% accuracy
 - ➤ Can be done in magnetized tracking detectors
- ➤ Intrinsic electron (anti)neutrino & NC background must be measured with 3% accuracy

NC AND INTRINSIC ELECTRON (ANTI)NEUTRINO BACKGROUND

- ➤ Total and electron (anti)neutrino fluxes are nearly identical at IWCD and far detector
- ➤ Can directly measure these backgrounds at the IWCD
- ➤ We can achieve 3% statistical precision with 6 m diameter inner detector, even better with larger diameter





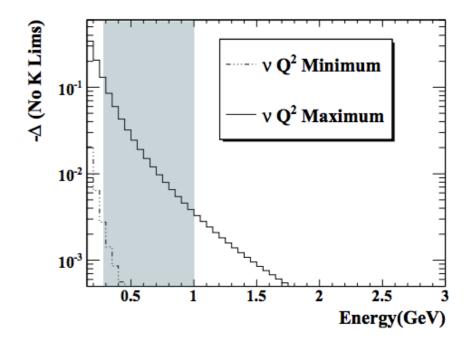


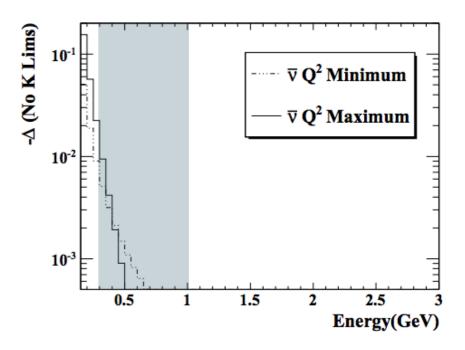
ELECTRON (ANTI)NEUTRINO CROSS SECTION

- ➤ Source of systematic error on the signal contribution from the electron (anti)neutrino cross section
- ➤ Uncertainties on the relative cross section ratios propagate directly into the analyses

$$\sigma_{v_e}/\sigma_{v_\mu}$$
 $\sigma_{ar{v_e}}/\sigma_{ar{v_\mu}}$

- Sources of errors are:
 - Phase space difference between interactions involving muons and electrons
 - ➤ Radiative corrections
 - ➤ Form factor uncertainties for sub-leading terms in cross section that depend on lepton mass
- ➤ We should make direct and indirect measurements to address these uncertainties





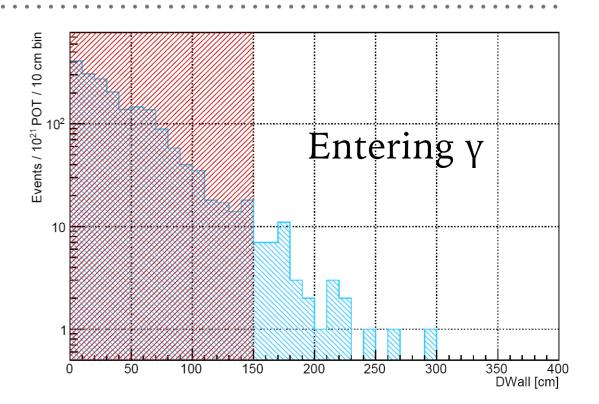
Phys.Rev. D86 (2012) 053003

ELECTRON (ANTI)NEUTRINO CROSS SECTION MEASUREMENT

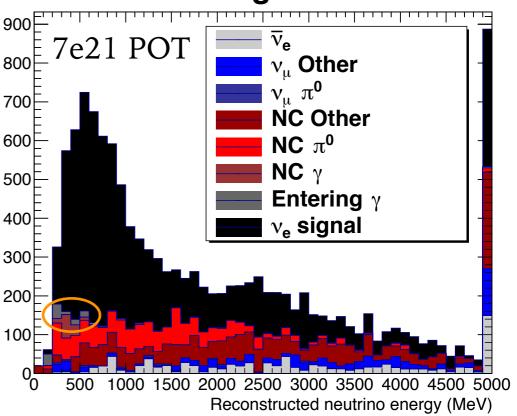


- ➤ Measure in the intermediate water Cherenkov detector
- ➤ The fraction of electron (anti)neutrinos increases at further off-axis angles

- The entering γ background is controlled by outer detector veto and fiducial cut
 (>1.5 m from wall)
- ➤ The cross section is measured relative to the muon neutrino candidates to produce a cross section ratio

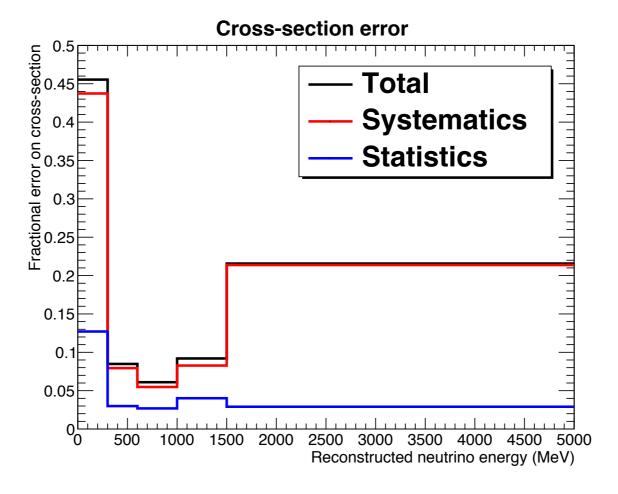


Selected 1-ring e-like events



UNCERTAINTY ON ELECTRON NEUTRINO CROSS SECTION

➤ The electron neutrino cross section relative to the muon neutrino cross section is extracted

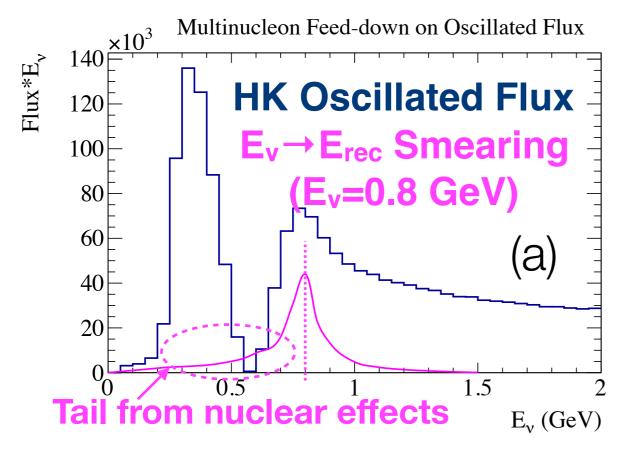


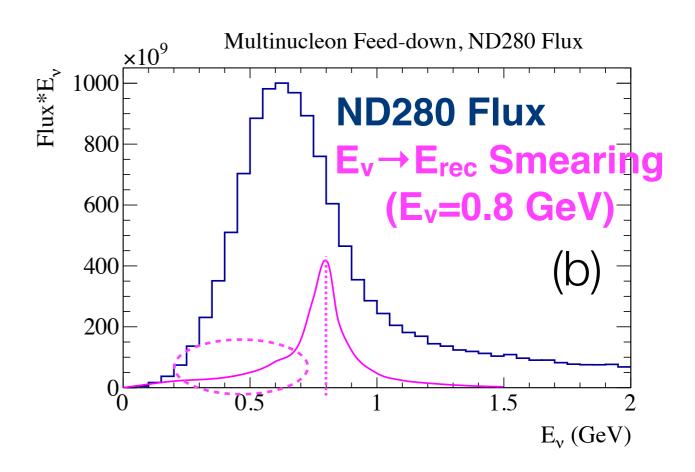
- > Statistical error of 3% can be achieved in relevant bins
- > Systematic error should be reduced with update to new flux errors and better use of control samples to constrain backgrounds
 - ➤ May need additional measurements to control relative flux error
- ➤ Analysis will be extended to electron (anti)neutrinos

SPECTRUM UNCERTAINTIES

- Previous discussion covered normalization type errors.
- Uncertainties that impact the shape of the reconstructed spectra are also important

