



THE UNIVERSITY  
OF BRITISH COLUMBIA



Natural Sciences and Engineering  
Research Council of Canada

Conseil de recherches en sciences  
naturelles et en génie du Canada

Canada

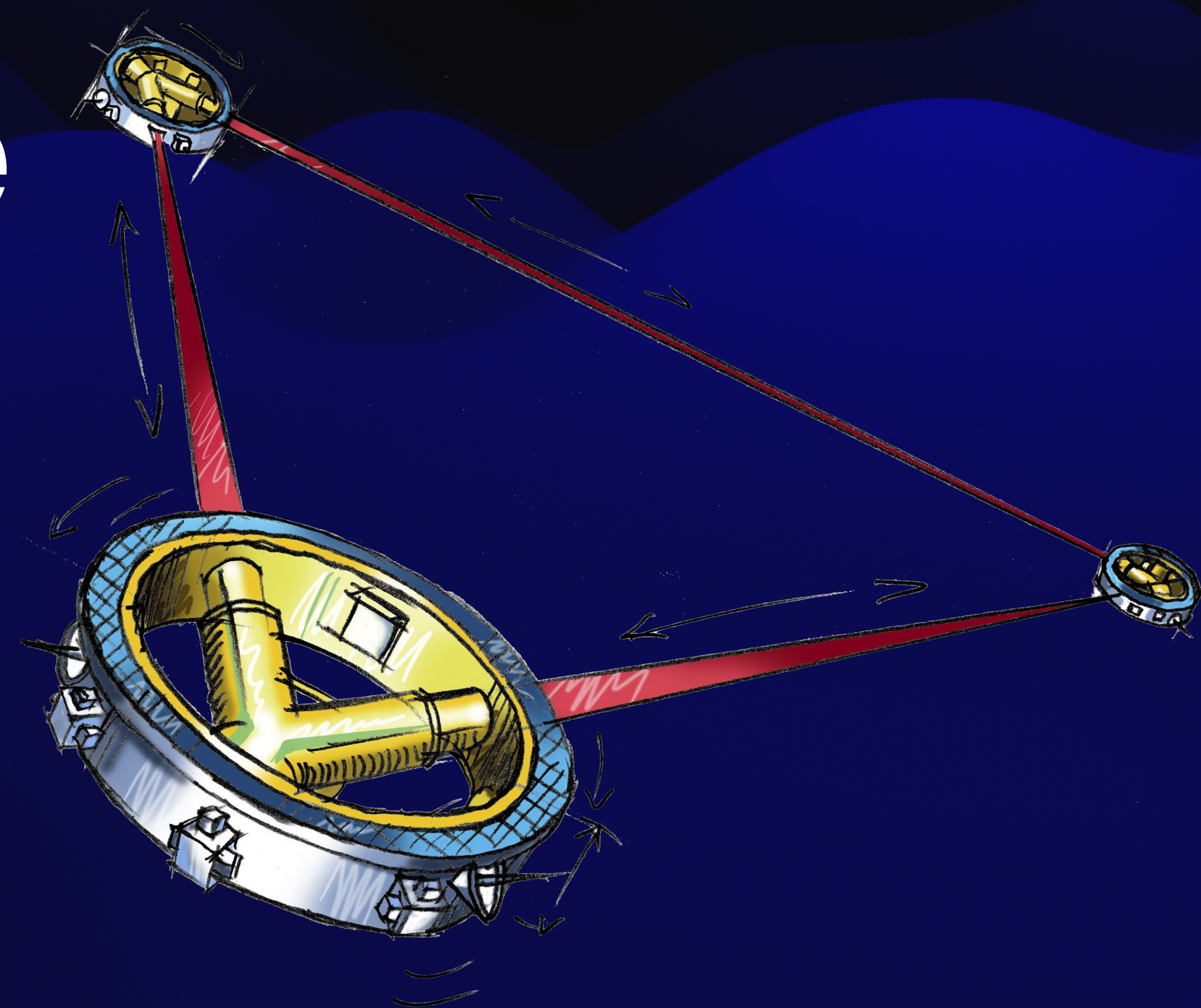
We acknowledge the support of  
the Natural Sciences and  
Engineering Research Council  
of Canada (NSERC).

Nous remercions le Conseil de  
recherches en sciences naturelles  
et en génie du Canada (CRSNG)  
de son soutien.

# LISA Data Challenge

## Glitch Mitigation in Verification Binaries

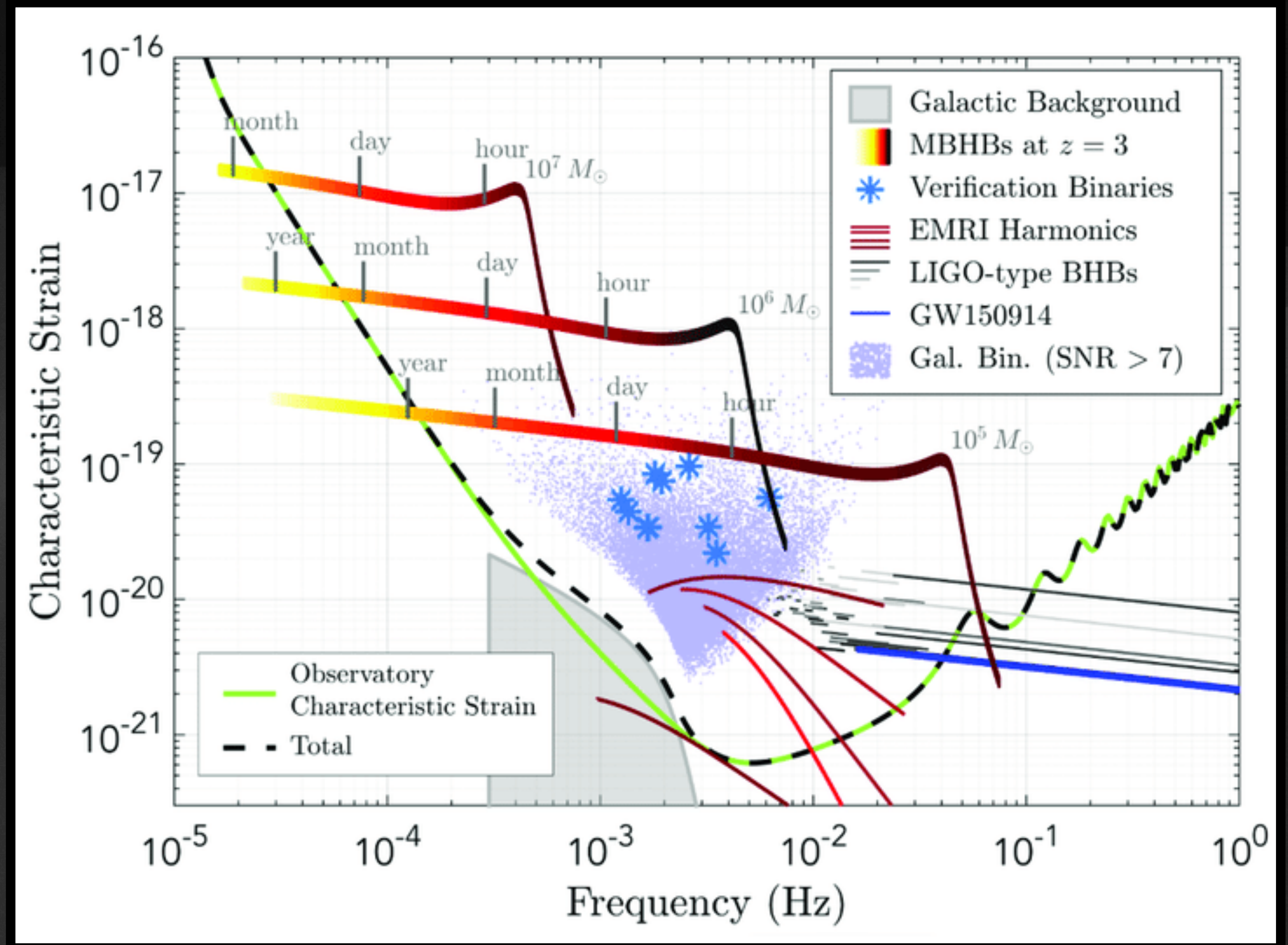
Kye Emond, Jess McIver, Scott Oser





# Verification Binaries

What are they?

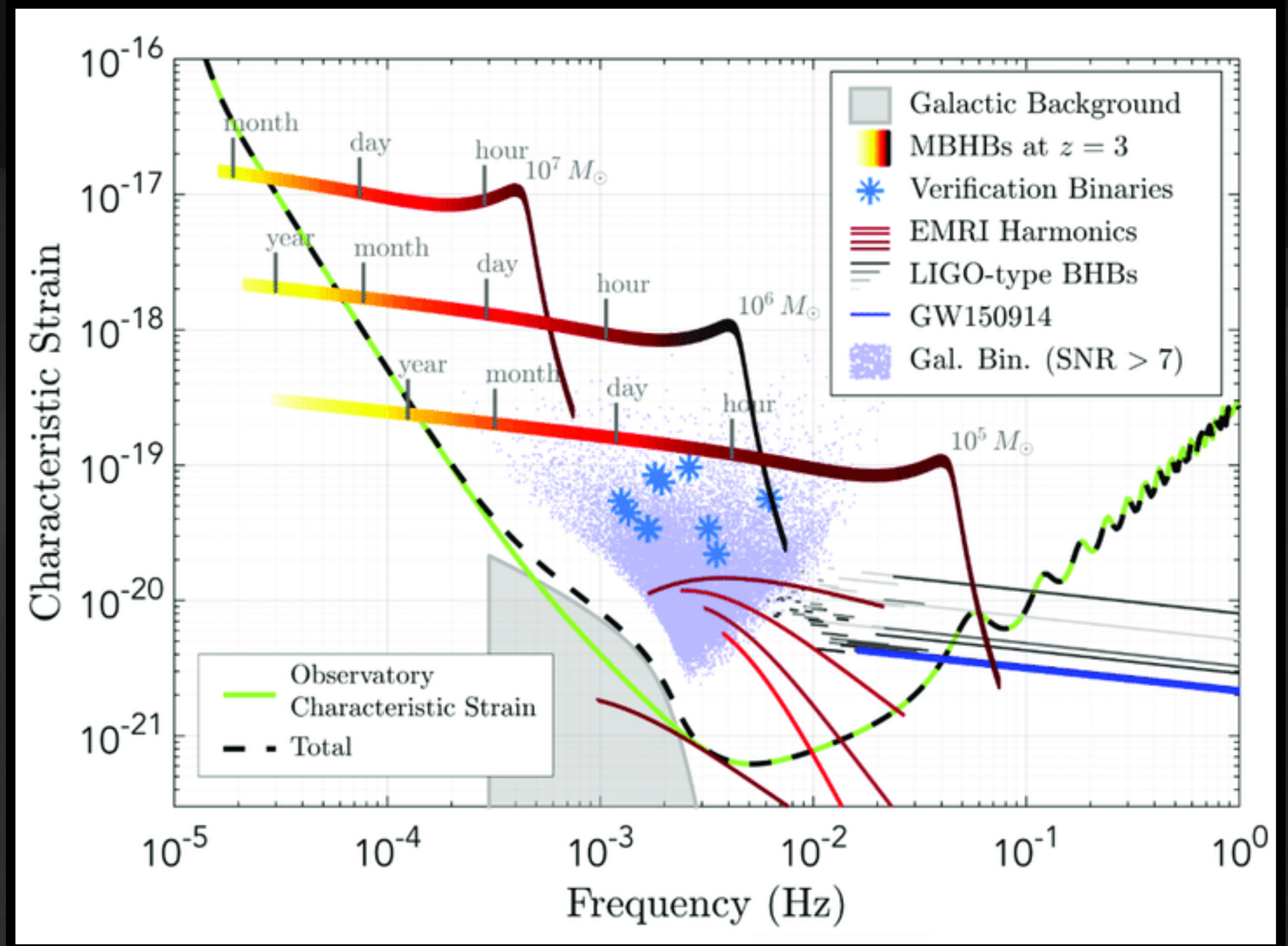




# Verification Binaries

## What are they?

- LISA can detect GWs in the  $10^{-4}$  to 1 Hz range

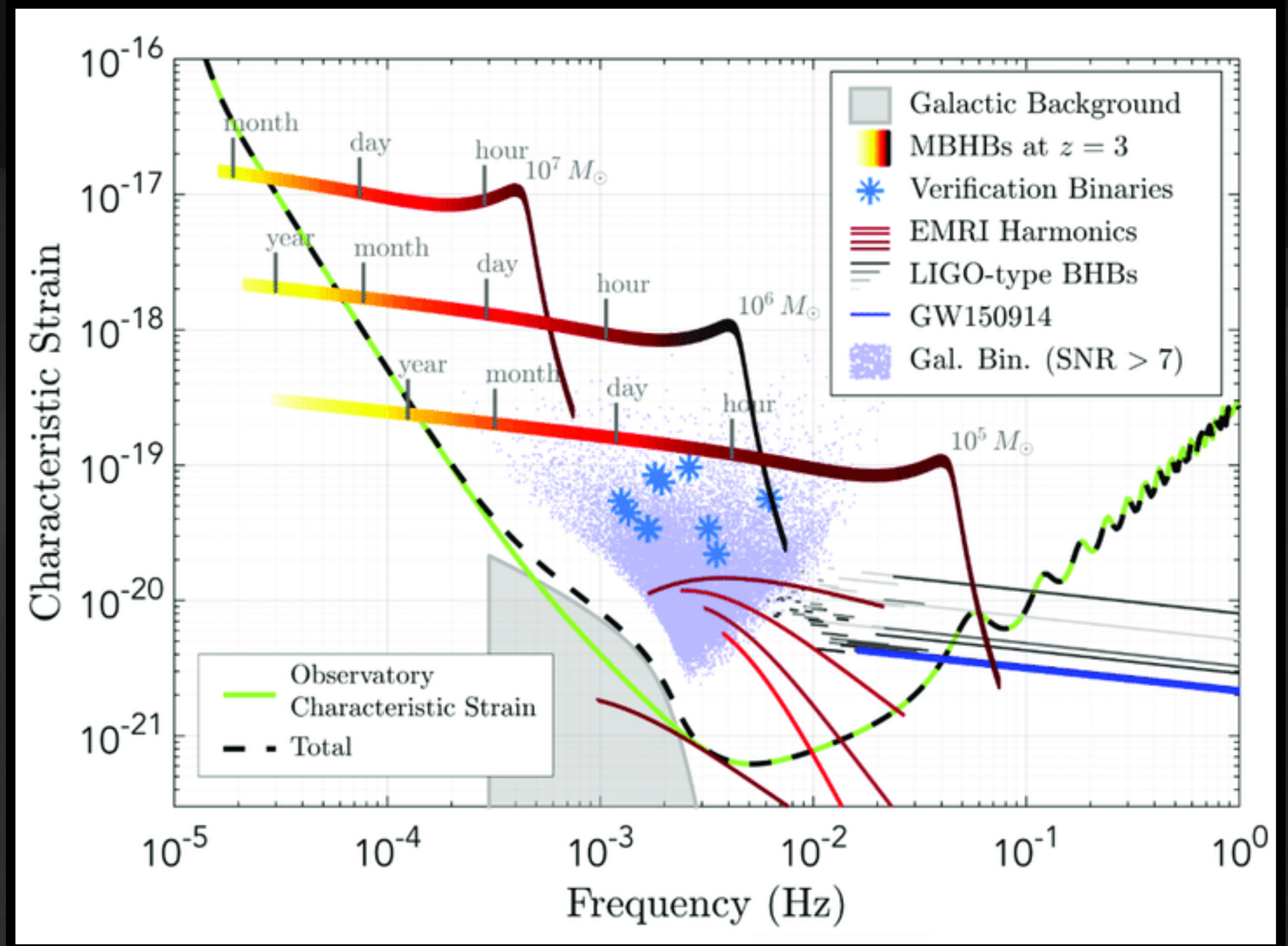




# Verification Binaries

## What are they?

- LISA can detect GWs in the  $10^{-4}$  to 1 Hz range
- Includes galactic binaries

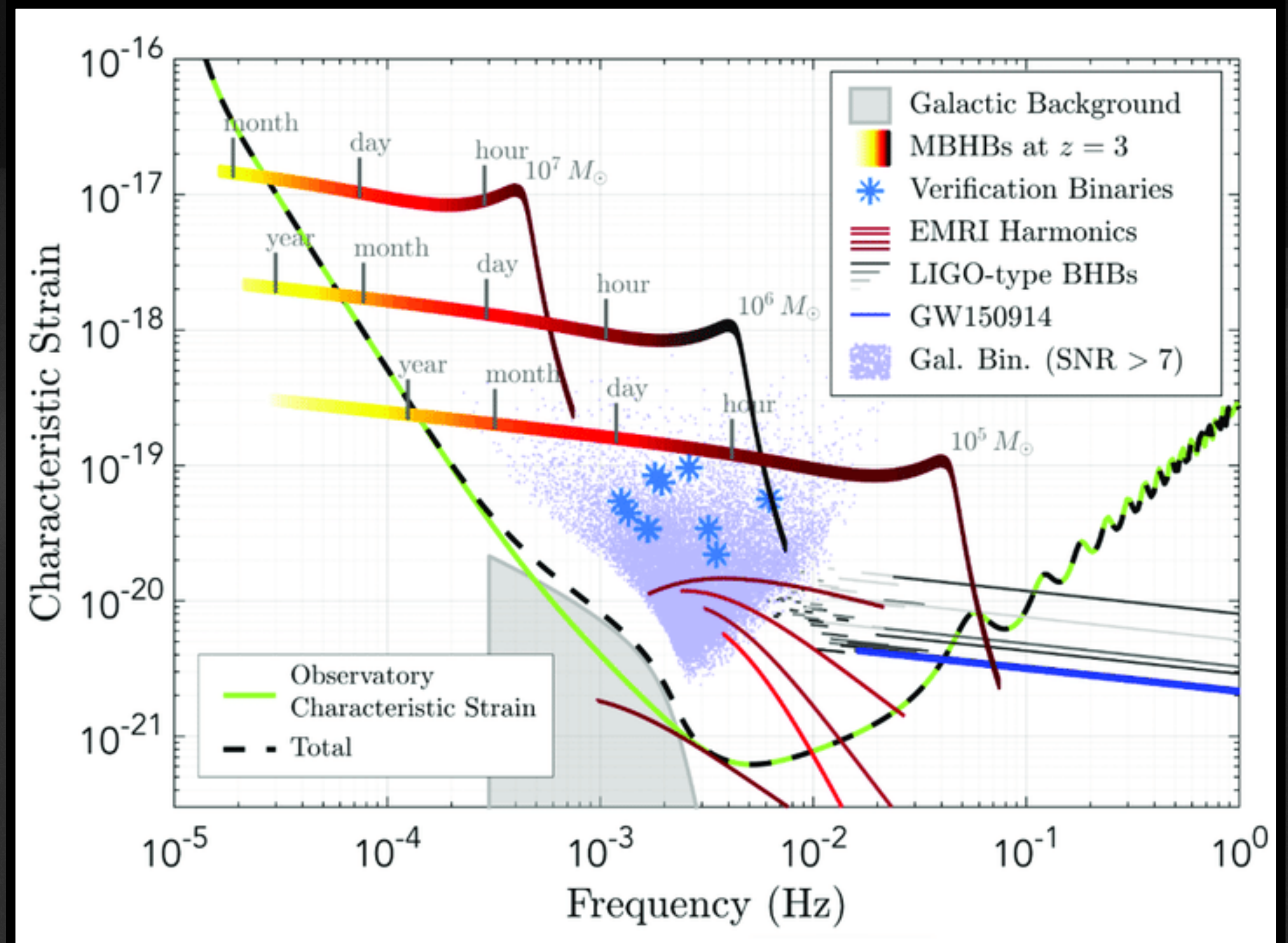




# Verification Binaries

## What are they?

- LISA can detect GWs in the  $10^{-4}$  to 1 Hz range
  - Includes galactic binaries
- We know parameters of  $\sim 30$  binaries from other observations

