



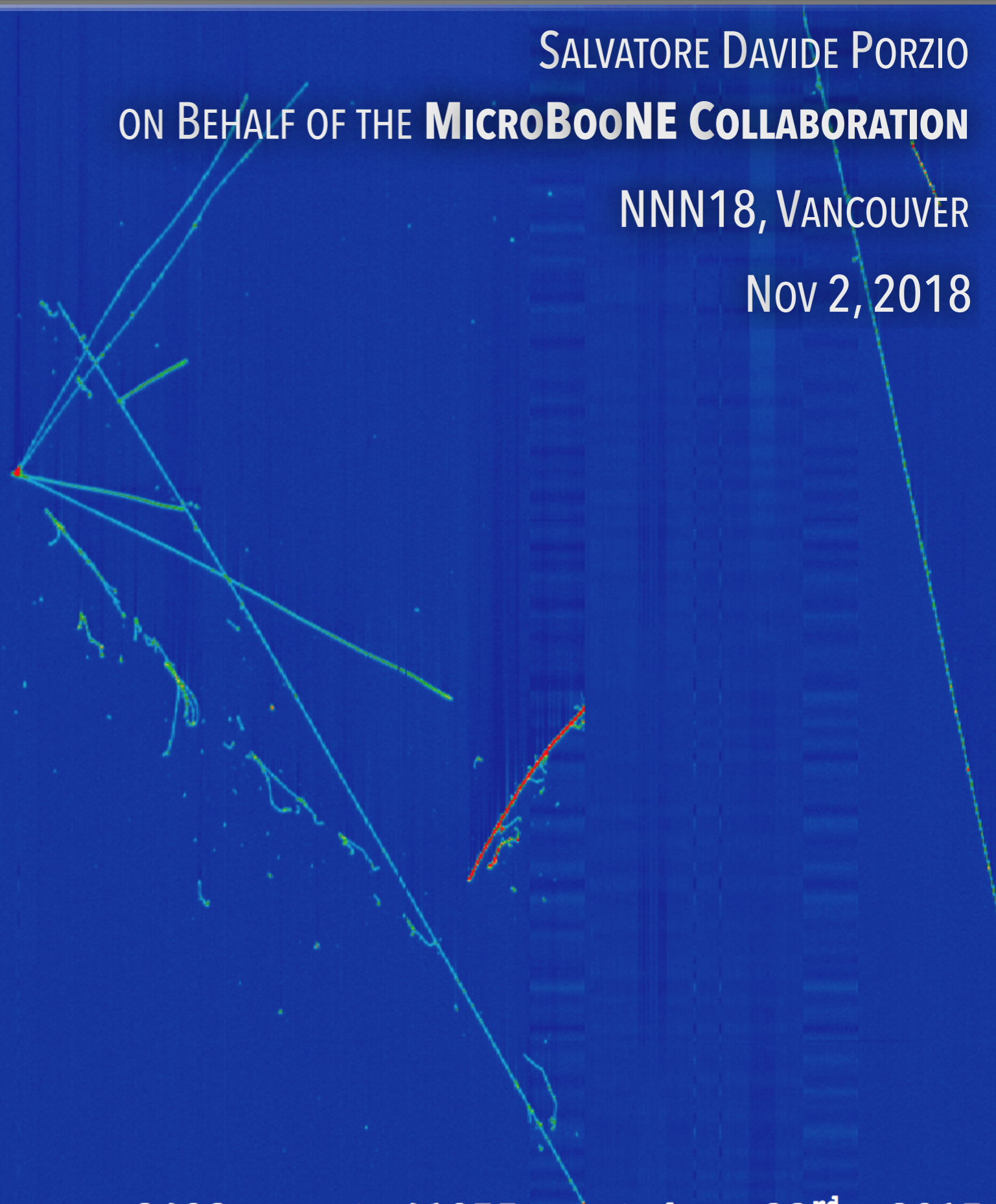
SALVATORE DAVIDE PORZIO

ON BEHALF OF THE **MICROBooNE COLLABORATION**

NNN18, VANCOUVER

Nov 2, 2018

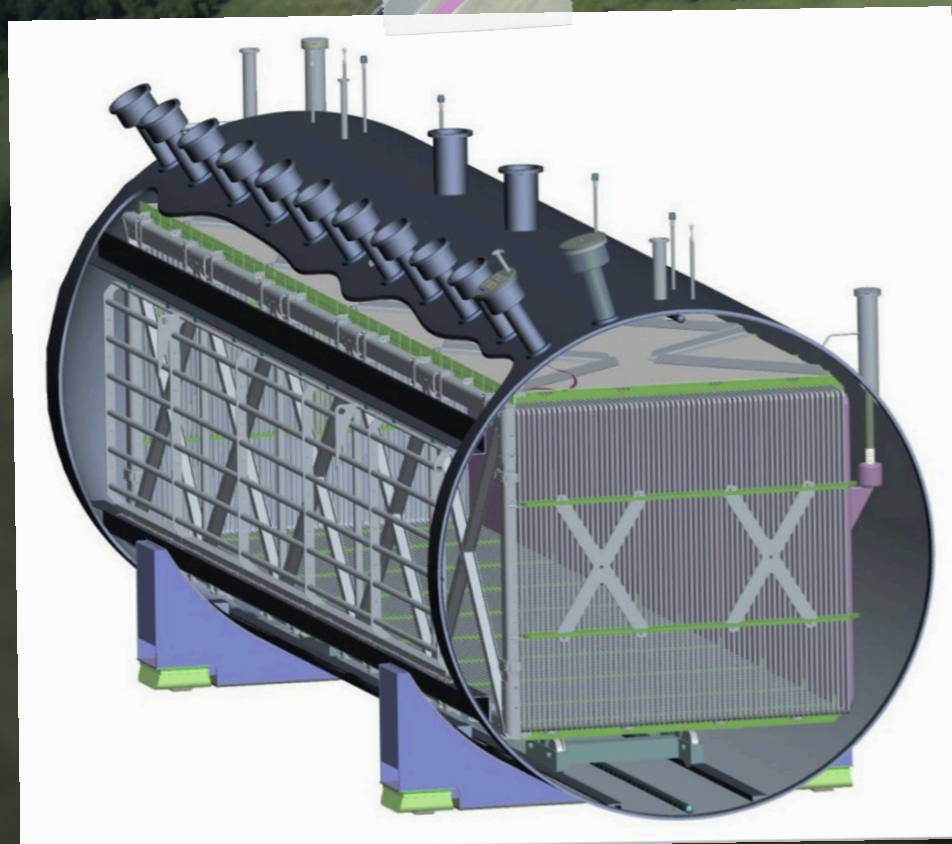
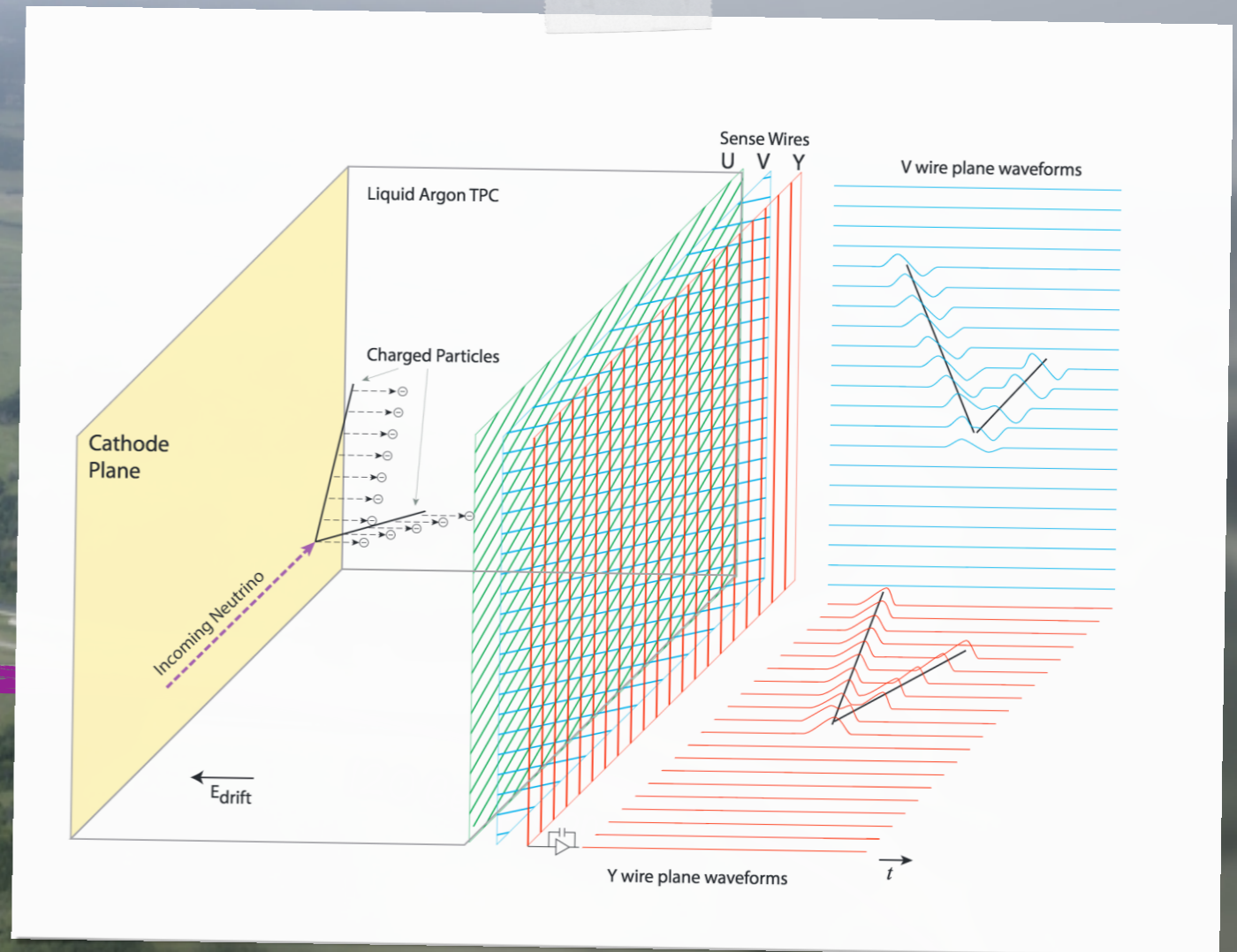
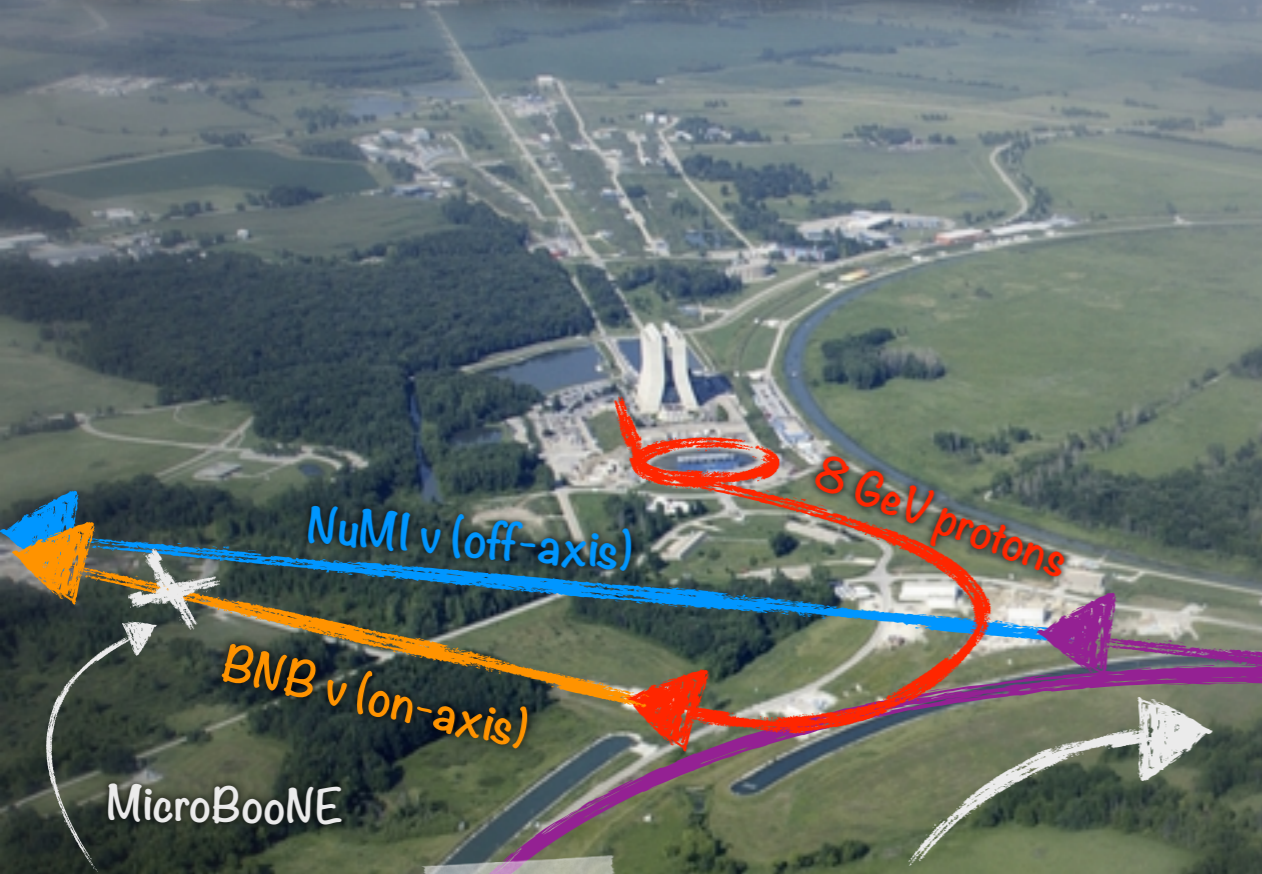
**ANALYSIS AND
SYSTEMATIC
UNCERTAINTY
EXPERIENCE FROM
MICROBooNE**



75 cm

Run 3493 Event 41075, October 23rd, 2015

MICROBOONE

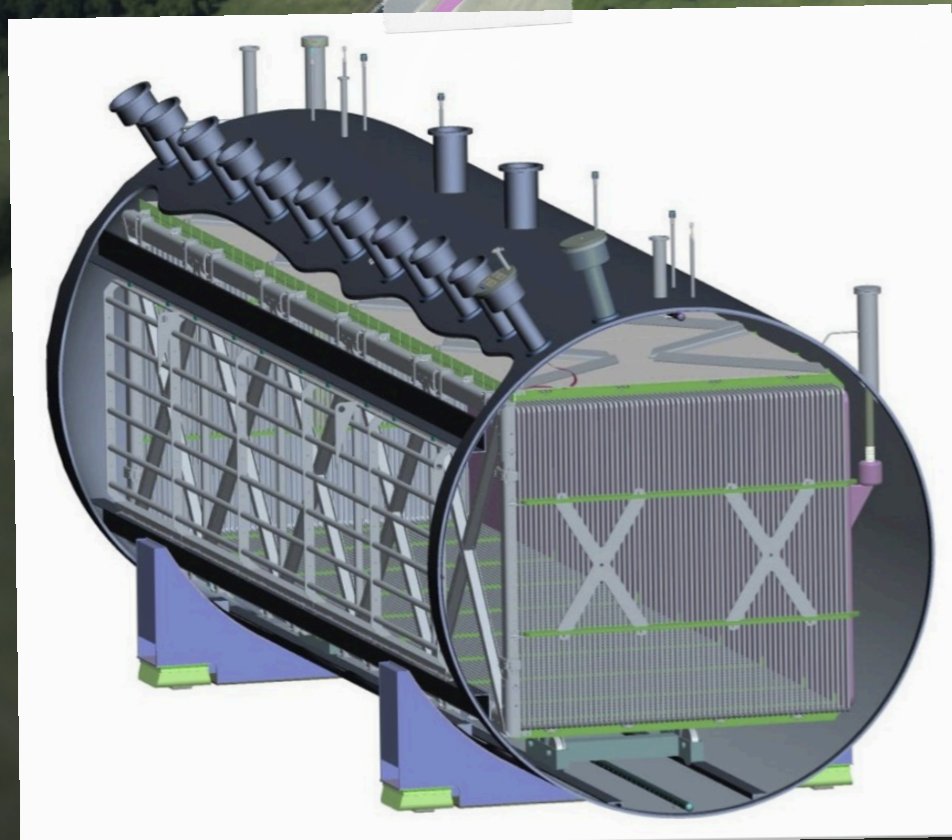
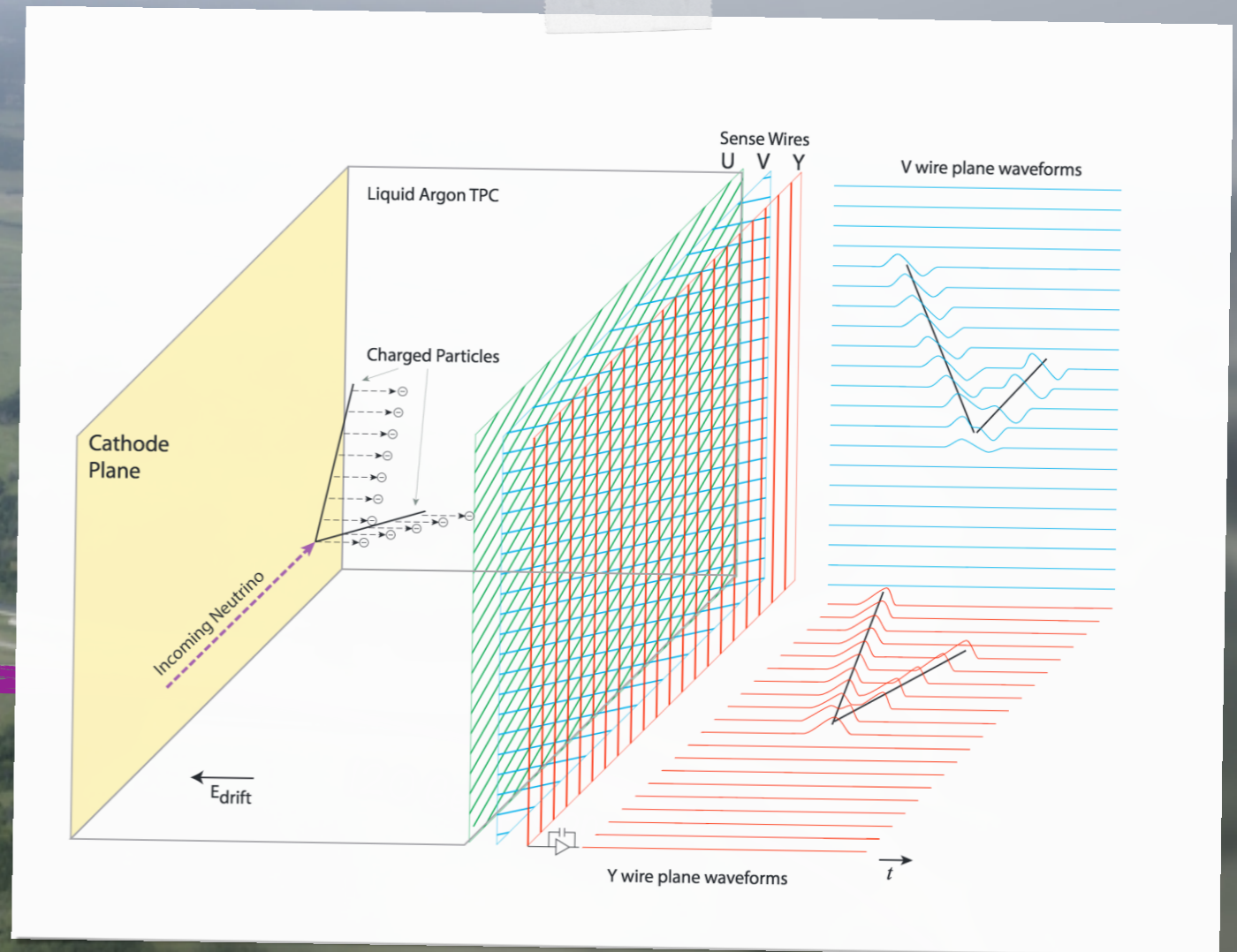