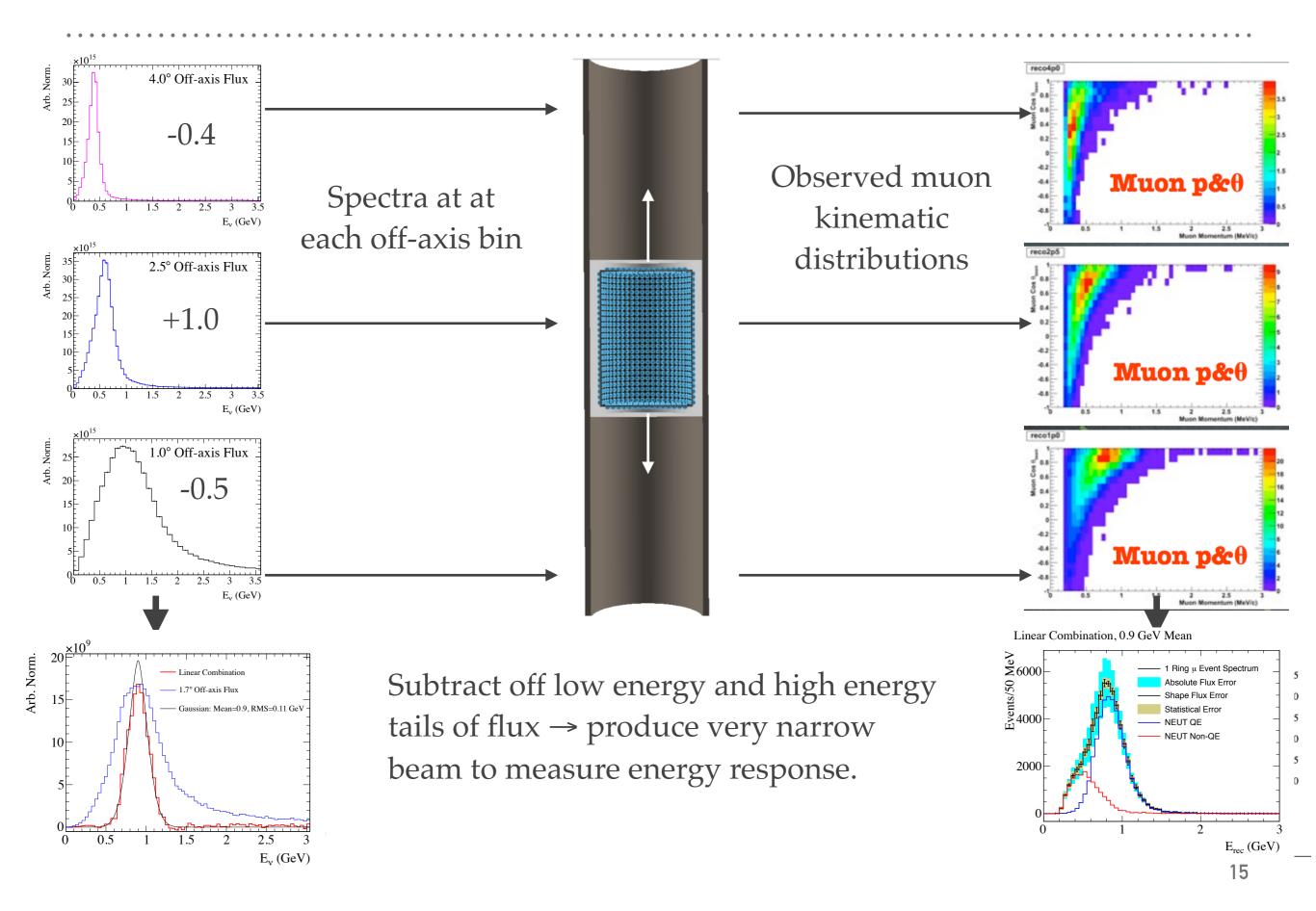
➤ Example:





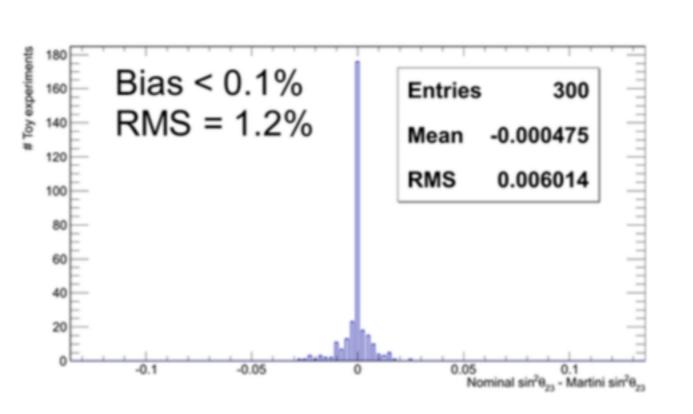
- ➤ Modeling the feed-down component is critical for disappearance measurement
- ➤ Difficult to measure in the near detector, where it resides under the flux peak

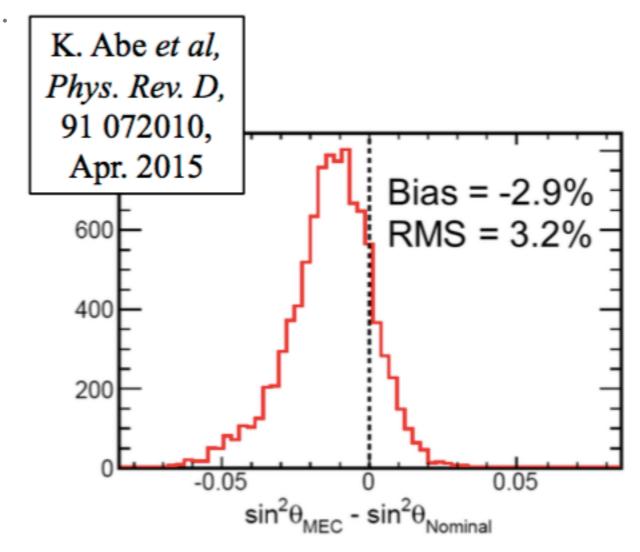
OFF-AXIS MEASUREMENTS IN IWCD



OFF-AXIS MEASUREMENTS IN IWCD

- ➤ Mock data analysis at T2K
 - ➤ Addition of multi-nucleon events to mock data
 - ➤ Analyze MC without multi-nucleon events
 - \triangleright Biased values of θ_{23} measured

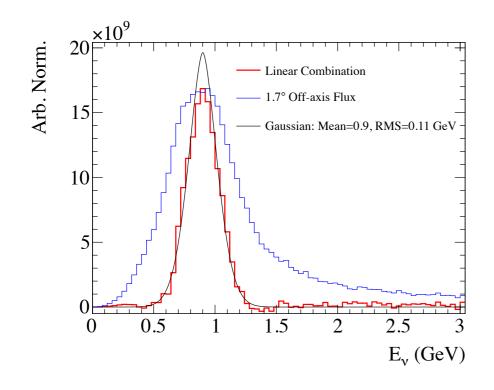


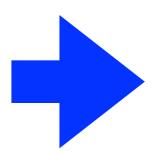


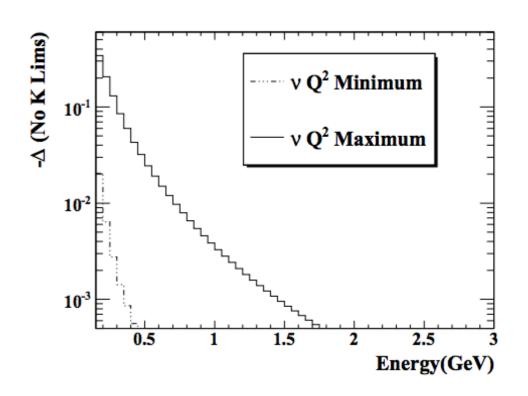
- ➤ Identical analysis with IWCD
 - Multi-nucleon events added to mock data
 - ➤ Not added to MC
 - Linear combination applied
 - ightharpoonup Measured θ_{23} unbiased

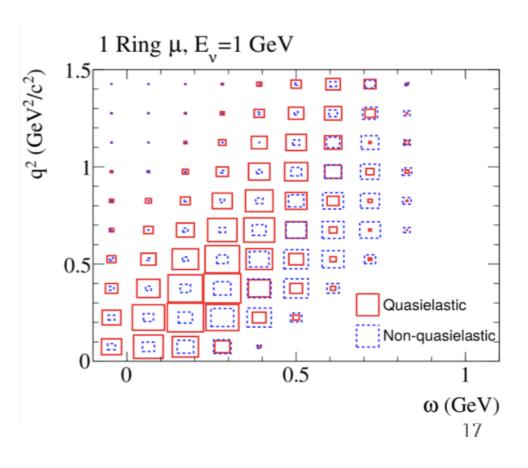
ANOTHER USE FOR NARROW BAND BEAMS

- Source of error on the relative v_e to v_μ cross section: phase space difference
- ➤ Use linear combination method to produce narrow band beam
- Can reconstruction the momentum and energy transfer
- ➤ Study the regions of phase space where the electron neutrino and muon neutrino differ