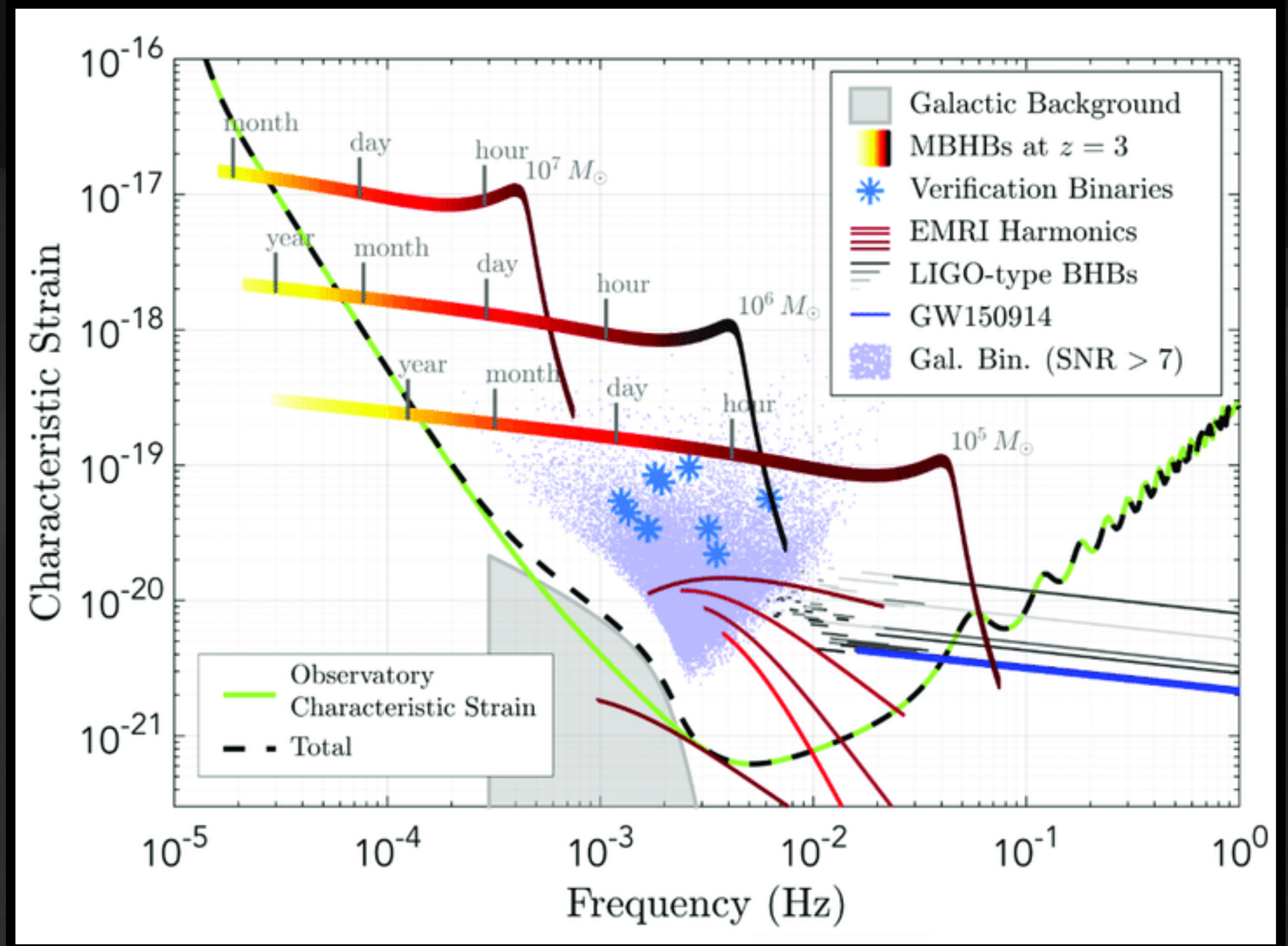




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- LISA can detect GWs in the  $10^{-4}$  to 1 Hz range
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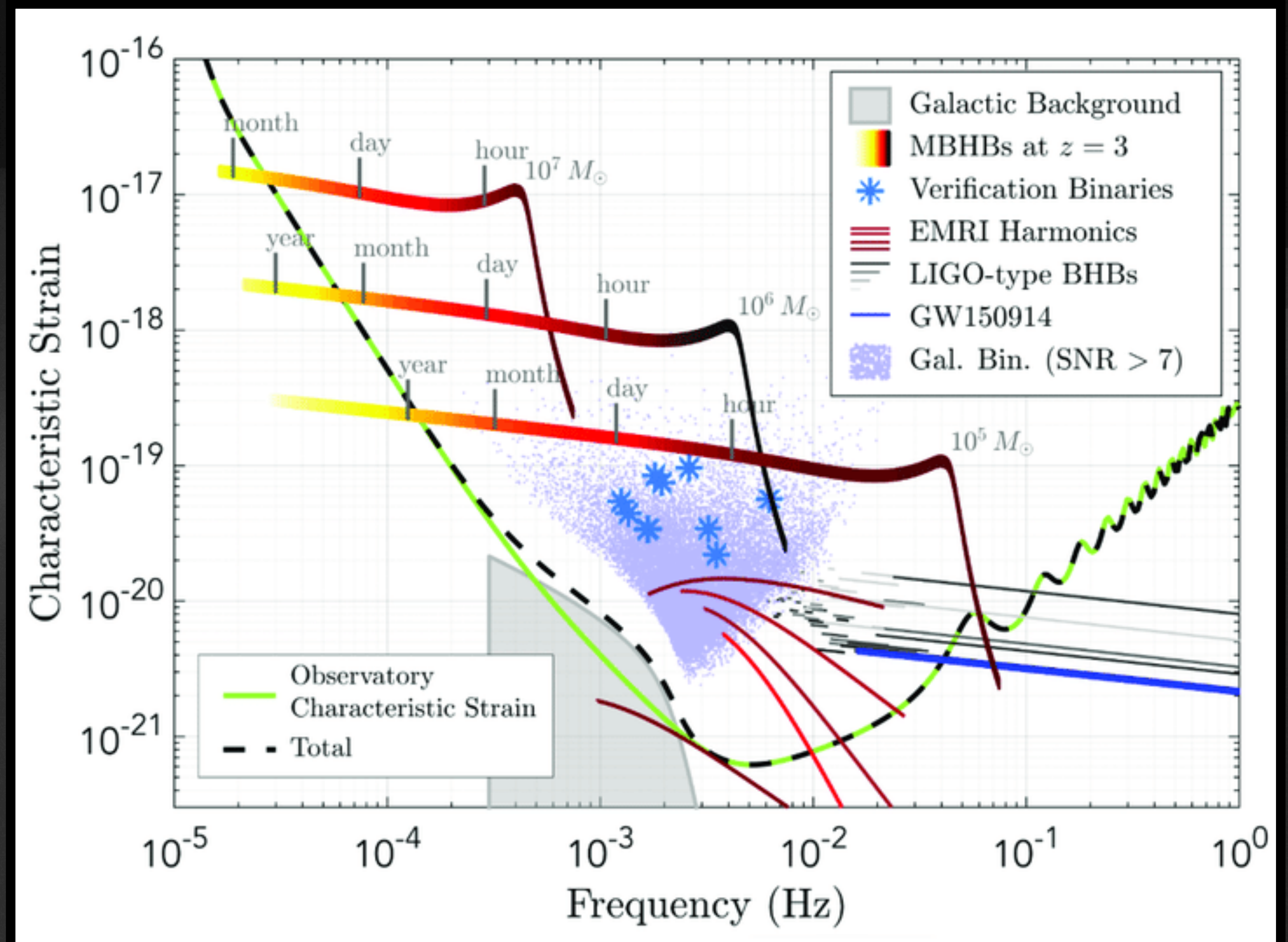




# Verification Binaries

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  - Includes galactic binaries
- We know parameters of  $\sim 30$  binaries from other observations
- We can *verify* LISA is working by checking that we get predicted GWs
- These are *verification binaries*





# LISA Data Challenges

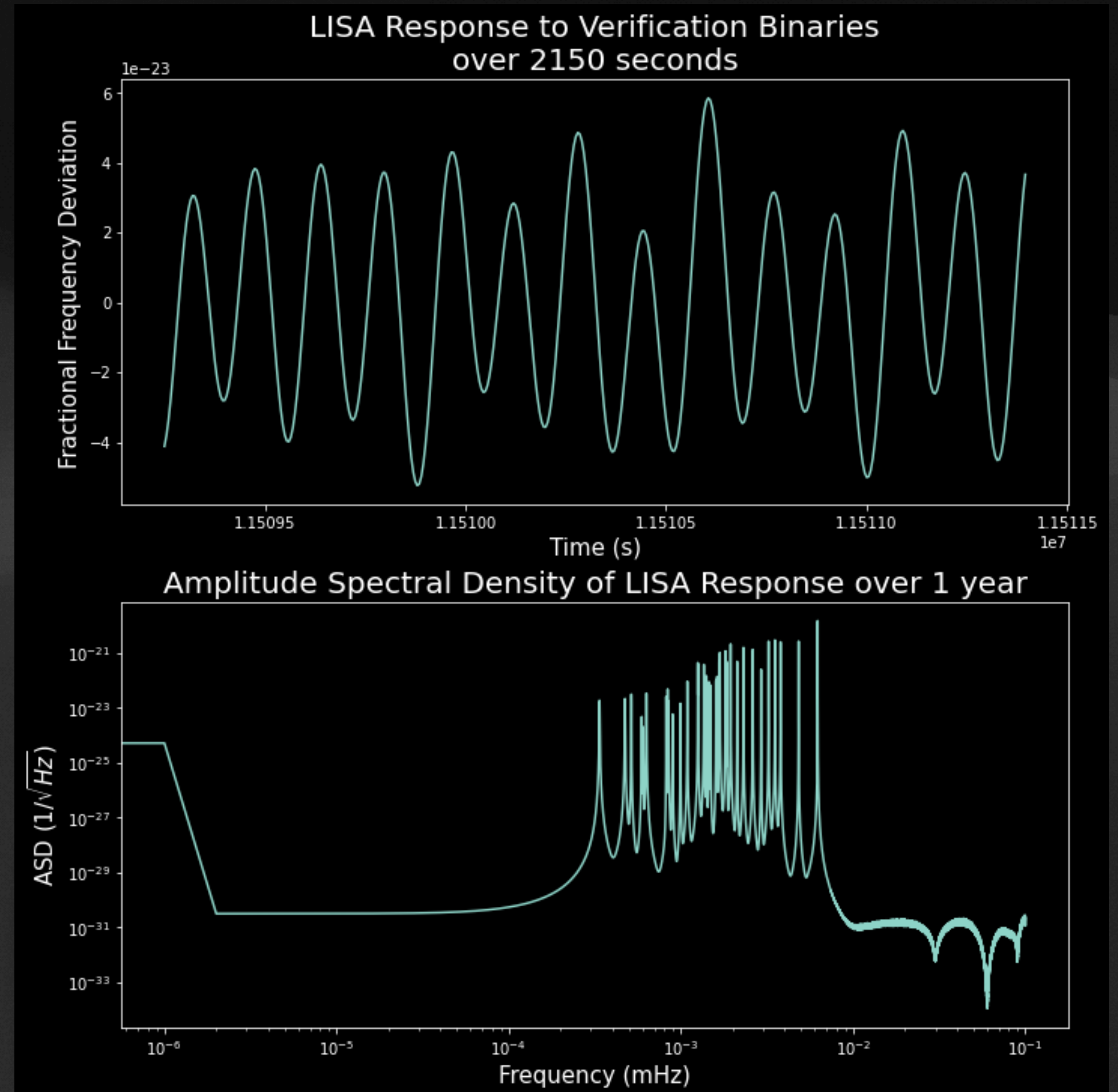
**Spritz: Verification Binaries**



# LISA Data Challenges

## Spritz: Verification Binaries

- GW signal from verification binaries

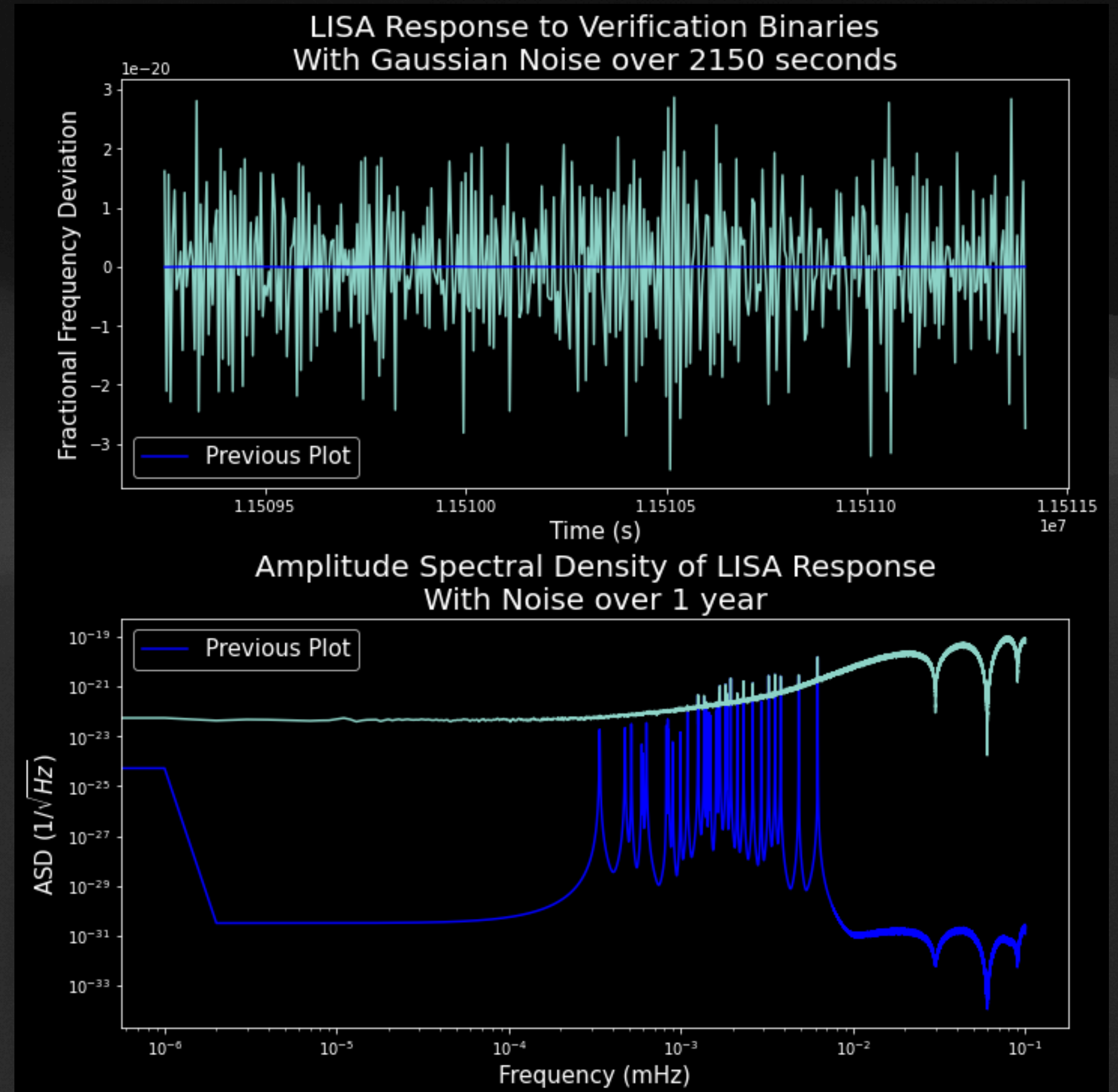




# LISA Data Challenges

## Spritz: Verification Binaries

- GW signal from verification binaries
- Plus instrumental noise

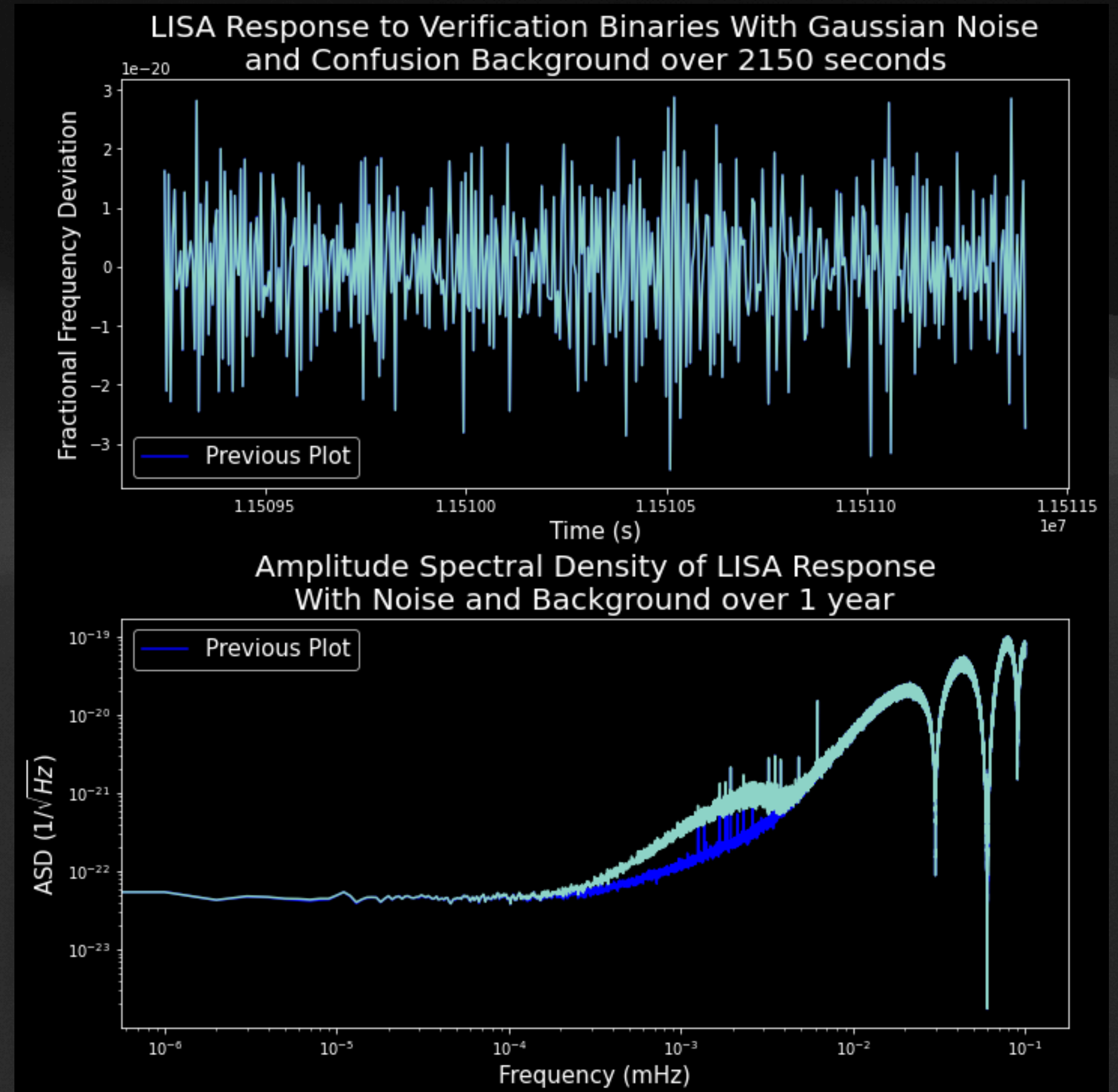




# LISA Data Challenges

## Spritz: Verification Binaries

- GW signal from verification binaries
- Plus instrumental noise
- Plus a lot of low-amplitude, unresolvable GWs from events in galaxy

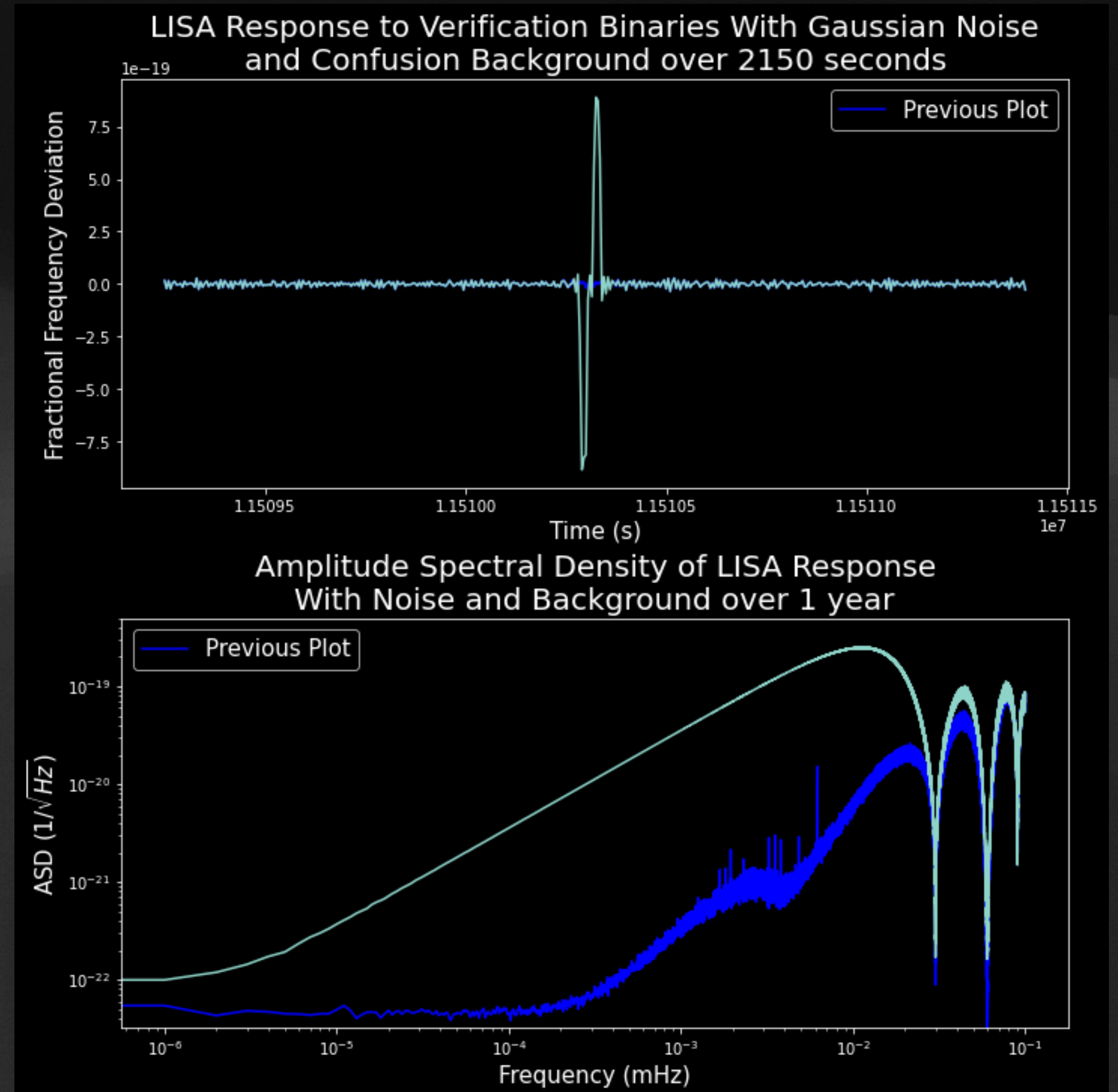




# LISA Data Challenges

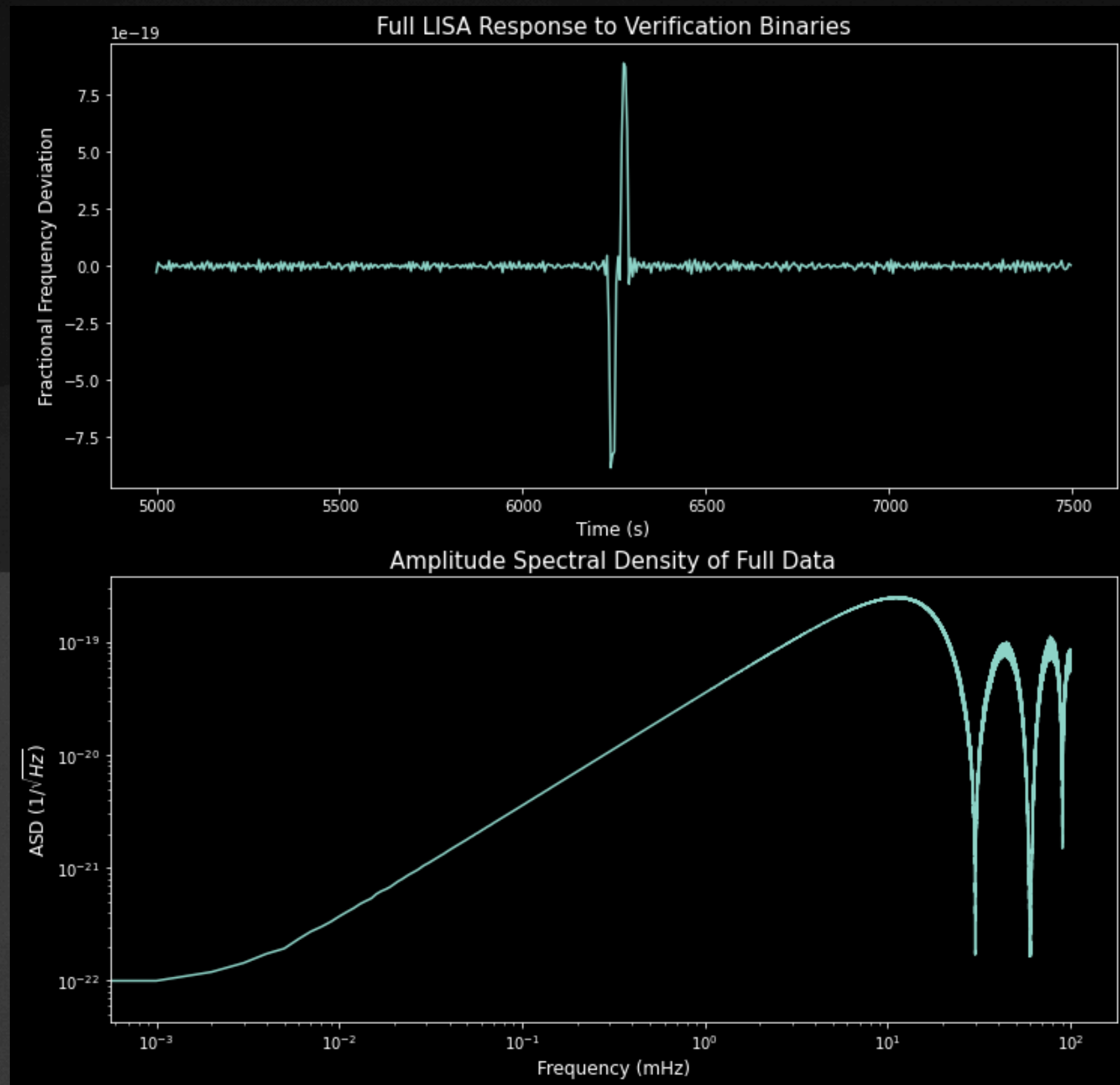
## Spritz: Verification Binaries

- GW signal from verification binaries
- Plus instrumental noise
- Plus a lot of low-amplitude, unresolvable GWs from events in galaxy
- Plus glitches and gaps (from LISA Pathfinder mission)