- **Liquid Argon Time Projection Chamber**
- **Short-Baseline Oscillation¹ Experiment (470 m baseline), along with SBND and ICARUS**
- **Downstream of BNB and NuMI**

- **87-ton active volume LAr**
- **Drift chamber with E-field at 273 V/cm**
- **Three wire planes (8192 wires):**
 - 2 induction, 1 collection
 - 3 mm wire pitch
 - 3 mm plane spacing
- **32 8" Cryogenic PMTs**

MICROBOONE

See Brooke Russell talk tomorrow morning:
Short Baseline Neutrino Experiments

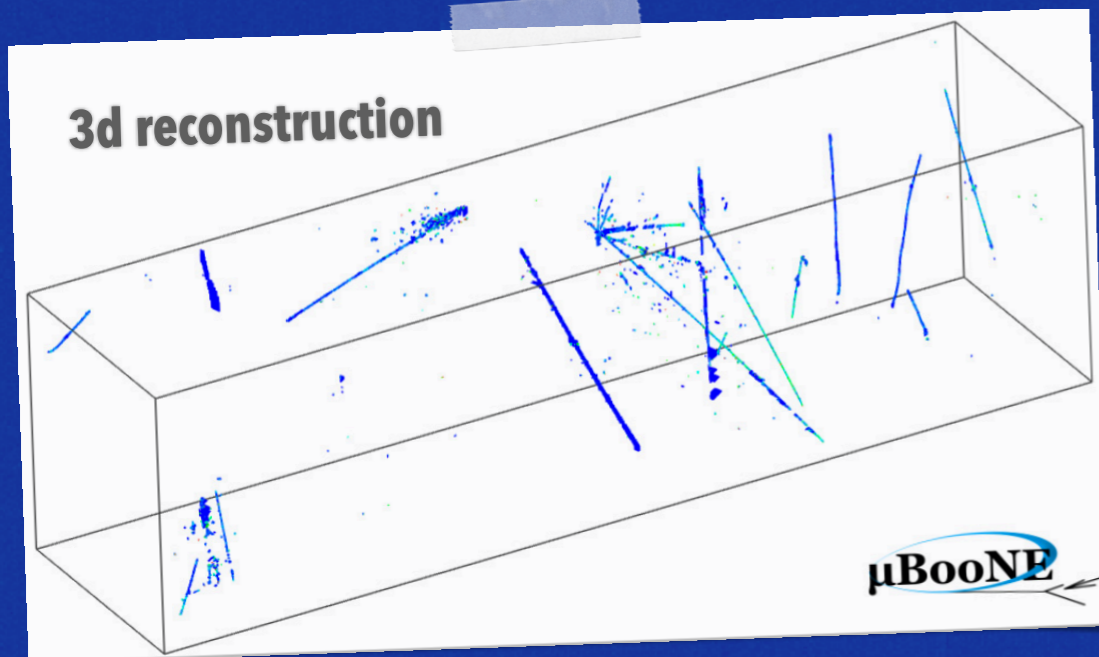


- Liquid Argon Time Projection Chamber
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μ BooNE

MICROBOONE



- Science goals:
- Resolve MiniBooNE low energy excess
 - Study ν -argon cross-sections
 - LArTPC R&D
 - Astrophysics and Exotica Physics

Bubble chamber quality images
+ calorimetry and automated reconstruction!

Excellent resolution (~3mm)

75 cm

Run 3493 Event 41075, October 23rd, 2015

EXPERIENCE FROM PHYSICS ANALYSIS

Also providing great input to:

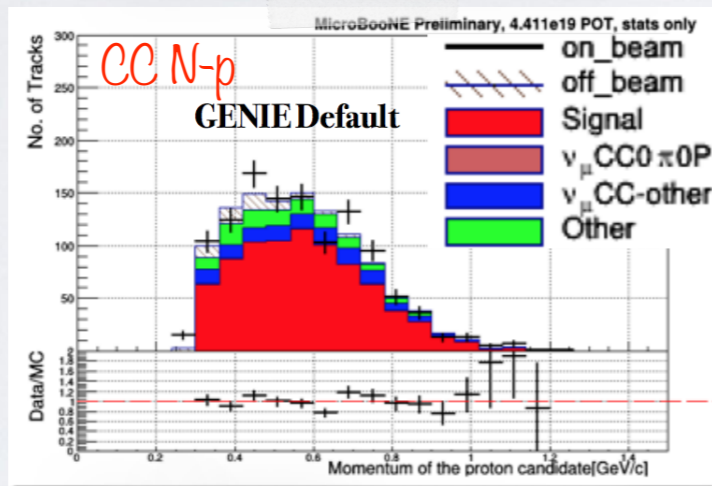
- Oscillation analysis
- Physics R&D

Invaluable experience towards understanding of LArTPC systematics!

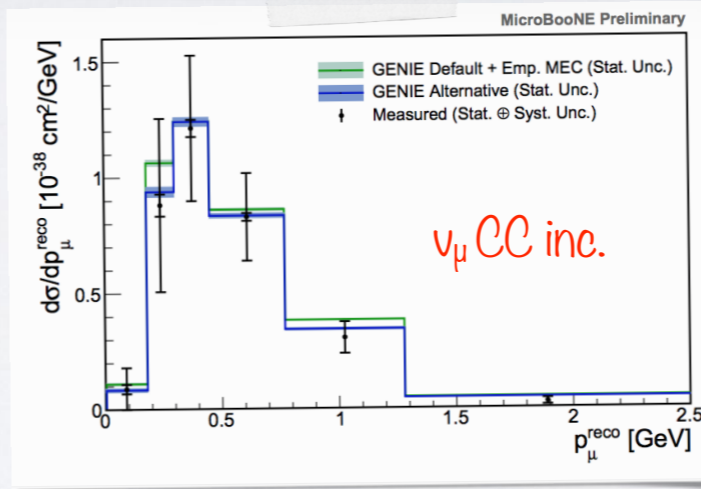
Building up momentum thanks to our first **interaction** and **cross sections measurements**:

- Charged Particle Multiplicity
- ν_μ CC inclusive
- CC π^0
- CC N-p
- NuMI CC ν_e

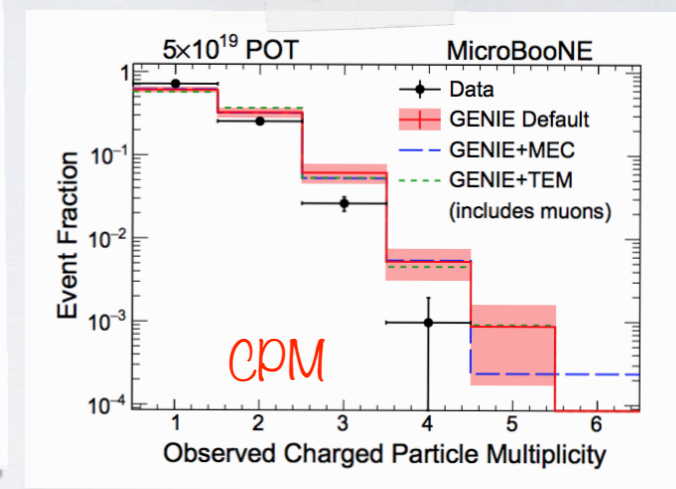
CC N-p NuInt18 talk:
https://indico.cern.ch/event/703880/contributions/3159004/attachments/1735013/2805892/Raquel_NuInt.pdf



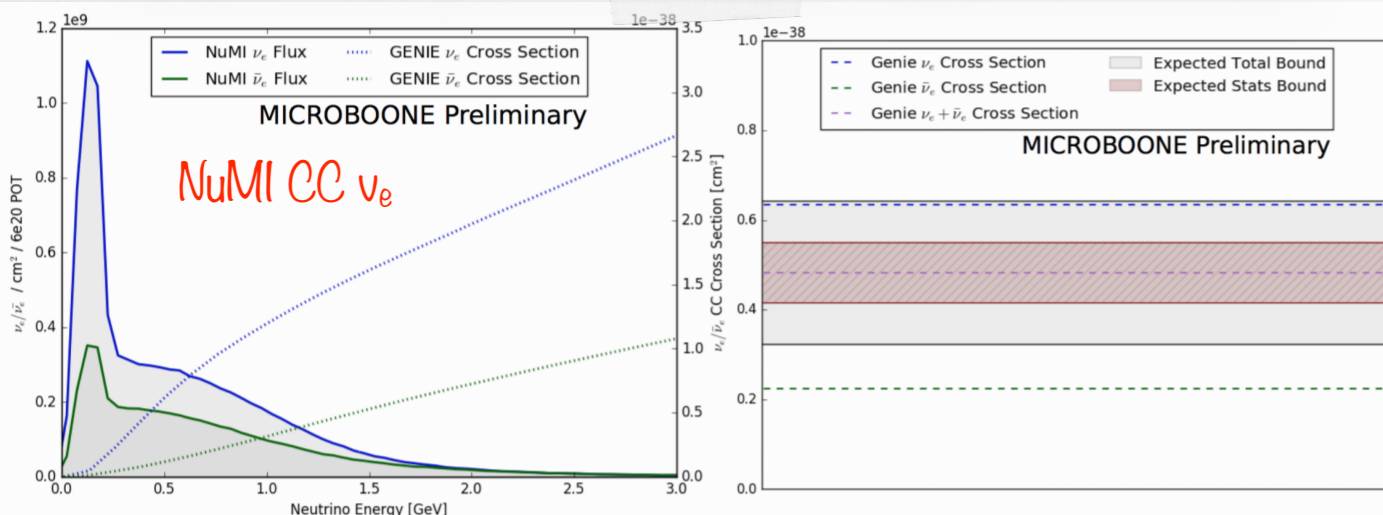
ν_μ CC Inclusive Cross Section MICROBOONE-NOTE-1045-PUB:
<http://microboone.fnal.gov/wp-content/uploads/MICROBOONE-NOTE-1045-PUB.pdf>



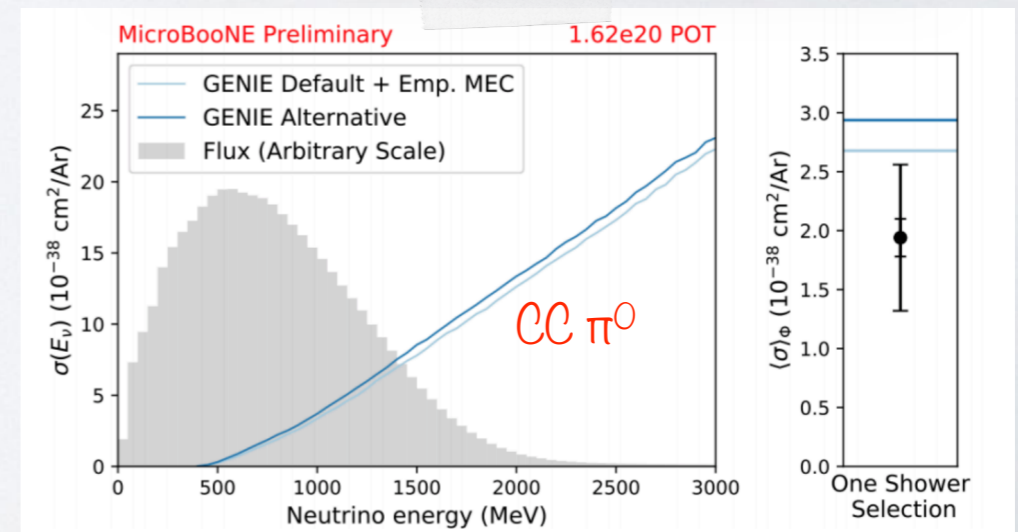
ν -Ar Multiplicity Distribution comparison to GENIE:
<https://arxiv.org/abs/1805.06887>



NuMI CC ν_e NuInt18 talk:
https://indico.cern.ch/event/703880/contributions/3159003/attachments/1734968/2805811/colton_hill_nuint_2018.pdf



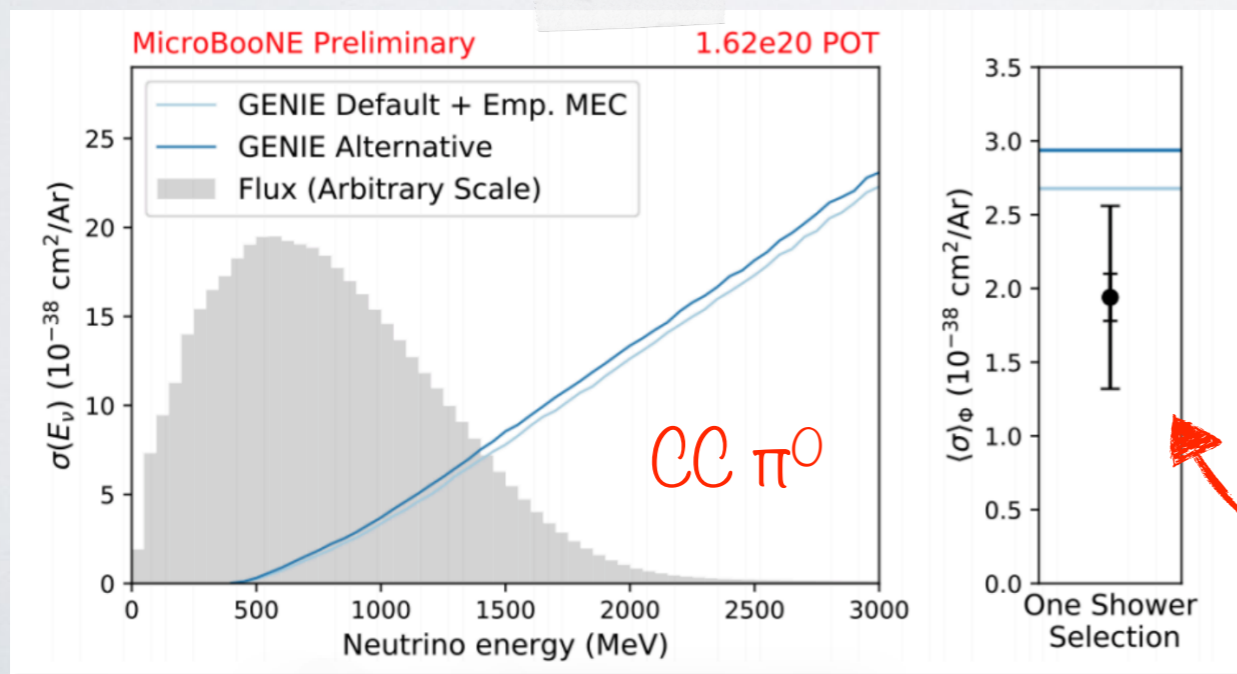
CC π^0 Cross Section MICROBOONE-NOTE-1032-PUB:
<http://microboone.fnal.gov/wp-content/uploads/MICROBOONE-NOTE-1032-PUB.pdf>



APPROACH TO SYSTEMATICS

- Approach based on:
 - **Simulating / re-weighting** parameters
 - **Propagating effects** through the full chain

- Use 1σ variations from central value, **constrained by data** when available or using an **alternative underlying model**.
- Many parameters have not been fully constrained yet by **internal or external measurements**.
- Very **high statistics**, need better understanding of systematics than previously ever needed.
- First iteration has **conservative estimates** on systematic uncertainties.
- But it will lead us to more **precise constraints** for the next iterations, providing us with **invaluable experience**.