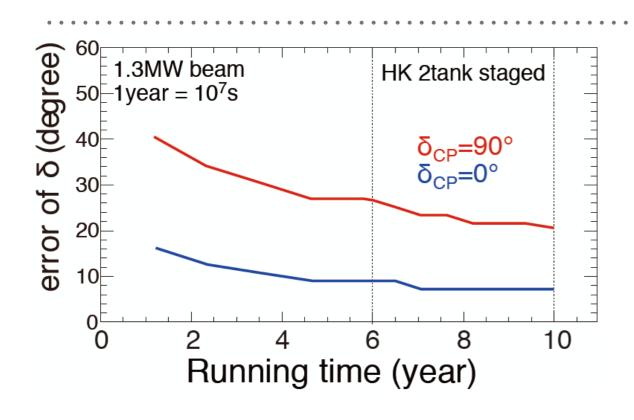




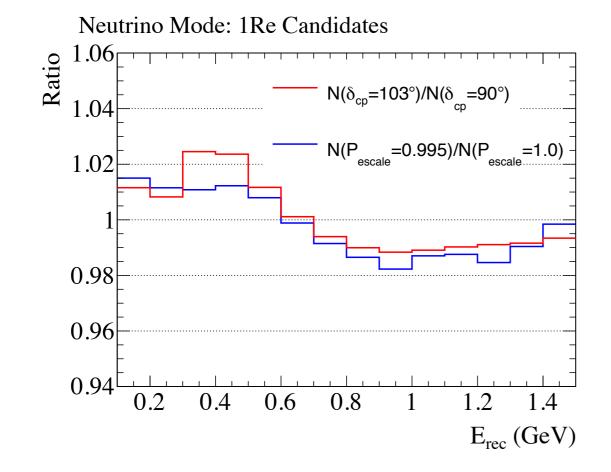




SHAPE MATTERS FOR CP PHASE TOO



- ► If δ_{cp} is near 90° or 270°, precision measurement becomes challenging
- Rely on spectrum variations introduced by $cos(\delta_{CD})$ term



- ightharpoonup A shift of δ_{cp} by 13° is nearly indistinguishable from a change of energy scale by 0.5%
- We also need precise calibration of the energy response in an our near/intermediate detectors and Hyper-K
- Similar precision is needed for Δm^2_{32} measurement

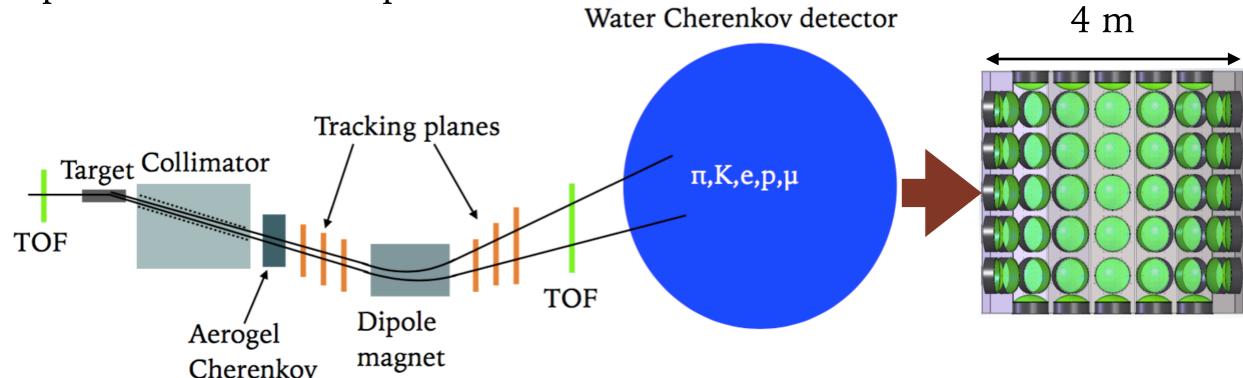
SOURCES OF SHAPE UNCERTAINTY

- ➤ Uncertainty on beam direction:
 - > Should be measured precisely by INGRID detector
- ➤ Uncertainty on feed-down contributions and removal energy from nuclear effects
 - ➤ Measure directly in IWCD
 - ➤ Study nuclear effects with hadronic recoil system measurements in magnetized tracker
- ➤ Uncertainty on energy scale in near/intermediate detectors and Hyper-K
 - ➤ Need precision energy scale calibration in all detectors
 - ➤ Work needed to improve on 2% energy scale uncertainty in T2K analyses

A TEST BEAM EXPERIMENT

- ➤ Calibrating water Cherenkov detectors with 1% level precision is necessary for Hyper-K
- ➤ Study detector response and performance of calibration systems in a test beam with known particle type and momenta

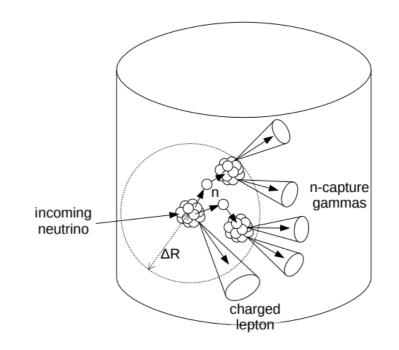
➤ Proposal is test beam experiment:



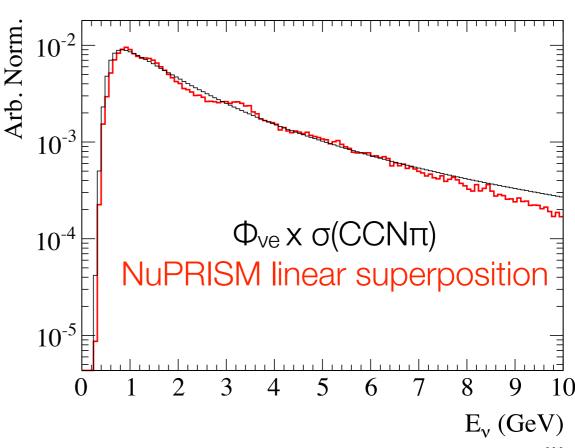
- Considering test beams at Fermilab and CERN
- Looking for collaborators

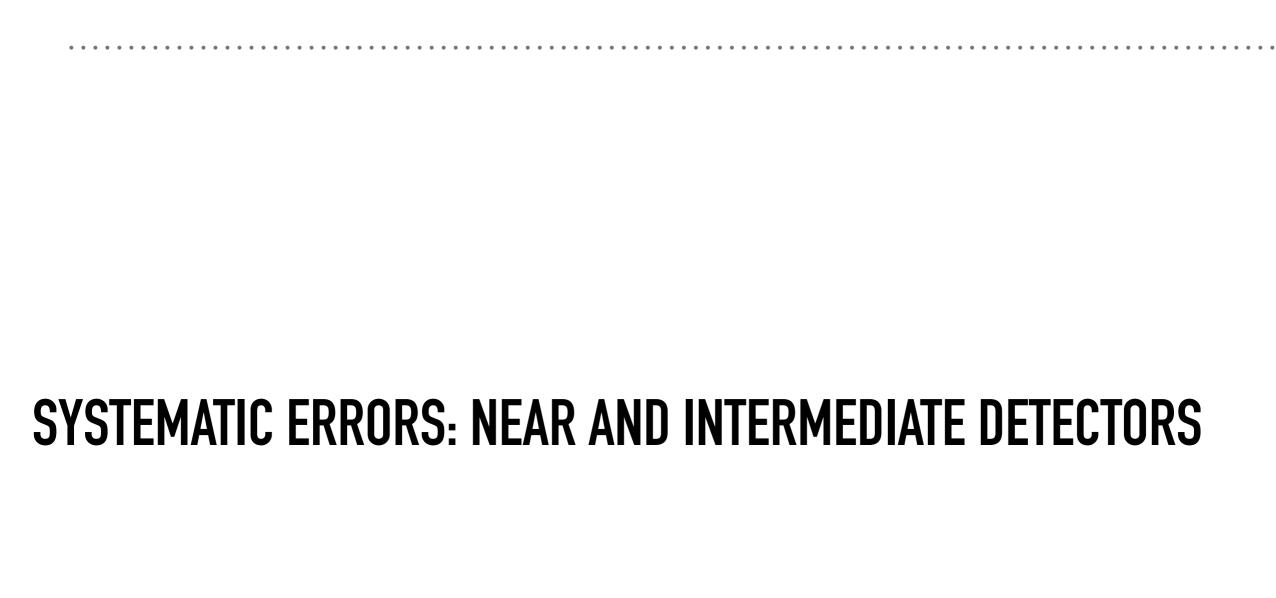
NUCLEON DECAY BACKGROUND AND NEUTRON DETECTION

- ► IWCD: measure the $\mu^+\pi^0$ background from neutrino interactions, related to the p→e⁺ π^0 proton decay mode
- ➤ Will load with Gd for neutron detection
 - ➤ Simulation including neutron backgrounds shows 75% tagging efficiency with 92% purity can be achieved

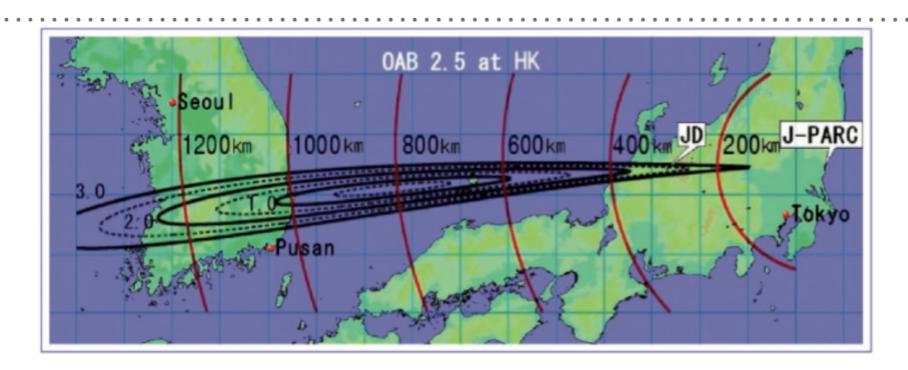


- Using the linear combination method, produce atmospheric spectrum for CCNπ production
 - ➤ Can measure the neutron production with the correct input spectrum