# Goal



Parameter	Value	Unit
Amplitude	1.6e-22	1
EclipticLatitude	0.087	rad
EclipticLongitude	4.1	rad
Frequency	0.0018	Hz
FrequencyDerivative	2.5e-18	Hz <sup>2</sup>
Inclination	0.52	rad
InitialPhase	5.1	rad
Polarization	6.1	rad

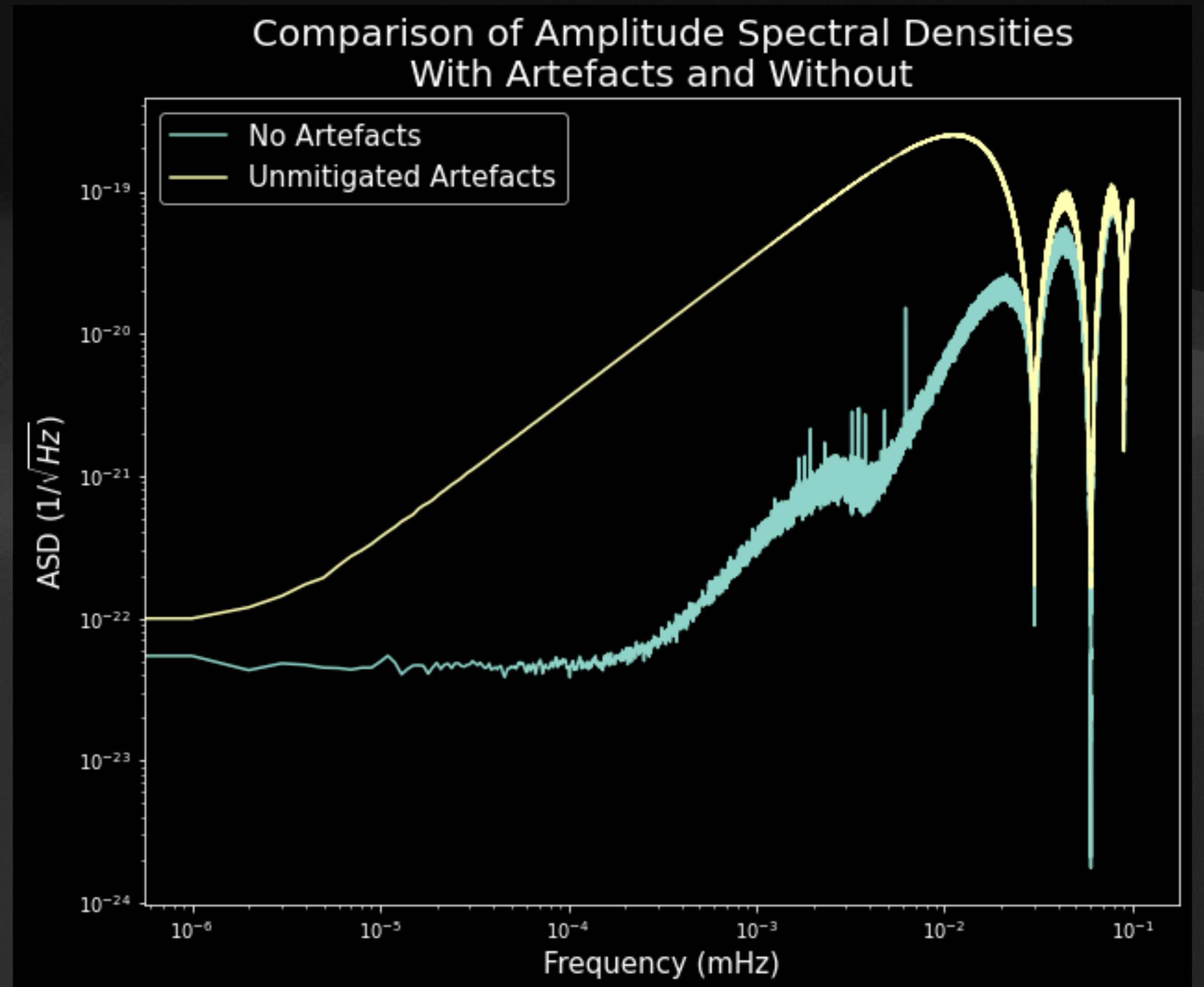
× 36

Go from full data to accurate parameter estimations



# Artefact Mitigation

## Dealing With Artefacts

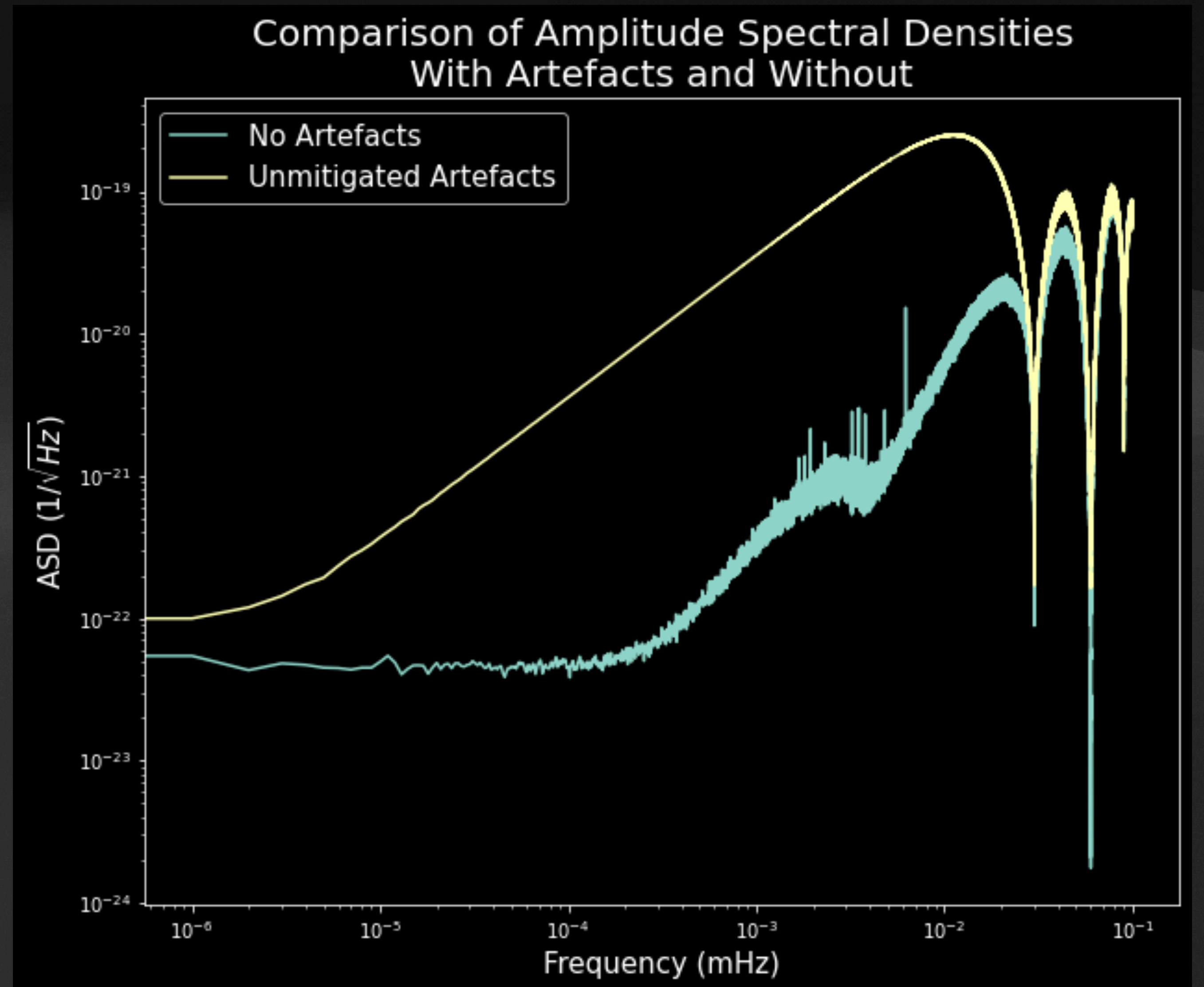




# Artefact Mitigation

## Dealing With Artefacts

- Main issue to solve is artefacts  
completely drown out binaries

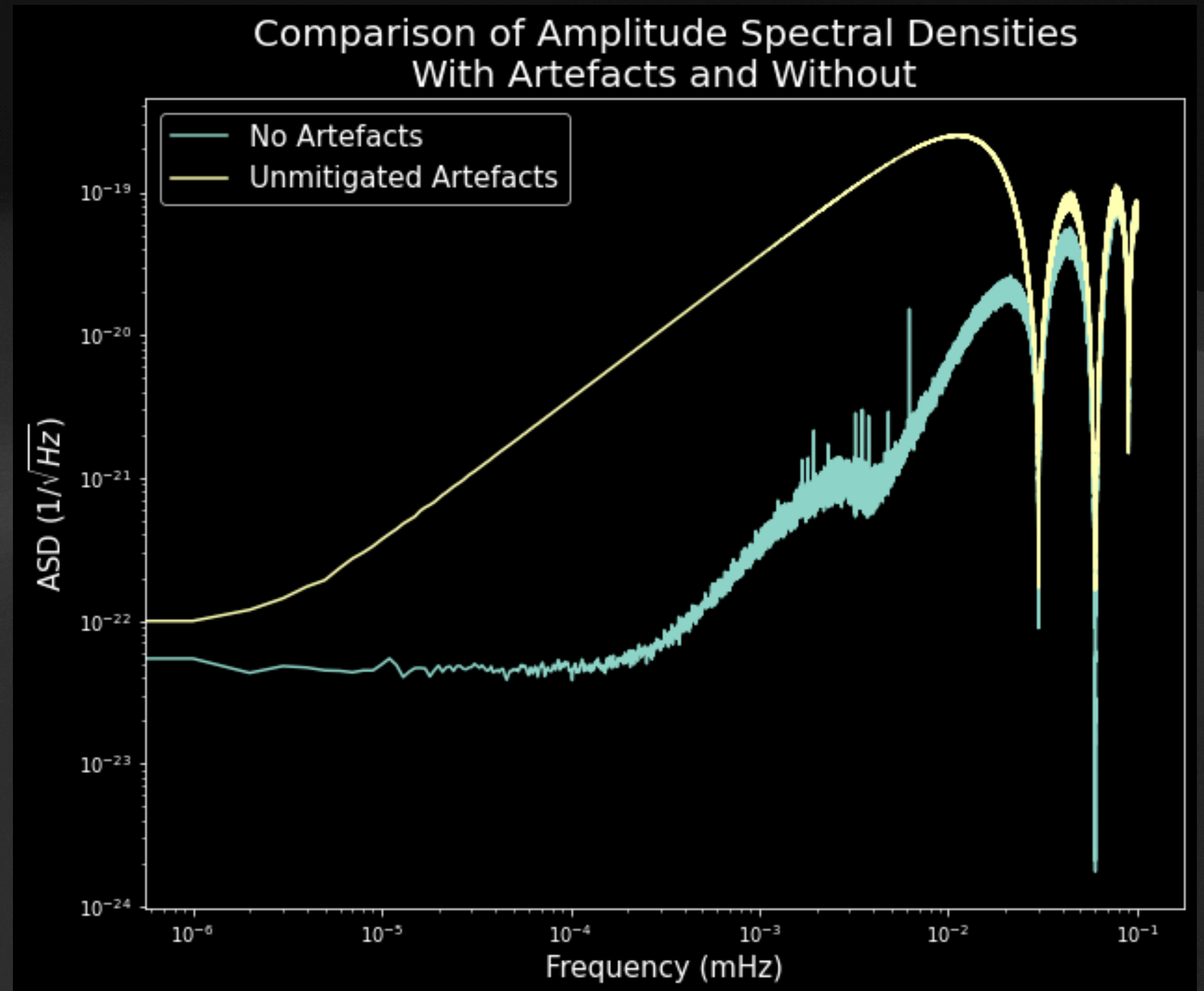




# Artefact Mitigation

## Dealing With Artefacts

- Main issue to solve is artefacts  
completely drown out binaries
- Need to mitigate that





# Artefact Mitigation

## Gaps



# Artefact Mitigation

## Gaps

- Detect gaps by looking for missing data



# Artefact Mitigation

## Gaps

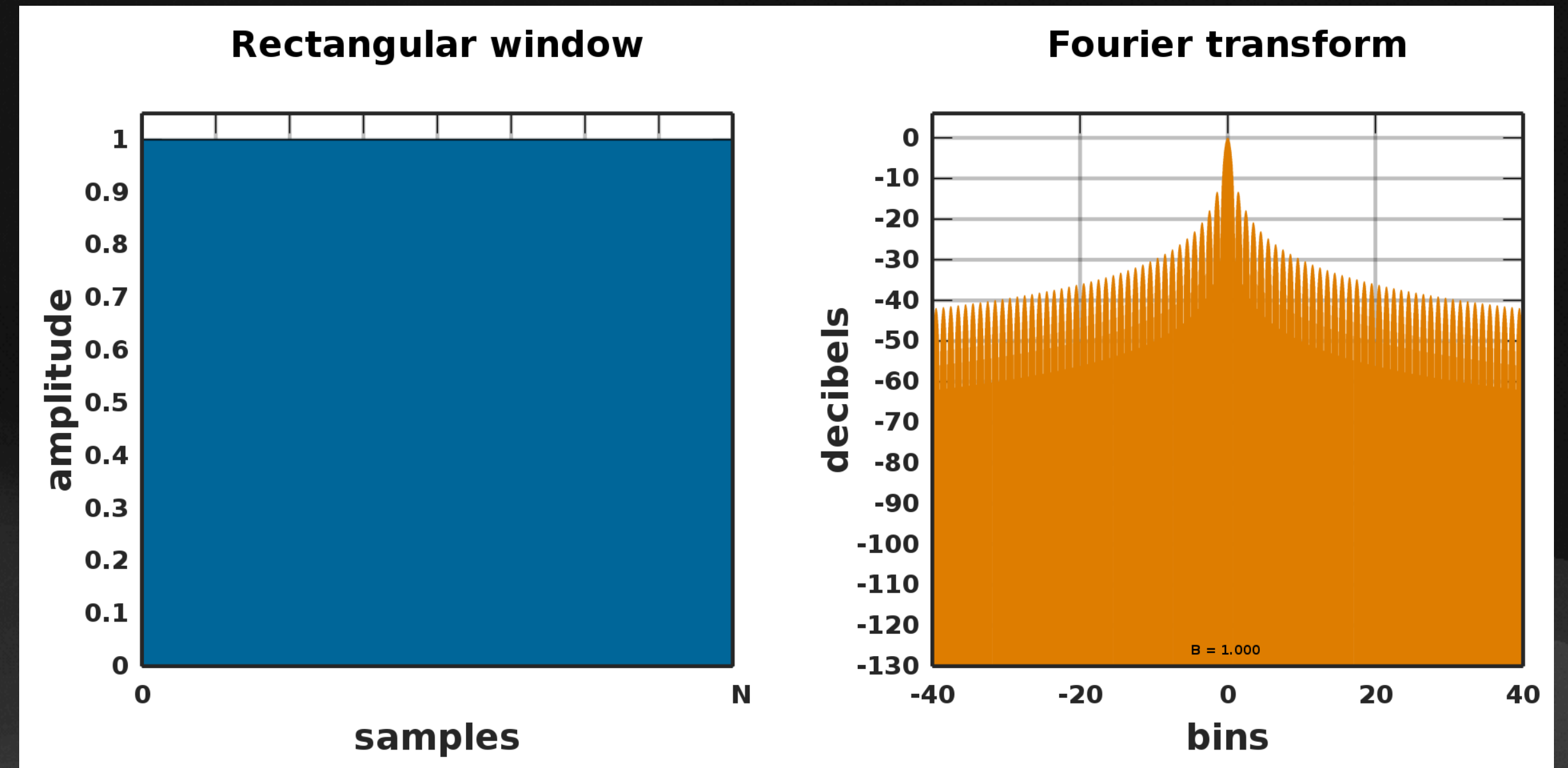
- Detect gaps by looking for missing data
- Set missing data to 0



# Artefact Mitigation

## Gaps

- Detect gaps by looking for missing data
- Set missing data to 0
- Results in spectral leakage

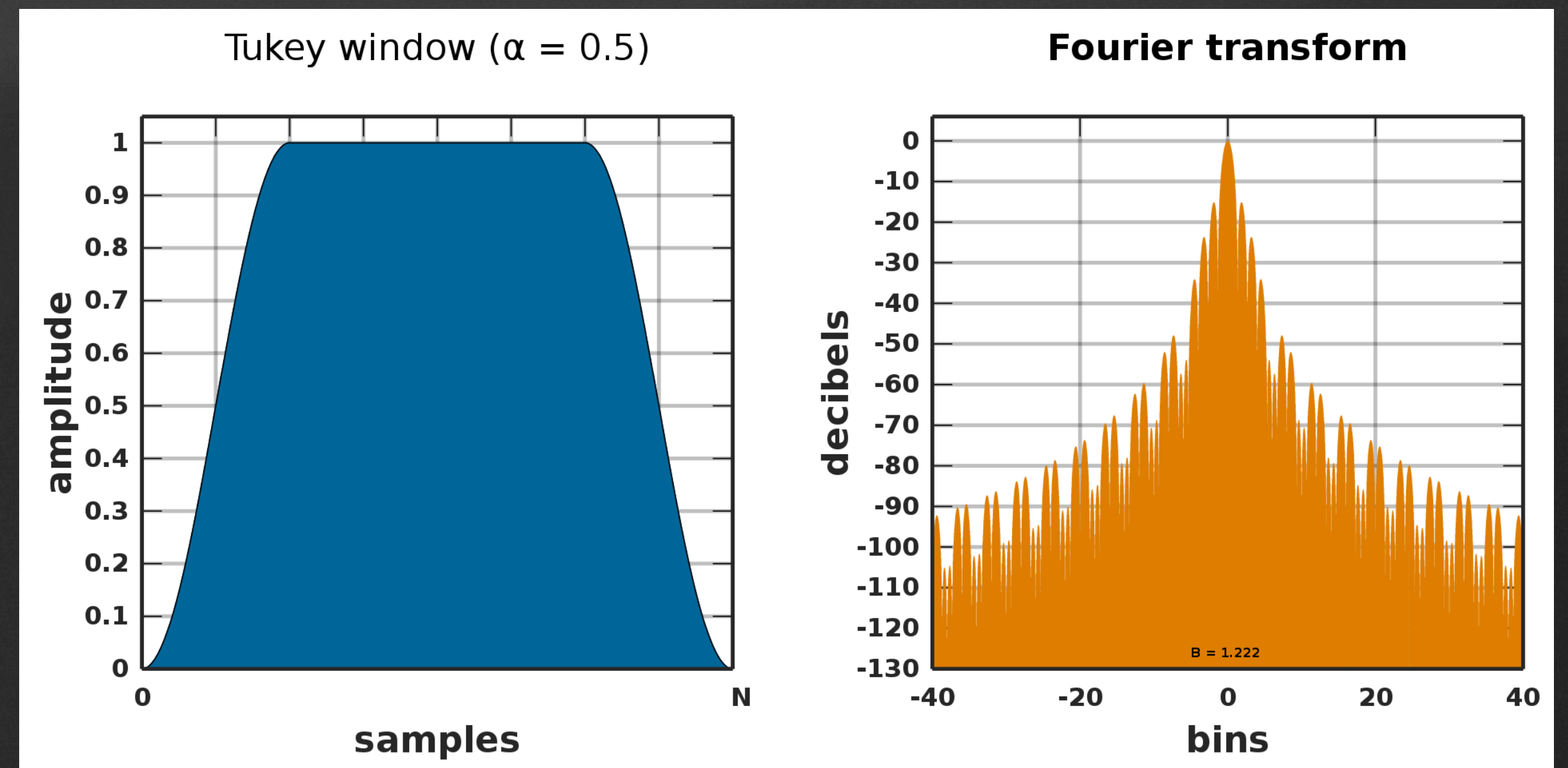
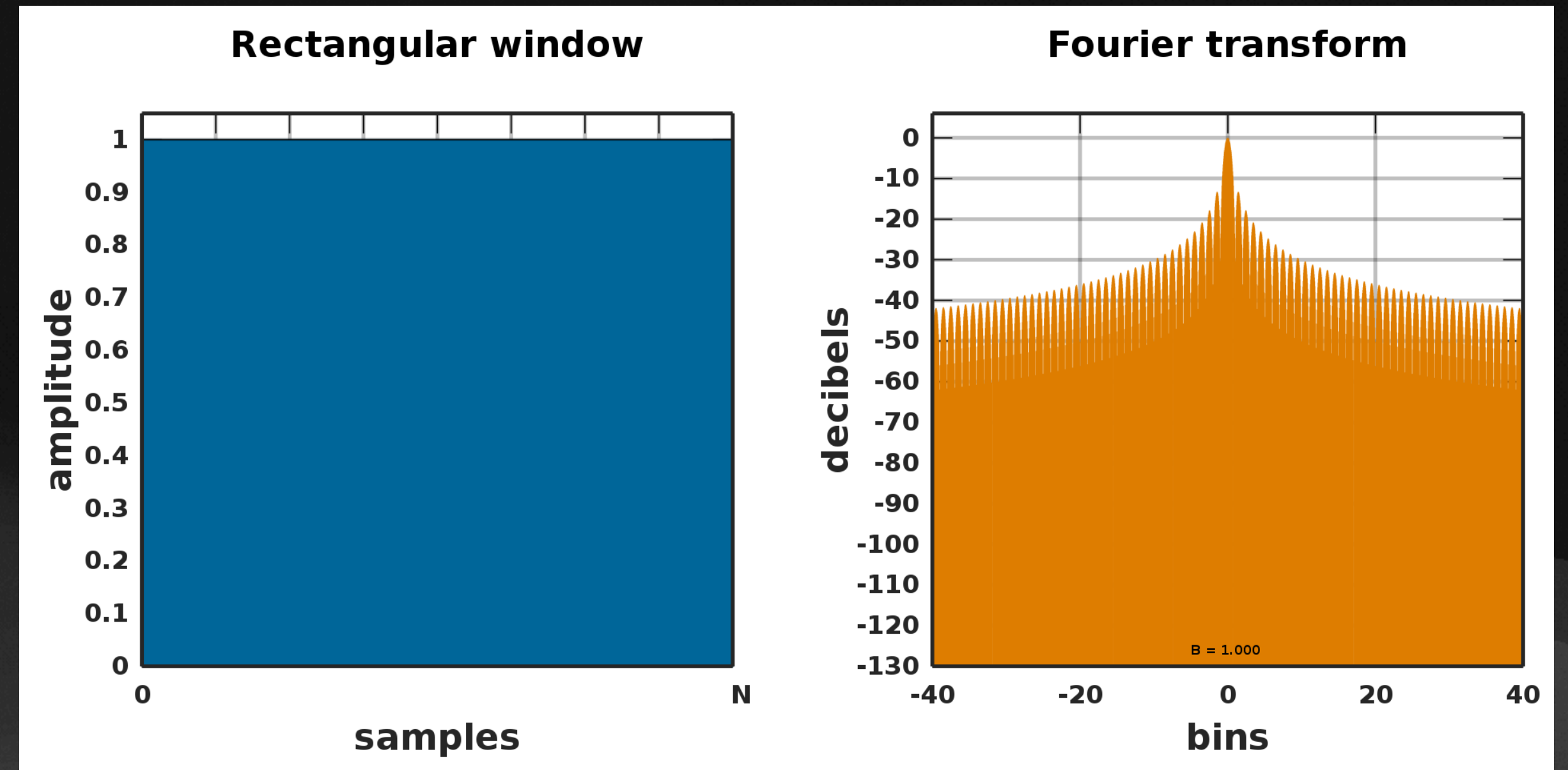




