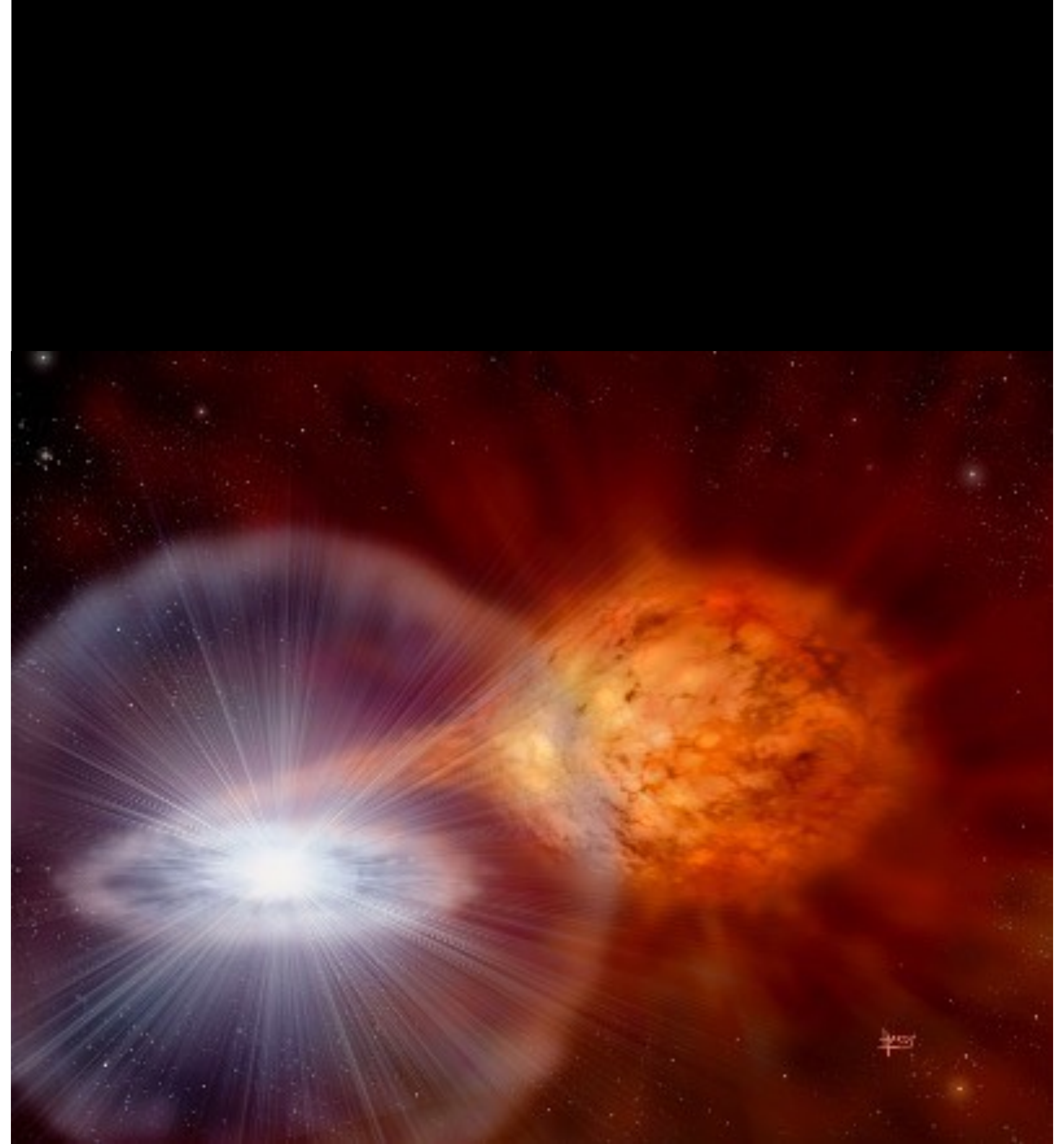


Precision mass measurements of neutron- deficient strontium

Implications for the rp-process and isospin
symmetry

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WNPPC
Feb 16th-19th, 2023



Type-I X-ray bursts

Type-I X-ray bursts: periodic thermonuclear explosions on the surfaces of accreting neutron stars