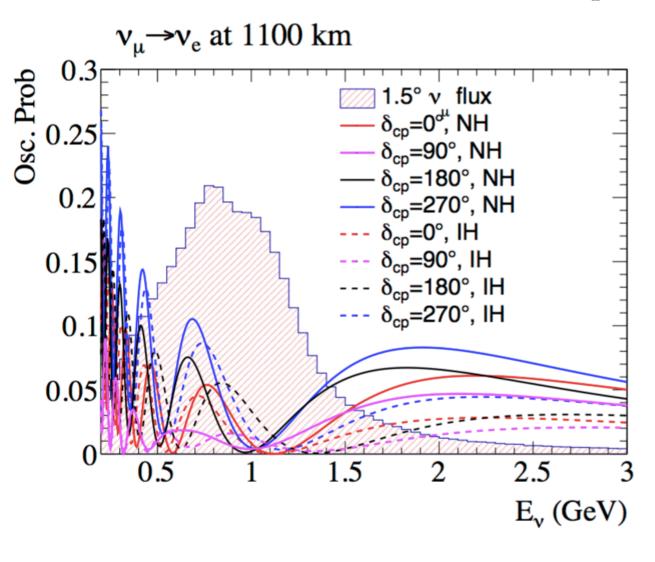


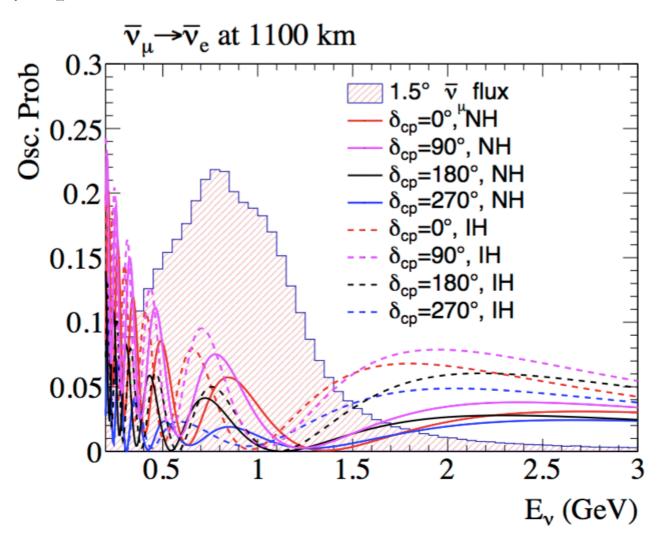
ADDITION OF THE KOREAN DETECTOR



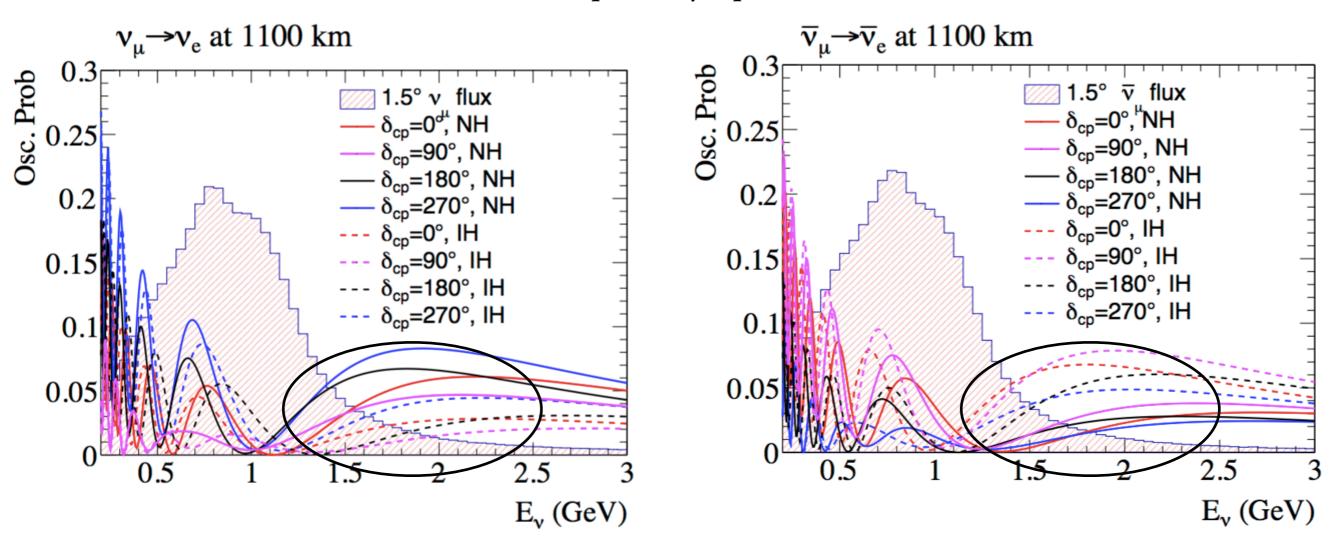
- ➤ Plans for a second detector in Korea are being developed
 - ➤ Physics studies of possible baselines/off-axis angles
 - > Selection of two candidate sites:
 - ➤ Mt. Bisul: 1.3° off-axis, 1088 km baseline, 1084 m overburden
 - ➤ Mt. Bohyun: 2.3° off-axis, 1043 km baseline, 1124 m overburden
 - ➤ Preliminary geological survey and cost estimation have been carried out for candidate sites

➤ Consider Mt. Bisul site at 1.3° as primary option



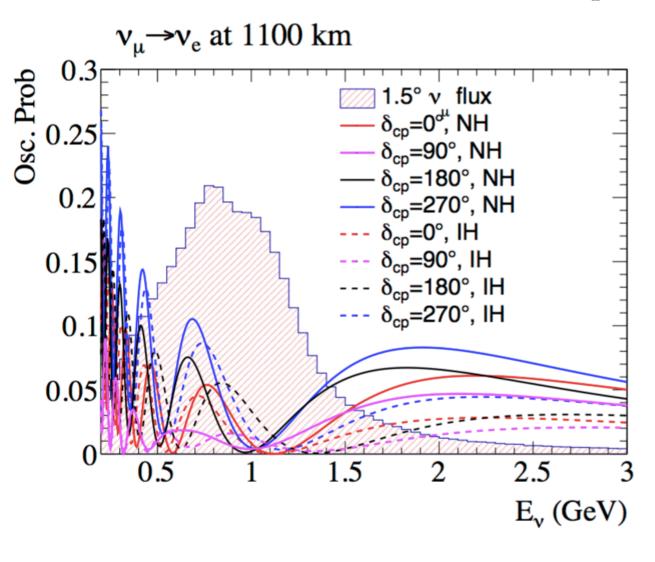


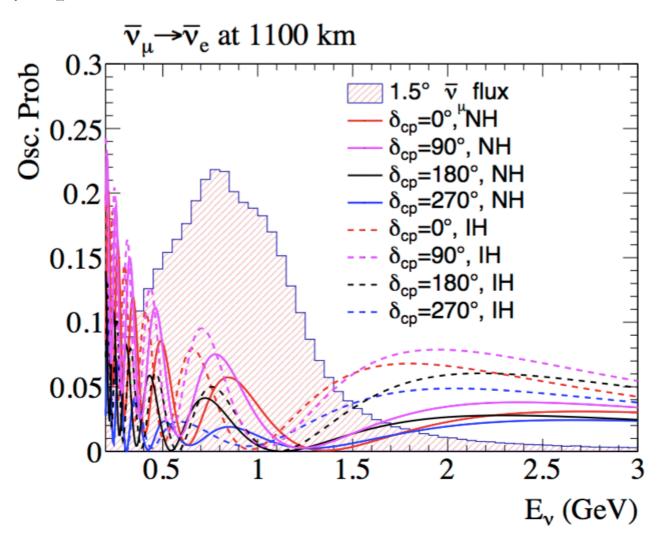
► Consider Mt. Bisul site at 1.3° as primary option