# Artefact Mitigation

## Gaps

- Detect gaps by looking for missing data
- Set missing data to 0
  - Results in spectral leakage
- Window the data smoothly in/out

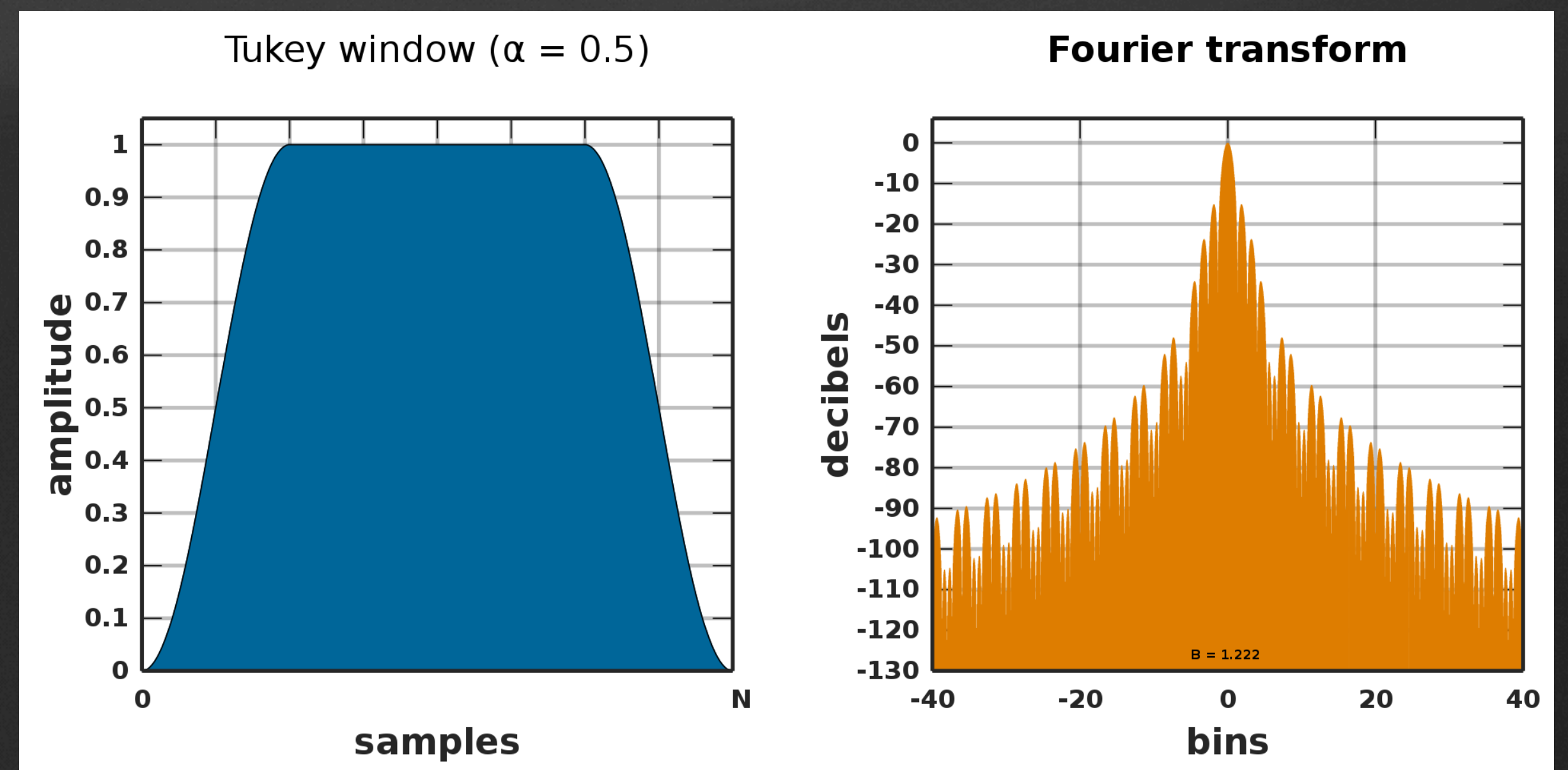
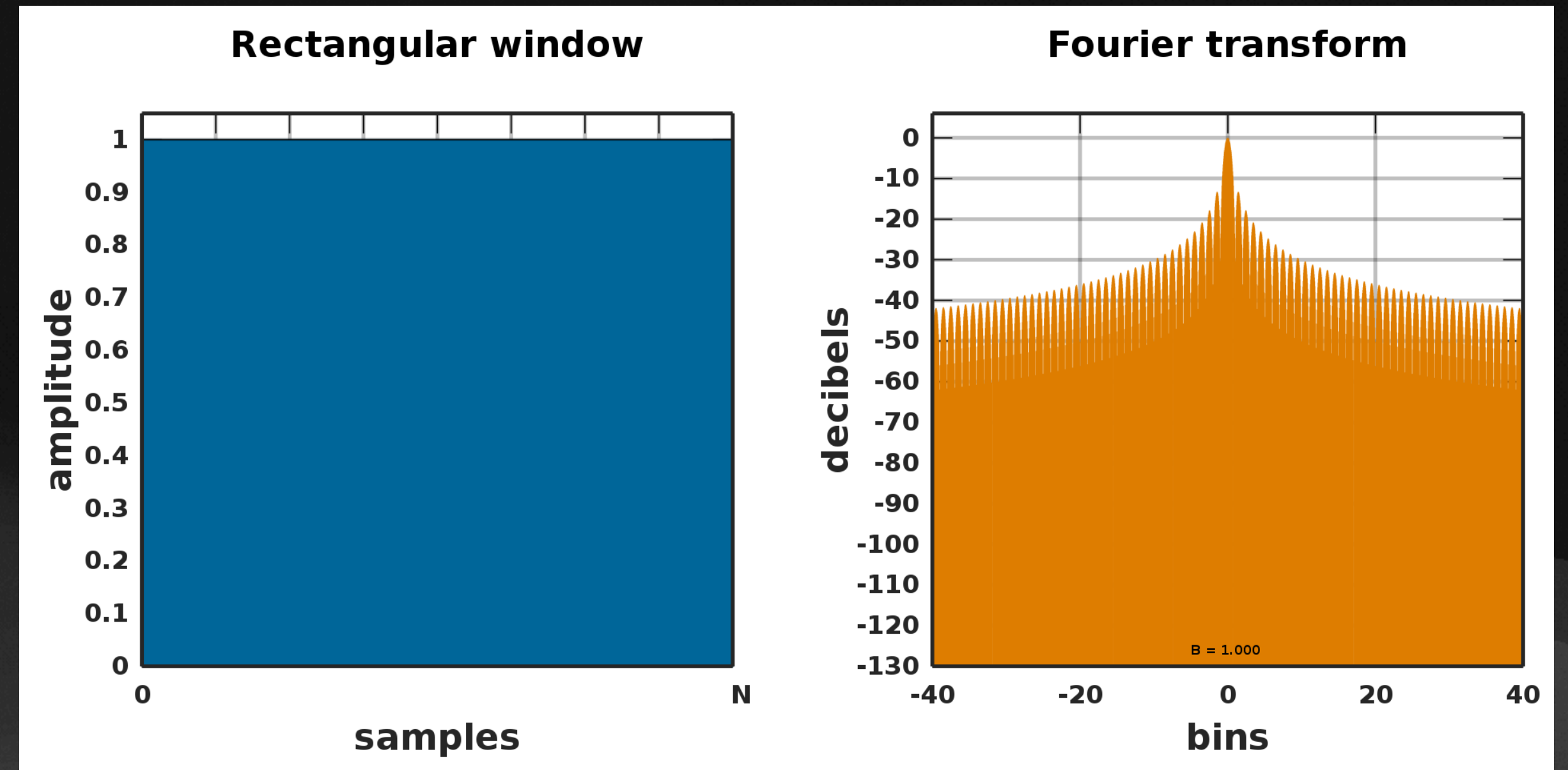




# Artefact Mitigation

## Gaps

- Detect gaps by looking for missing data
- Set missing data to 0
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- Window the data smoothly in/out
  - Multiply data by half of a very long cosine

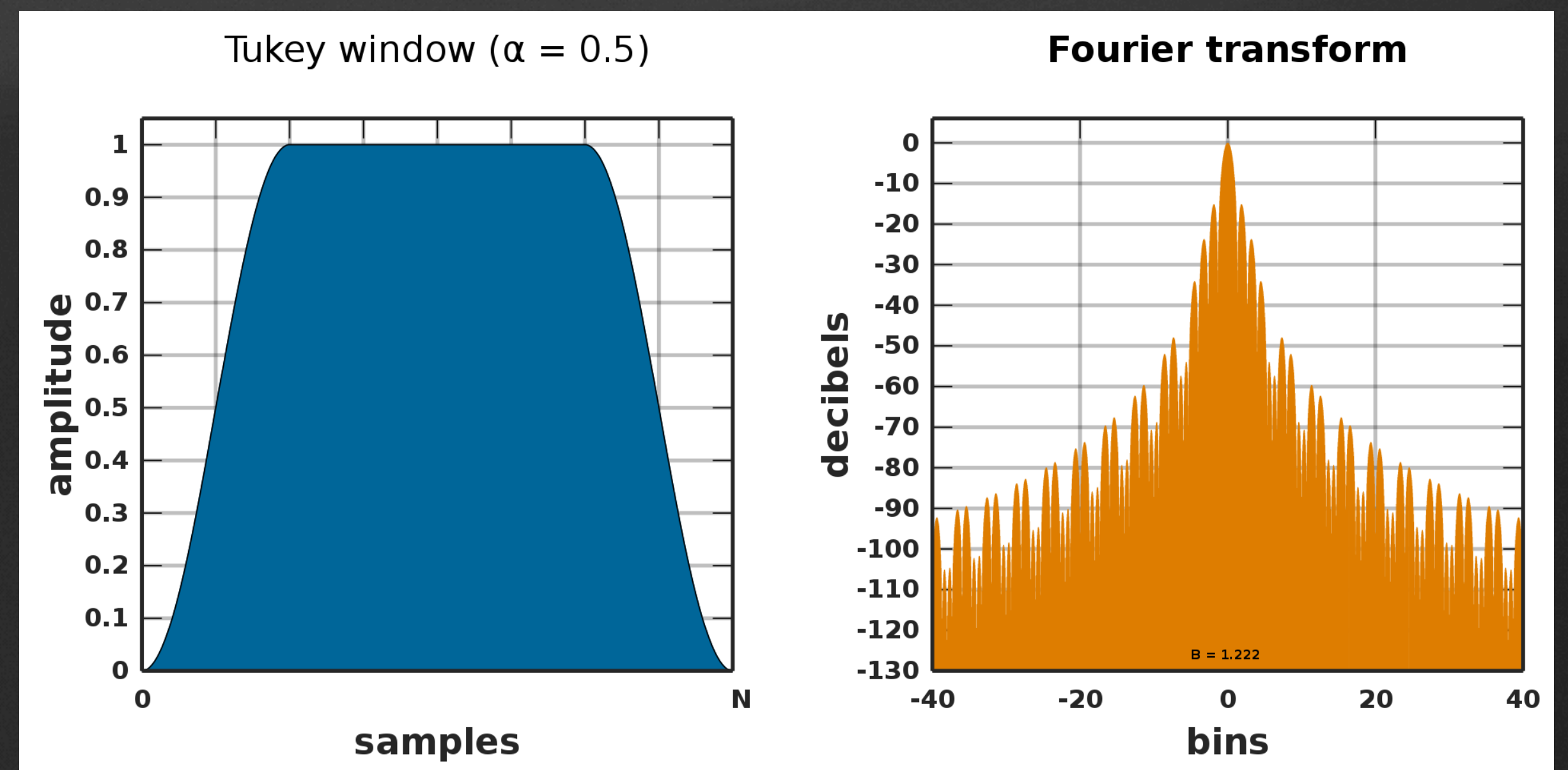
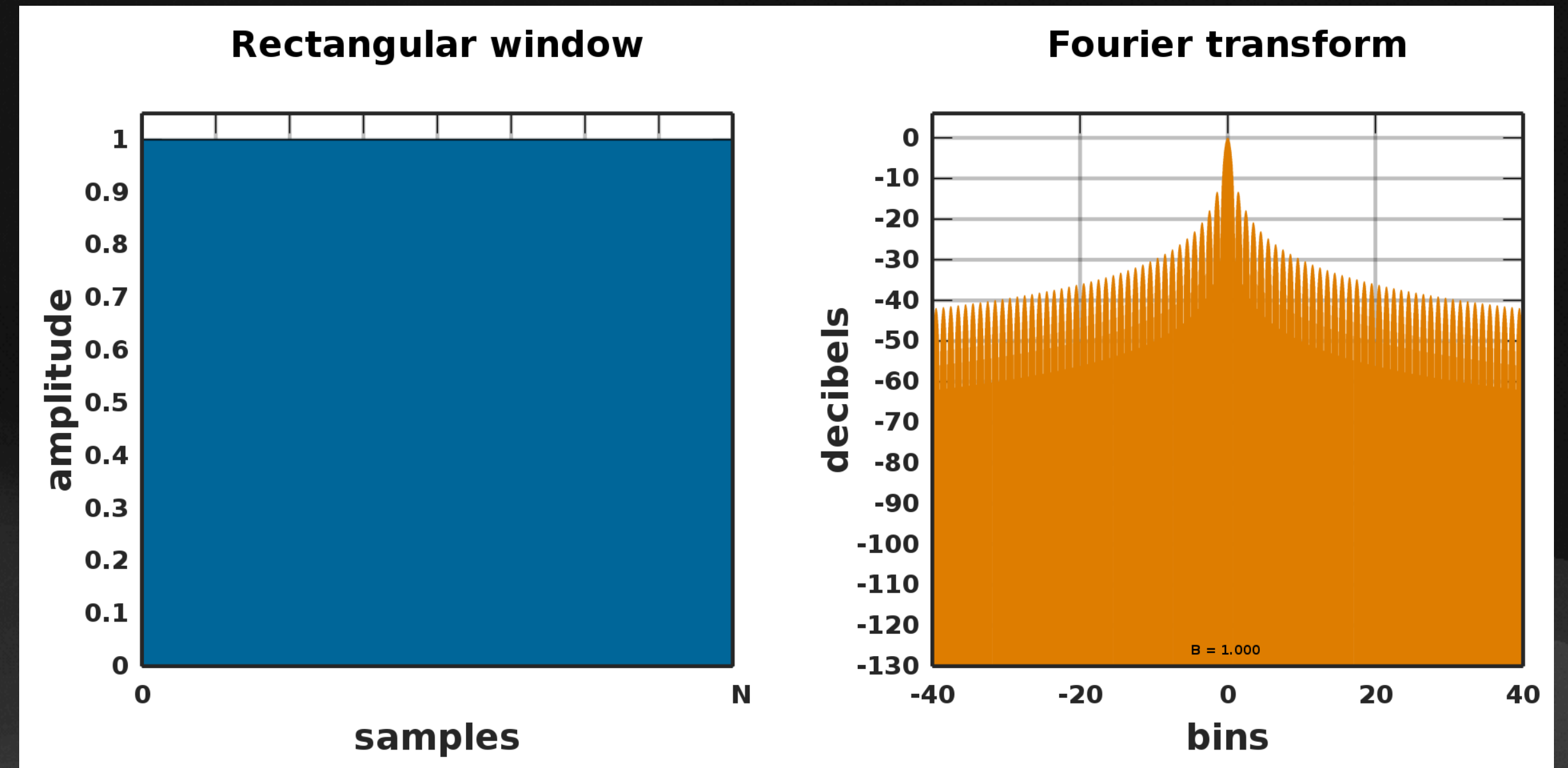




# Artefact Mitigation

## Gaps

- Detect gaps by looking for missing data
- Set missing data to 0
  - Results in spectral leakage
- Window the data smoothly in/out
  - Multiply data by half of a very long cosine
  - Reduces leakage





# Artefact Mitigation

## Glitches



# Artefact Mitigation

## Glitches

- Can remove glitches the same way as gaps



# Artefact Mitigation

## Glitches

- Can remove glitches the same way as gaps
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# Artefact Mitigation

## Glitches

- Can remove glitches the same way as gaps
  - Set to 0
  - Window surrounding data



# Artefact Mitigation

## Glitches

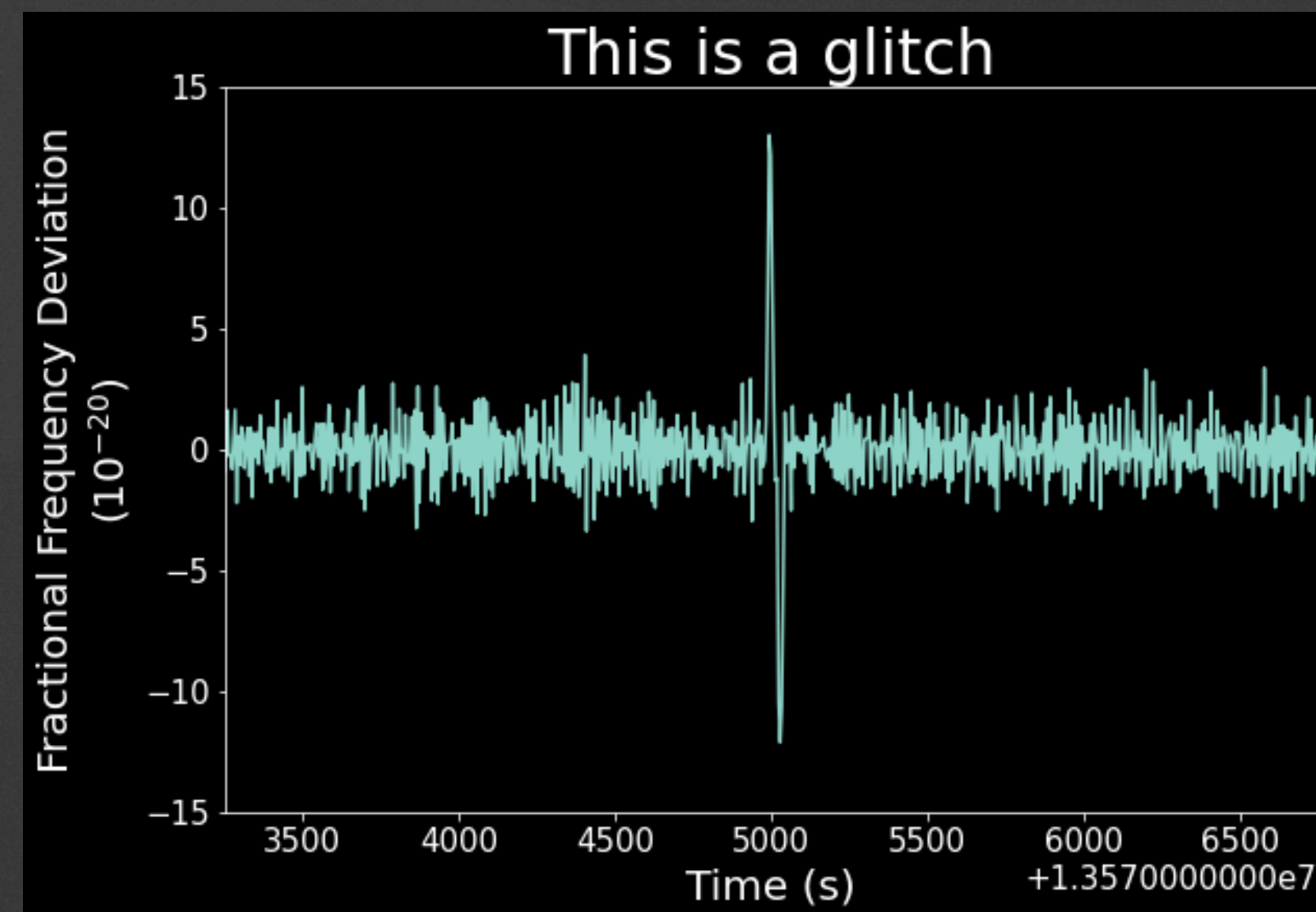
- Can remove glitches the same way as gaps
  - Set to 0
  - Window surrounding data
- Much harder to detect glitches