CC π^0 Cross Section MICROBONE-NOTE-1032-PUB

<http://microboone.fnal.gov/wp-content/uploads/MICROBOONE-NOTE-1032-PUB.pdf>

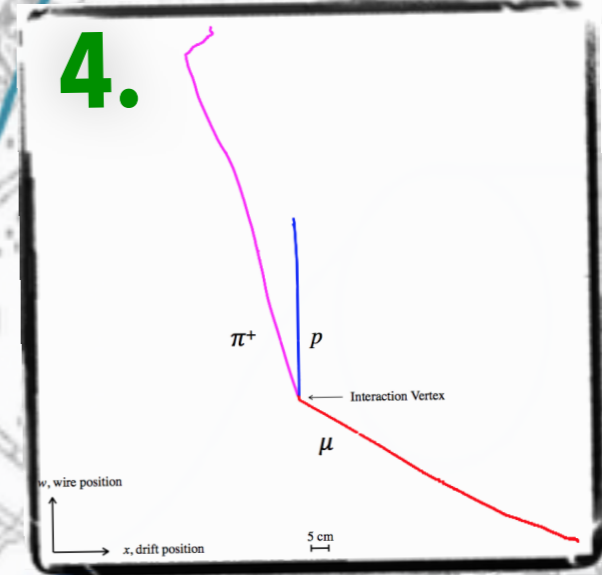
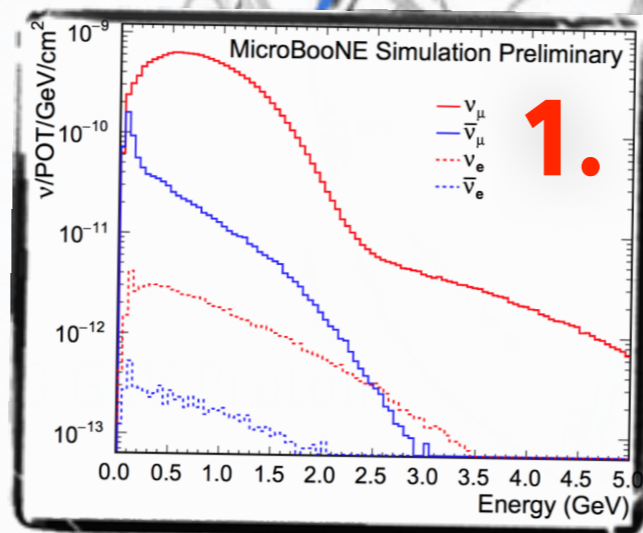
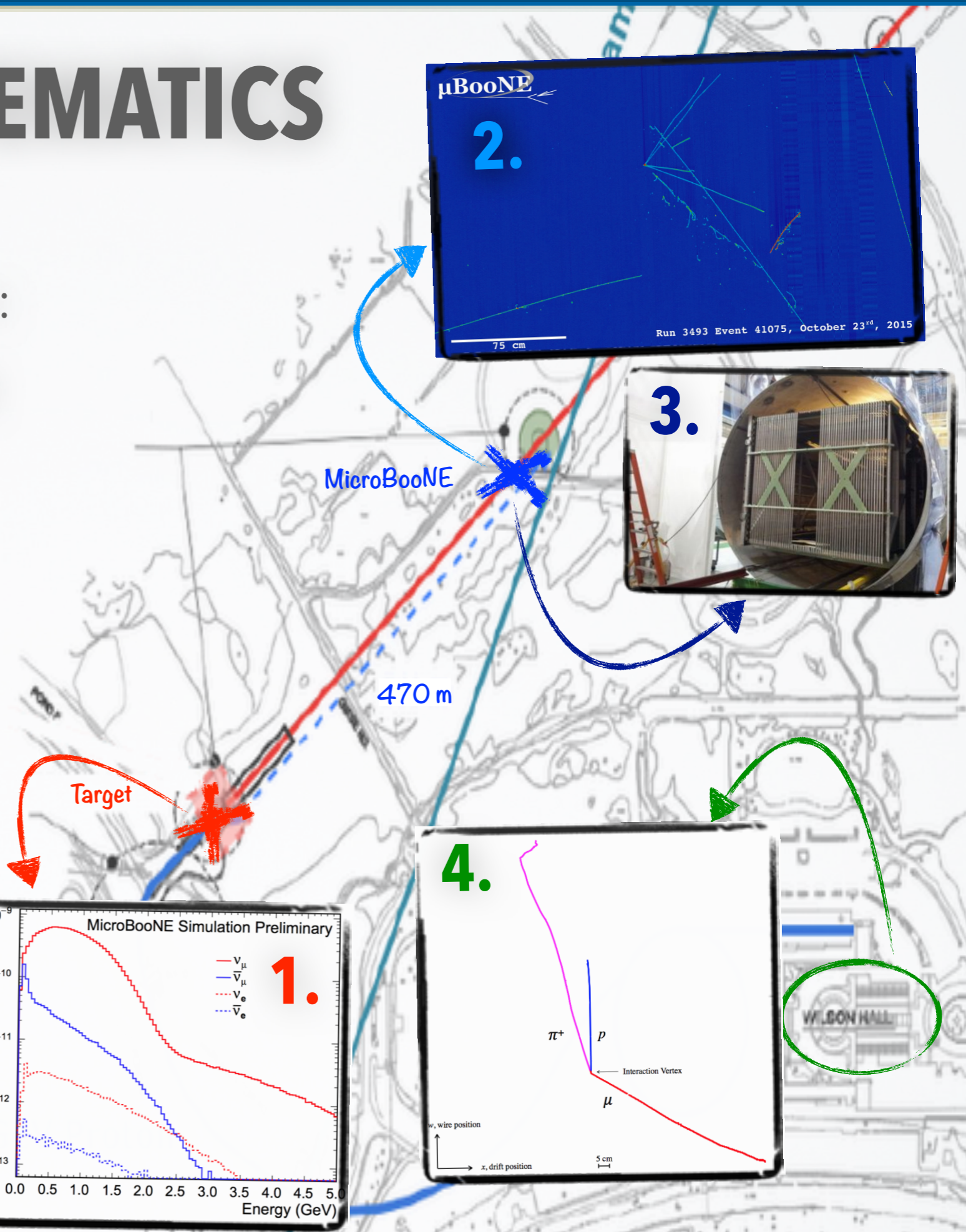
Example from CC π^0

$$\sigma_{\text{syst}} \sim 4 \cdot \sigma_{\text{stat}}$$

SOURCE OF SYSTEMATICS

Main sources of **uncertainties** from:

1. Beam flux and POT counting
2. Cross section modelling
3. Detector response
4. Reconstruction



1. BEAM FLUX

- **BNB:** Using **MiniBooNE / SciBooNE** techniques
 - **>15 years of experience** running neutrino experiments in BNB.

Hadron production uncertainties

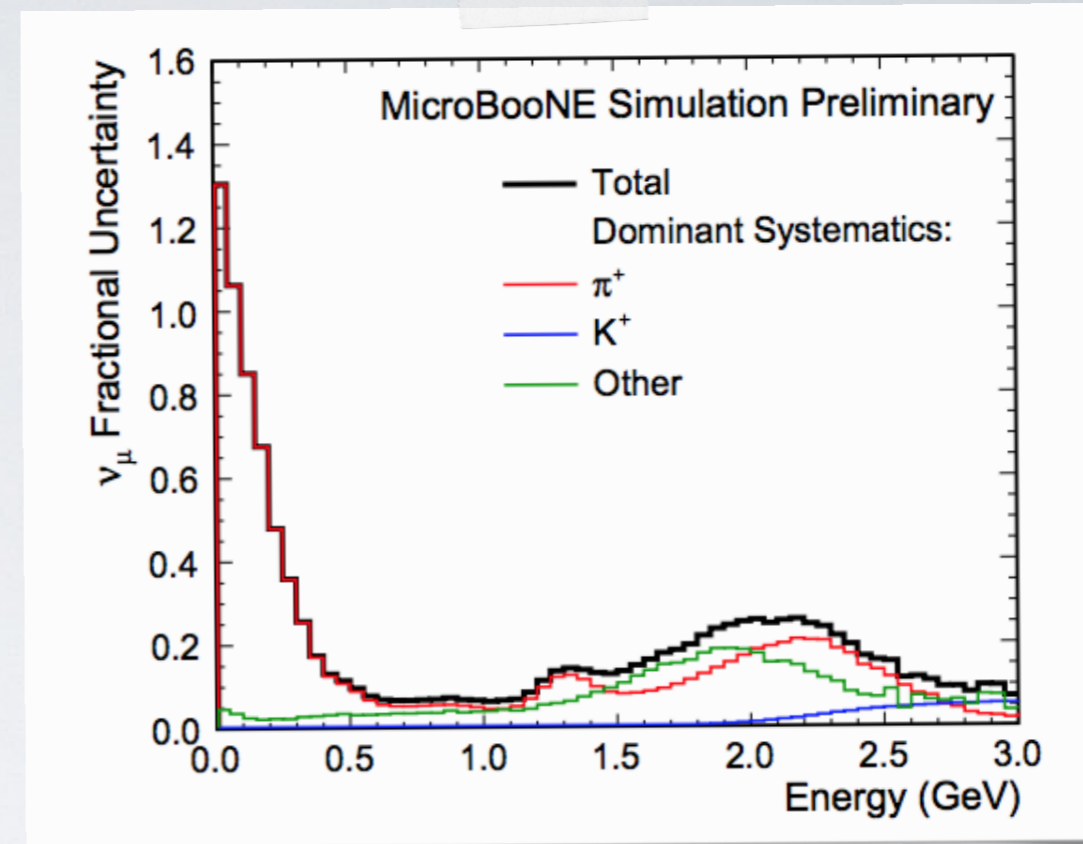
- Uncertainties in the production of secondary particles after proton collision
 - $\pi^+, \pi^-, K^+, K^-, K_L^0$
 - Based on fits to HARP data and Feynman scaling for cross-section measurements at different energies

Non-Hadron production uncertainties

- Mismodelling of horn current distribution
- Horn current miscalibration
- Pion and nucleon scattering cross sections on Be and Al

Strategy: **CORRELATED UNIVERSES SIMULATION** via reweighting

- Additional uncertainty on **normalization** from **proton delivery**, measured independently by two toroids.



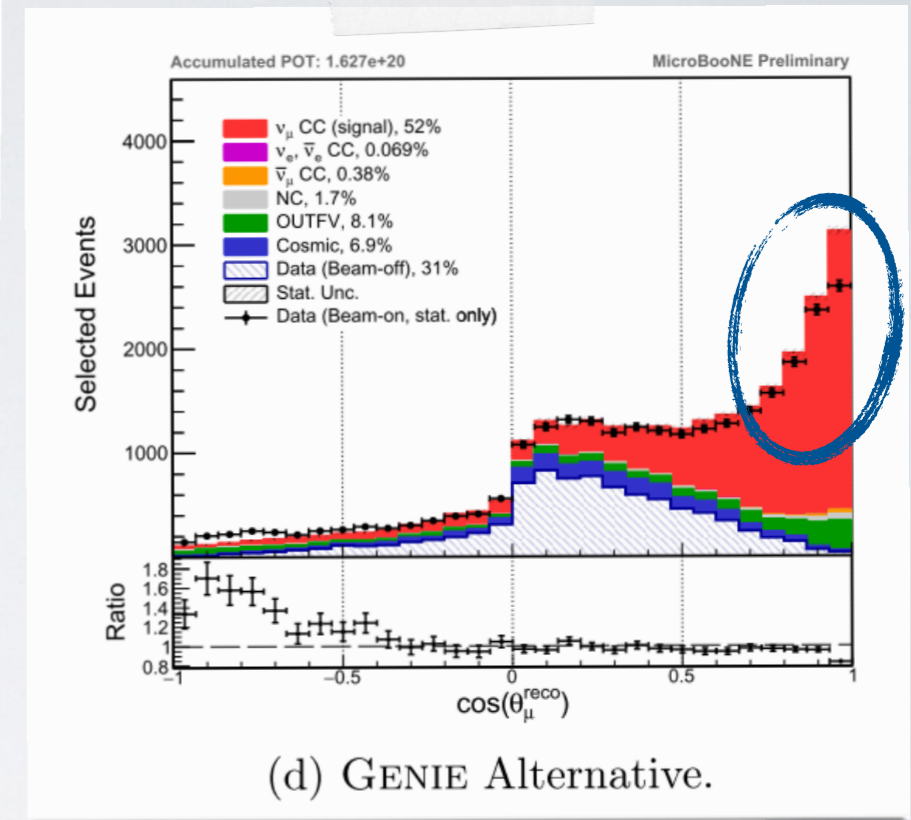
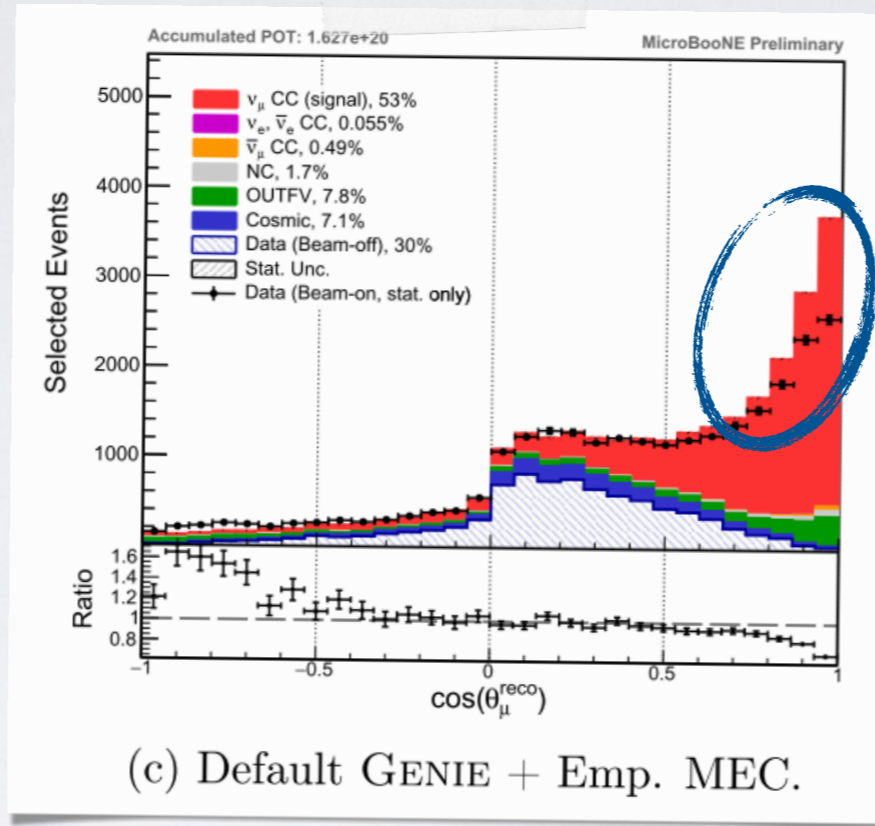
BNB Flux Prediction Public Note **MICROBOONE-NOTE-1031-PUB**:
<http://microboone.fnal.gov/wp-content/uploads/MICROBOONE-NOTE-1031-PUB.pdf>

| Systematic | $\nu_\mu/\%$ | $\bar{\nu}_\mu/\%$ | $\nu_e/\%$ | $\bar{\nu}_e/\%$ |
|-----------------|--------------|--------------------|------------|------------------|
| Proton delivery | 2.0 | 2.0 | 2.0 | 2.0 |
| π^+ | 11.7 | 1.0 | 10.7 | 0.03 |
| π^- | 0.0 | 11.6 | 0.0 | 3.0 |
| K^+ | 0.2 | 0.1 | 2.0 | 0.1 |
| K^- | 0.0 | 0.4 | 0.0 | 3.0 |
| K_L^0 | 0.0 | 0.3 | 2.3 | 21.4 |
| Other | 3.9 | 6.6 | 3.2 | 5.3 |
| Total | 12.5 | 13.5 | 11.7 | 22.6 |

- **NuMI:** Similar approach as BNB but taking care of off-axis angle. Following MINERVA and NOVA techniques.