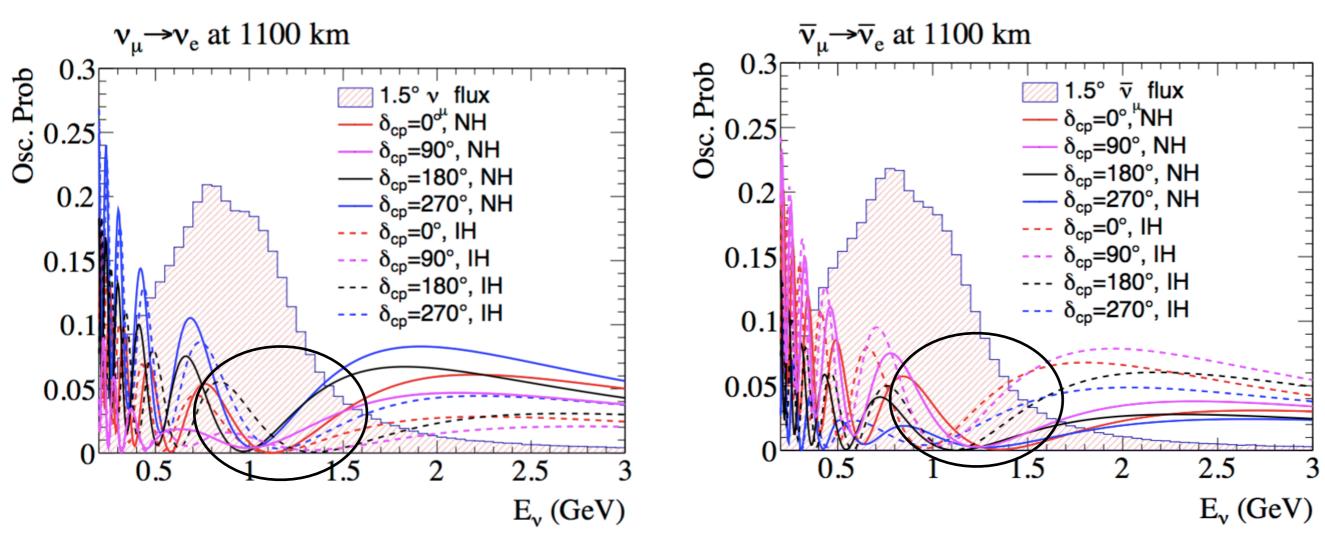
➤ Low energy part of first oscillation maximum: sensitivity to mass hierarchy

➤ Consider Mt. Bisul site at 1.3° as primary option



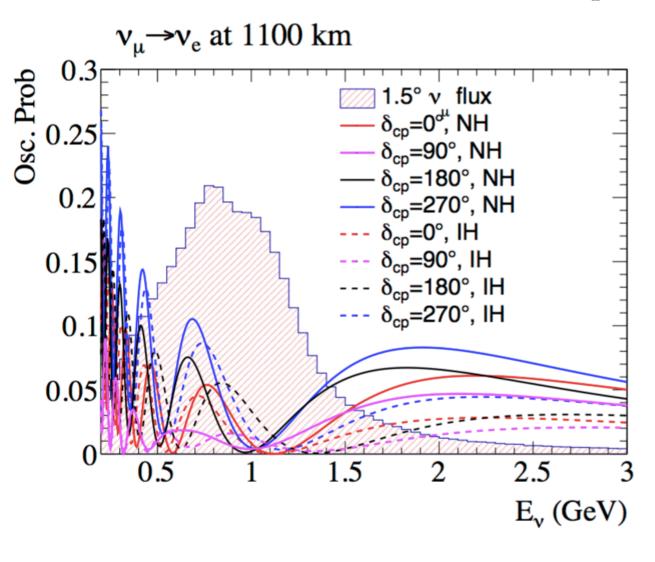


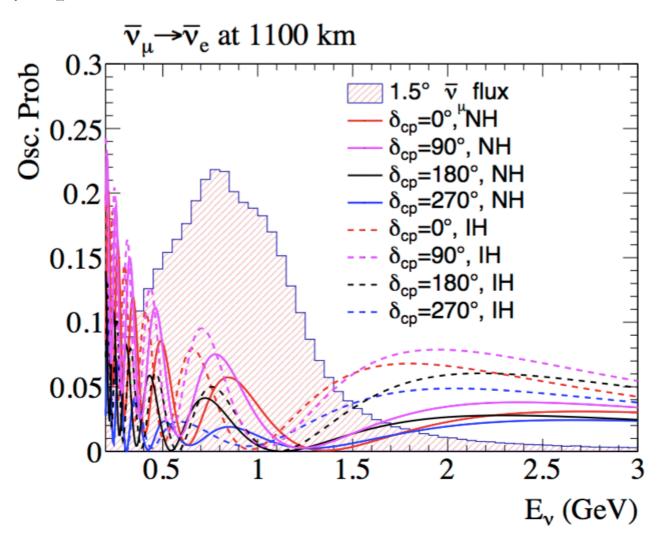
➤ Consider Mt. Bisul site at 1.3° as primary option



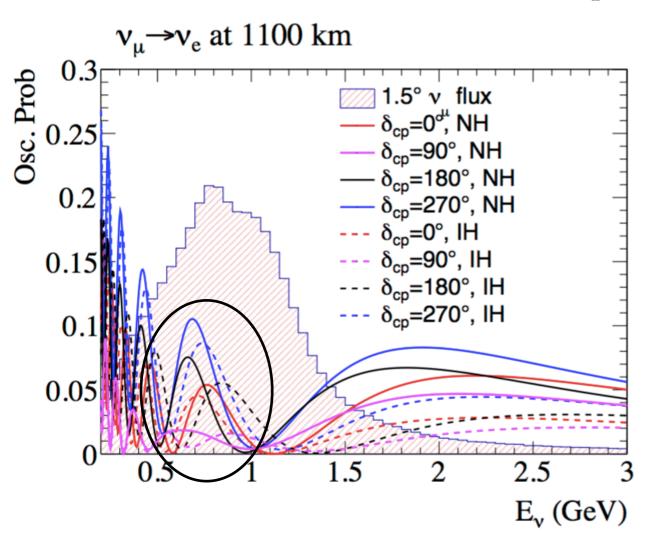
> Region between first and second oscillation maximum: sensitive to cosδ term

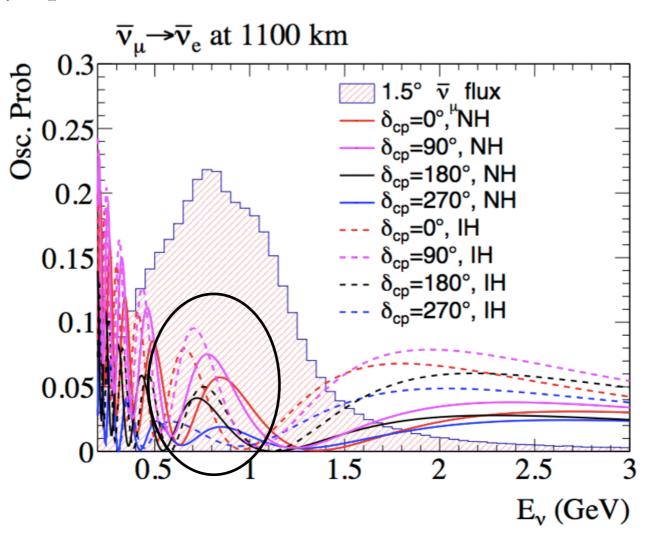
➤ Consider Mt. Bisul site at 1.3° as primary option





➤ Consider Mt. Bisul site at 1.3° as primary option

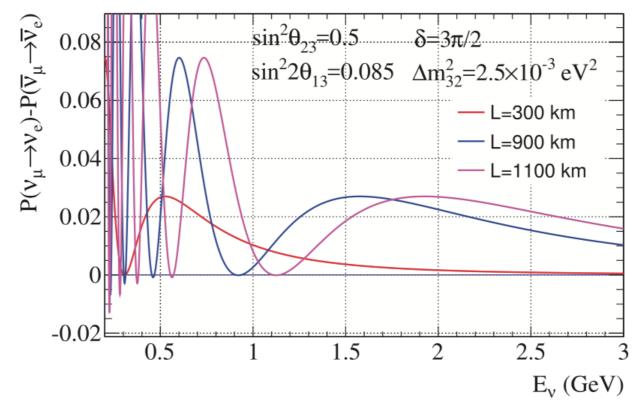




 \triangleright Second oscillation maximum: sensitive to sin δ term

ADVANTAGE 1: LARGER CP EFFECT

➤ For a fixed energy, the CP effect is three times larger at the second oscillation maximum (in vacuum)



- ➤ Accounting for actual baseline, spectrum and matter effects:
 - ➤ Maximum CP asymmetry at Japanese detector: 20%
 - ➤ Maximum CP asymmetry at Korean detector: 37%-44%
- ➤ Since the CP effect is twice as large, can survive with twice larger systematic normalization errors in the Korean detector
- ➤ Rate of accumulating statistics is lower due to the longer baseline

