## Theia



International Workshop on Next-Generation Nucleon Decay and Neutrino Detectors 3rd November, 2018

Gabriel D. Orebi Gann UC Berkeley & LBNL

BERKELEY LAB



Development of new scintillators e.g. WbLS



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#### Fully-equipped, deep underground labs (+ beam)



#### Theia at NNN, G. D. Orebi Gann



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**Development of new** scintillators e.g. WbLS

top window

mcp 1

mcp 2

anode readout-

photocathode (pc)

Advanced computing & reconstruction methods J. Inst. 9, Po6012 (2014), Nucl. Inst. Methods A849, 102 (2017).**Fully-equipped**, deep Newunderground labs (+ beam) generation of large-scale, Fast, efficient low-threshold, photodetectors directional V,V detectors 4850 Level (4300 n **Proposed Laboratories** incoming photon Experiment Hall UX/ZEPLIN **R&D** opportunities LBNE Davis Campus pc gap evel 10 kT and 24 kT liquid are LUX MID inter-mcp gap **BHSU Underground Campus R&D** opportunities CASPAR anode gap **Ross Campus** 

## Overview

- Detector concept
- Physics Program
- Development of Detector Capabilities

Theia Detector Concept

## Cherenkov / Scintillation Separation

### Separation in charge, time, wavelength Methods to enhance separation:

- Ultra-fast photon detection (LAPPDs)
- Delay scintillation light
- Optimize cocktail: scintillation fraction & spectrum (fluor)
- Readout sensitivity



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# Cherenkov / Scintillation Separation



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## Theia

- Large-scale detector (50-100 kton)
- Water-based LS target
- Fast, high-efficiency photon detection with high coverage
- Deep underground (e.g. Homestake)
- Isotope loading (Gd, Te, Li...)
- *Flexible*! Target, loading, configuration

Broad physics program!





Concept paper - <u>arXiv:1409.5864</u>

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## Theia Physics Program

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- I. Neutrinoless double beta decay
- 2. Solar neutrinos (solar metallicity, luminosity)
- 3. Geo-neutrinos
- 4. Supernova burst neutrinos & DSNB
- 5. Source-based sterile searches
- 6. Nucleon decay
- 7. Long-baseline physics (mass hierarchy, CP violation)

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High-

Energy

Physics

Physics over 5 orders of magnitude

Nuclear

High-

Energy

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Remarkably, the same detector could show that neutrinos and antineutrinos are the same, **and** that "neutrinos" and "antineutrinos" oscillate differently



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Physics over 5 orders of magnitude

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**Physics** 

Physics

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50 kton water-based liquid scintillator detector High coverage with fast photon detectors Deep underground 8-m radius balloon with high-LY LS and isotope 7-m fiducial, 3% <sup>nat</sup>Te or <sup>enr</sup>Xe, 10 years Builds on critical developments by KLZ & SNO+ collaborations

Evonts / BOLy

		· 1		Livenus	/ ICOL y
			$\mathbf{Signal}$	Te Loading $e$	$^{nr}$ Xe Loading
		Cosmogenic	$0\nu\beta\beta$ (10 meV)	65.4	116.4
		$2\nu\beta\beta$	2 uetaeta	48.0	38.2
	着 40 m	<sup>8</sup> B ν ES (α, n) External γ	$^{8}$ B Solar ES (50%)	138.5	138.4
			$^{10}C$ (92.5%)	24.6	25.4
40 m	- And - And		<sup>130</sup> I	48.3	
			$^{130m}\mathrm{I}$	1.7	
			$^{136}Cs$		0.57
			$^{208}$ Tl	0.02	0.002
		Internal U cha	<sup>214</sup> Bi (99.9%)	4.0	4.4
		Internal Th chain	Balloon <sup>214</sup> Bi (50%)	24.0	27.4
			Balloon $^{208}$ Tl (50%)	0.25	0.14
	Miller .	SNO+ Collaboration	Total	289.5	234.5
	3 - 4				

Phys.Rev.Lett. 110:062502 (2013); Adv.High Energy Phys. 2016 (2016) 6194250; Phys. Rev. D 87 no. 7:071301 (2013)







# Long-Baseline Program

- Large-scale detector at Homestake, in the LBNF beam
- Complementary program to LArTPC (DUNE)
- Build on WCD studies (arXiv:1204.2295)

40 kt LArTPC

100 kt WCD

----- 50 kt WCD

Plus advantages from low-threshold scintillation

- Assumes 75% reduction in NC background relative to SK-I
- Uses only single-ring samples



#### Mass Hierarchy Sensitivity

,**30**, √×

20

15

10

Normal Hierarchy

3.5+3.5 v+v years

 $sin^2 2\theta_{13} = 0.085$  $sin^2 \theta_{23} = 0.45$ 

Testing the existence of GUTs with THEIA:

- Large size (statistics), deep location, very clean
- n tagging (low threshold plus potential isotope loading)
- Sub-Cherenkov threshold detection

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Figs from arXiv:1409.5864, assume 100t FV; studies based on Phys. Rev. D 72,075014 (2005); LAr from JHEP 0704:041,2007

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## Solar Neutrinos with Theia



- Dominant background to CNO v measurement: <sup>210</sup>Bi
- Theia offers unique low-threshold, directional detection

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Signal	Normalization sensitivity (%)
<sup>8</sup> Β ν	0.4
<sup>7</sup> Be $\nu$	0.4
pep v	3.8
CNO v	5.3
<sup>210</sup> Bi	0.1
<sup>11</sup> C	11.5
<sup>85</sup> Kr	10.5
<sup>40</sup> K	0.04
<sup>39</sup> Ar/ <sup>210</sup> Po	21.9
<sup>238</sup> U chain	0.02
<sup>232</sup> Th chain	0.05