# Artefact Mitigation

## Glitches

- Can remove glitches the same way as gaps
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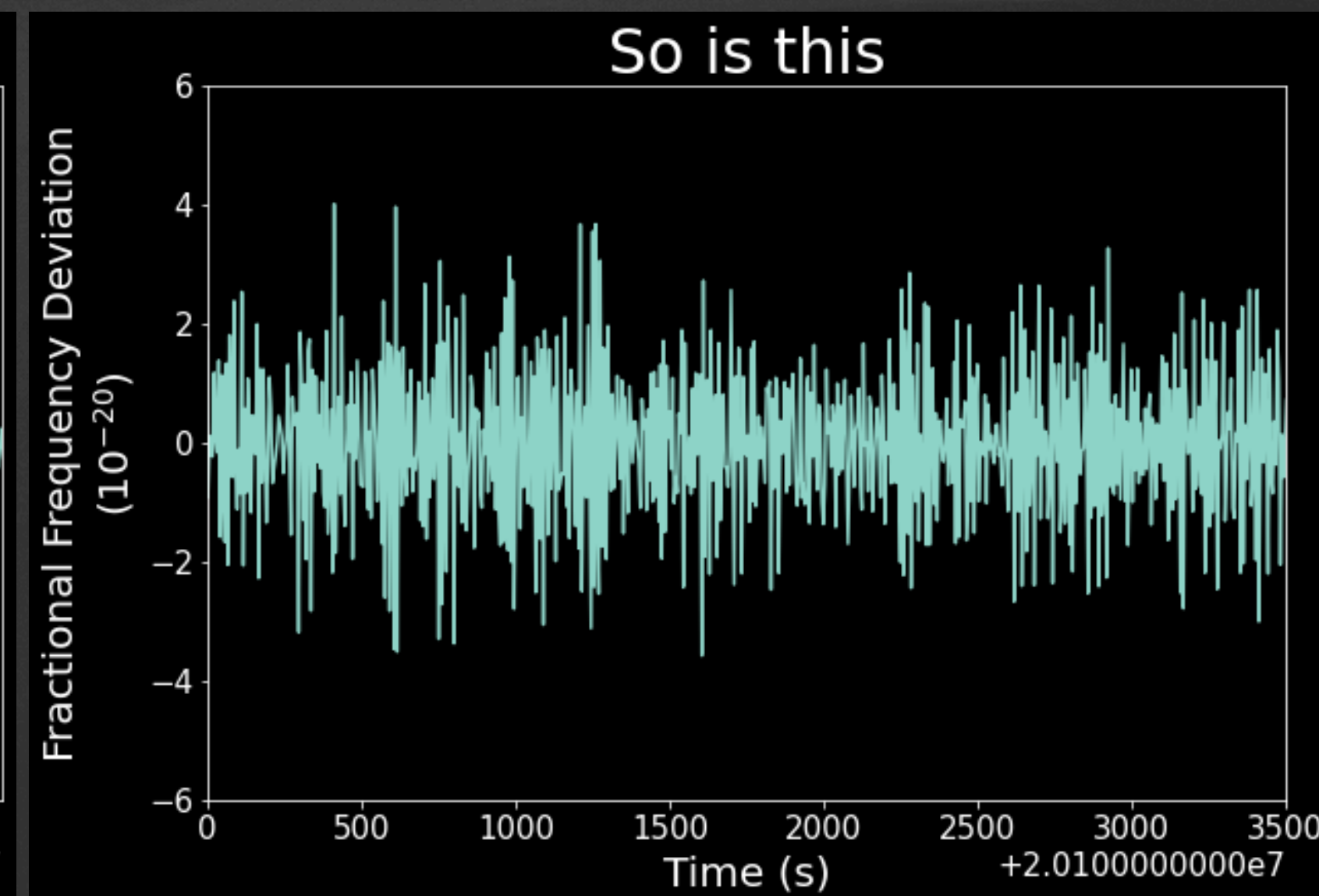
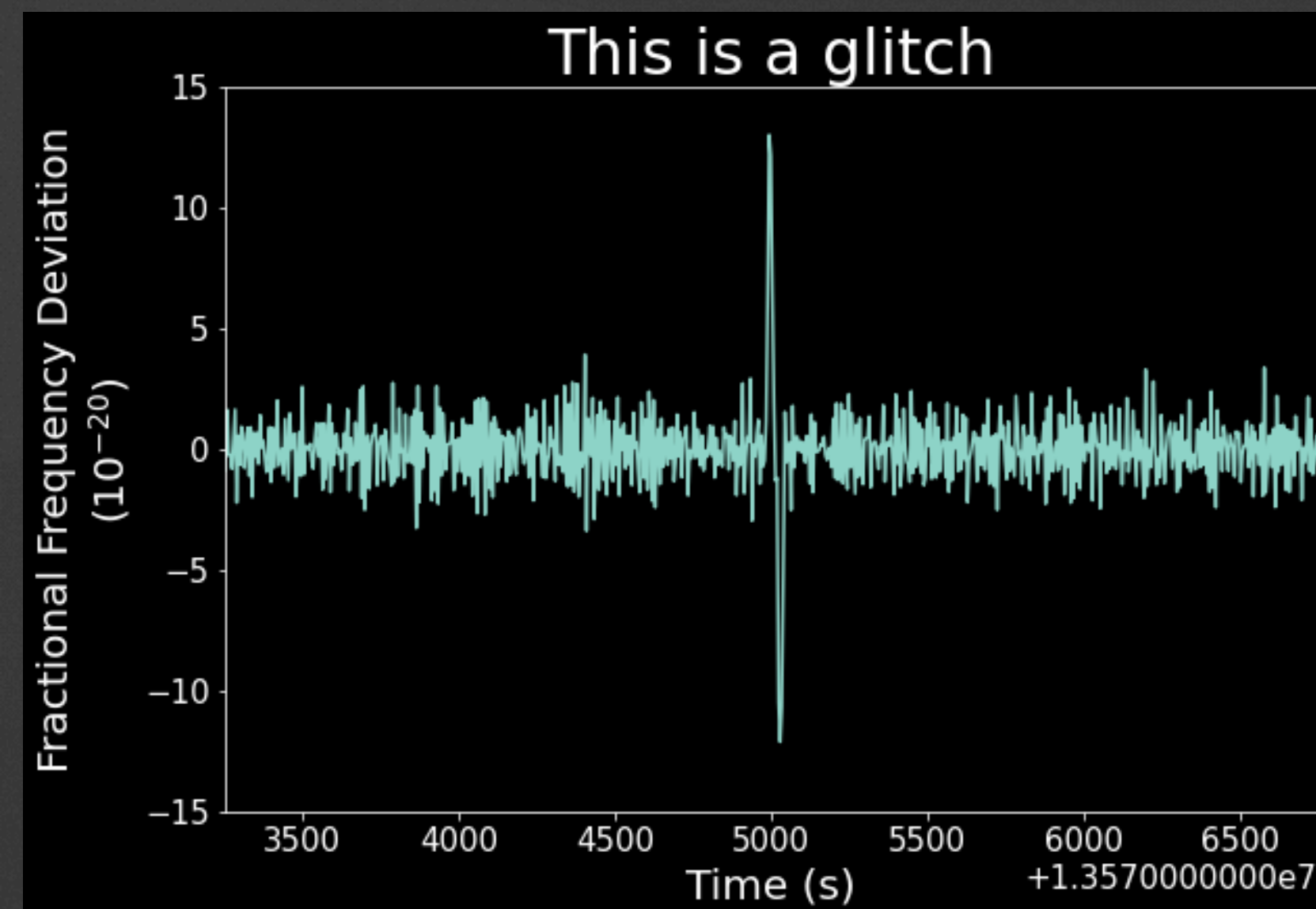




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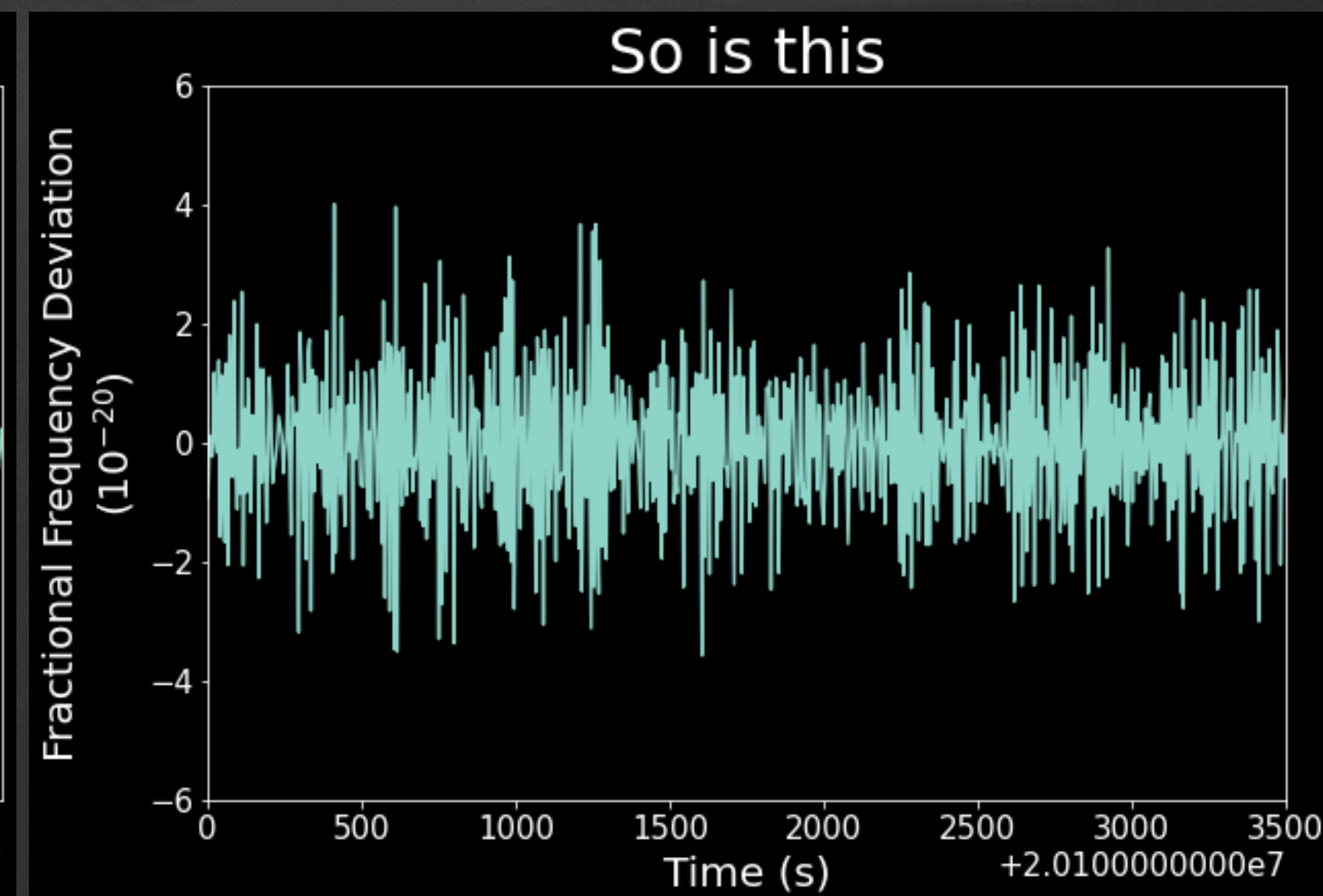
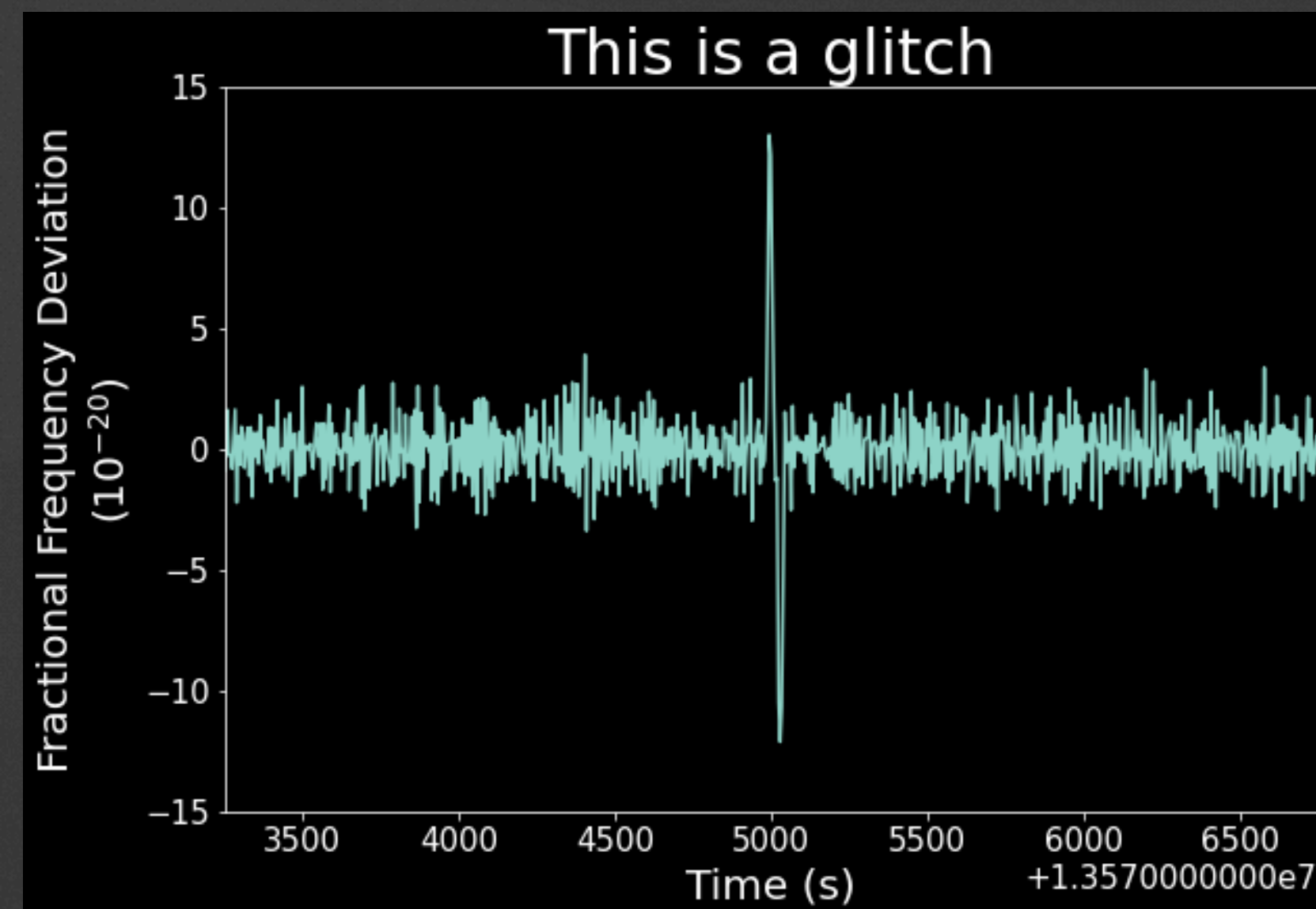




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## Glitches

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# Artefact Mitigation

## Detection Methods



# Artefact Mitigation

## Detection Methods

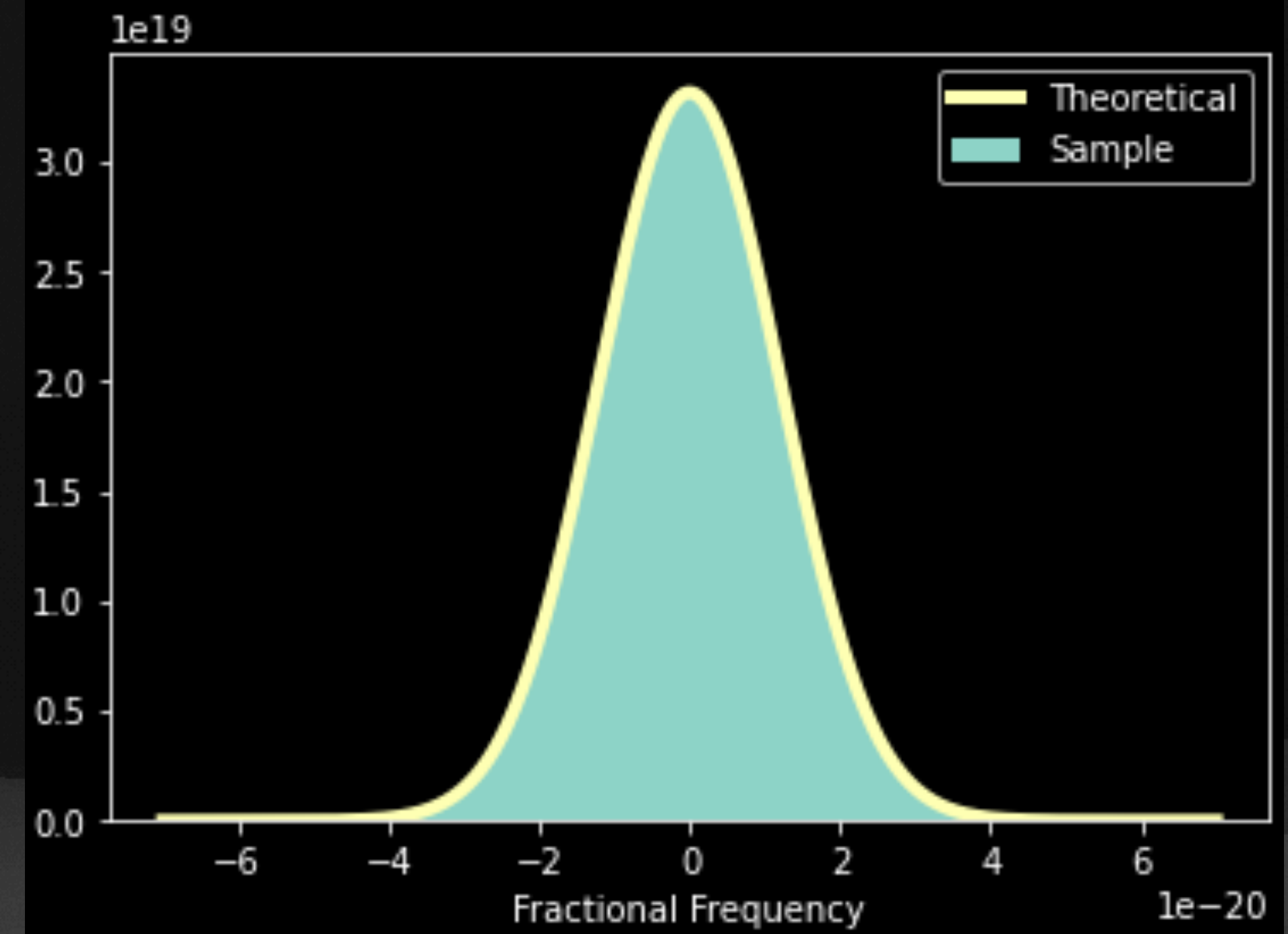
- Method I use is cut out glitches based on:



# Artefact Mitigation Detection Methods

- Method I use is cut out glitches based on:
  - Data amplitude ( $A$ )

Probability Density



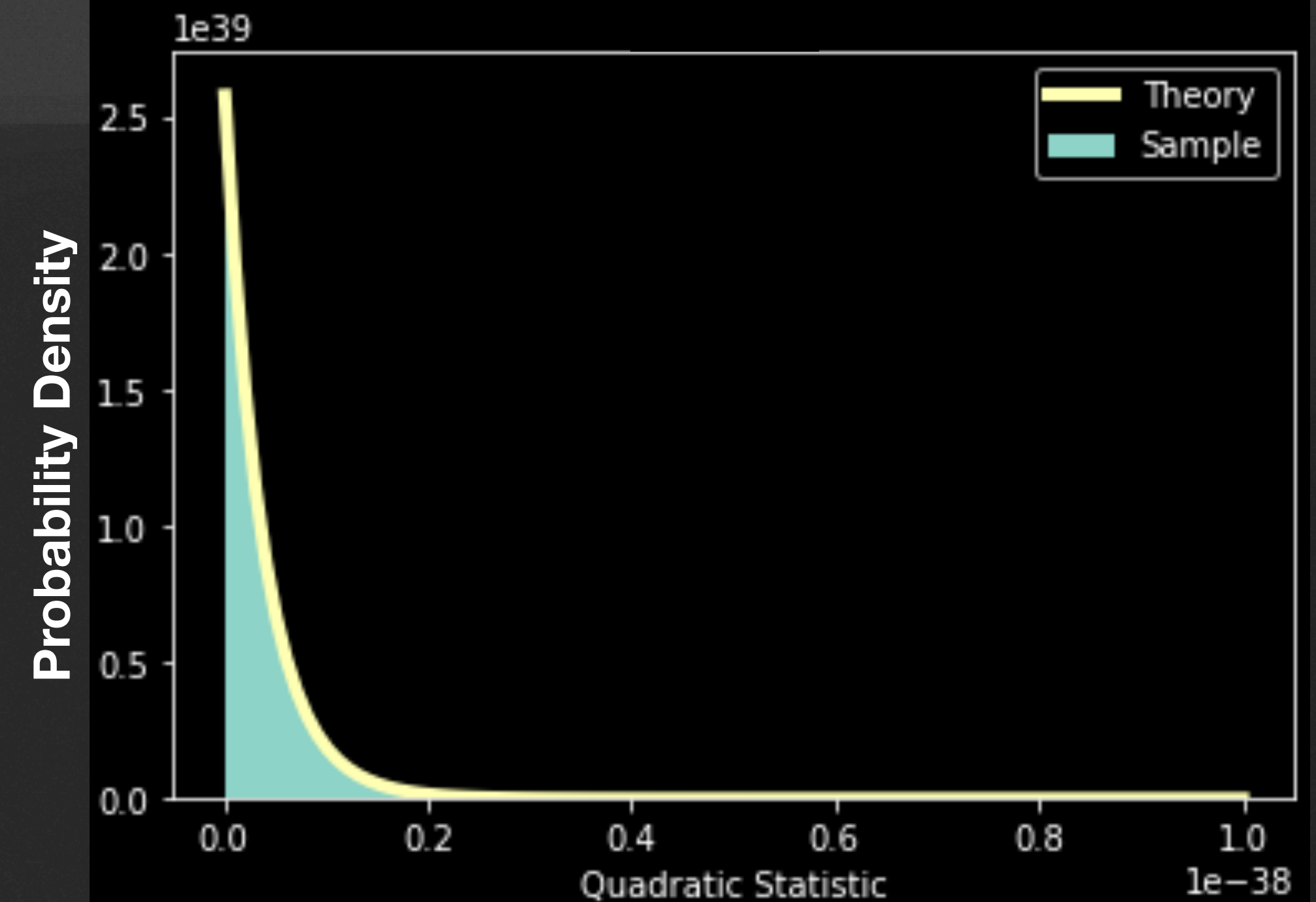
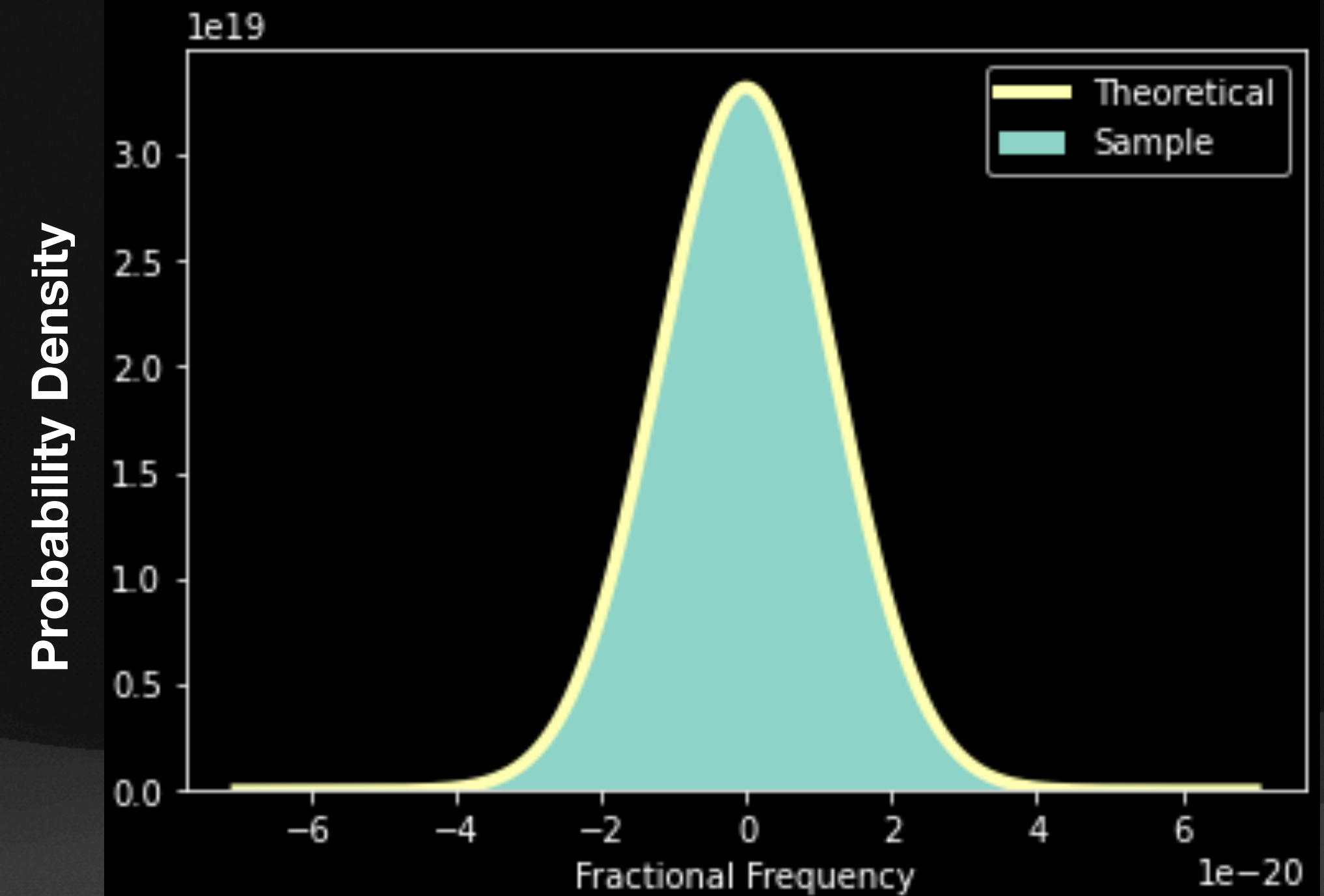