KOREAN DETECTOR SENSITIVITIES - CP VIOLATION

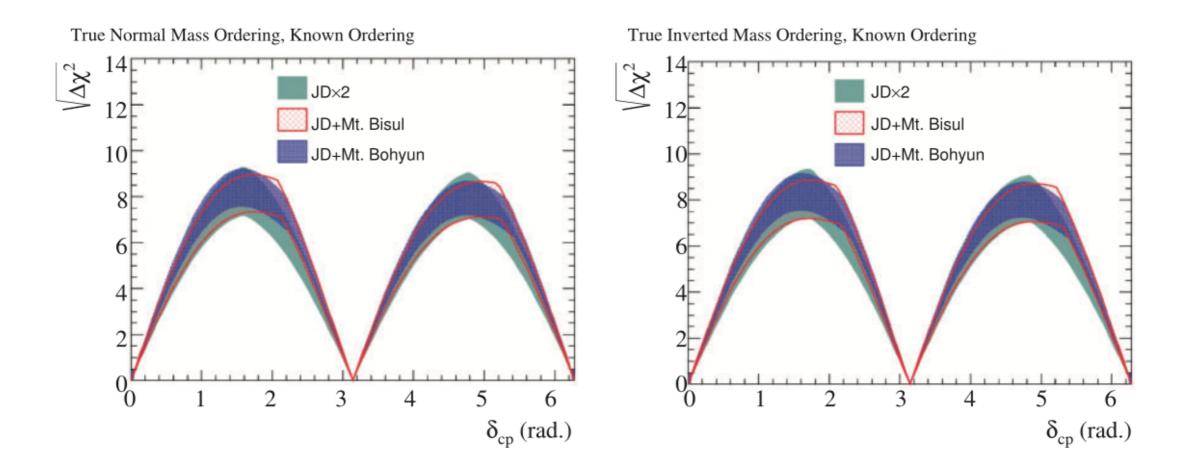
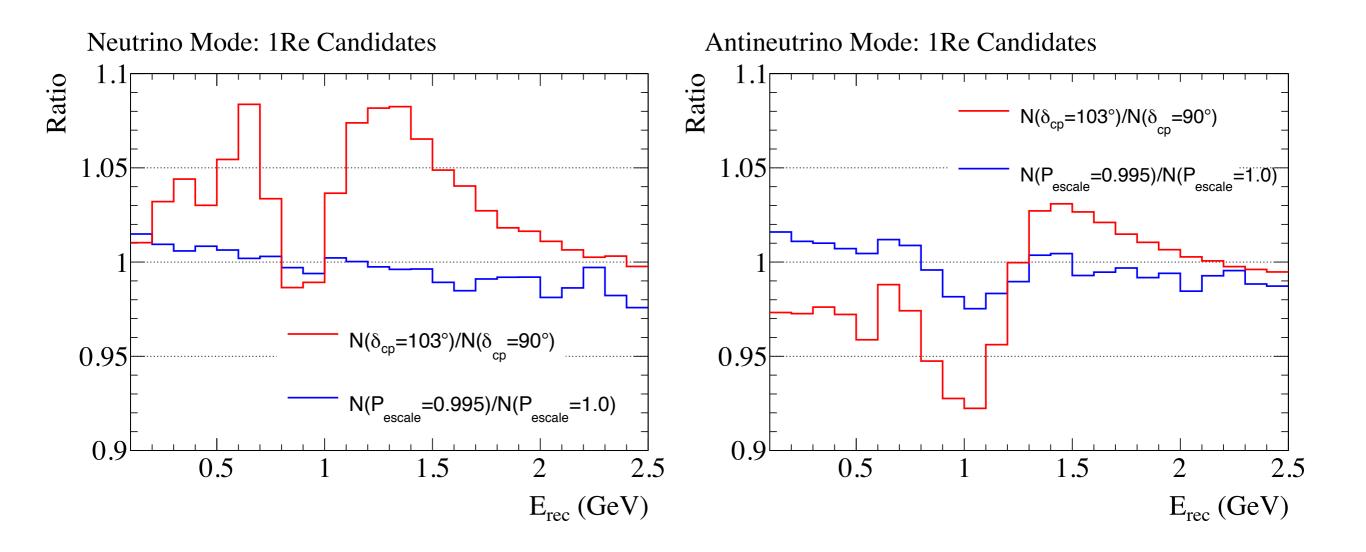


Table 7. The fraction of true δ_{CP} values for which CP violation can be discovered at 3σ or 5σ .

	True NH, known		True IH, known		True NH, unknown		True IH, unknown	
	3σ	5σ	3σ	5σ	3σ	5σ	3σ	5σ
$JD \times 1$	0.70	0.47	0.71	0.48	0.43	0.23	0.41	0.24
$JD\times 2$	0.74	0.55	0.74	0.55	0.52	0.27	0.50	0.28
JD+KD at 2.5°	0.76	0.58	0.76	0.59	0.76	0.48	0.72	0.30
JD+KD at 2.0°	0.78	0.61	0.78	0.61	0.77	0.55	0.79	0.51
JD+KD at 1.5°	0.77	0.59	0.77	0.59	0.77	0.59	0.77	0.59

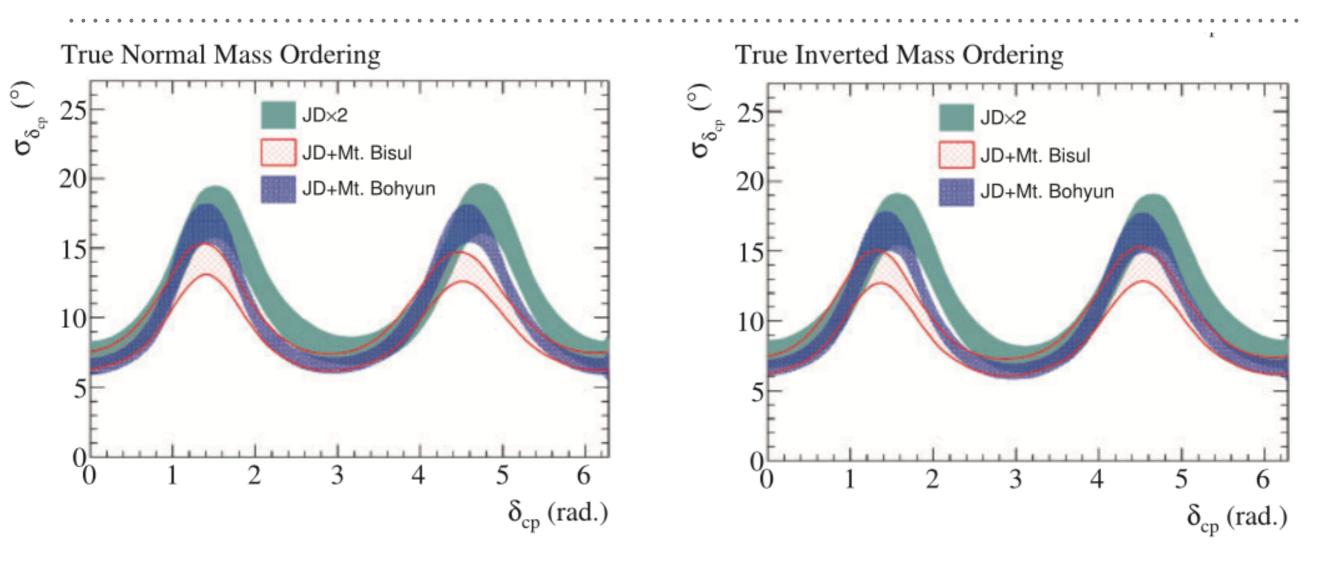
➤ The Korean detector improves the CP violation discovery sensitivity relative to 1 or 2 detectors in Japan

ADVANTAGE 2: cos & SENSITIVE REGION



- \blacktriangleright By probing the region between the first and second oscillation maximum, the degeneracy between energy scale and the δ_{cp} variation is removed
- \triangleright Effect of $\cos \delta_{cp}$ interference term is largest between oscillation maxima

KOREAN DETECTOR SENSITIVITIES - CP PRECISION



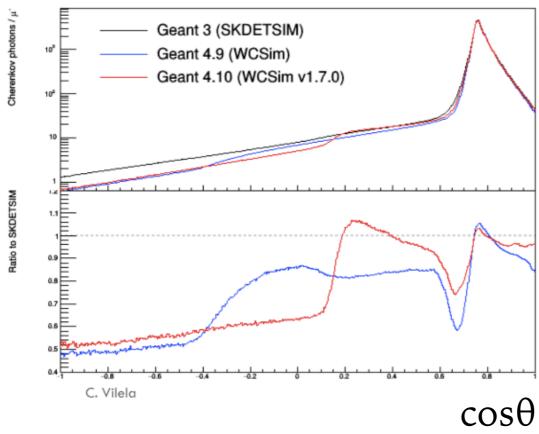
- ➤ With the Mt. Bisul location, the phase precision near maximal CP violation values is improved to better than 15 degrees
- ➤ Longer running time or increased beam power can lead to further improvement of the precision