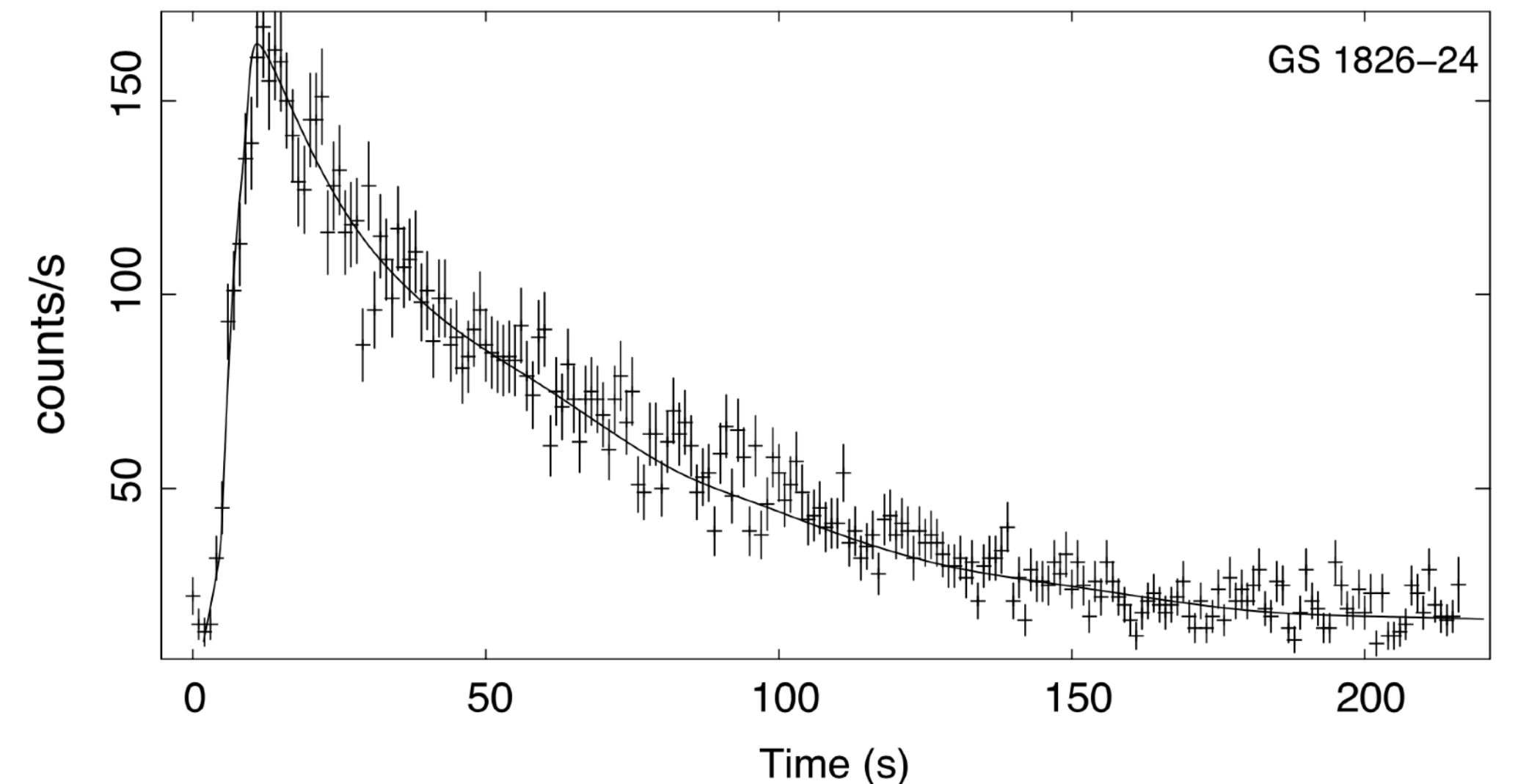
- Over 100 X-ray bursters are identified in our galaxy [1]
- Burst produces a light curve that is observable to space-based X-ray telescopes

Reasons to study:

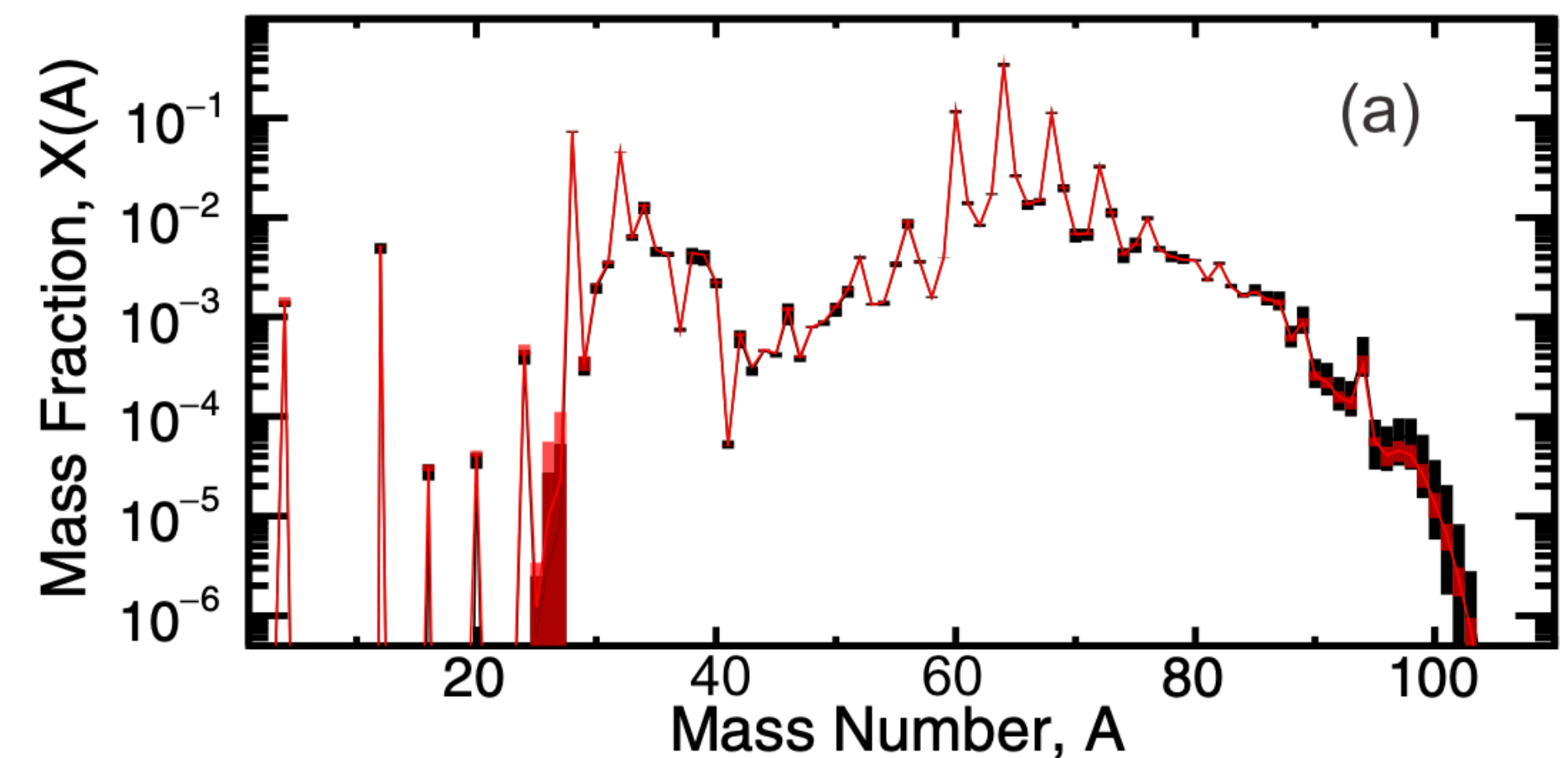
- Study nuclear processes which power the explosion
- Study chemical composition of neutron star surface [2]
- Source of nucleosynthesis?