2. CROSS SECTION MODELLING

- ν -Ar Cross-sections difficult to predict.
 - Large nucleus, complicated nuclear effects
- First generation analyses using default **GENIE v2** with **empirical MEC**,
 - Using also **alternative tune** with Valencia QE/MEC.
- Recent results from ν_{μ} CC have shown **better agreement** with alternative tune.



| Model element | Default GENIE + Emp. MEC | GENIE Alternative |
|-------------------------|--------------------------|-------------------|
| Nuclear Model | Bodek-Ritchie Fermi Gas | Local Fermi Gas |
| Quasi-Elastic | Llewellyn-Smith | Nieves |
| Meson-exchange Currents | Empirical | Nieves |
| Resonant | Rein-Seghal | Berger-Seghal |
| Coherent | Rein-Seghal | Berger-Seghal |
| FSI | hA | hA2014 |

2. CROSS SECTION MODELLING

- For the future planning transition to **GENIE v3** using either:
 - **G18_10a**
 - Theory-driven comprehensive model
 - **G18_10i**
 - Like G18_10a
 - Using z expansion in place of the dipole model for the QE axial form factor.
- **State of the art** generators.
- Expected to lead to **further reduction** in uncertainties

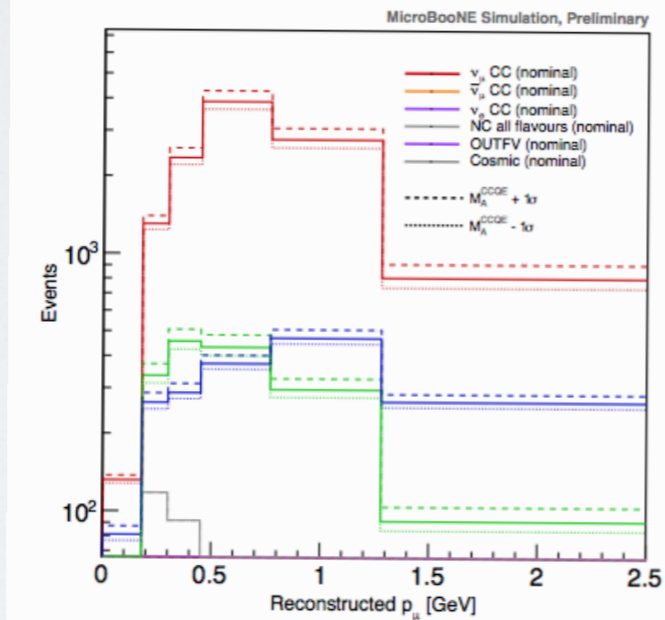
| Model Element | G18_10a |
|-------------------------|-----------------|
| Nuclear Model | Local Fermi Gas |
| Quasi-Elastic (CC) | Nieves |
| Quasi-Elastic (NC) | Empirical |
| Meson-exchange Currents | Nieves |
| Resonant | Berger-Sehgal |
| Coherent | Berger-Sehgal |
| FSI | hA2018 |

Strategy: **CORRELATED UNIVERSES SIMULATION** via reweighting

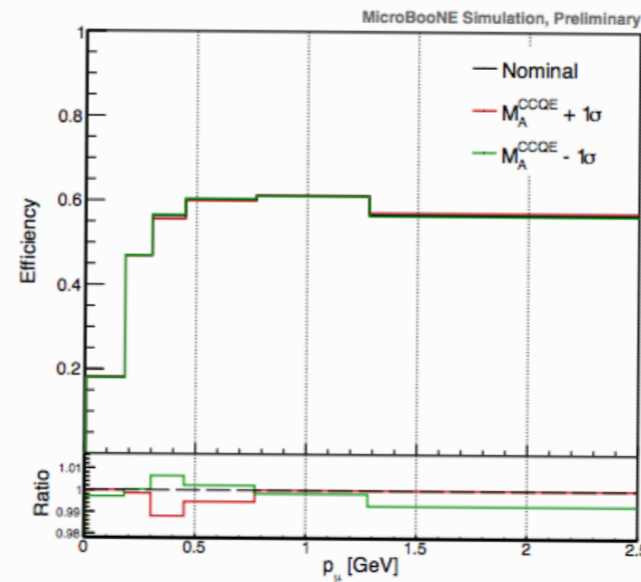
2. CROSS SECTION MODELLING

- Initial estimates varying parameters within standard 1σ estimates from GENIE manual.
- Developing in addition a re-weighting strategy for **QE (RPA)** and **MEC** based on differences between **GENIE default** and **Valencia QE/MEC** models.
 - Planning on importing **MINERvA/NOvA** RPA/MEC uncertainties treatment in next-generation analyses.

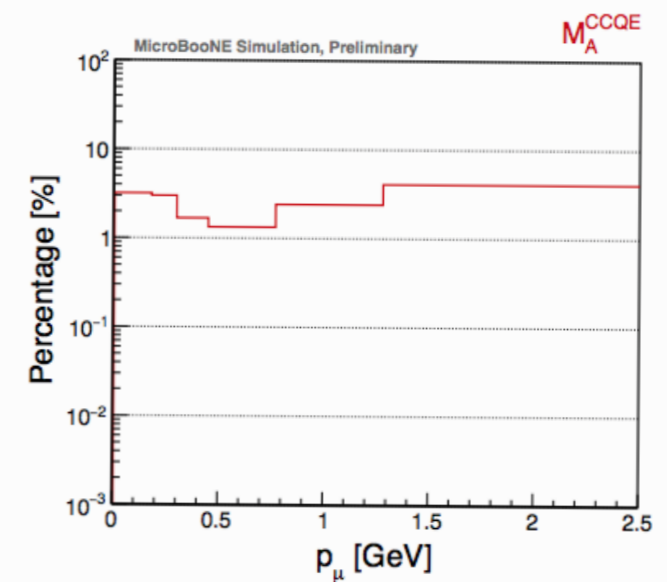
Example of the impact of the variation of M_A^{CCQE} axial mass parameter on ν_μ CC inclusive



(a) Number of events



(b) Selection efficiency



(c) Cross section percentual difference

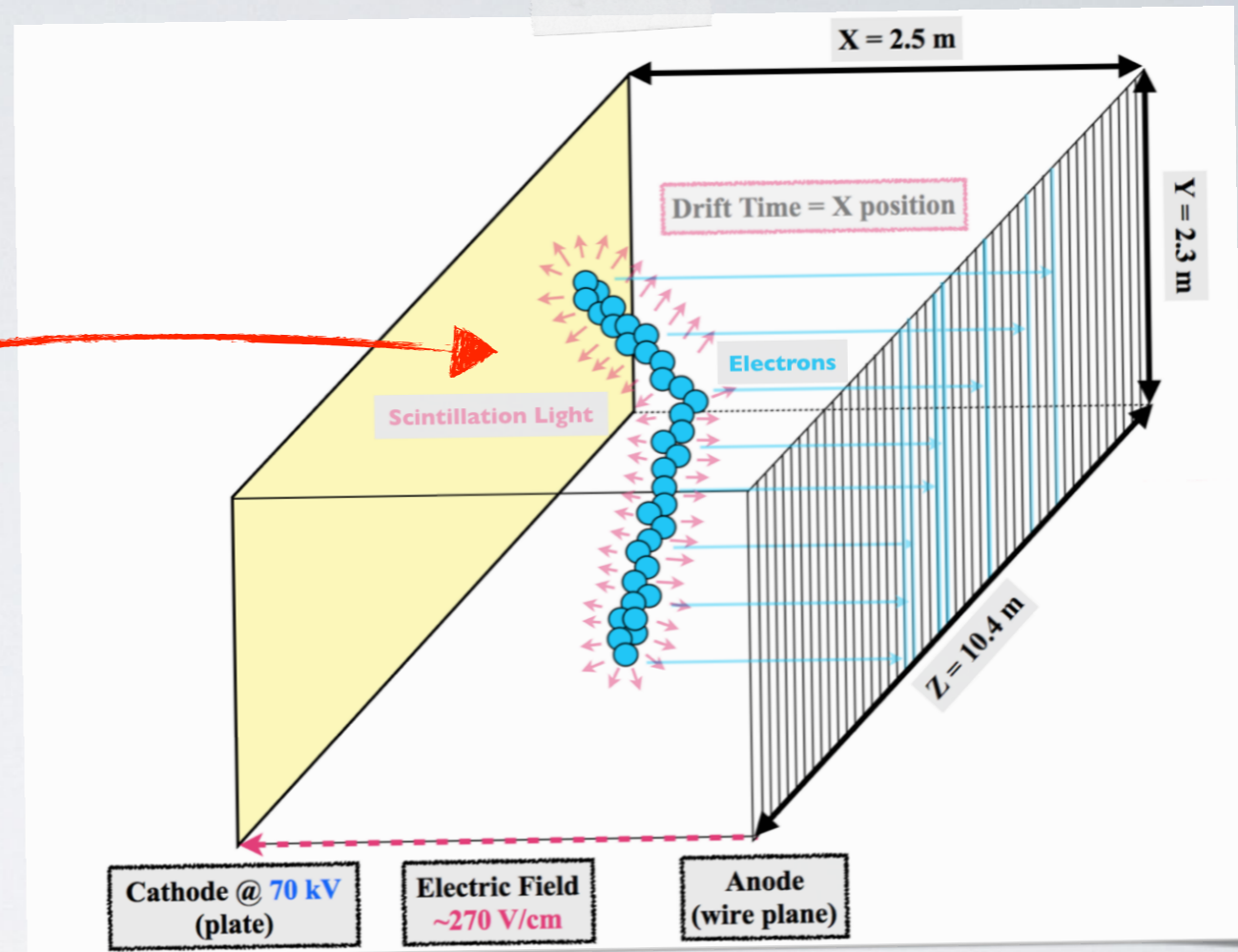
- Progress in the implementation of GEANT4 uncertainty on **hadron reinteraction** with bulk Argon.

RPA: Random Phase Approximation (*screening effect due to W polarization leading to cross-section suppression at low q^0*)

MEC: Meson Exchange Current (*interaction with correlated nucleon states populating multi-nucleon final states*)

3. DETECTOR RESPONSE

- Two main sub-categories:
 - Propagation of scintillation **light** and **electrons** through the detector and towards PMTs and wire plane
 - Readout response from *PMTs* and *wires*



| Analysis | Inclusive | ccPi0 |
|-----------------------------------|-----------|-------|
| Uncertainties | | |
| Cross section model uncertainties | 3.63% | 16% |
| Flux | 11.93% | 17% |
| Detector | 18.90% | 21% |



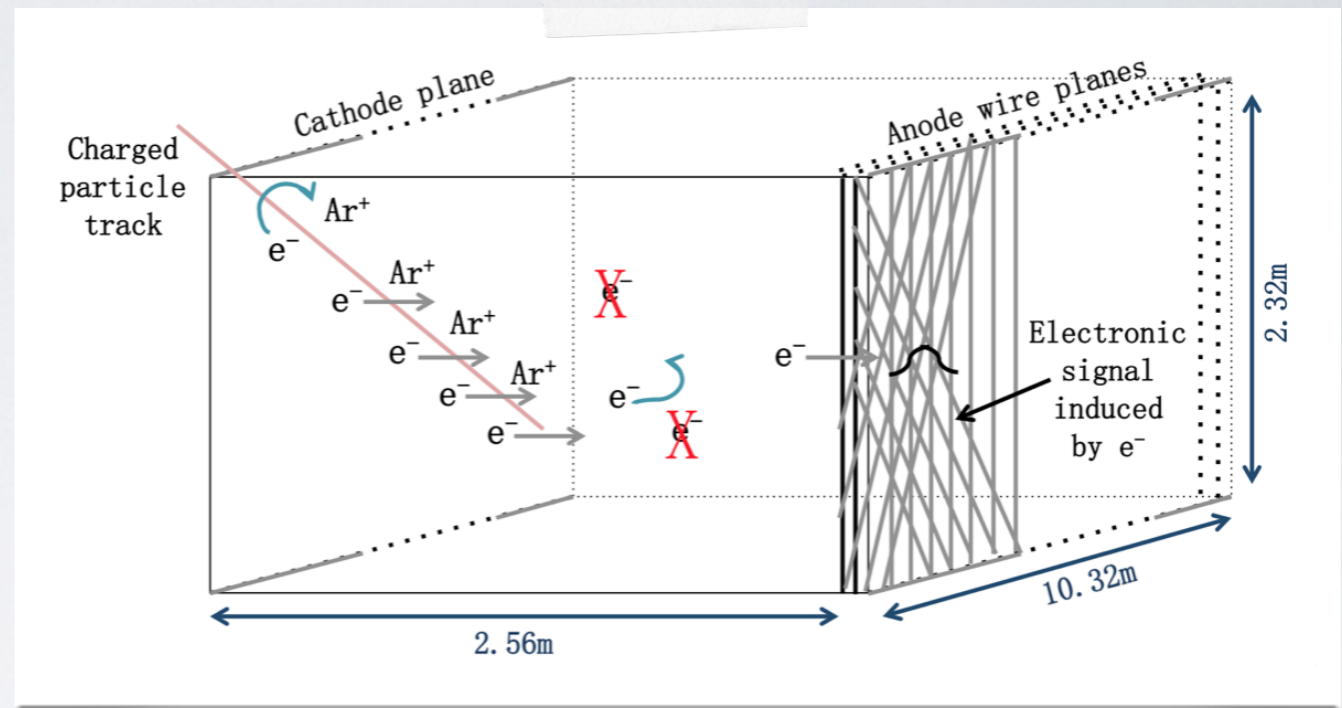
- Largest contribution so far to total uncertainty.

Uncertainty strategy: **UNCORRELATED UNIVERSES SIMULATION*** via whole resimulation

*all data samples use identical generated interactions

3A. LIGHT AND ELECTRON PROPAGATION

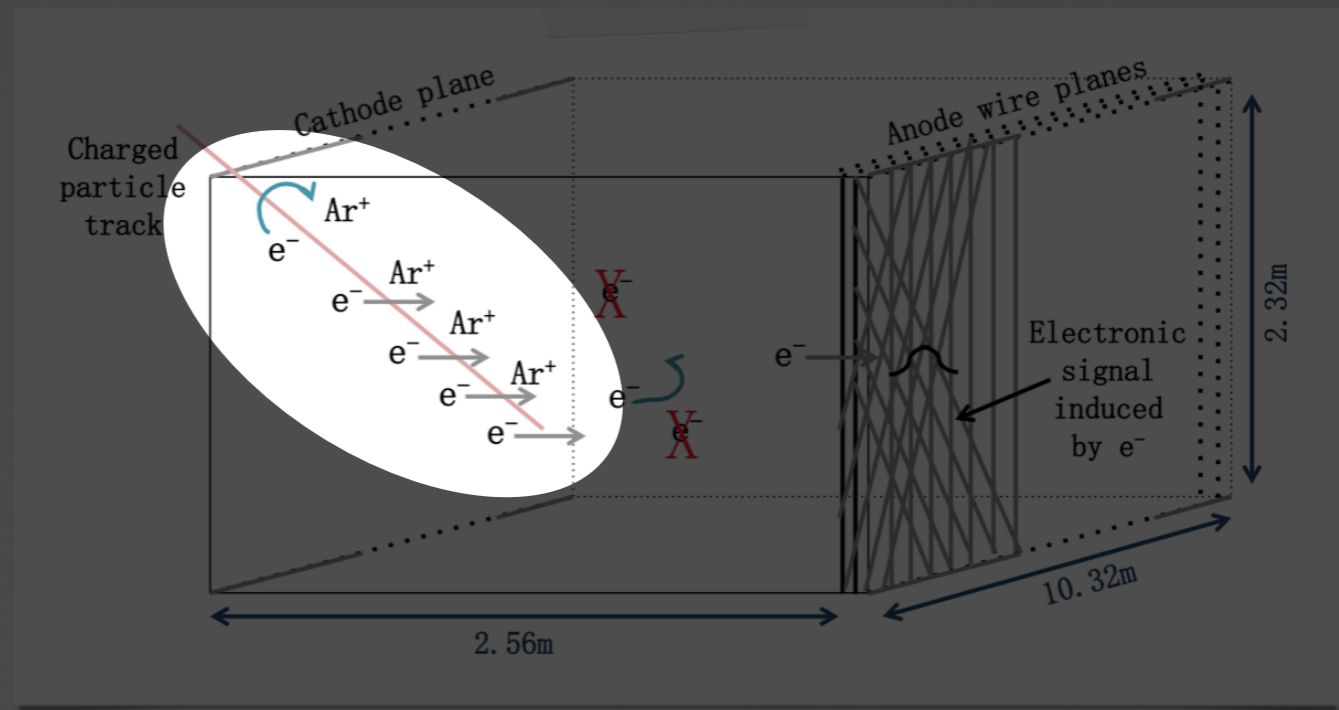
- Assuming a constant field, when propagating toward the wire plane, ionization electrons go through **multiple effects**.



$$\frac{dE}{dX} \times \frac{1}{W} \times R \times L \times D \times C = \frac{dQ}{dX}$$

3A. LIGHT AND ELECTRON PROPAGATION

- Assuming a constant field, when propagating toward the wire plane, ionization electrons go through **multiple effects**.



Energy lost by charged particle through ionization

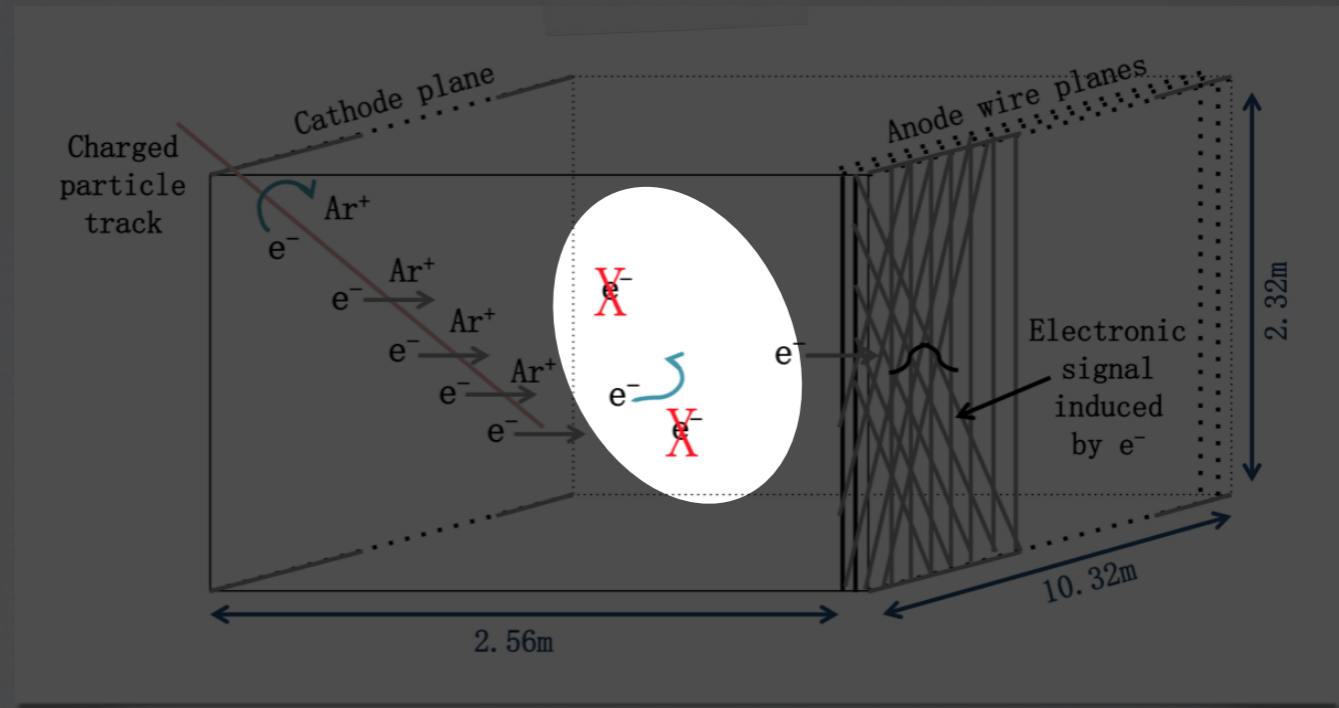
Recombination

$$\frac{dE}{dX} \times \frac{1}{W} \times R \times L \times D \times C = \frac{dQ}{dX}$$

Energy needed to free an e⁻

3A. LIGHT AND ELECTRON PROPAGATION

- Assuming a constant field, when propagating toward the wire plane, ionization electrons go through **multiple effects**.