# Artefact Mitigation Detection Methods

- Method I use is cut out glitches based on:

- Data amplitude ( $A$ )

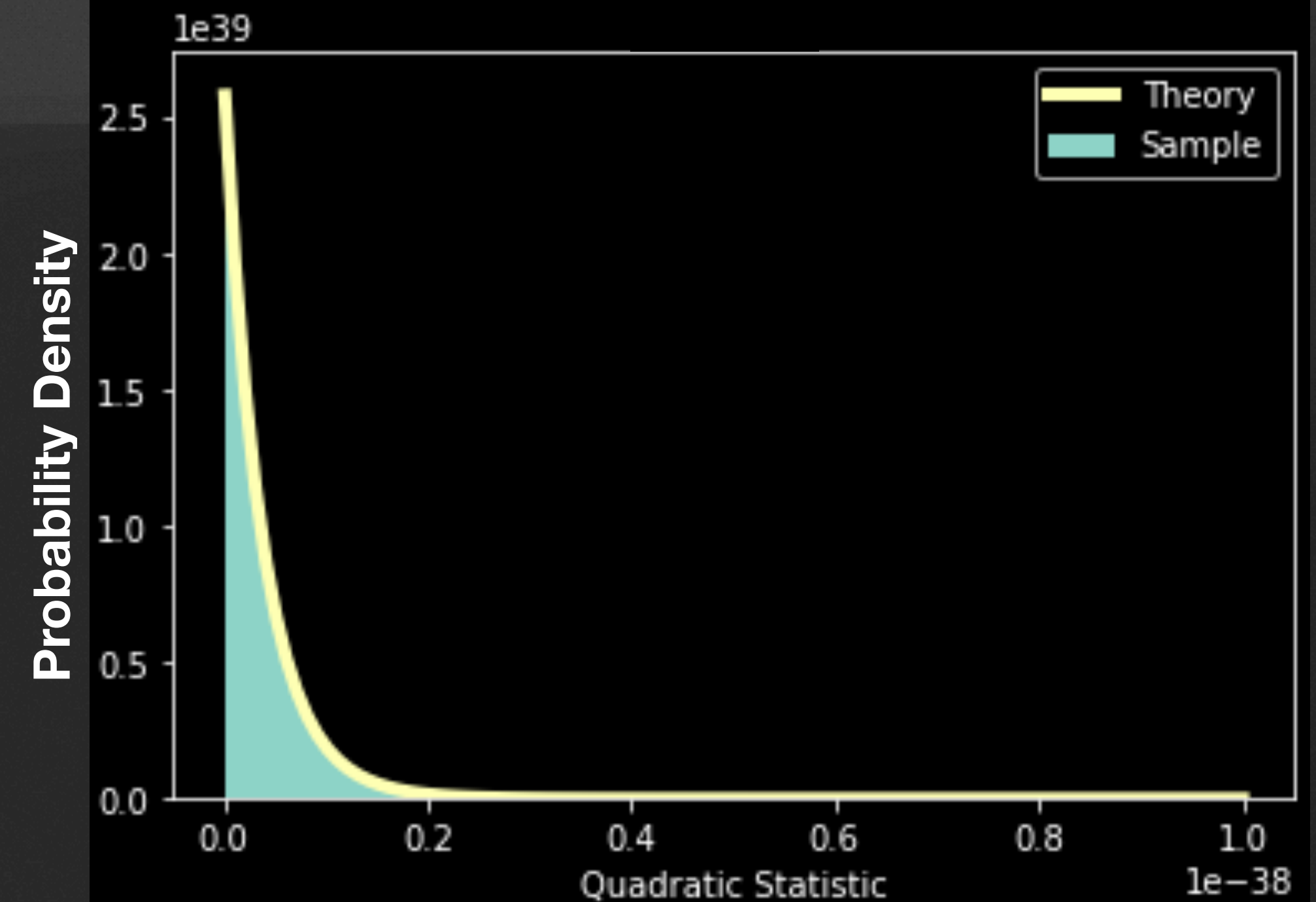
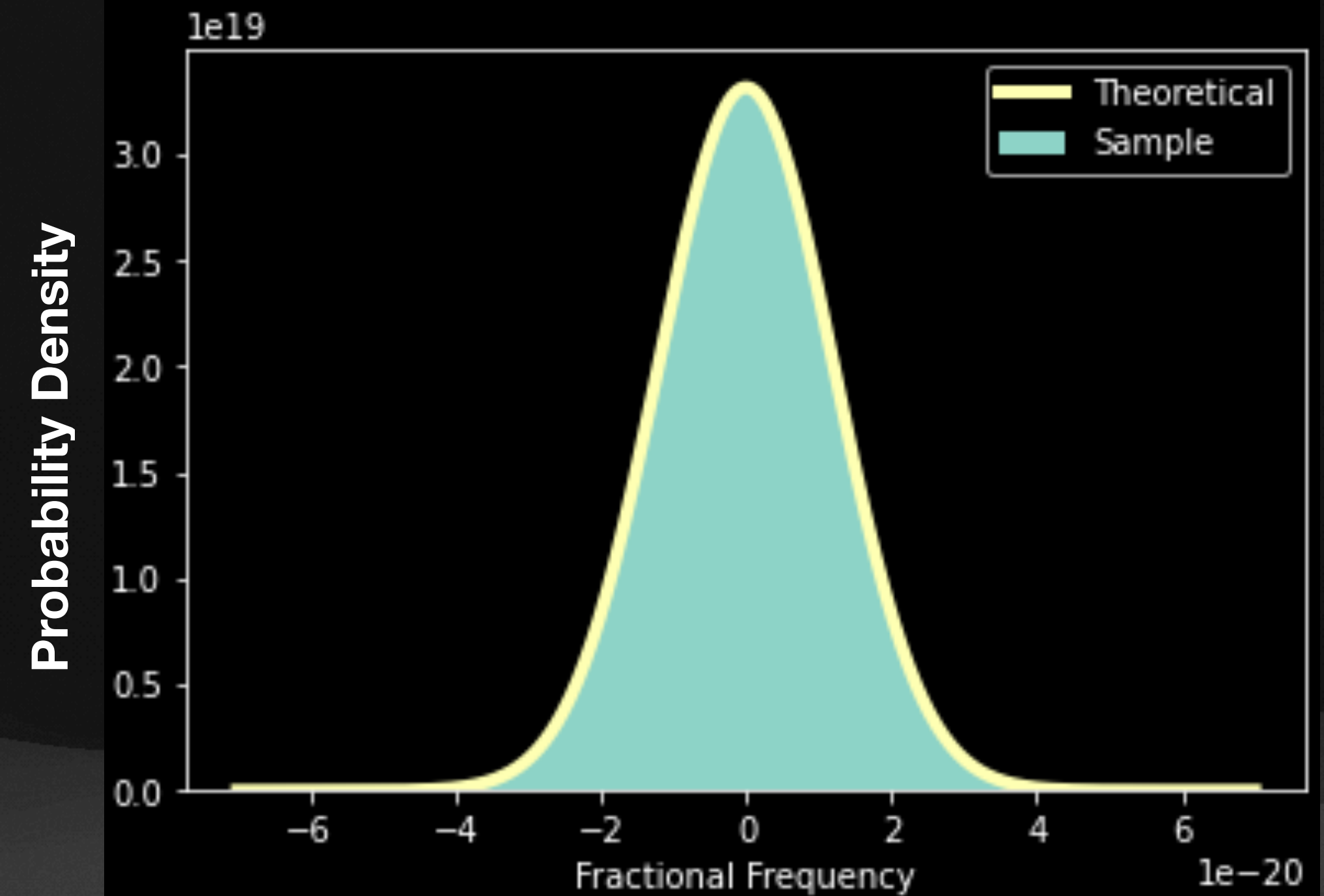
- “Quadratic Statistic”  $A^2 + \left(\frac{d}{dt}A\right)^2$





# Artefact Mitigation Detection Methods

- Method I use is cut out glitches based on:
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- Can derive expressions for PDFs in both regular and whitened data

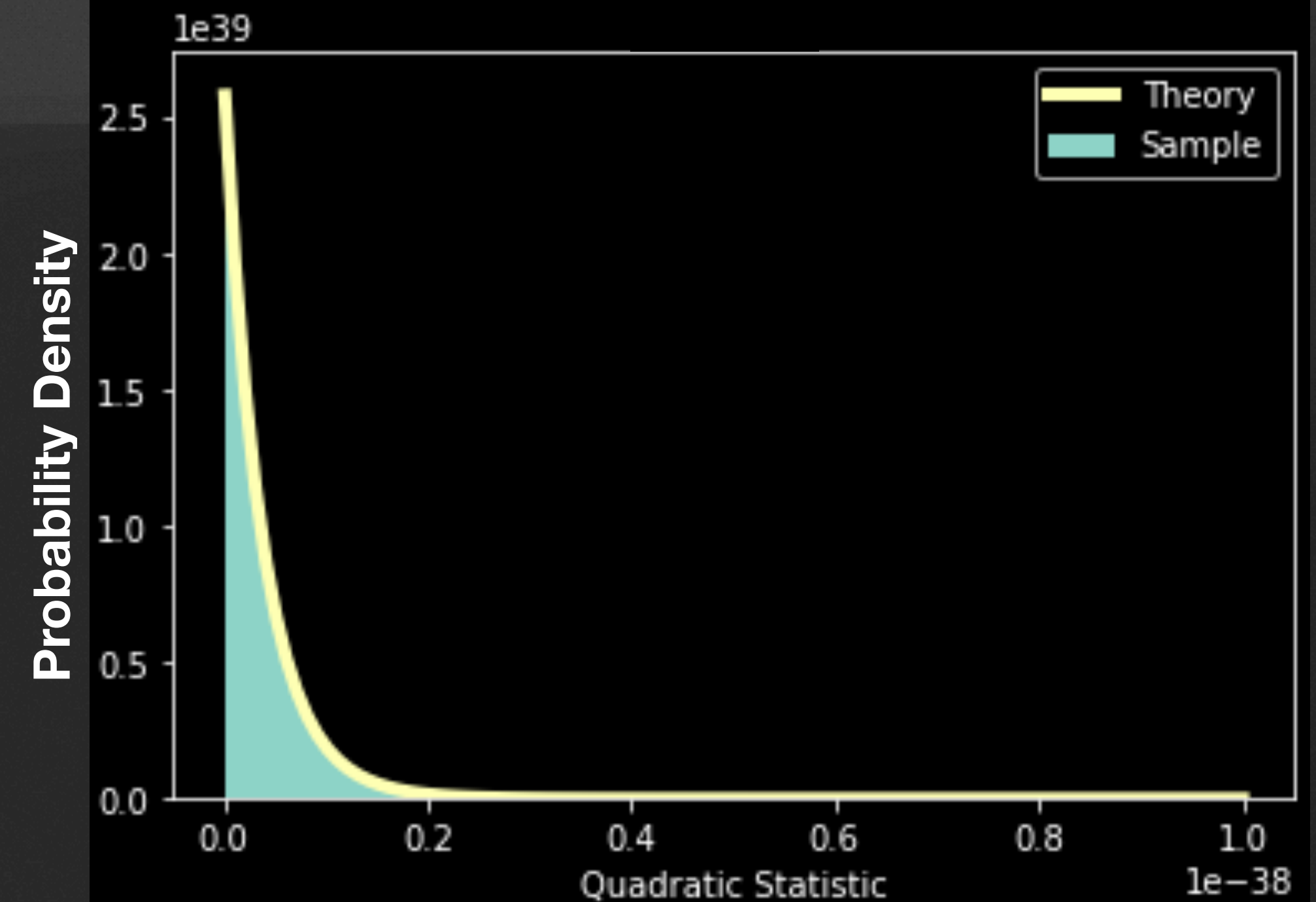
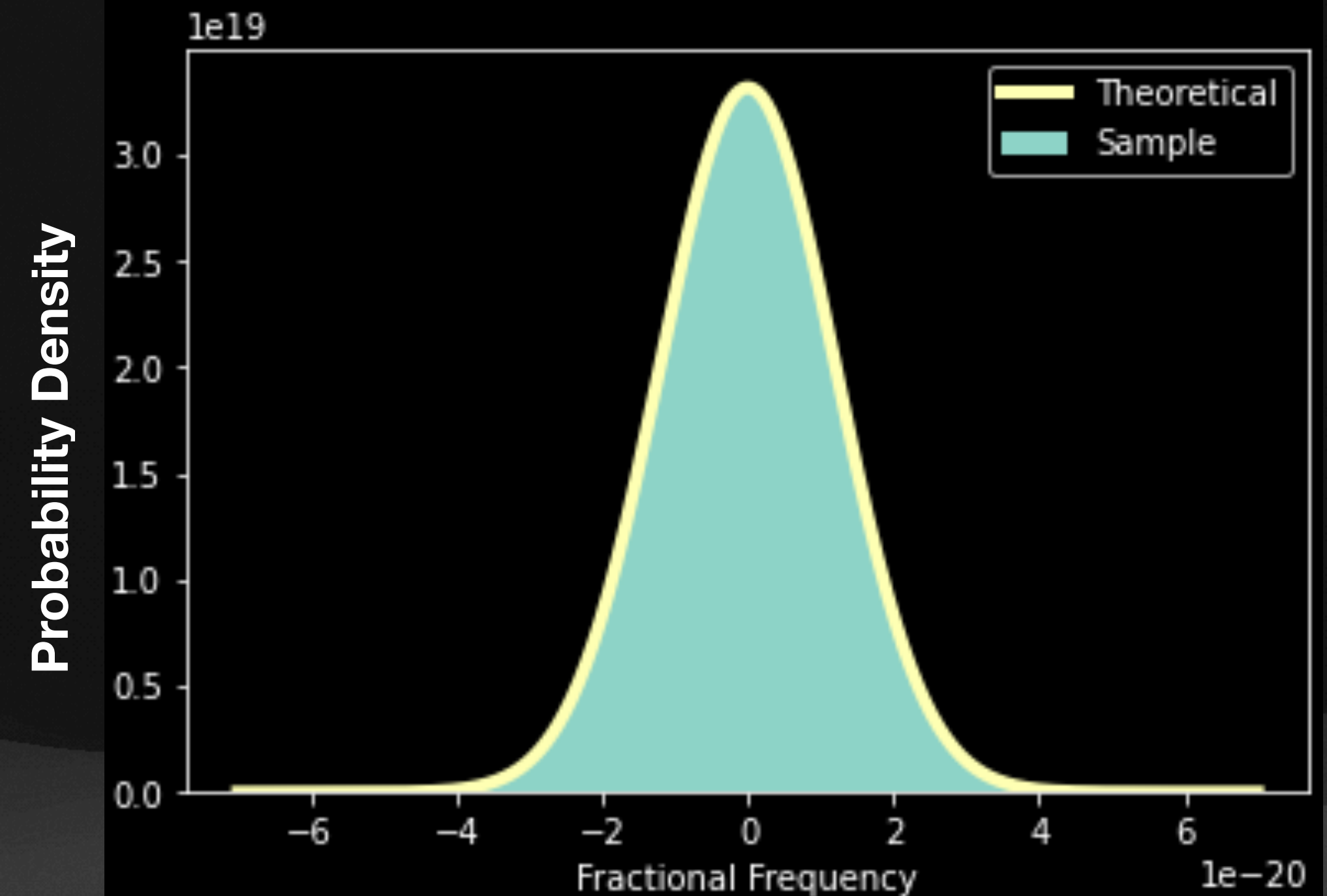




# Artefact Mitigation

## Detection Methods

- Method I use is cut out glitches based on:
  - Data amplitude ( $A$ )
  - “Quadratic Statistic”  $A^2 + \left(\frac{d}{dt}A\right)^2$
- Can derive expressions for PDFs in both regular and whitened data
- Choose how much data to lose, cut glitches based on choice





# Artefact Mitigation

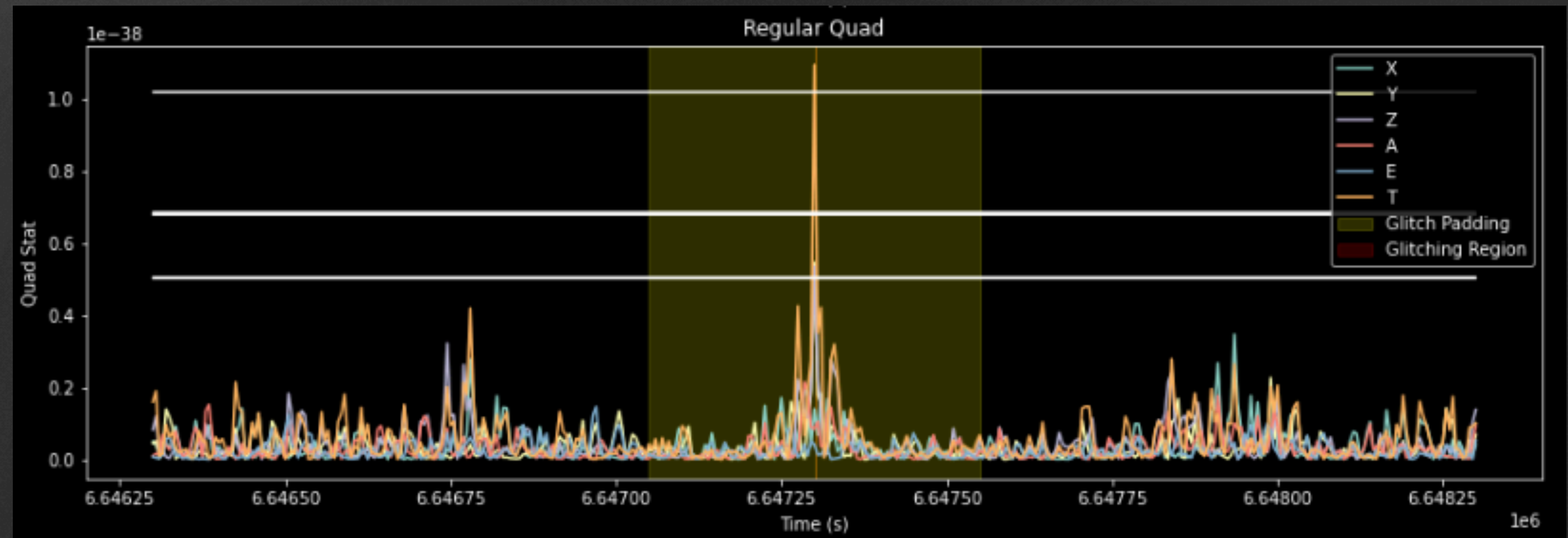
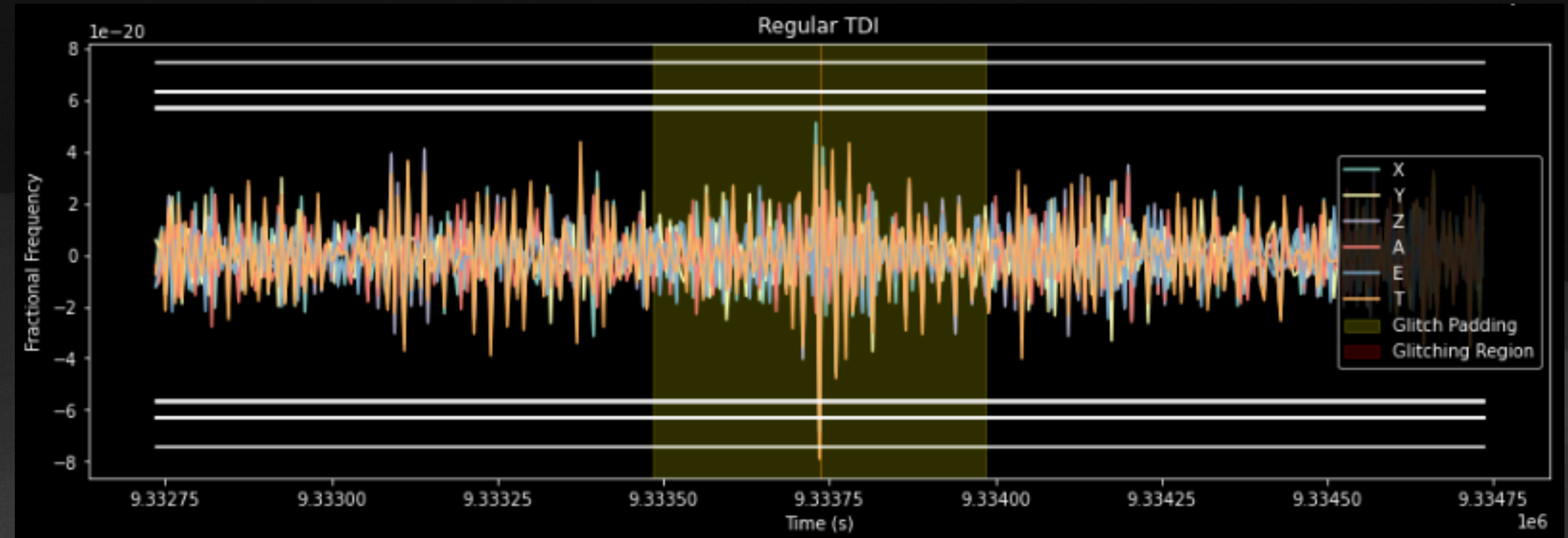
## Detection Methods