1. Jean in 't Zand, <https://personal.sron.nl/~jeanz/bursterlist.html>

2. Meisel Z *et. al.* Journal of Physics G: Nuclear and Particle Physics. 2018 Jul 25;45(9):093001.



Example light curve (Parikh, A., et al. *Progress in Particle and Nuclear Physics* 69 (2013))

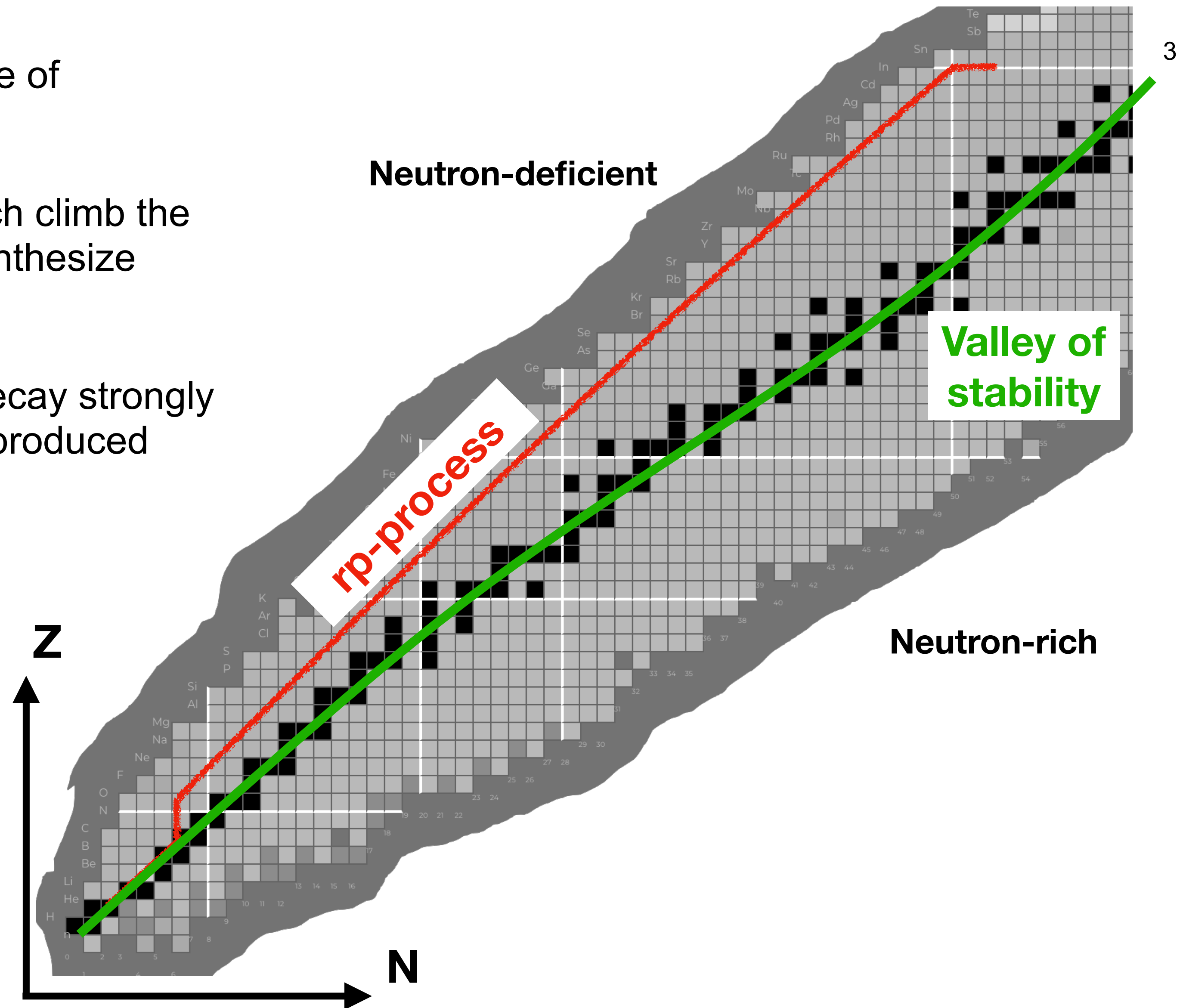


Simulated burst ashes (Hoff, D. E. M., et al. *Physical Review C* 102.4 (2020): 045810.)