Energy lost by charged particle through ionization

Recombination

Diffusion

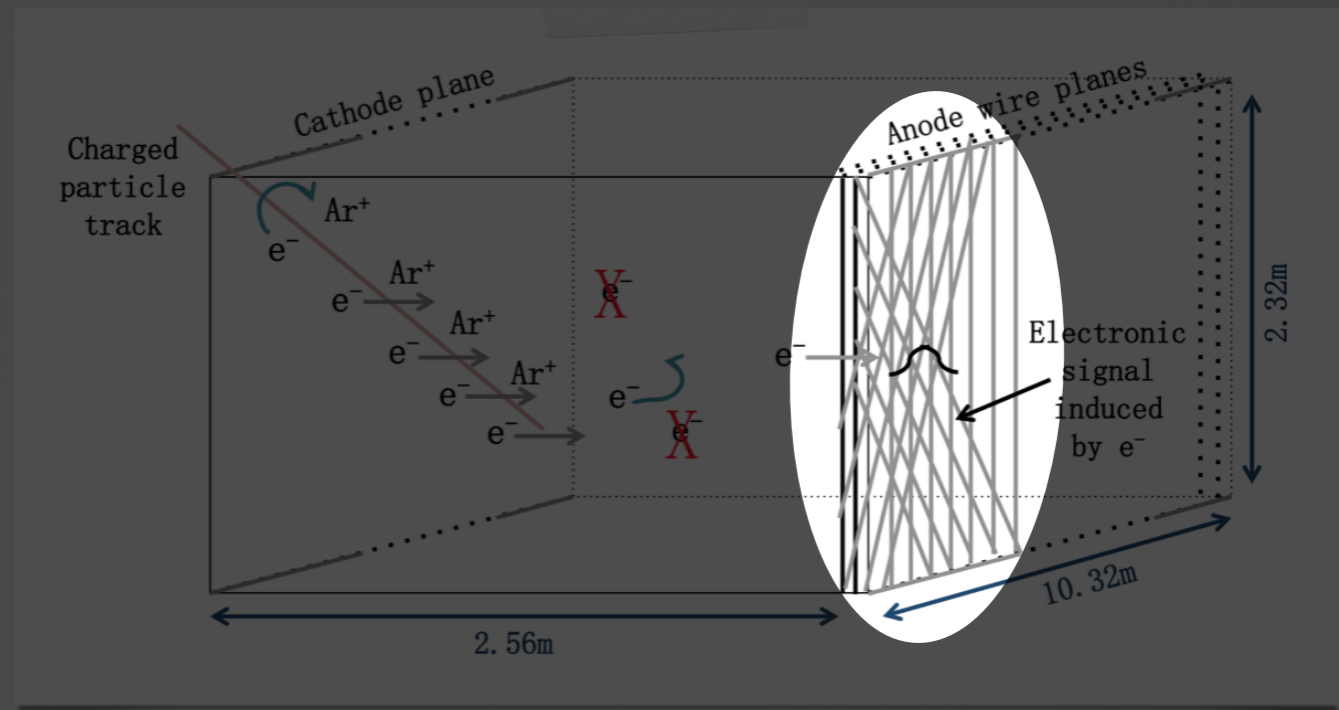
$$\frac{dE}{dX} \times \frac{1}{W} \times R \times L \times D \times C = \frac{dQ}{dX}$$

Energy needed to free an e^-

Attenuation

3A. LIGHT AND ELECTRON PROPAGATION

- Assuming a constant field, when propagating toward the wire plane, ionization electrons go through **multiple effects**.



Energy lost by charged particle through ionization Recombination Diffusion Calibrated signal on the anode

$$\frac{dE}{dX} \times \frac{1}{W} \times R \times L \times D \times C = \frac{dQ}{dX}$$

Energy needed to free an e^- Attenuation Calibration of electronic signal

3A. LIGHT AND ELECTRON PROPAGATION

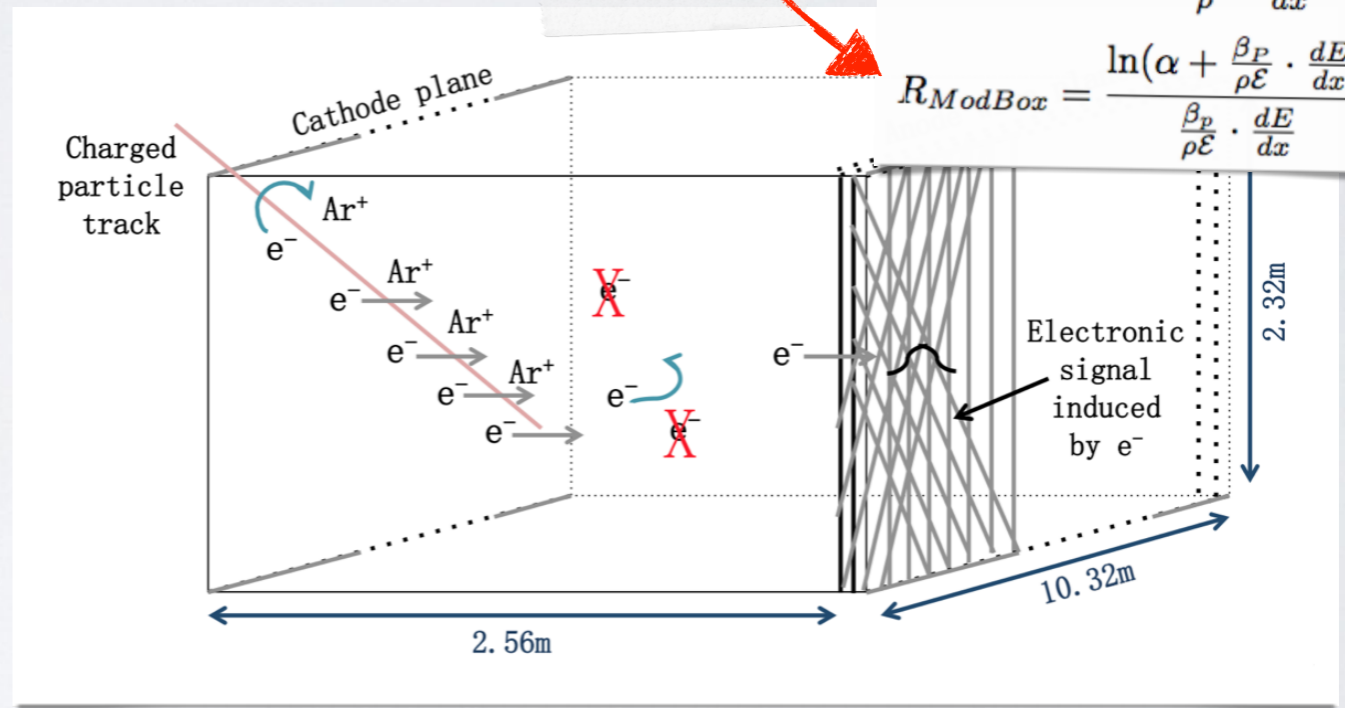
RECOMBINATION

- Before transport, the electron can **recombine** with the **Ar⁺ ions**.
- Effect depends on **density of Ar⁺ and e⁻**.
 - Affected by dE/dX and E-field strength.
- Using
 - Modified **box model** with parameters fit to **ArgoNeuT** data
 - **Birks model** fitted to **ICARUS** data for the systematic variations.

$$\frac{dE}{dX} \times \frac{1}{W} \times R \times L \times D \times C = \frac{dQ}{dX}$$

$$R_{Birk's} = \frac{A_B}{1 + \frac{k_B}{\rho} \cdot \frac{dE}{dx} \cdot \frac{1}{E}}$$

$$R_{ModBox} = \frac{\ln(\alpha + \frac{\beta_P}{\rho E} \cdot \frac{dE}{dx})}{\frac{\beta_P}{\rho E} \cdot \frac{dE}{dx}}$$



Current implementation:
Using alternative model externally constrained with other argon detectors measurements

Future implementation:
Internal constraints from ongoing MicroBooNE calibration measurements

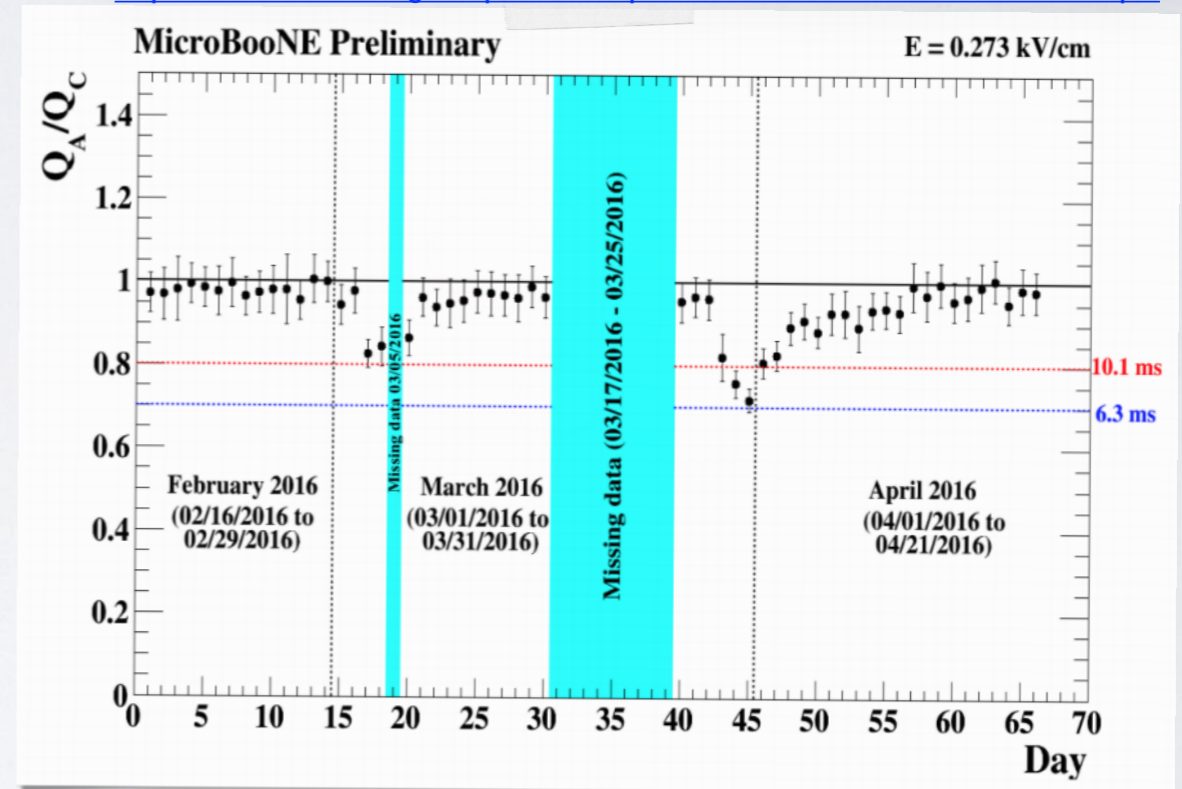
3A. LIGHT AND ELECTRON PROPAGATION

ELECTRON ATTENUATION

$$\frac{dE}{dX} \times \frac{1}{W} \times R \times L \times D \times C = \frac{dQ}{dX}$$

MicroBooNE Attenuation measurements Public Note MICROBOONE-NOTE-1026-PUB:
<http://microboone.fnal.gov/wp-content/uploads/MICROBOONE-NOTE-1026-PUB.pdf>

- **Contaminants** in the liquid argon, such as **oxygen** and **water** can **capture** the drifting electron.
 - Causing a drift-distance dependent **reduction** of the size of the **signal**.
- Using **electron drift lifetime** to characterize attenuation.
- MicroBooNE design electron lifetime: **3ms**
- We're currently well **above 10ms**.
- Simulating most **extreme value** allowed by run quality selection (10 ms, only ~10% of data at this purity) for systematic variations.



Current implementation:

Infinite lifetime for default
 "Extreme case" simulation (10 ms) for uncertainty

Future implementation:

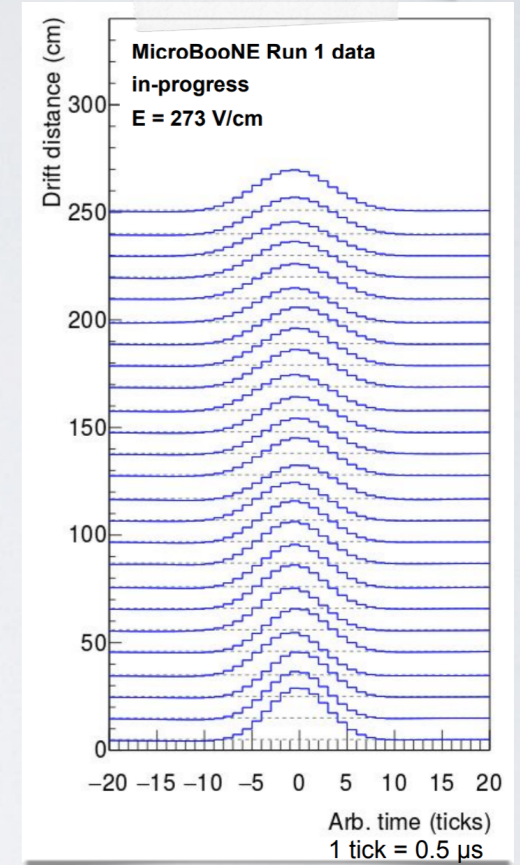
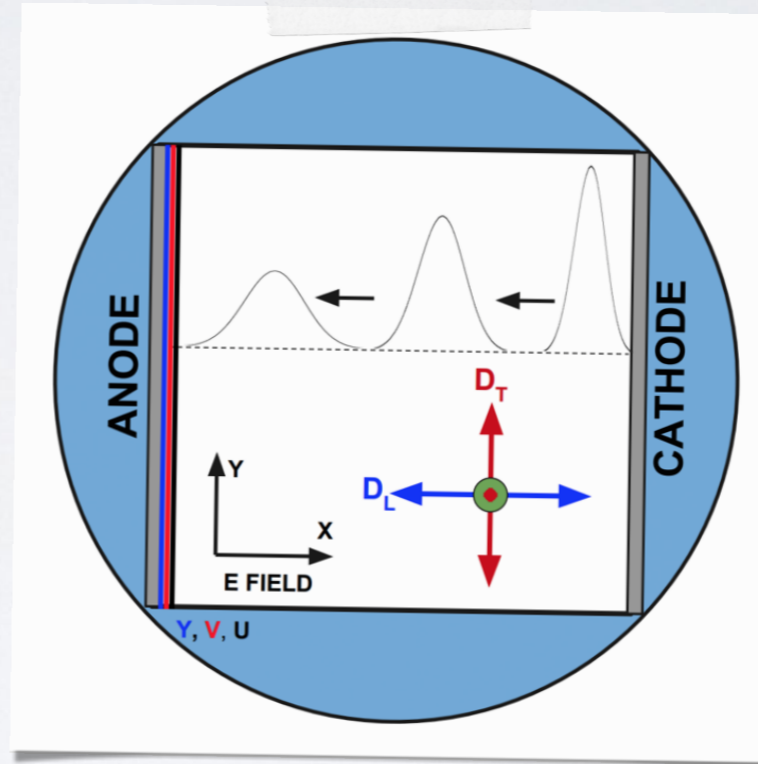
- Ongoing MicroBooNE calibration measurements to disentangle it from other effects.