# Artefact Mitigation

## Detection Methods

- Both methods incorrectly identify lots of data as glitches

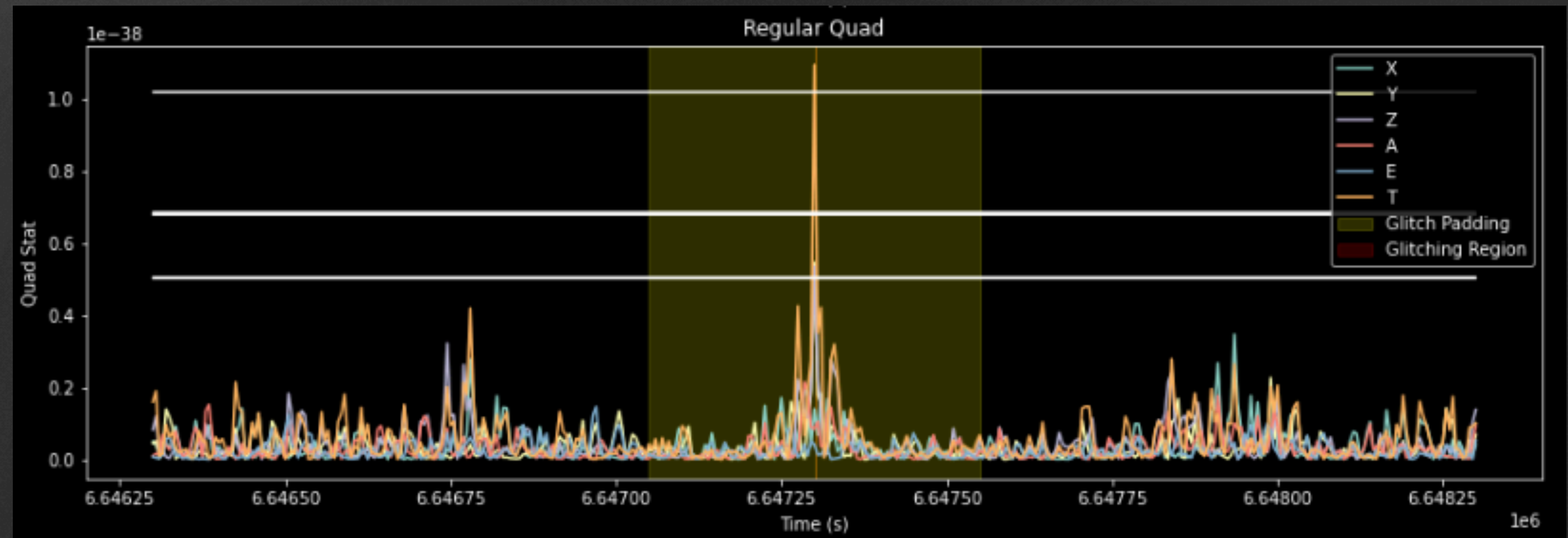
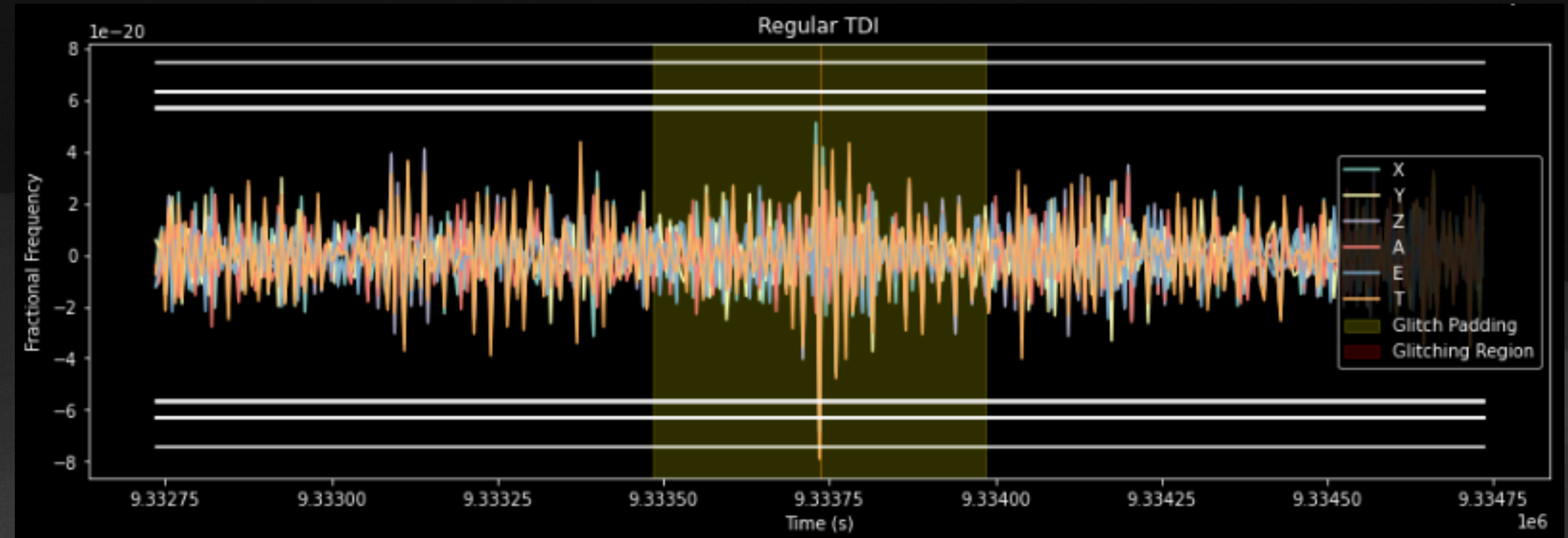




# Artefact Mitigation

## Detection Methods

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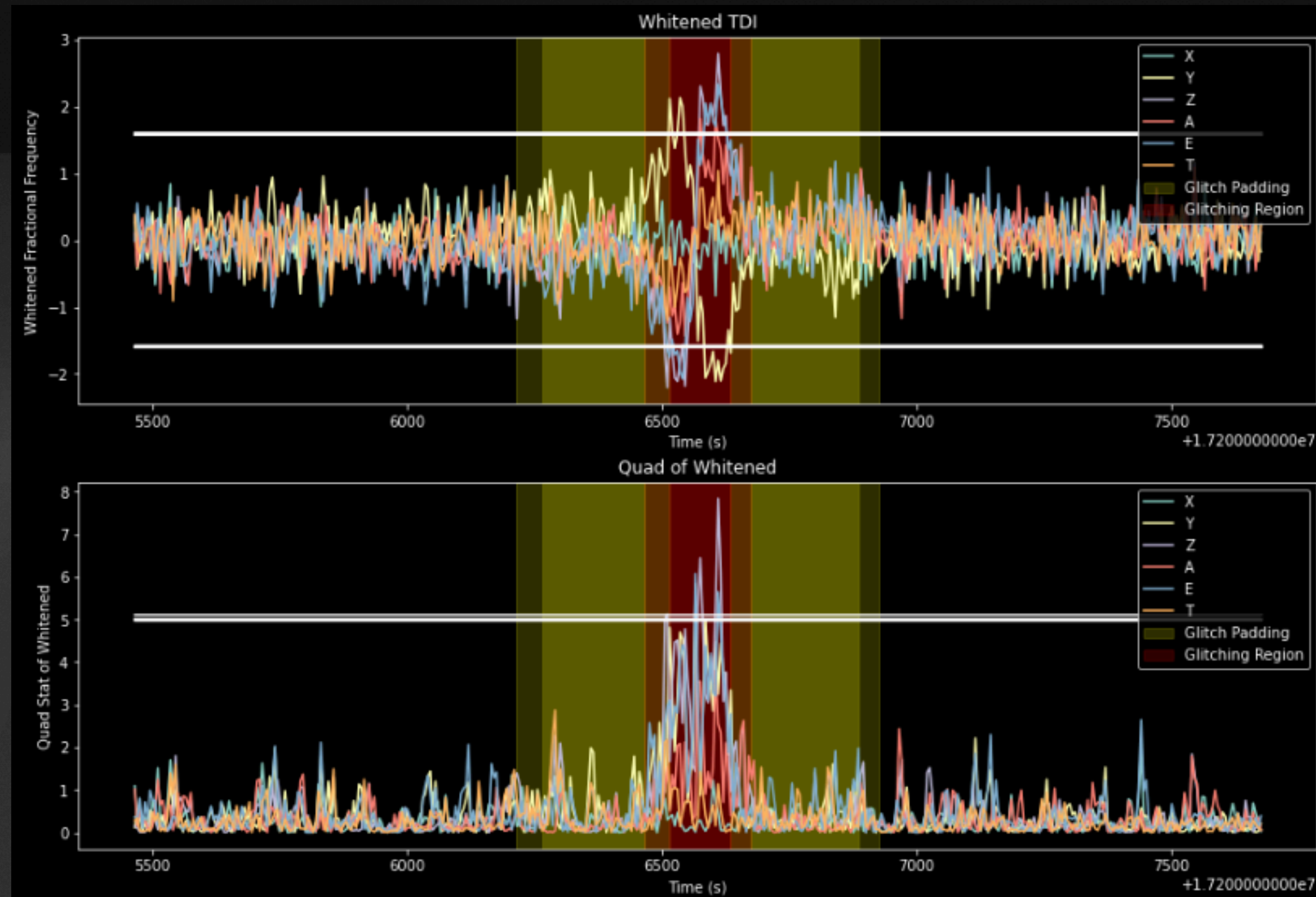




# Artefact Mitigation

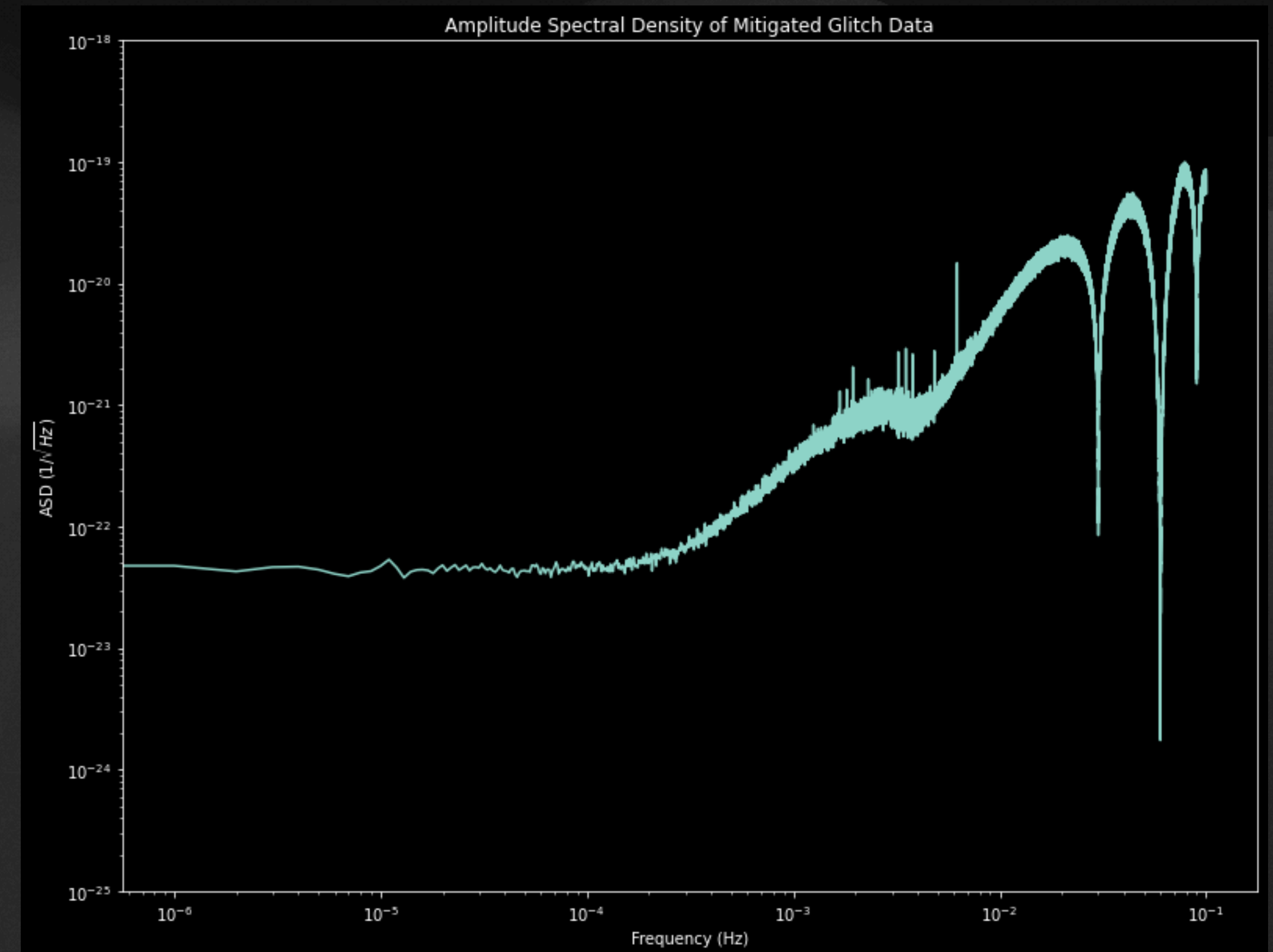
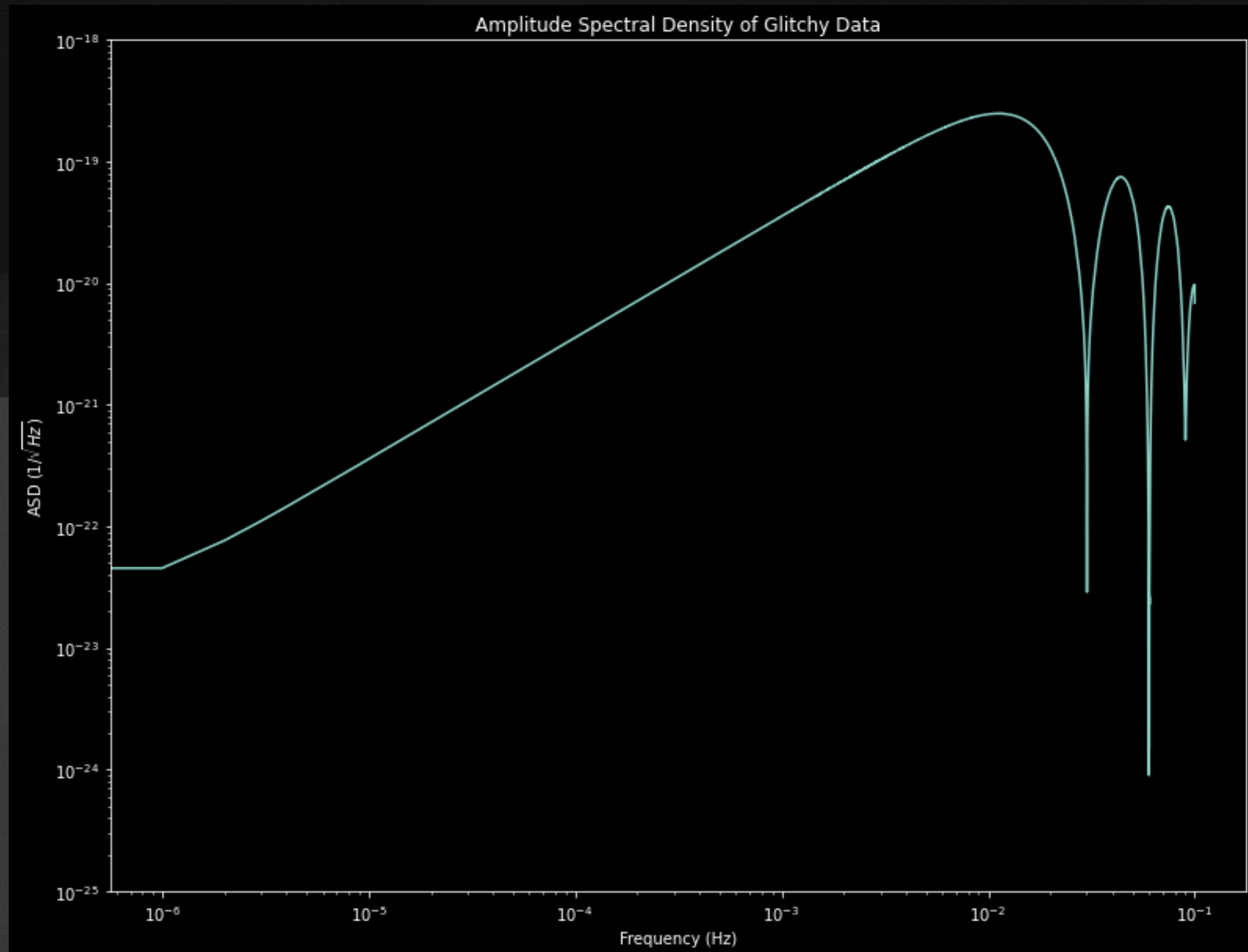
## Detection Methods

- Both methods incorrectly identify lots of data as glitches
- BUT
- Intersection of the two reduces false alarm rate drastically without affecting glitch detection





# Artefact Mitigation Results



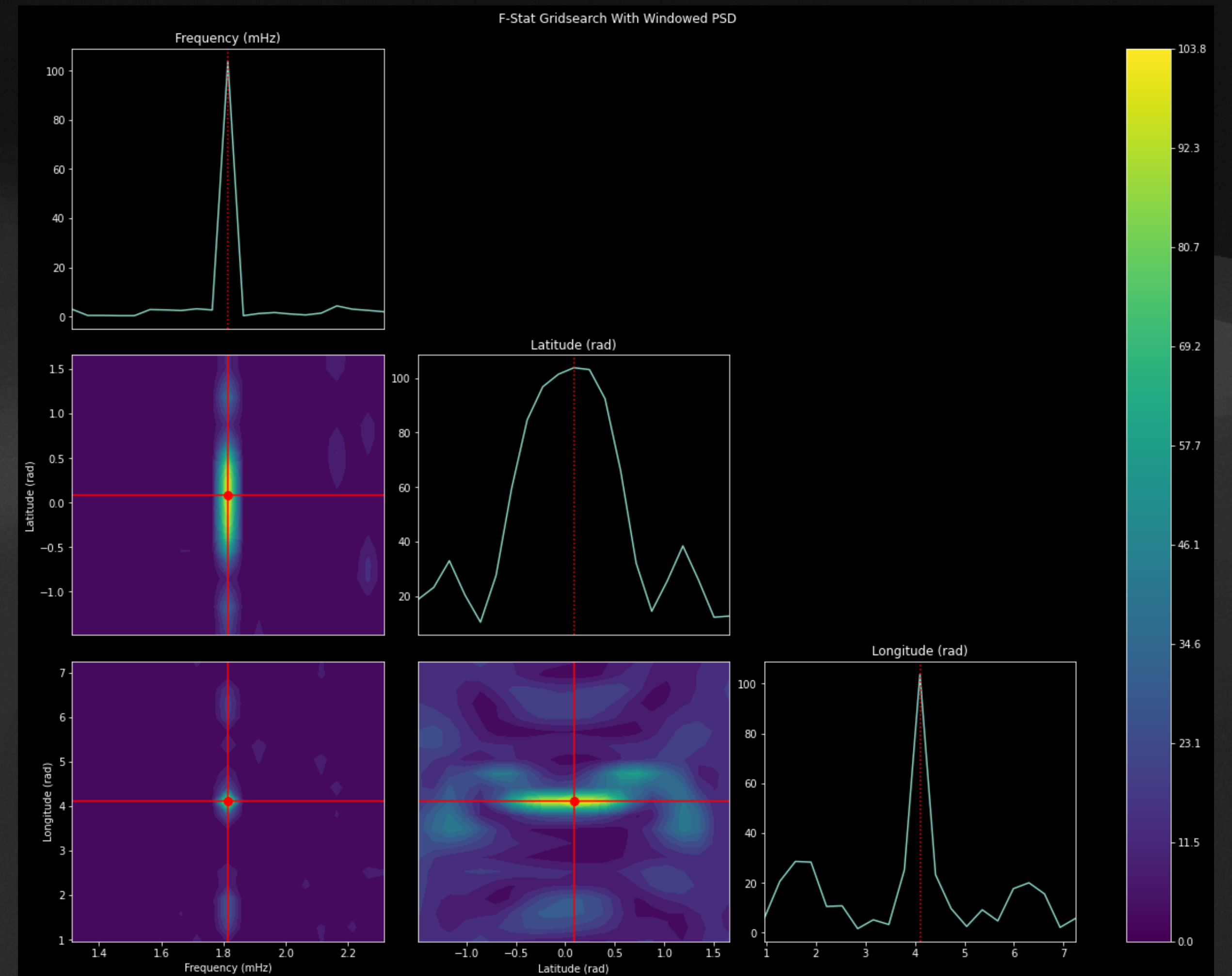


# Parameter Estimation



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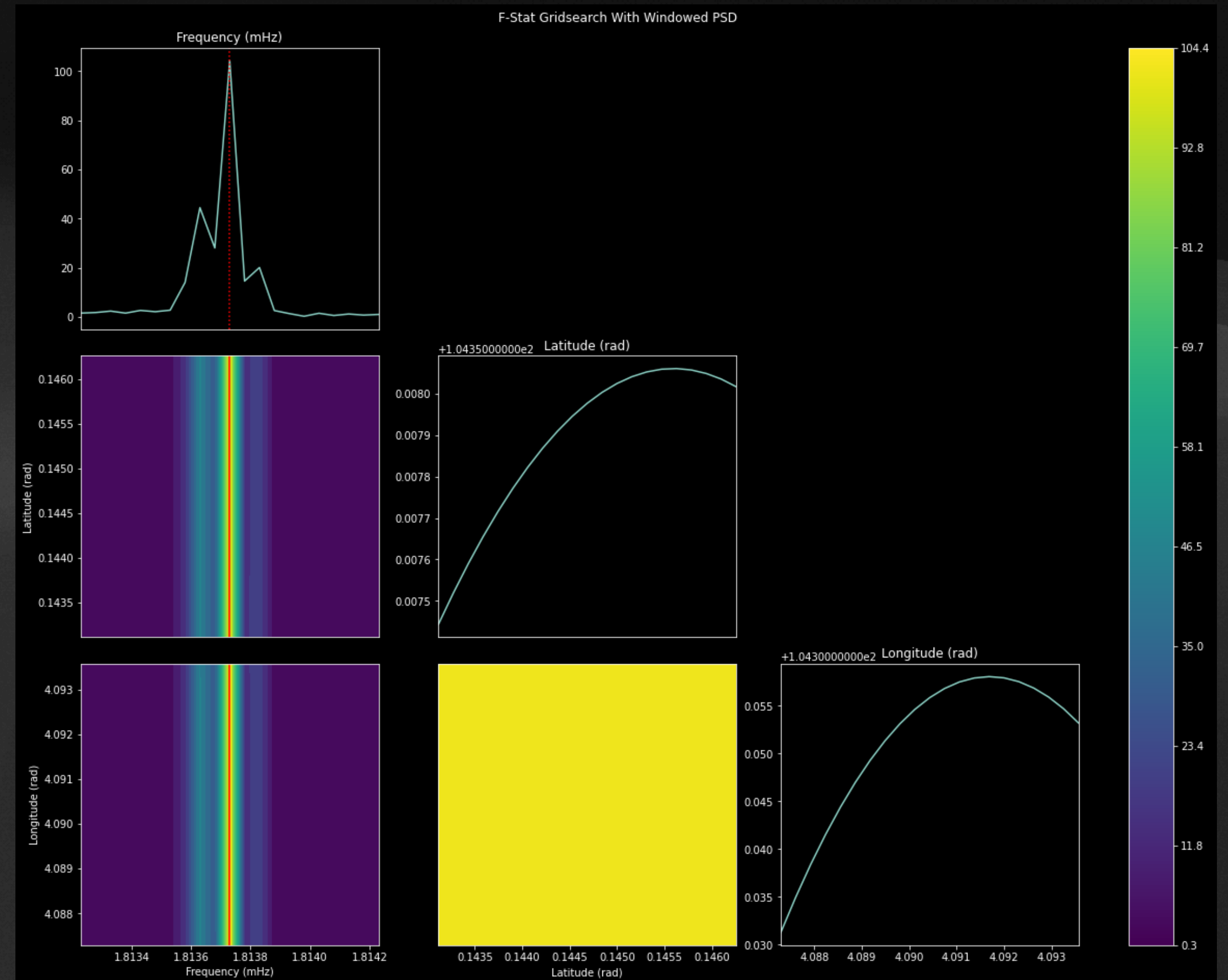
- Analytically maximize likelihood of four parameters, gridsearch over remainder