Type-I X-ray bursts

Rapid proton capture process (rp-process): source of nucleosynthesis for Type-I X-ray bursts

- **Chain of proton captures and β^+ decays** which climb the neutron-deficient side of the nuclear chart to synthesize increasingly heavier isotopes
- Interplay of charged particle reactions and β^+ decay strongly influence overall mass flow and nuclear ashes produced
- **Precise mass values important component for accurate calculation** of astrophysical reaction rates involved in rp-process [1]
- **Many $A=60-100$ masses along the rp-process are not accurately measured** [2]



1. Schatz H. International Journal of Mass Spectrometry. 2013 Sep 1;349:181-6.
2. Schatz H, Ong WJ. The Astrophysical Journal. 2017 Aug 1;844(2):139.

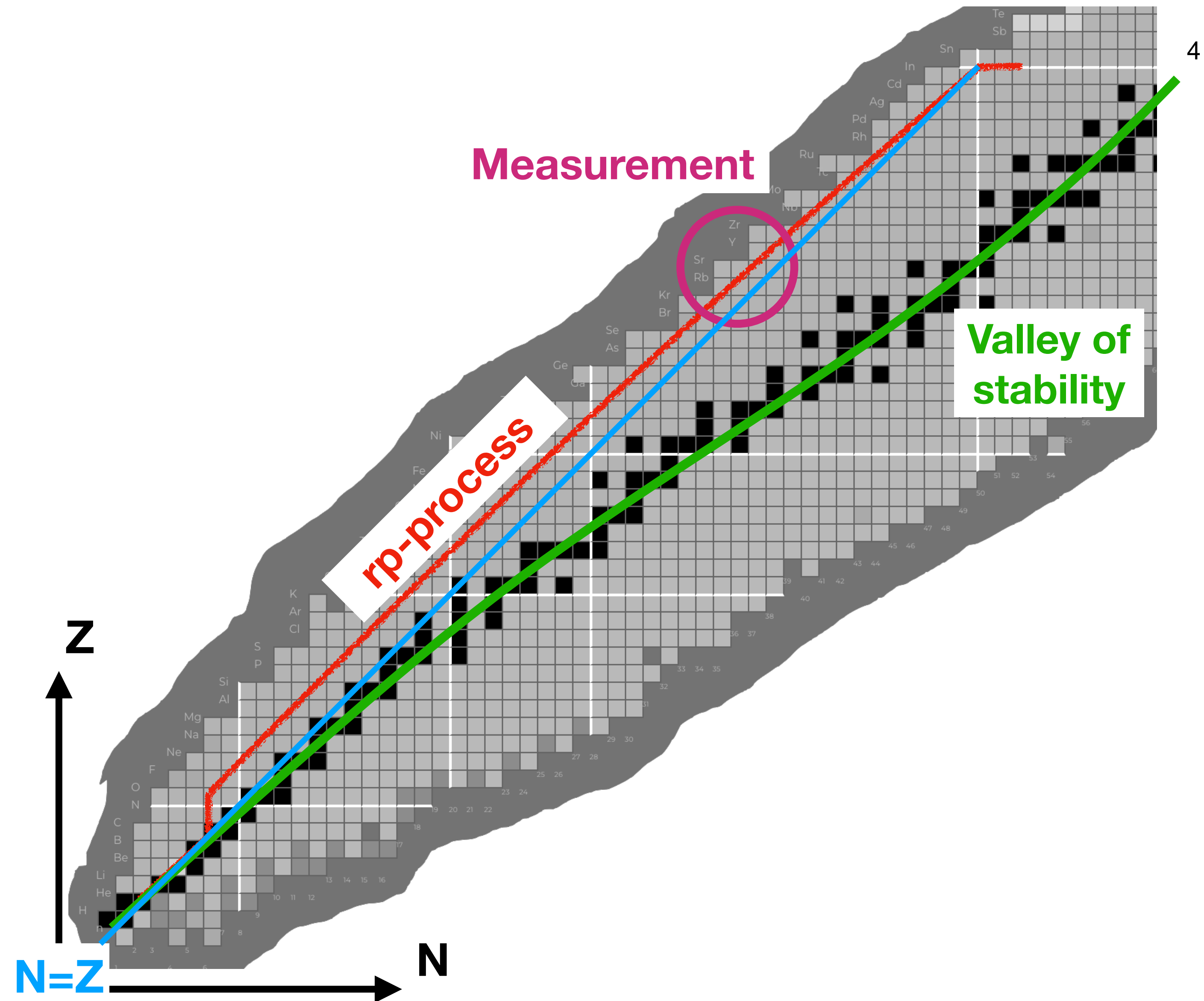
Neutron-deficient nuclei near $N=Z$

Measurement:

- $^{74,75,76}\text{Sr}$ ($Z=38$)
- Medium mass region near $N=Z$

Experimentally difficult to access:

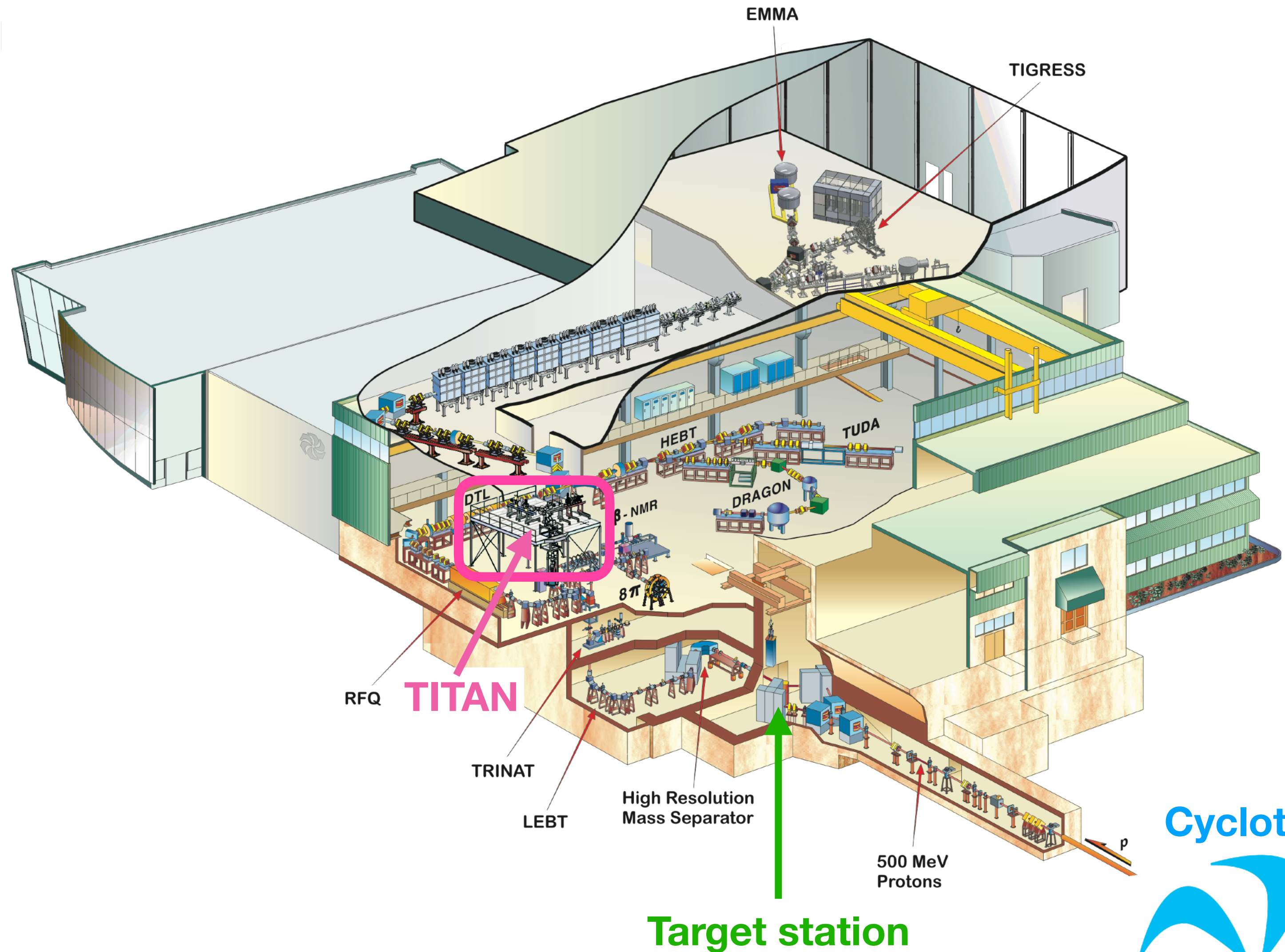
- Low cross-section for production
- Excessive in-beam isobaric contamination (alkalis and lanthanides) [1]