3A. LIGHT AND ELECTRON PROPAGATION

LONGITUDINAL & TRANSVERSE DIFFUSION

$$\frac{dE}{dX} \times \frac{1}{W} \times R \times L \times D \times C = \frac{dQ}{dX}$$

- Ionisation electrons travel due to electric field.
- During transport, the **shape** of the electron cloud is **smear**ed.
 - Depending on **distance** to wire plane.
 - Separated in:
 - **Longitudinal** component
(parallel to travel direction)
 - **Transverse** component
(perpendicular to travel direction)



Current implementation: $I\sigma$

- External constraints on argon for L
- T less studied in literature, using external measurements and theoretical extrapolation to MicroBooNE field strength

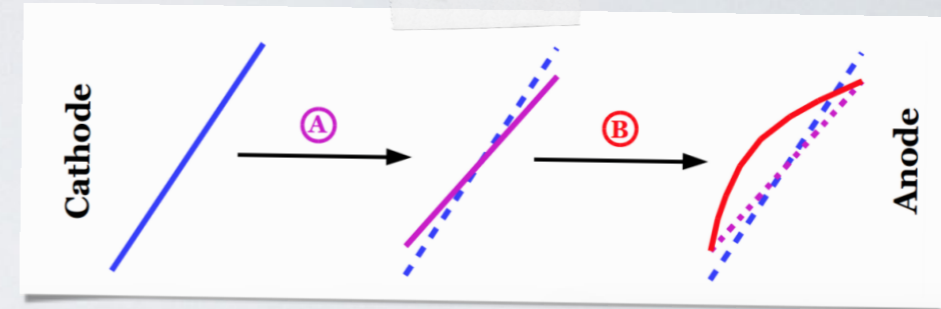
Possible future implementation:

- Internal constraints from ongoing MicroBooNE calibration measurements
- ^{39}Ar to disentangle L/T components

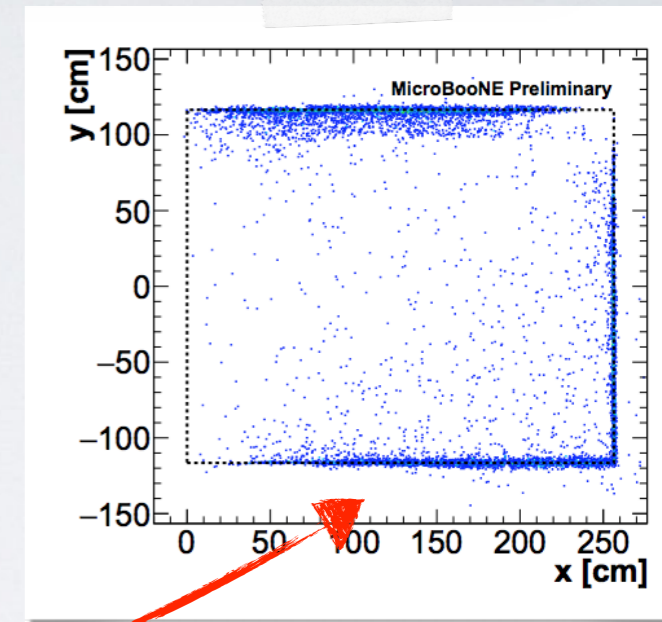
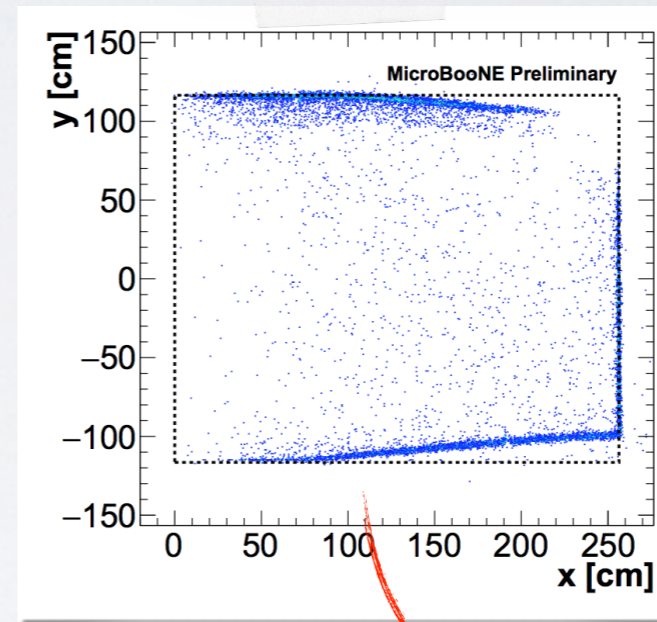
3A. LIGHT AND ELECTRON PROPAGATION

SPACE CHARGE

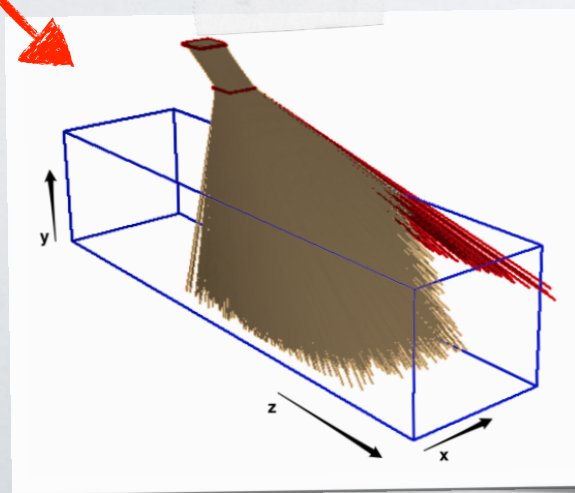
- Positive ions drift **105 times** slower than e^- in LAr
- **Build-up of Ar^+ ions** in steady-state configuration due to ionization **from cosmic rays** (3kHz).
- Leads to a **distortion of the electric field** within the detector.
 - **Displacement** in the reconstructed position of signal ionization electrons in LArTPC detectors (spatial distortion)
 - Variation in E-field strength (calorimetry distortion)
- Using a **data-driven correction** determined with **through-going cosmic muons** measured with an **external muon telescope** (MuCS).
- We built a map of **electric field** and obtain variations accordingly



MicroBooNE Space-Charge Effects Public Note **MICROBOONE-NOTE-1018-PUB**:
<http://microboone.fnal.gov/wp-content/uploads/MICROBOONE-NOTE-1018-PUB.pdf>



After correction



Current implementation:
 MicroBooNE data-driven calibration
 Calibration varied by 70% for systematic treatment

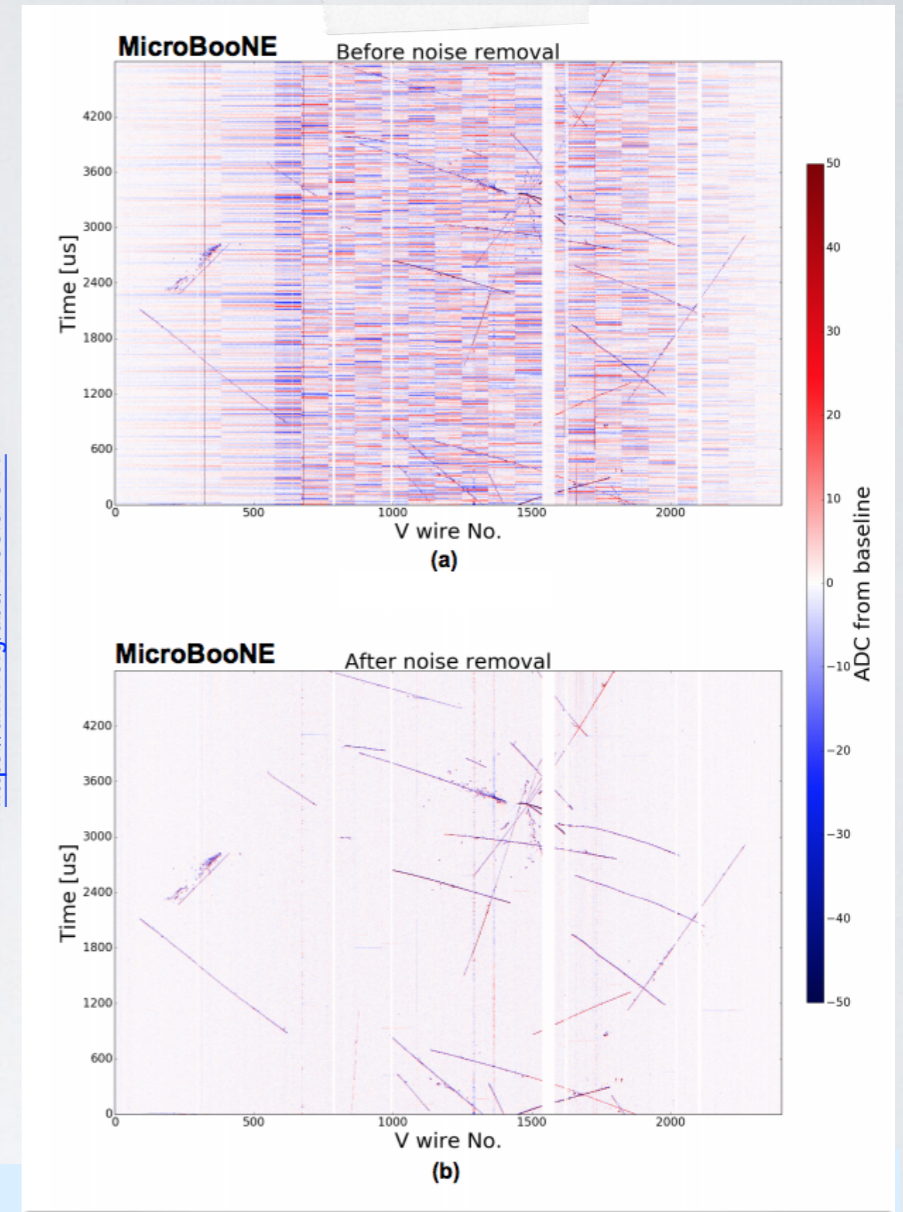
Possible future implementation:
 Improving mapping by combining laser induced tracks + cosmic measurement

3B. READOUT RESPONSE

NOISE, RESPONSE FUNCTION AND SATURATED/MISCONFIGURED CHANNELS

- Using MicroBooNE measured **noise spectrum** in data.
- Determined list of channels associated with:
 - Cold ASICs circuits that have a **different gain** and **shaping time** than desired. (*misconfigured channels*).
 - Prone to have cold **ASIC's circuits saturated** as charge builds up on capacitors due to wire motion. (*saturated channels*)
- Determined uncertainty on **response functions** from MicroBooNE data via *narrower/wider* response function motivated by data measurements.

MicroBooNE noise characterization and filtering [arXiv:1705.07341](https://arxiv.org/abs/1705.07341)
<https://arxiv.org/abs/1705.07341>



Current implementation:

- “Extreme case” simulation for saturated/misconfigured channels, turning off channel list.
- 1σ constrained by MicroBooNE data measurements for PMT/Wire noise and response functions.

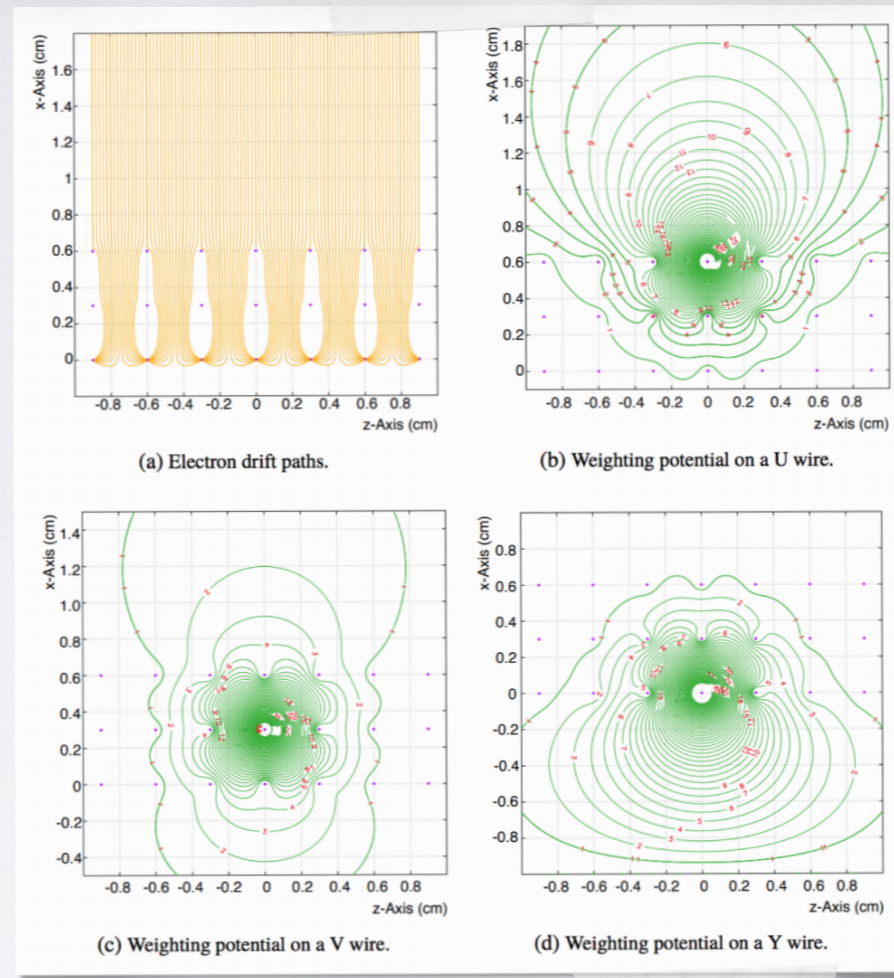
Future implementation:

- Ongoing MicroBooNE calibration measurements to improve constraints and capture an overall uncertainty
- Use of background data cosmics instead of CORSIKA simulated samples in next-generation analyses, is expected to reduce the impact of these uncertainties

3B. READOUT RESPONSE

DYNAMIC INDUCED CHARGE

- **Electrons** collected by wire **induce charge** also on **neighbouring wires**.
- Effect **observed in data**, but not currently modelled in default simulation.
 - Effect strongest for **tracks travelling towards the wire plane**.
- Simulation of dynamic induced charge (DIC) **improves data/MC agreement**.
- Determining the impact of switching DIC on/off and use as exaggerated effect for systematic variation.
- So far **largest contributor** to systematic uncertainty.



First LArTPC experiment to take this effect into account

MicroBooNE Signal processing:
 arXiv:1802.08709
 arXiv:1804.02583
<https://arxiv.org/abs/1802.08709>
<https://arxiv.org/abs/1804.02583>

| | ν_μ CC incl. | ν_μ CC π^0 |
|------------------------------------|--------------------|----------------------|
| Total systematic uncertainty | 25% | 31% |
| Detector response uncertainty | 19% | 21% |
| Dynamic induced charge uncertainty | 15% | $\approx 15\%$ |

Current implementation:

Updated vs. current simulation

Future implementation:

New simulation and reconstruction handling dynamic induced charge effects becomes default

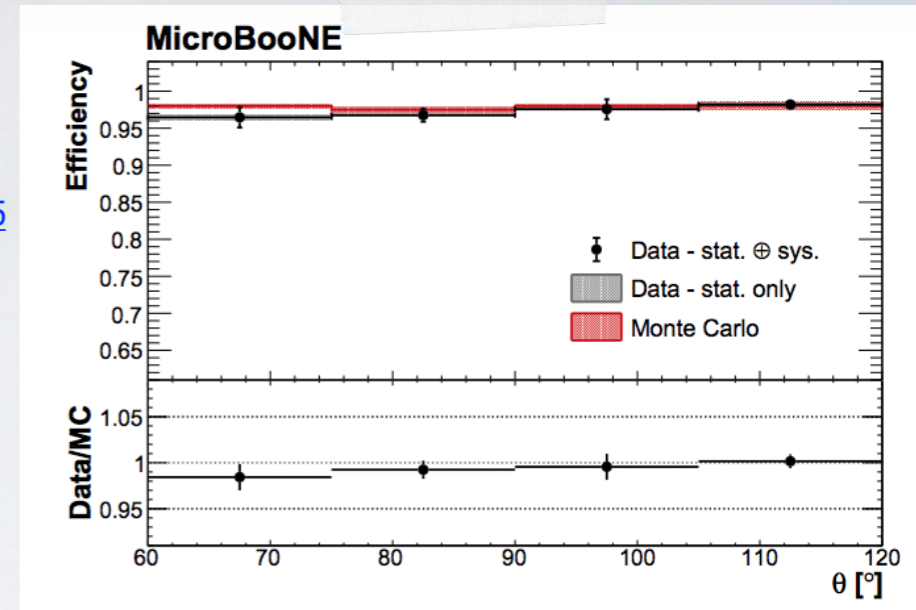
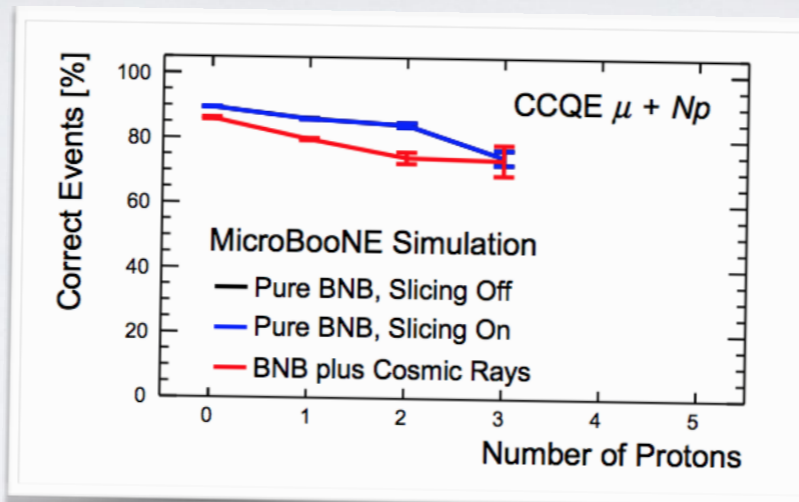
4. RECONSTRUCTION