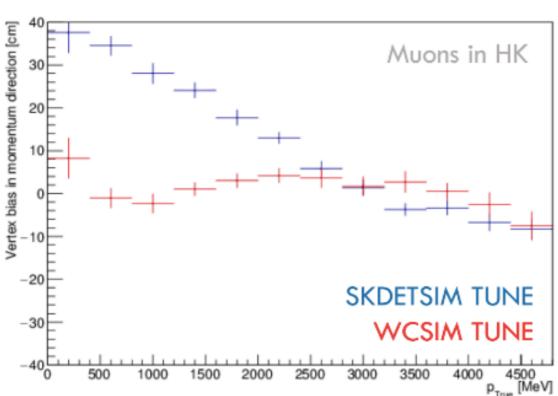
ANALYSIS SOFTWARE

ANALYSIS SOFTWARE

- ➤ Analysis software tools used in Hyper-K:
 - ➤ JNUBEAM: simulates neutrino production. Should be migrated from GEANT3 to GEANT4
 - ➤ **NEUT**: generates neutrino interactions in the detectors
 - ➤ WCSim: simulate the detector response for water Cherenkov detectors
 - ➤ fiTQun: likelihood based event reconstruction for high energy events
 - ➤ BONSAI: time residual vertex likelihood fit for low energy events

WCSIM PHYSICS





- ➤ We can benefit from work on the physics models in WCSim
- ➤ Even Cherenkov photons from muons show differences in different simulation code
 - ➤ Impact on reconstructed quantities (vertex) can be significant
- May be resolved with literature review, or may need new measurements
 - Possible measurement in water
 Cherenkov test experiment

FITQUN RECONSTRUCTION

FiTQun is originally developed for T2K experiment, and is based on maximum likelihood method.

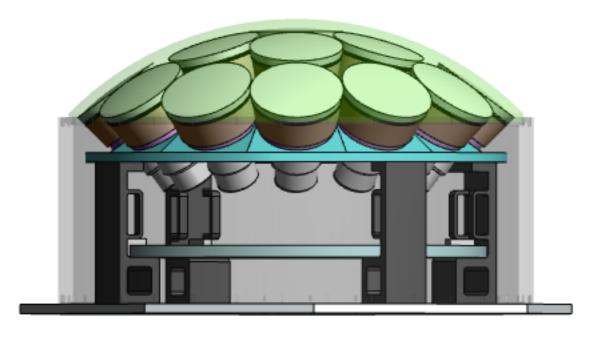
PMT unhit probability PMT charge pdf

 $L(\mathbf{x}) = \prod_{j}^{\text{unhit}} P_j(\text{unhit}|\mu_j) \prod_{i}^{\text{hit}} \{1 - P_i(\text{unhit}|\mu_i)\} f_q(q_i|\mu_i) f_t(t_i|\mathbf{x})$ PMT timing pdf

 $\mathbf{x} = (\boldsymbol{x}, t, p, \theta, \phi)$: Particle hypothesis

 $\mu = \mu^{
m dir} + \mu^{
m sct}$: Poisson mean of predicted charge detected by each PMT, which is also a function of ${f x}$

- Current optimization of fiTQun to handle multi-PMTs is ongoing
 - ➤ Code modification to account for the angular response and orientation of each 3-inch PMT in the module



FITQUN RECONSTRUCTION

FiTQun is originally developed for T2K experiment, and is based on maximum likelihood method.

PMT unhit probability PMT hit probability PMT charge pdf

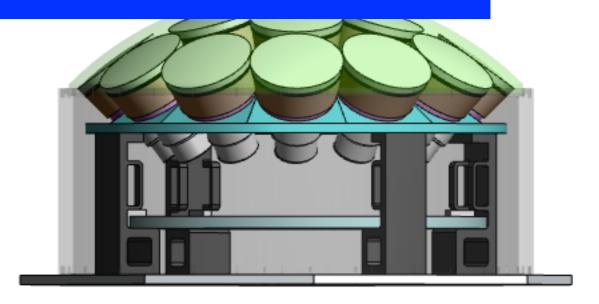
 $L(\mathbf{x}) = \prod P_i(\text{unhit}|\mu_i) \prod \{1 - P_i(\text{unhit}|\mu_i)\} f_q(q_i|\mu_i) f_t(t_i|\mathbf{x})$

Further improvements to the high energy reconstruction?

Build correlations between PMTs into the charge PDF

Machine learning approaches to improved event selection and particle identification

- Current optimization of fiTQun to handle multi-PMTs is ongoing
 - ➤ Code modification to account for the angular response and orientation of each 3-inch PMT in the module



LOW ENERGY RECONSTRUCTION IMPROVEMENTS

- ➤ Significant work is ongoing to develop low energy tagging of 2.2 MeV from neutron captures
- Likelihoods that include the angular acceptance of the PMT are being used

$$L_{ ext{signal}} = \sum_{i}^{ ext{all hits}} \log\Bigl(rac{rac{\epsilon(\cos\delta_i)}{r_i^2}e^{-rac{r_i}{\lambda}}}{\sum_{j}^{ ext{all PMT}}\bigl(rac{\epsilon(\cos\delta_j)}{r_j^2}e^{-rac{r_j}{\lambda}}\bigr)}\Bigr), \ \ L_{ ext{dark}} \ \ = \sum_{i}^{ ext{all hits}} \log\Bigl(rac{1}{N_{ ext{PMT}}}\Bigr)$$
 vertex Gamma direction

- ➤ PMT angular acceptance likelihood, opening angle, neutron capture time, number of hits in 10 ns window and goodness of vertex fit combined in TMVA
- ➤ Tagging efficiency for 8.4 kHz dark noise is improved from 42.5% to 52.9%
- ➤ Addition of multi-PMTs for neutron tagging is open topic

SUMMARY

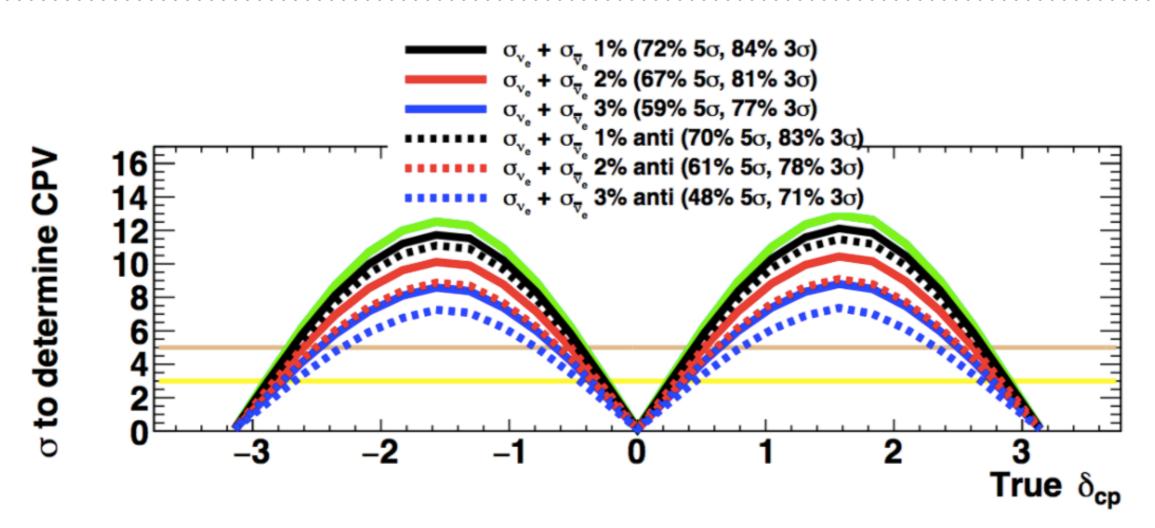
- ➤ There are many areas to contribute to Hyper-K to improve physics results including:
 - ➤ Development of near and intermediate detectors to control systematic errors
 - ➤ Intermediate water Cherenkov detector
 - ➤ Off-axis magnetized detector
 - > Development of calibration systems at near and far detectors
 - ➤ Chance to test calibration systems in test beam experiment
 - > Development of new analysis software tools
 - ➤ Improvement of simulation physics models and upgrades to modern tools
 - ➤ Improving reconstruction by adding new PMT information (angular acceptance) and developing new techniques (multi-variate analysis and machine learning)

THANK YOU

FURTHER IMPROVEMENTS TO HIGH ENERGY RECONSTRUCTION?