1. Gallant A T *et. al.* Physical Review Letters. 2014 Aug 19;113(8):082501.

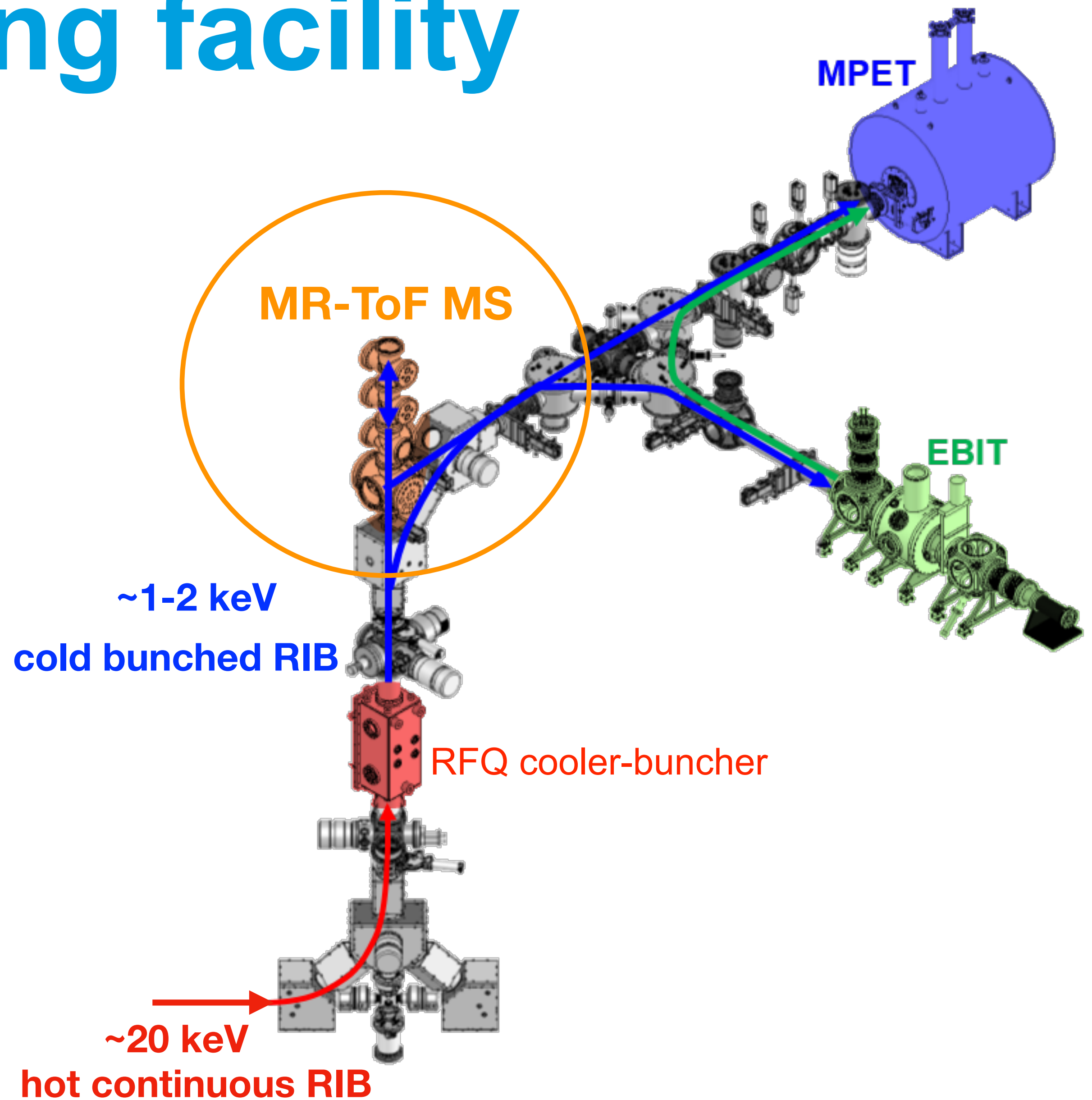
Radioisotope production at TRIUMF-ISAC

- **TRIUMF Cyclotron:** 480 MeV, 50 μ A, p^+ beam
- **ISAC target station:**
 - p^+ beam impinged onto niobium target to produce radioisotope soup
 - Selectively ionize strontium using resonant ionization laser
 - Formation into radioactive ion beam (RIB) and transport to TITAN
- measurement performed at **TITAN ion trapping facility**



TITAN ion trapping facility

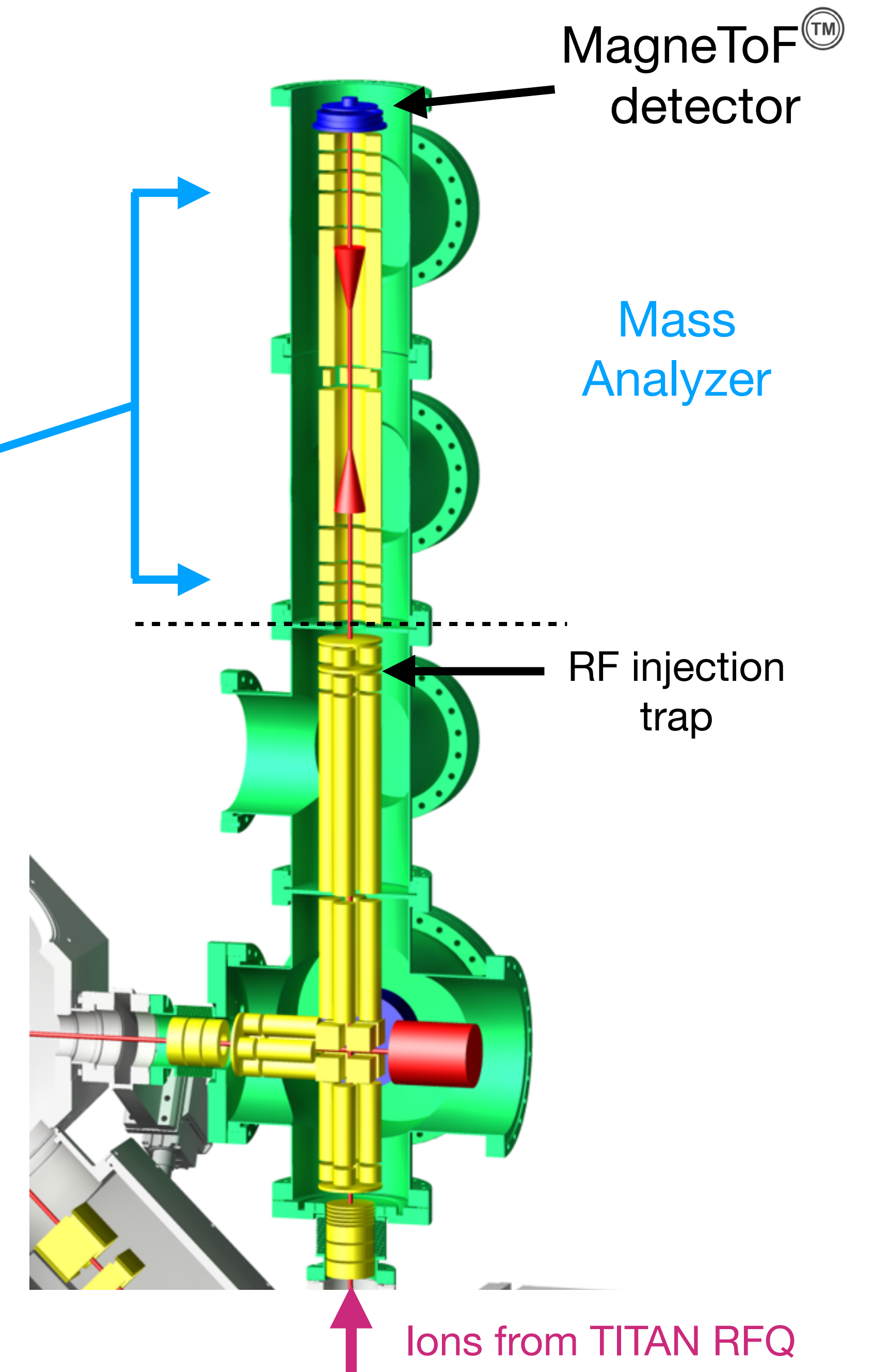
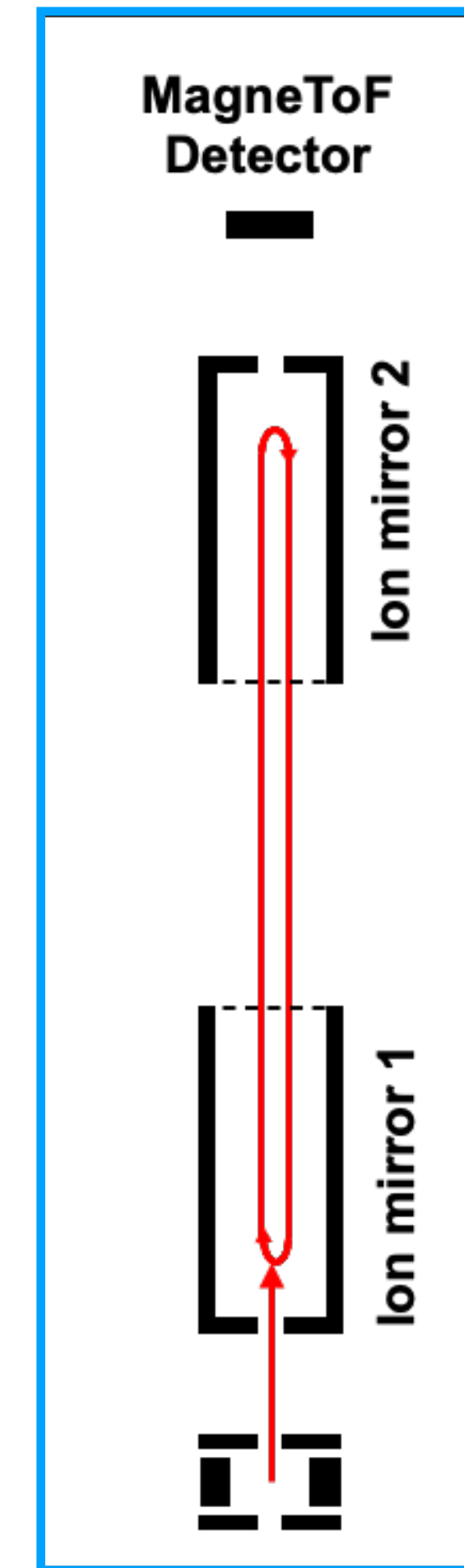
- **Beam preparation:** TITAN RFQ cooler-buncher
 - Converts the continuous RIB from ISAC into bunches
 - Buffer gas cooling of beam for injection into downstream traps
- **3 measurement traps:**
 - EBIT (charge state breeding and in-trap decay spectroscopy)
 - MPET (Penning trap mass measurements)
 - MR-ToF MS (multi-reflection time-of-flight mass spectrometer)