Eur. Phys. J. C (2018) 78: 435

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# Theia Spectral Sensitivity

1996, W.C. Haxton: isotope loading for CC interaction (water)

"Salty water Cherenkov detectors" W.C. Haxton PRL 76 (1996) 10 CC detection in WbLS: high-precision spectral measurement to low energy ⇒ search for new physics, solar metallicity, MSW effect



Detector: 30kt fiducial 1% 7Li by mass Conservative 100 pe/MeV



arXiv:1409.5864

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## Antineutrino Detection

- Detect via IBD
- High light yield allows enhanced n tag : 2.2 MeV  $\gamma$  from <sup>1</sup>H
  - Suppress single-event background that limits water Cherenkov
- Higher detection efficiency than Gd-H<sub>2</sub>O due to high scint. yield
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- Current total geo-v exposure: < 10kt-yr (KL + Borexino)
  - **THEIA:** large statistics in a complementary geographical location

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DSNB

- Enhanced n tag
- Reduced NC background
- Most sensitive search to-date
- Plus NaCl for v signal

## Supernova

## Neutrinos

Neutrino	Percentage of	Type of
Reaction	Total Events	Interaction
$\overline{\nu}_e + p \to n + e^+$	88%	Inverse Beta
$\nu_e + e^- \rightarrow \nu_e + e^-$	1.5%	Elastic Scattering
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- ~15k events for SN at 10 kpc (50 kt volume)
- ~90% events are IBD

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Highly complementary to v<sub>e</sub>-dominated LAr signal

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Early warning (PR value)



### Talk by M.Yeh [Thurs pm] WbLS Development



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Poster by V. Fishcer

### Talks by V. Fischer, S. Qian, L. Wen, A. Lyashenko, J. Kameda, T. Lindner Photon Sensor Development



Nucl. Inst. Meth. Phys. Res. A. Volume 814, 19-32, (April 2016); Nucl. Inst. Meth. Phys. Res. A. (Oct. 2016)

#### PRC 95 055801 (2017)

## CHESS: CHErenkov-Scintillation Separation

### **Ring-imaging experiment**



### Time- and charge-based separation in LAB/PPO



#### Eur. Phys. J. C (2017) 77:811

	LAB (time)	LAB (charge)	LAB/PPO (time)	LAB/PPO (charge)
nerenkov etection fficiency	83 ± 3 %	96 ± 2 %	70 ± 3 %	63 ± 8 %
intillation tamination	11 ± 1 %	6±3%	36 ± 5 %	38 ± 4 %

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# CHESS Results: WbLS



## Time Profile

### PRELIMINARY



## Time Profile

#### Extract microphysical parameters by fitting to MC model Cerenkov $10^{-1}$ 90Sr source, single pe regime, detailed MC Scintillation Reemission $10^{-2}$ Calibrate method using well-understood LAB/PPO target DATA Time profile model: 3 exp. decay + rise time Preliminary $\rho(t) \propto (1 - e^{-t/\tau_r}) \times \sum_{i}^{3} A_i e^{-t/\tau_i} \begin{cases} \tau_r = 0.7 \text{ ns} \\ \tau_1 = 4.3 \text{ ns} \\ \tau_2 = 16 \text{ ns} \\ \tau_3 = 166 \text{ ns} \end{cases}^{\text{H. M. O'Keeffe et al.}}_{\text{Add0, 119 (2011)}}$ $10^{-3}$ $10^{-1}$ Good agreement between data and model "out of the box" 0 20 40 Time (ns)

#### Fit for scintillation time profile of WbLS



PRELIMINARY

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 $10^{-3}$ 

 $10^{-1}$ 

0

Time profile model: 3 exp. decay + rise time  $\rho(t) \propto (1 - e^{-t/\tau_r}) \times \sum_{i}^{3} A_i e^{-t/\tau_i} \begin{cases} \tau_r = 0.7 \text{ ns} \\ \tau_1 = 4.3 \text{ ns} \\ \tau_2 = 16 \text{ ns} \\ \tau_3 = 166 \text{ ns} \end{cases}^{\text{H. M. O'Keeffe et al.}}_{\text{A640, 119 (2011)}}$ 

Good agreement between data and model "out of the box"



PRELIMINARY

20

40

Time (ns)

# Preliminary light yield

Note: assumes LAB/PPO wvl emission profile

Fit for WbLS time profile



Method: define LAB/PPO LY Calibrate setup to LAB/PPO charge collection Determine LY of WbLS cocktail (data/MC fit)



# Signal Separation in Theia

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# Signal Separation in Theia



## Ring Imaging

## Ring Imaging



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## Ring Imaging



Talks by M. Askins, V. Fishcer, C. Mauger

## Community Interest

Site	Scale	Target	Measurements	Timescale
UChicago	bench top		fast photodetectors	Exists
CHIPS	10 kton	H2O	electronics, readout, mechanical infrastructure	2019
EGADS	200 ton			Exists
ANNIE	30 ton	H2O+Gd	isotope loading, fast photodetectors	Exists
WATCHMAN	l kton			2020
NuDot	l ton	LS	directionality	2018
Penn	30 L	(Wb)LS	light yield, timing, loading	Exists
SNO+	<b>780</b> ton			2018
CHESS (LBNL)	bench top	WbLS	signal separation, tracking, reconstruction / light yield, loading, attenuation	Exists
BNL	l ton			Exists



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## **THEIA** Collaboration



## Summary

 THEIA: broad program of compelling science

• Flexibility to adapt to new directions in the scientific program as the field evolves

Powerful instrument of discovery

Rich, exciting program of ongoing R&D