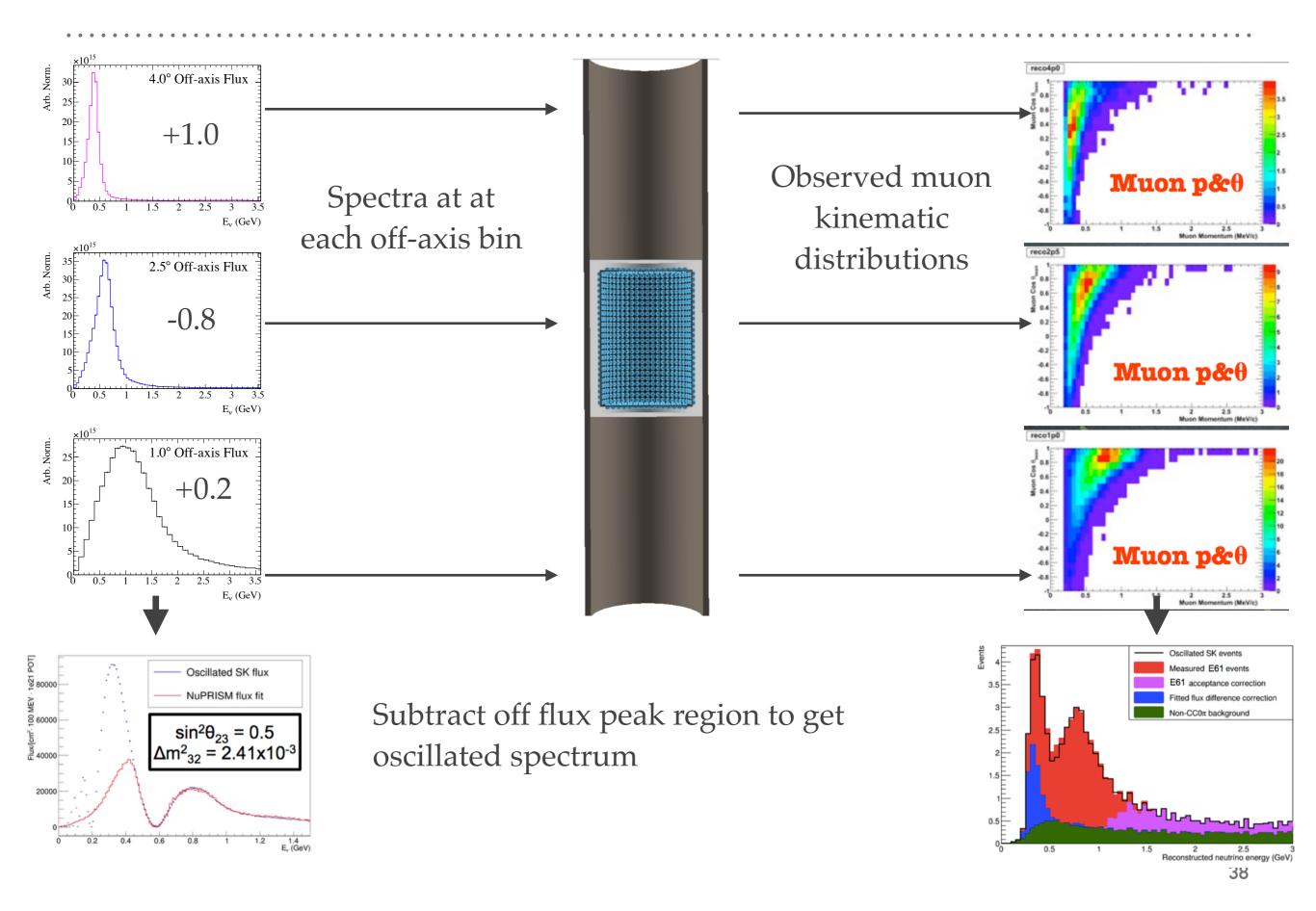
- ➤ The charge PDF used in fiTQun is an average over many events
 - ➤ Only see the average effects of lepton and photon scattering processes?
 - ➤ Does this limit ability to separate electron and photons or other aspect of particle ID performance?
- ➤ To account for structure within the ring due to stochastic scattering processes two approaches may be considered
 - ➤ Calculate correlations between PMTs when calculating the charge PDF
 - ➤ Apply machine learning methods that train on collections of single electron, muon, photon, pion, etc. events

NORMALIZATION ERROR FOR THE CPV MEASUREMENT



- ➤ Aiming to achieve 3% error or better on the relative rate of electron neutrino and electron (anti)neutrino rates
- ➤ Total systematic error is the quadrature sum of all systematic errors
 - ➤ Should aim to keep each individual error source below 1% if possible
- ➤ Three examples on the following slides

OFF-AXIS MEASUREMENTS IN IWCD



FURTHER COMMENTS ON SYSTEMATIC ERRORS

- ➤ In addition to near/intermediate detector measurements, calibration of the far detector is critical
- ➤ More precise and extensive measurements of hadron production will help extract the critical neutrino interaction information from the data observed in near and intermediate detectors
- ➤ Modeling of the atmospheric neutrino backgrounds are important for nucleon decay measurements
 - ➤ These should be measured in the near and intermediate detectors where possible
- ➤ Neutrino interaction measurements in the near and intermediate detectors an also constrain systematic errors relevant for the atmospheric neutrino measurements

NEUTRON DETECTION - NUCLEON DECAY

- ➤ Neutron detection plays a key role in a number of Hyper-K physics analyses
- ➤ In the baseline design, the neutrons are detected by the captures on H that produce a 2.2 MeV gamma
 - ➤ Gd loading in Hyper-K can increase the neutron detection efficiency

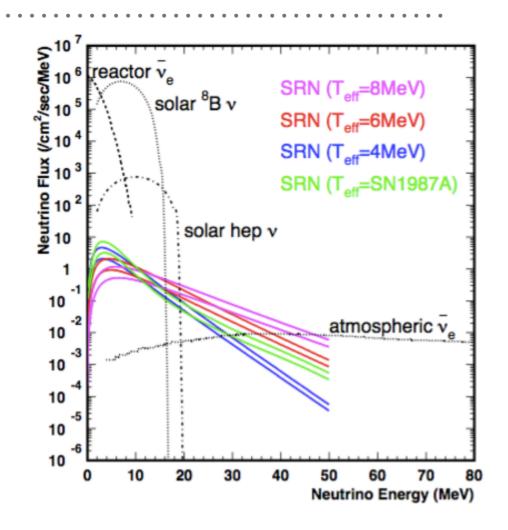
Proton decay:

- ► In the p→e⁺ π^0 channel, about 3 atmospheric neutrino background events after 10 years
- ➤ According to simulation, 80% of atmospheric background interactions eject at least one neutron
- > 70% neutron tagging efficiency leads to 70% reduction of atmospheric background
- ➤ Measurements of the (anti)neutrino background and neutron production are critical for estimating the background
- ➤ Analysis improvements that improve the neutron tagging efficiency can have a large impact

NEUTRON DETECTION - SUPERNOVA NEUTRINOS

Supernova relic neutrino detection

- ➤ Detection of the relic supernova neutrino background
- ➤ Detection of the anti-electron neutrino through the IBD channel
- ➤ IBD channel can tagged with detection of neutron, removing solar background and lowering threshold



Supernova burst neutrino detection

- ➤ Pointing accuracy of supernova neutrino detection is important to direct multimessenger observation of the supernova
- ➤ Directional information comes from neutrino-electron scattering events
 - ➤ They sit on a large background of IBD events
- ➤ If IBD events are removed with neutron tagging, direction accuracy can be improved

HYPER-K PHYSICS - CP VIOLATION

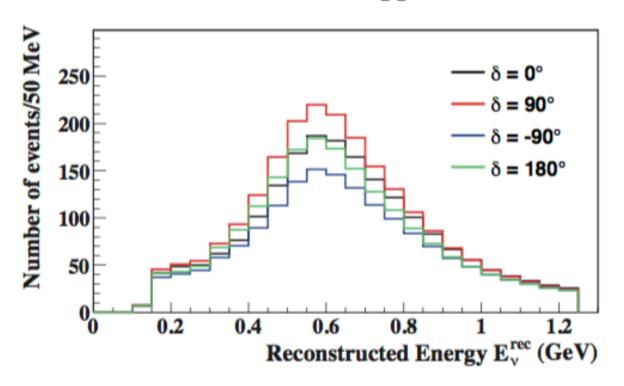
➤ Search for an asymmetry in the muon (anti)neutrino to electron (anti)neutrino oscillation probabilities

$$P_{\mu \to e} = \sin^2 \theta_{23} \sin^2 2\theta_{13} \sin^2 \left(\frac{\Delta m_{31}^2 L}{4E_{\nu}}\right) \mp \frac{\sin 2\theta_{12} \sin 2\theta_{23}}{2\sin \theta_{13}} \sin^2 2\theta_{13} \sin \left(\frac{\Delta m_{21}^2 L}{4E_{\nu}}\right) \sin^2 \left(\frac{\Delta m_{31}^2 L}{4E_{\nu}}\right) \sin \delta_{CP} + \dots$$

Neutrino mode: appearance

300 250 200 150 100 50 0 0.2 0.4 0.6 0.8 1 1.2 Reconstructed Energy E^{rec} (GeV)

Antineutrino mode: appearance



➤ With about 2000 candidates in each channel, the statistical error on the relative neutrino/antineutrino rate is about 3% - systematic errors should be smaller