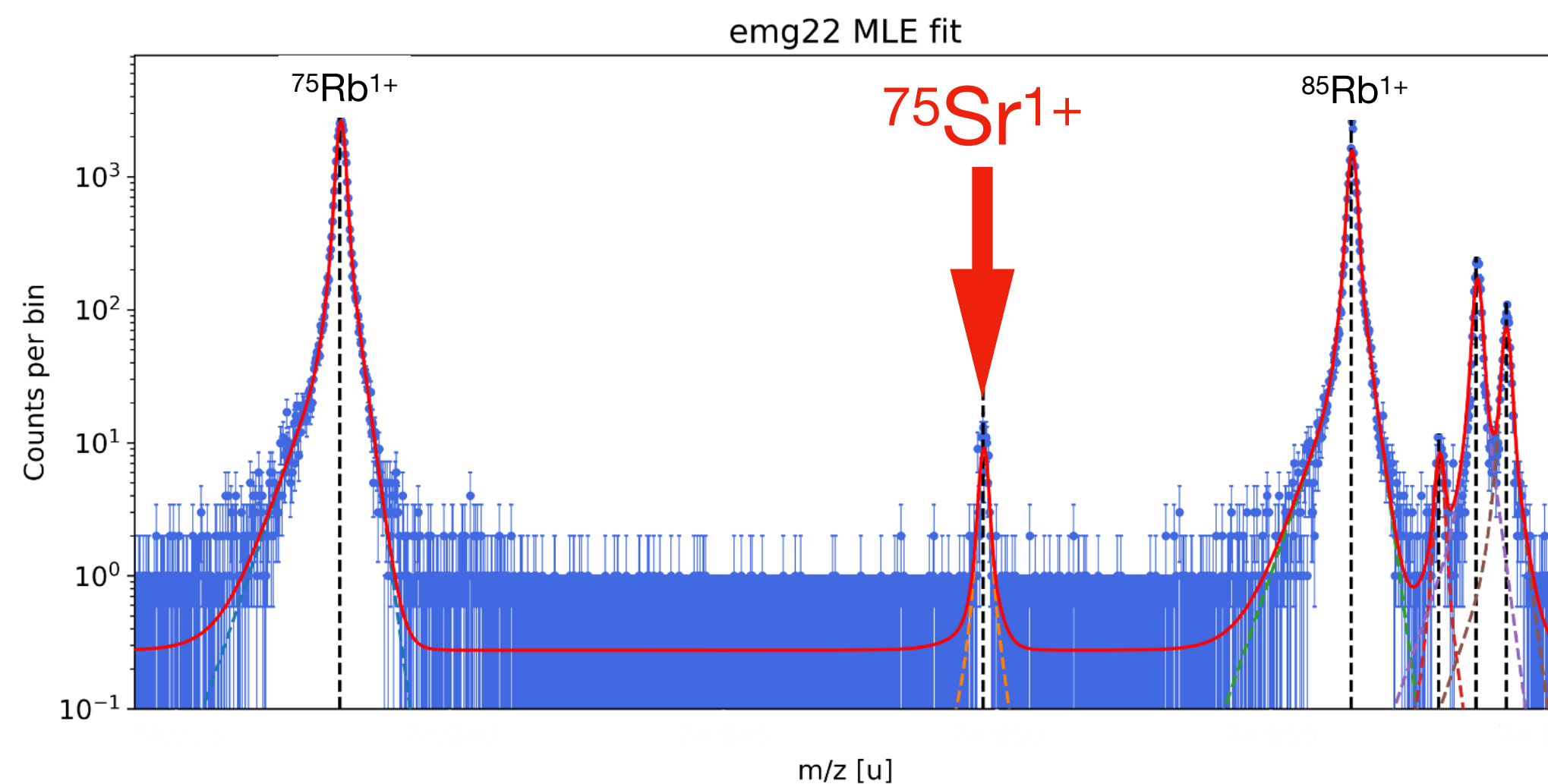
Mass Measurements with Multi-Reflection Time-of-Flight Mass Spectrometer (MR-ToF MS)