DATA-DRIVEN RECONSTRUCTION EFFICIENCIES

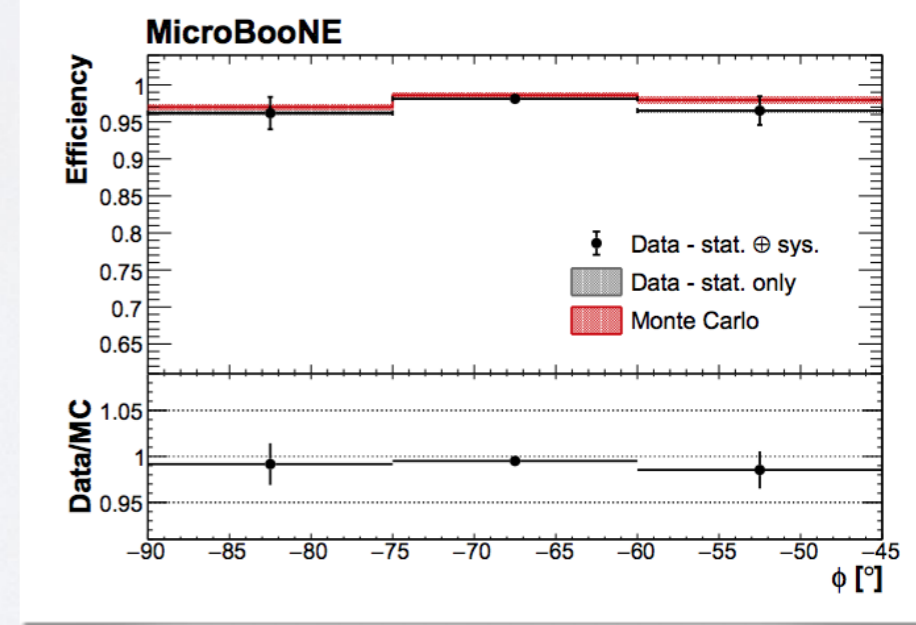
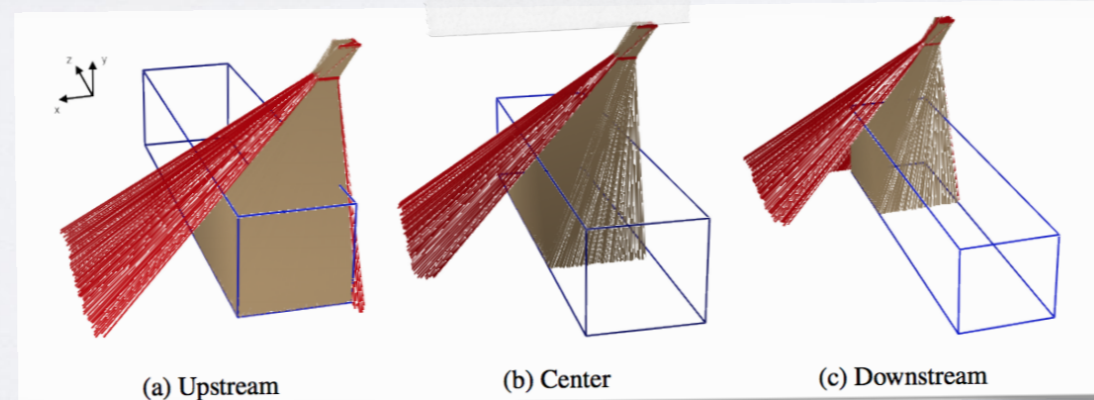
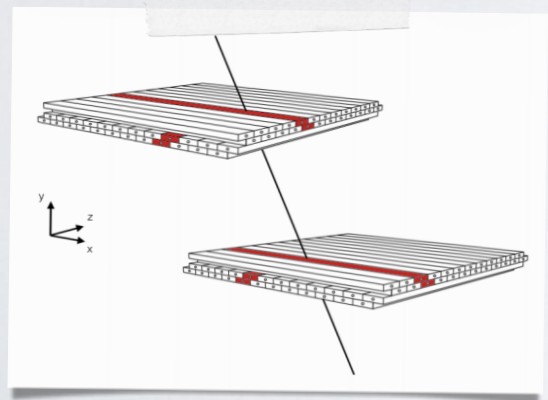
MicroBooNE reconstruction performance studies MICROBOONE-NOTE-1049-PUB:
<http://microboone.fnal.gov/wp-content/uploads/MICROBOONE-NOTE-1049-PUB.pdf>

MicroBooNE Pandora approach to pattern recognition
 Eur. Phys. J. C78 (2018) no.1, 82: <https://arxiv.org/abs/1708.03135>

- Using **Pandora** pattern recognition toolkit including multiple algorithms
- Good **reconstruction efficiency** in the presence of cosmics.
- Using an external movable **cosmic-ray telescope** to measure **data-driven efficiency** of **track reconstruction** algorithm with cosmics.



(a) θ



- Overall **high reconstruction** efficiency: $(97.1 \pm 0.1 \text{ (stat)} \pm 1.4 \text{ (sys)}) \%$
- Measured reconstruction efficiency from **data agrees with the predicted efficiency** in the simulation
 - **Confirmation** of our **simulation** and **reconstruction** chain.

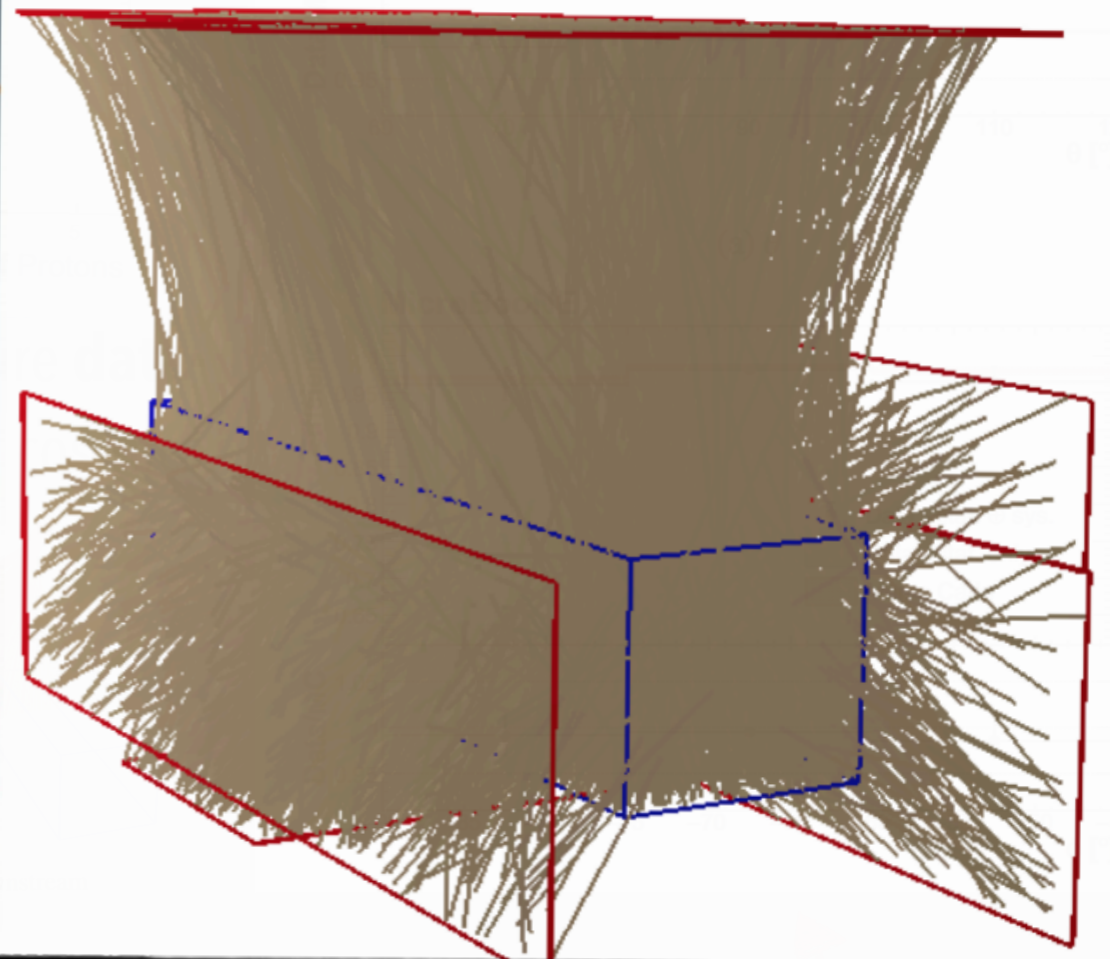
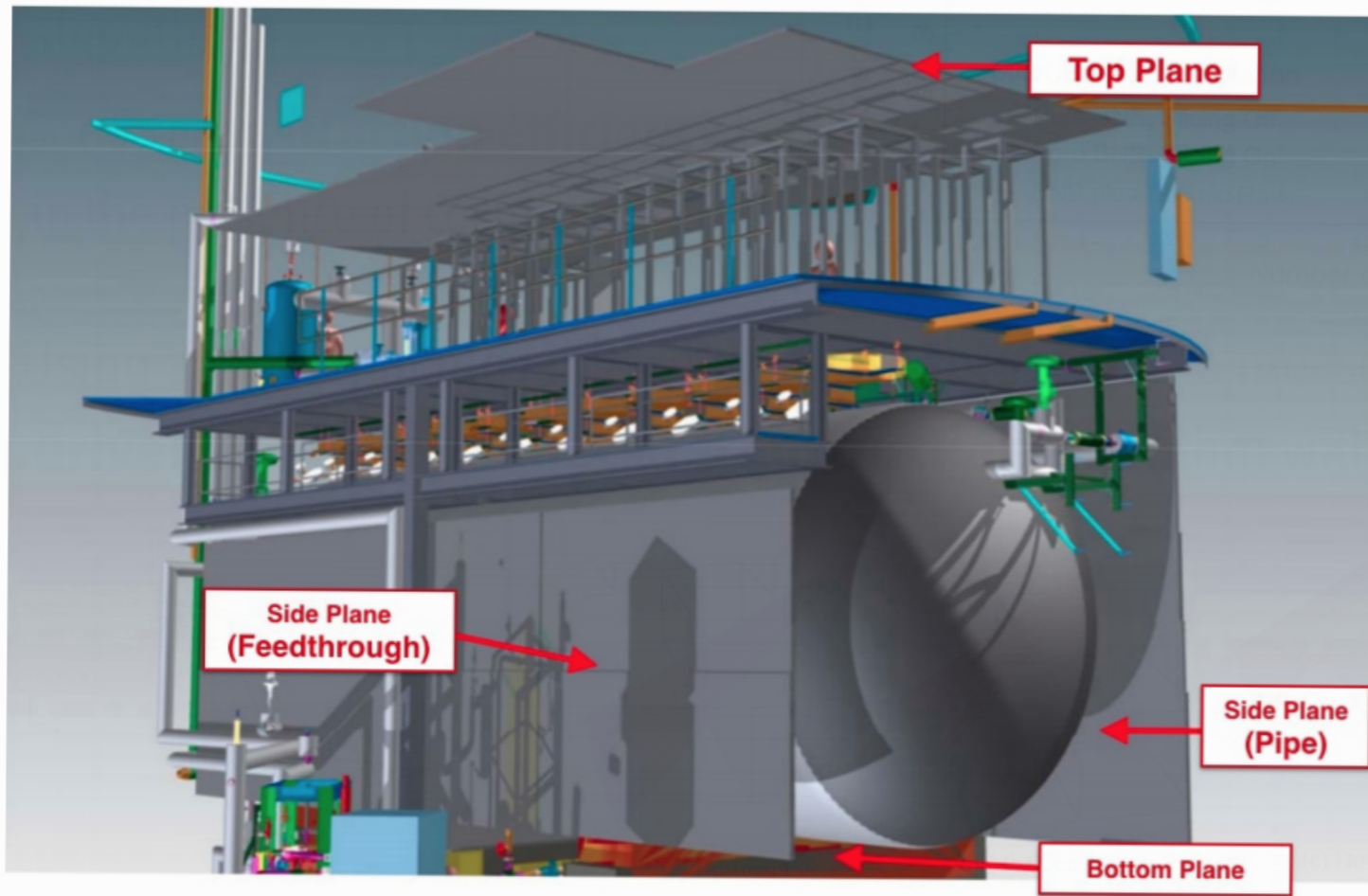
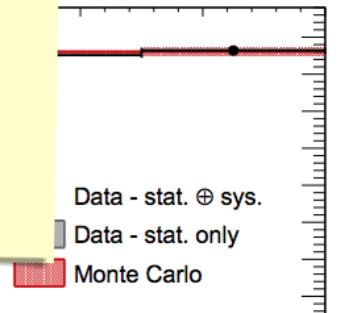
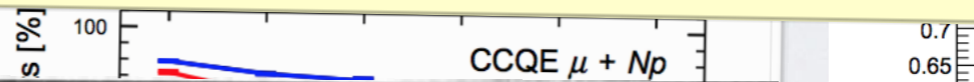


COSMIC RAY TAGGER

85% through-going muon coverage

BOONE-NOTE-1049-PUB:
BOONE-NOTE-1049-PUB.pdf

- Using μ BooNE pattern recognition toolkit including multiple



- Overall **high reconstruction** efficiency: $(97.1 \pm 0.1 \text{ (stat)} \pm 1.4 \text{ (sys)}) \%$
- Measured reconstruction efficiency from **data agrees with the predicted efficiency** in the simulation
 - Confirmation** of our **simulation** and **reconstruction** chain.

THE PATH FORWARD

- MicroBooNE is the first stage LArTPC **exposed by high intensity neutrino beam**:
 - **Large statistics**, allowing us to measure ν -Ar **cross-sections precisely**.
- First LArTPC on **surface**.
 - **Many cosmic rays**, measurements more challenging.
 - But also useful tools to **calibrate** and **understand** our TPC
- **Many other analyses** currently in progress:
 - Neutral-Current elastic scattering, charged-current 0π , $1\mu+1p$ channel
 - Charged pion production, CC and NC neutral pion production, Coherent pion production
 - Kaon production
 - Exotic physics: Heavy Sterile Neutrinos & SuperNova ν
- Many **calibration analyses** are maturing.
 - Allowing us to set better **data-driven constraints** on our **systematics**.
 - Getting ready for the **next generation of analysis**.



THE PATH FORWARD

- MicroBooNE is the first stage LArTPC **exposed by high intensity neutrino beam**:
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• First LArTPC

-
-

• Many

-
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• ν production, ν production, ν and ν neutral pion production, Coherent pion production

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- Many **calibration analyses** are maturing.
 - Allowing us to set better **data-driven constraints** on our **systematics**.
 - Getting ready for the **next generation of analysis**.

Thank you for your attention



BACKUP

3. DETECTOR RESPONSE

| | Uncertainty source | Present implementation |
|----------------------------|-----------------------------------|---|
| Electron-Light propagation | Space charge | Modified MicroBooNE model |
| | Longitudinal/Transverse diffusion | 1σ (external constraints) |
| | Recombination | Alternative external models |
| | Electron attenuation | “Extreme case” simulation |
| | Outside-TPC light visibility | Conservative estimate |
| | Light production yield | Updated vs. current simulation |
| Readout response | Dynamic Induced Charge | Updated vs. current simulation |
| | Saturating channels | “Extreme case” simulation |
| | Misconfigured channels | “Extreme case” simulation |
| | Wire response function | 1σ (internal constraints from data) |
| | Wire noise | 1σ (internal constraints from data) |
| | PMT PE noise | 1σ (internal constraints from data) |

Ongoing MicroBooNE calibration measurements will constraint these uncertainties further

New simulation becomes default simulation

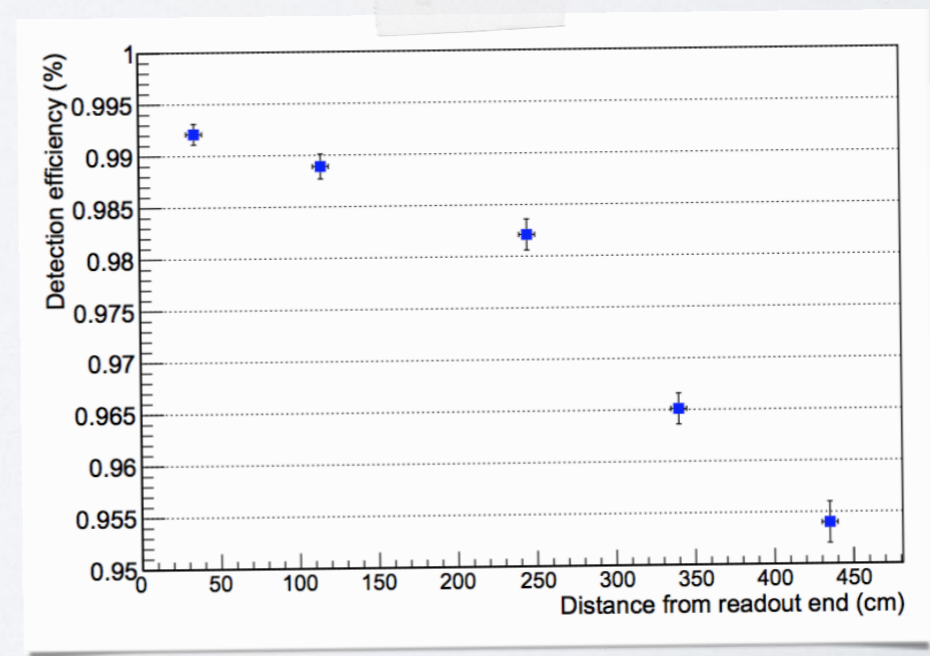
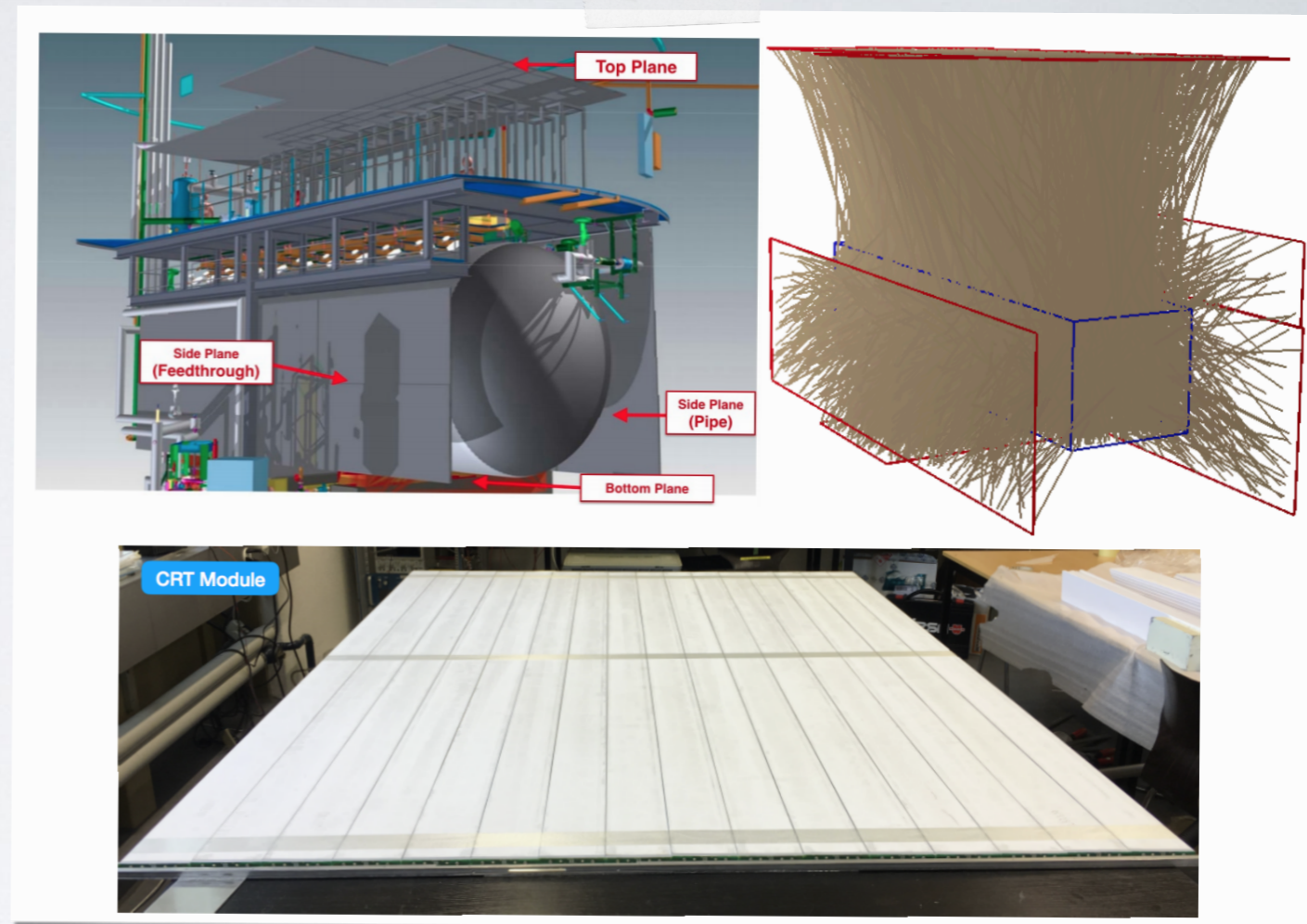
Ongoing work on more realistic treatment