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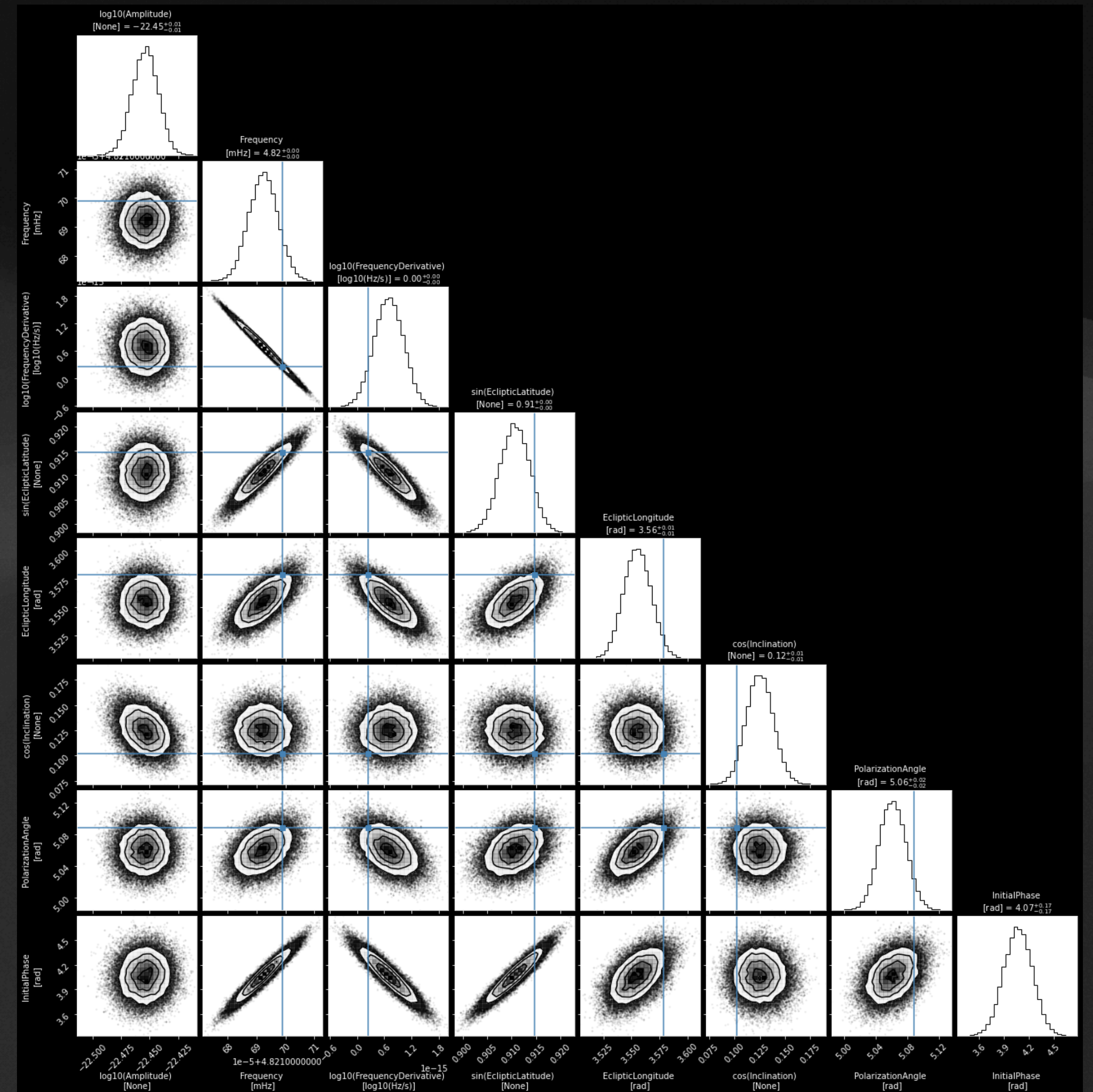
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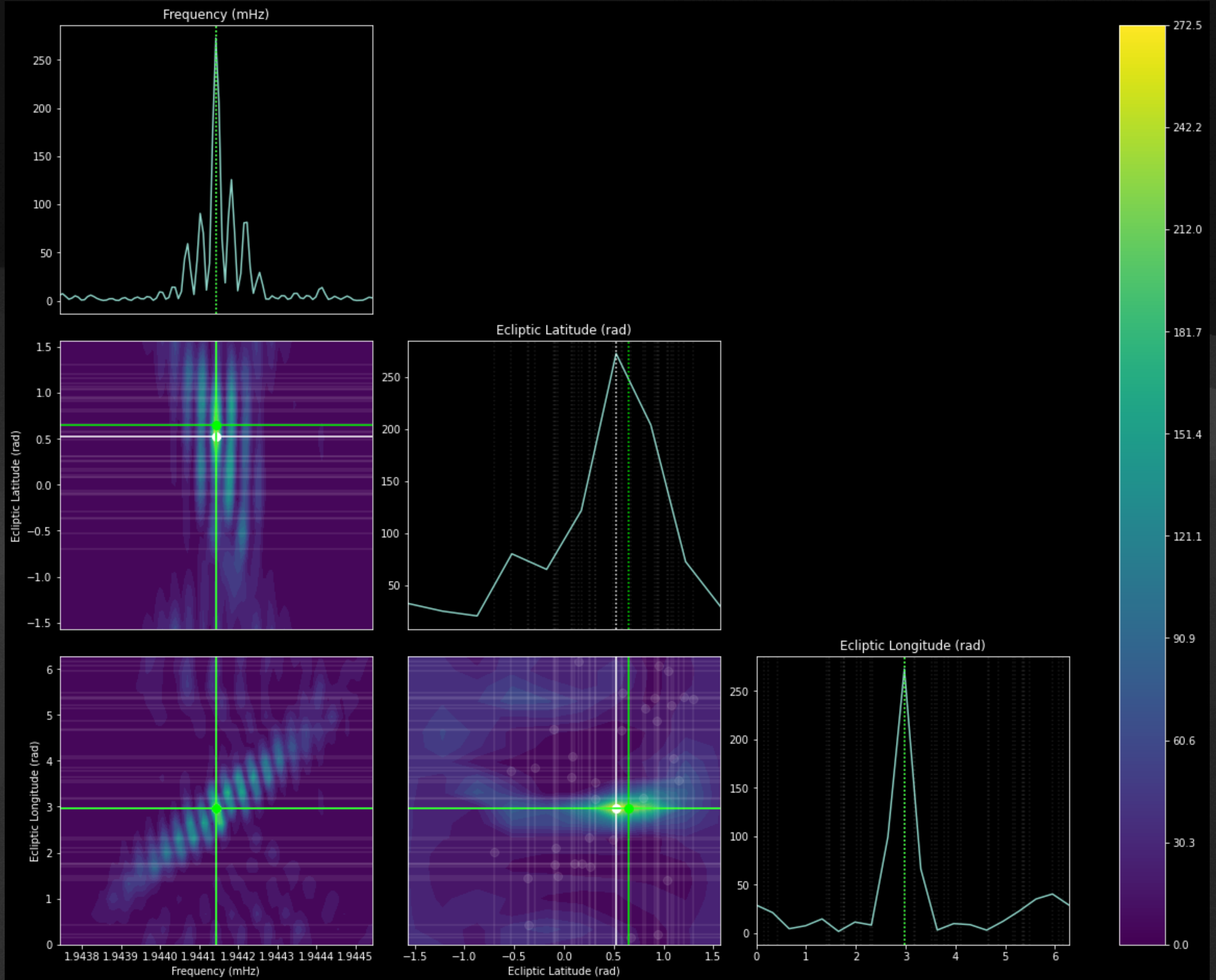
- Analytically maximize likelihood of four parameters, gridsearch over remainder
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- Start an MCMC sampler in the maxima and let it map out the posterior





# Example of Maximal Likelihood Gridsearch

## Results

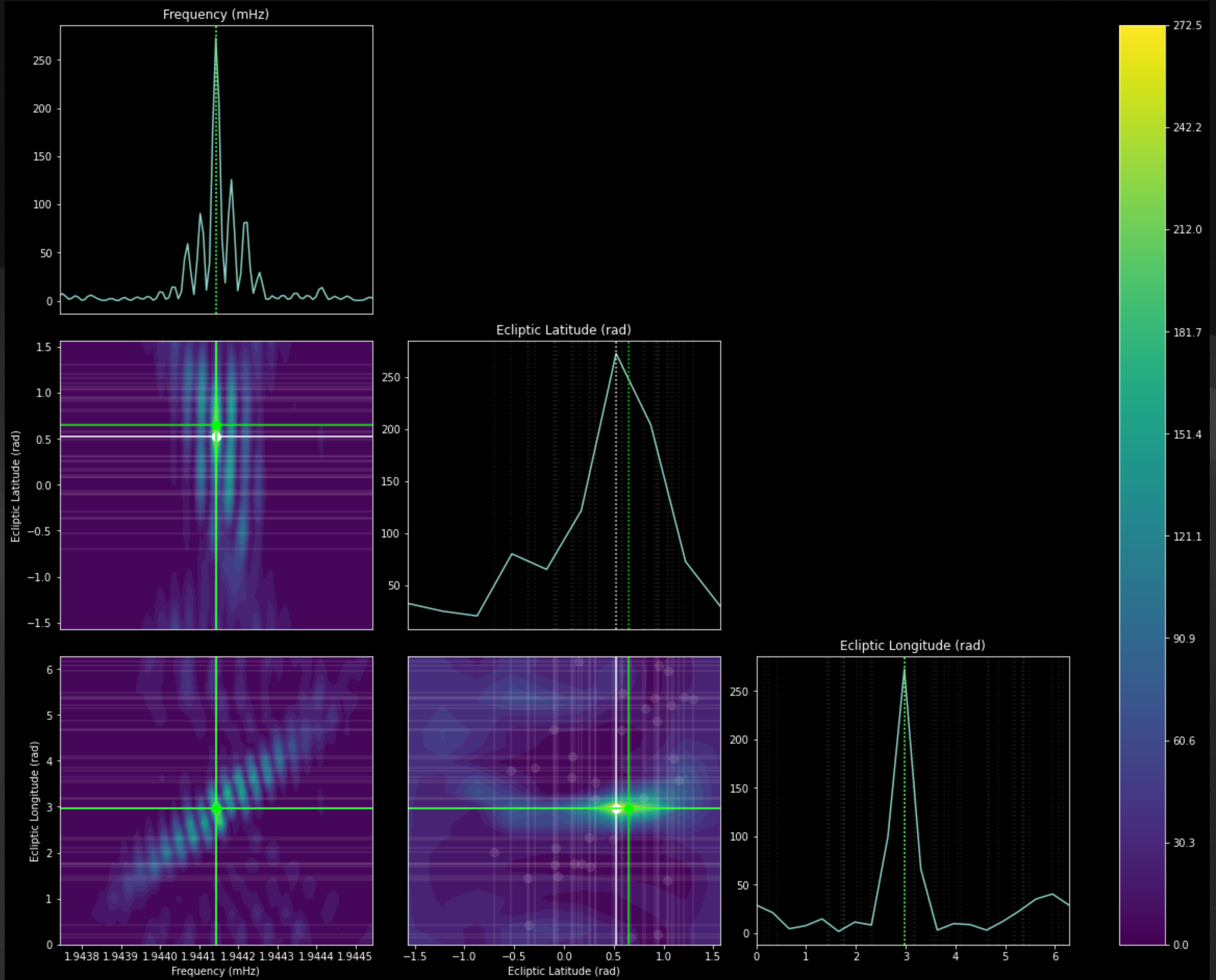




# Example of Maximal Likelihood Gridsearch

## Results

- All sources detected in provided artefact-free dataset also detected in full dataset using this procedure

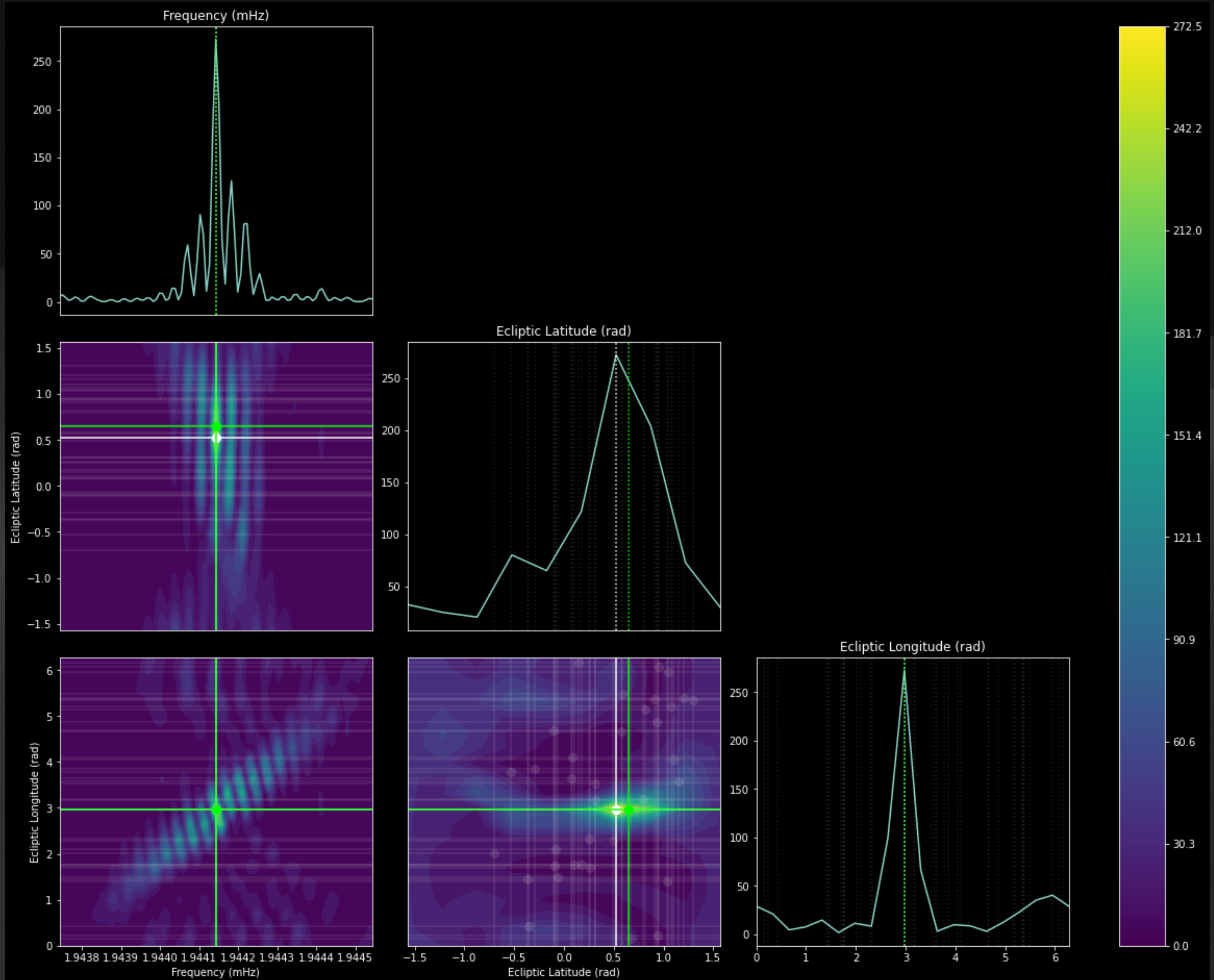




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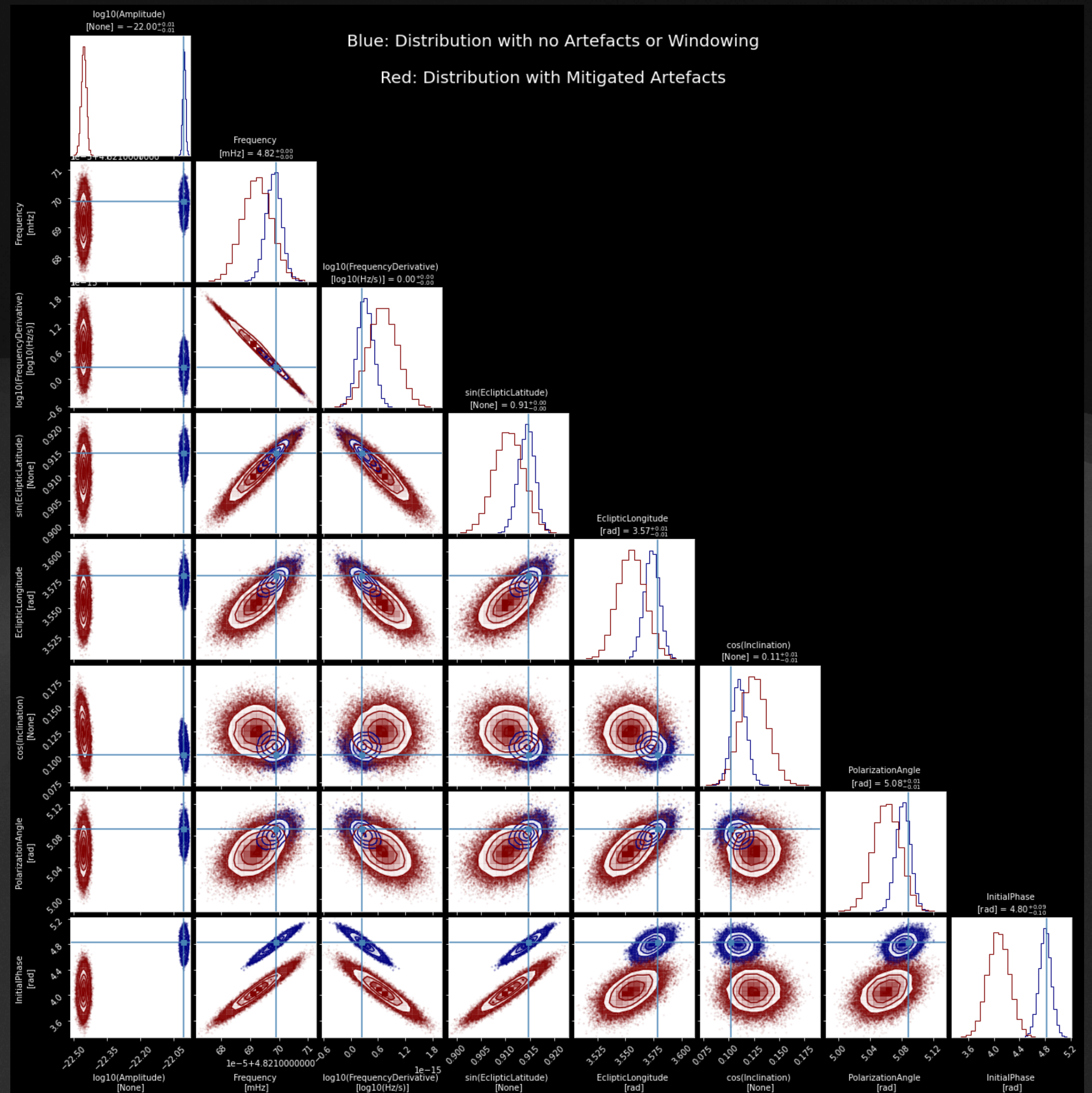
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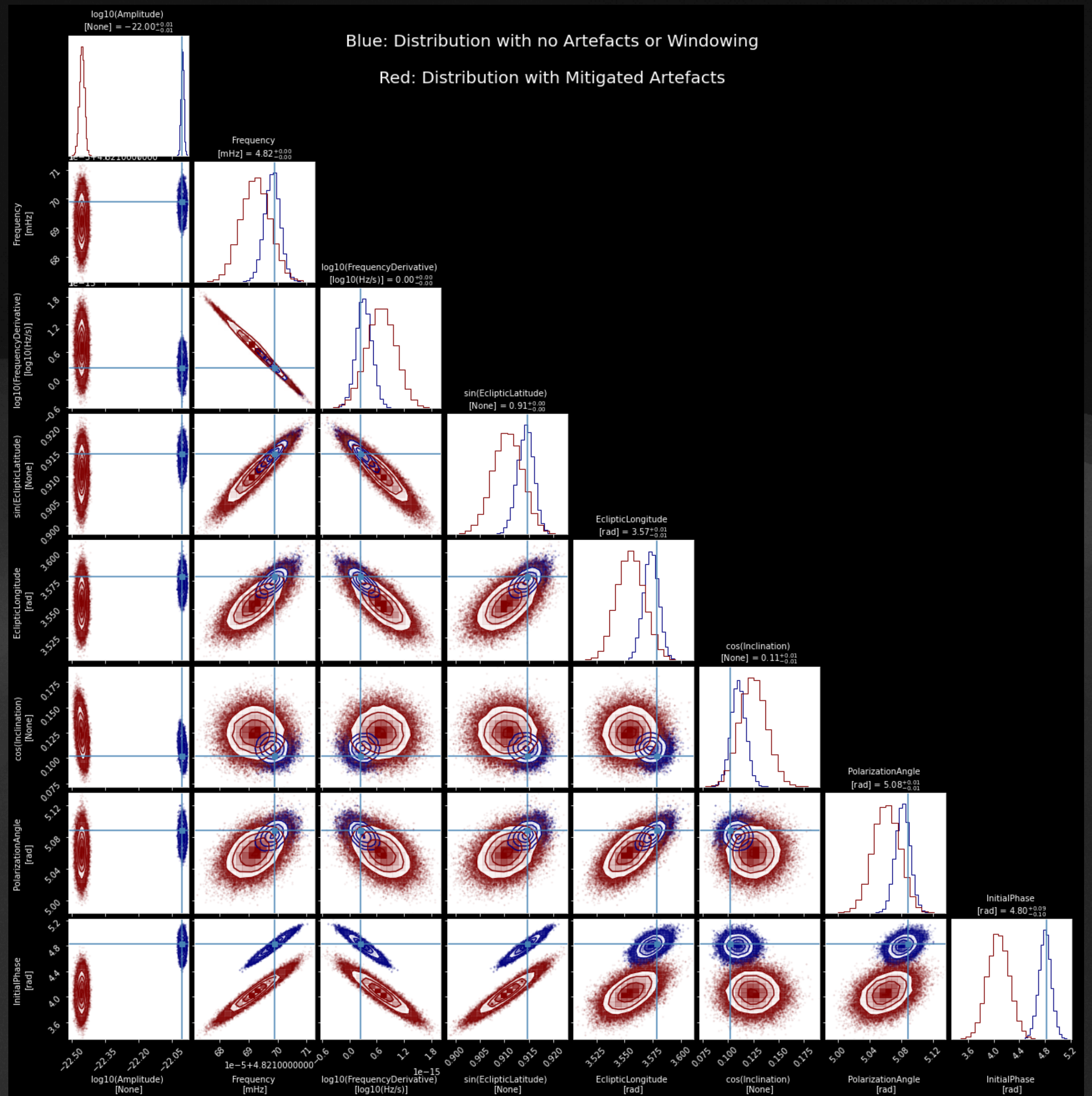
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- MCMC after artefact mitigation is biased, especially in amplitude and initial phase





# Results

- All sources detected in provided artefact-free dataset also detected in full dataset using this procedure
- Maximized likelihood gridsearch effective at narrowing in on true source parameters
- MCMC after artefact mitigation is biased, especially in amplitude and initial phase
- Including the windowing from mitigation in likelihood function probably helpful - Working on this now





# Questions