Improvement in noise filtering and signal processing and using cosmic data as background (overlay) expected to reduce impacts of these effects

4. RECONSTRUCTION

DATA-DRIVEN RECONSTRUCTION EFFICIENCIES

- The cosmic-ray telescope has been replaced by a **novel Cosmic Ray Tagger (CRT)** system.
- Based on **plastic scintillator modules** and **SiPMs readout**.
- **85% coverage** for through-going muons and **excellent detection efficiency**.
- Will allow to **expand** on the **reconstruction studies** and improve on them.





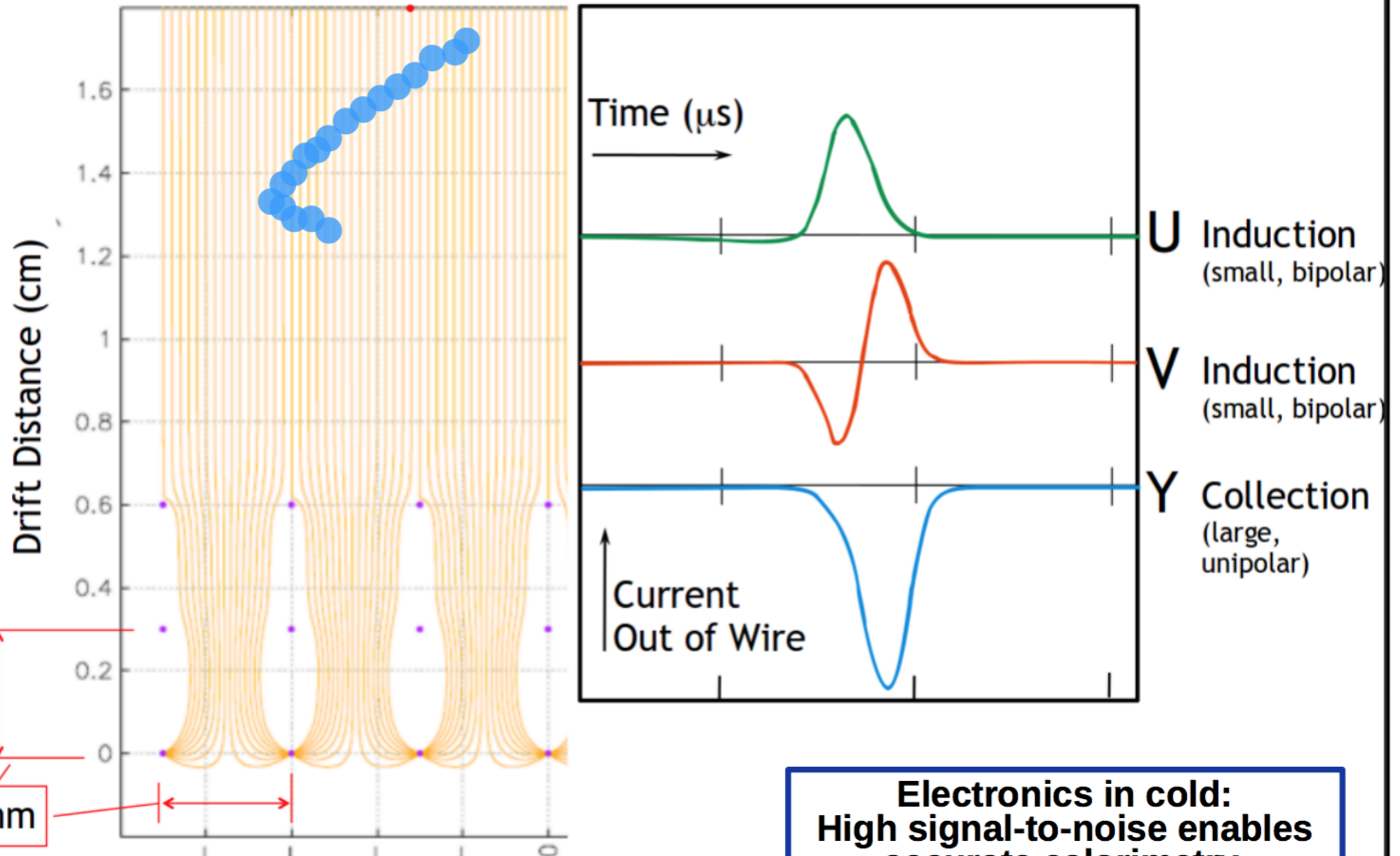
All unresponsive wires on all three planes (~10%)



All unresponsive wires with no redundancy (~3%)

Bo Yu (BNL)

Charge Signal Formation



**Electronics in cold:
High signal-to-noise enables
accurate calorimetry.**

CPM UNCERTAINTIES

- GENIE (short-track)

| Observed multiplicity | $\frac{\Delta P_n}{P_n}$ Default | $\frac{\Delta P_n}{P_n}$ MEC | $\frac{\Delta P_n}{P_n}$ TEM |
|-----------------------|----------------------------------|------------------------------|------------------------------|
| 1 | +7% | +7% | +8% |
| 2 | -11% | -12% | -12% |
| 3 | -25% | -25% | -25% |
| 4 | -33% | -36% | -39% |
| 5 | -44% | -48% | - |

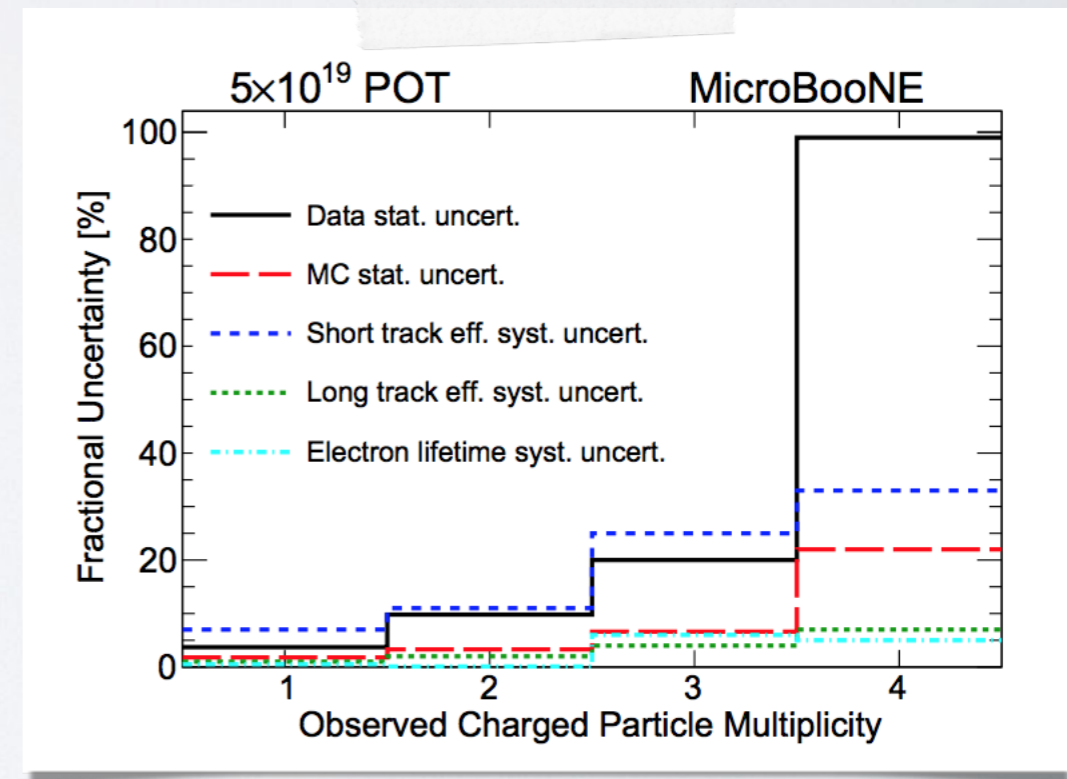
- GENIE (long-track)

| Observed multiplicity | $\frac{\Delta P_n}{P_n}$ Default | $\frac{\Delta P_n}{P_n}$ MEC | $\frac{\Delta P_n}{P_n}$ TEM |
|-----------------------|----------------------------------|------------------------------|------------------------------|
| 1 | -1% | -1% | -1% |
| 2 | +2% | +2% | +2% |
| 3 | +4% | +4% | +2% |
| 4 | +7% | +7% | +7% |
| 5 | +9% | +9% | - |

- Total

| Uncertainty Sources | Uncertainty Estimates | | | |
|-------------------------------|-----------------------|--------|--------|--------|
| | mult=1 | mult=2 | mult=3 | mult=4 |
| Data statistics | 4% | 10% | 20% | 99% |
| MC statistics | 2% | 3% | 7% | 22% |
| Short track efficiency | 7% | 11% | 25% | 33% |
| Long track efficiency | 1% | 2% | 4% | 7% |
| Background model systematics | 2% | 2% | 0% | 0% |
| Flux shape systematics | 0% | 0.4% | 0.2% | 0.5% |
| Electron lifetime systematics | 0.5% | 0.1% | 6% | 5% |

- Total (histogram)



CC π^0 UNCERTAINTIES

• Flux

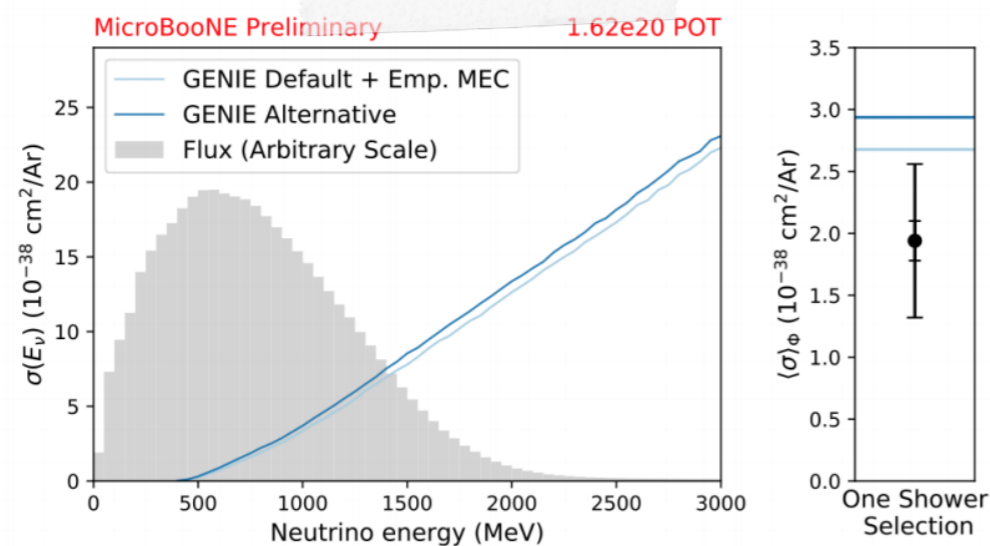
• Cross-sections

• Detector

| Variation | 1σ Uncertainty |
|--------------------------|-----------------------|
| p+Be $\rightarrow \pi^+$ | 11.5% |
| Beamline | 10.2% |
| p+Be $\rightarrow K^+$ | 1.4% |
| p+Be $\rightarrow K^-$ | 0.4% |
| p+Be $\rightarrow K^0$ | 0.4% |
| p+Be $\rightarrow \pi^-$ | 0.3% |
| Total Uncertainty | 15.5% |

| Variation | 1σ Uncertainty |
|--------------------------------|-----------------------|
| Cross Section Parameterization | 11.5% |
| Final State Interactions | 10.2% |
| Hadronization | 1.4% |
| Deep Inelastic Scattering | 0.0% |
| Total Uncertainty | 17.2% |

| Variation | 1σ Uncertainty |
|--------------------------|-----------------------|
| Micro-physics | 12.9% |
| Detector Response | 12.5% |
| Cosmic Simulation | 11.0% |
| Total Uncertainty | 21.1% |



CC INCLUSIVE UNCERTAINTIES

• Flux

| Parameter | Description | Total Cross Section Relative Uncertainty |
|--------------------|----------------------------------|--|
| Non-Hadron | Non-Hadron | 5.34% |
| K^- Production | K^- production cross section | 0.50% |
| K^+ Production | K^+ production cross section | 0.55% |
| K^0 Production | K^0 production cross section | 0.51% |
| π^- Production | π^- production cross section | 0.73% |
| π^+ Production | π^+ production cross section | 9.69% |
| Total | Combined uncertainty | 11.93% |

• Total (+ cross-section)

| Error Source | Method | Estimated Relative Uncertainty |
|------------------------|---|--------------------------------|
| Beam Flux | Estimated with multisim variations | 12% |
| Cross Section Modeling | Estimated with multisim variations | 4% |
| Detector Response | Estimated with unisim variations | 19% |
| POT Counting | Toroids Resolution | 2% |
| Cosmics (in-time) | Estimated from data-driven cosmic model | 7% |
| Cosmics (out-of-time) | Estimated from off-beam statistics | 1% |
| Beam Timing Jitter | Estimated from on- minus off-beam flashes | 4% |

• Detector

| Detector System-atic Sample | Description | Type | Total Cross Section Relative Uncertainty [%] |
|--|---|-----------------|--|
| Space Charge | A simple data-driven calibration is applied to the space charge simulation to make it better match measured space charge effects [29]. | Modified Model | 2.7 |
| Induced Charge | Charge induction is simulated on a longer spatial range than in the default MC, so that more distant wires see the effect of drifting charge. | Alternate Model | 15 |
| Light Yield | An improved light production simulation model is used. | Alternate Model | 3.7 |
| Remove Channels Prone to Saturating | Turning off channels that frequently become saturated as charge builds up on capacitors in the ASIC circuits, resulting in deadtime. | Alternate Model | 2.1 |
| Remove Misconfigured Channels | Turning off the misconfigured channels associated with ASICs that have a different gain and shaping time than desired | Modified Model | 2.1 |
| Wire Response Function | The wire response functions used during deconvolution are stretched by 20% based on MicroBooNE data. | $\pm 1\sigma$ | 1.4 |
| Longitudinal Diffusion | The amplitude of longitudinal diffusion is varied based on world data [32, 33]. | $\pm 1\sigma$ | 1.4 |
| Transverse Diffusion | The amplitude of transverse diffusion is varied based on world data [34, 35, 36]. | $\pm 1\sigma$ | 2.1 |
| Wire Noise | The amplitude of the wire noise model varied. | $\pm 1\sigma$ | 6.4 |
| PE Noise | The single-PE noise of the PMTs is varied. | $\pm 1\sigma$ | 2.1 |
| TPC Visibility | The light yield in the cryostat but outside the TPC is increased by 50%. | Alternate Model | 4.3 |
| Lifetime | The electron lifetime is reduced to 10 ms. (This condition affects only about ~10% of data taken with lower purity). | Alternate Model | 1.2 |
| Recombination | The Birks recombination model, with parameters derived from ICARUS, is used instead of the default modified box model, with parameters derived from Argoneut. | Alternate Model | 1.3 |
| Total combined relative uncertainty | | | 18.7 |