Mass measurement procedure:

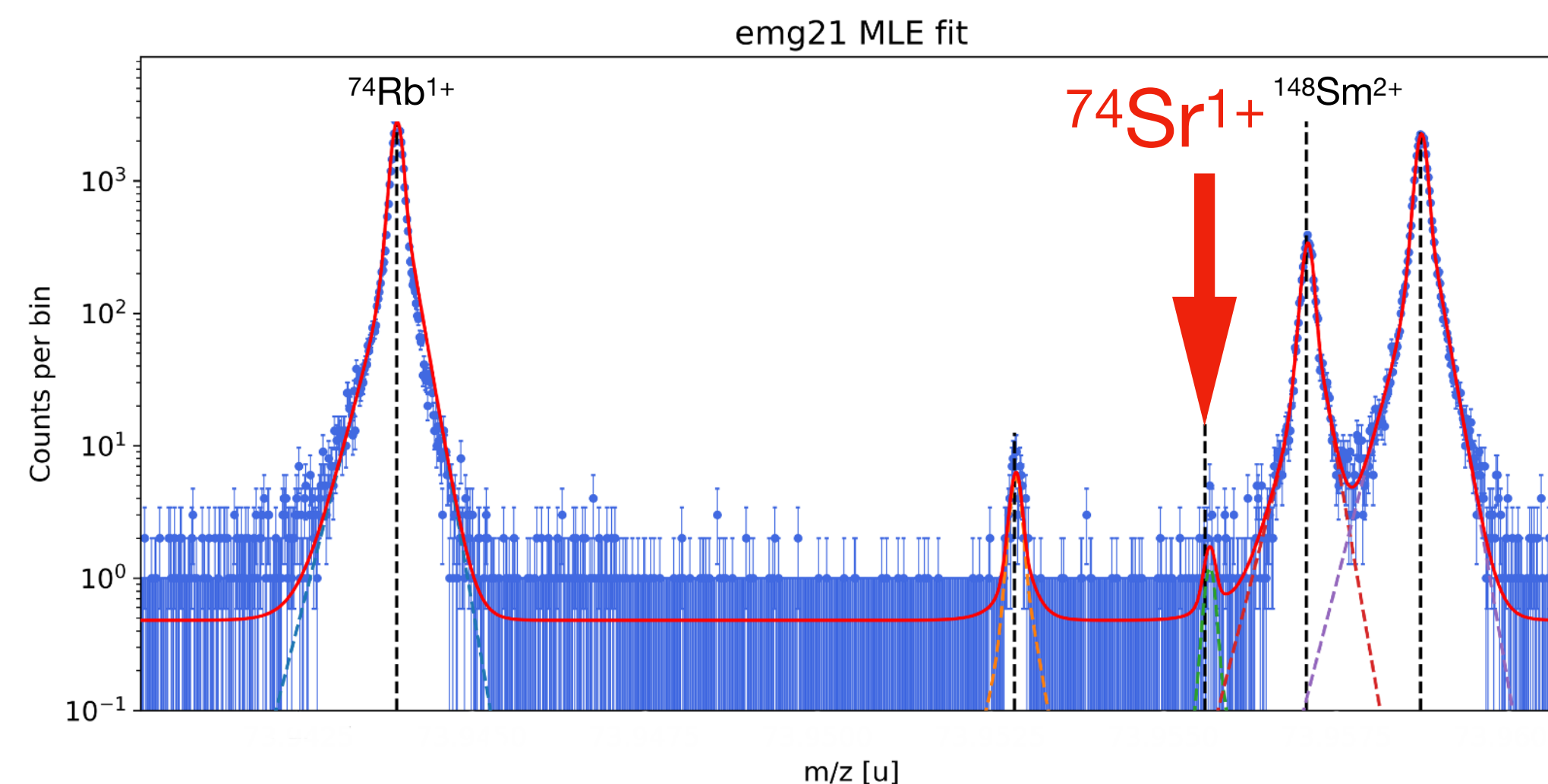
- Ions injected into Mass Analyzer
- Ions drift in a field free region cycled between electrostatic mirrors
- $t \propto \sqrt{m/q}$
- Released onto time-sensitive detector
- Build time-of-flight histogram with distinct peaks showing the beam composition



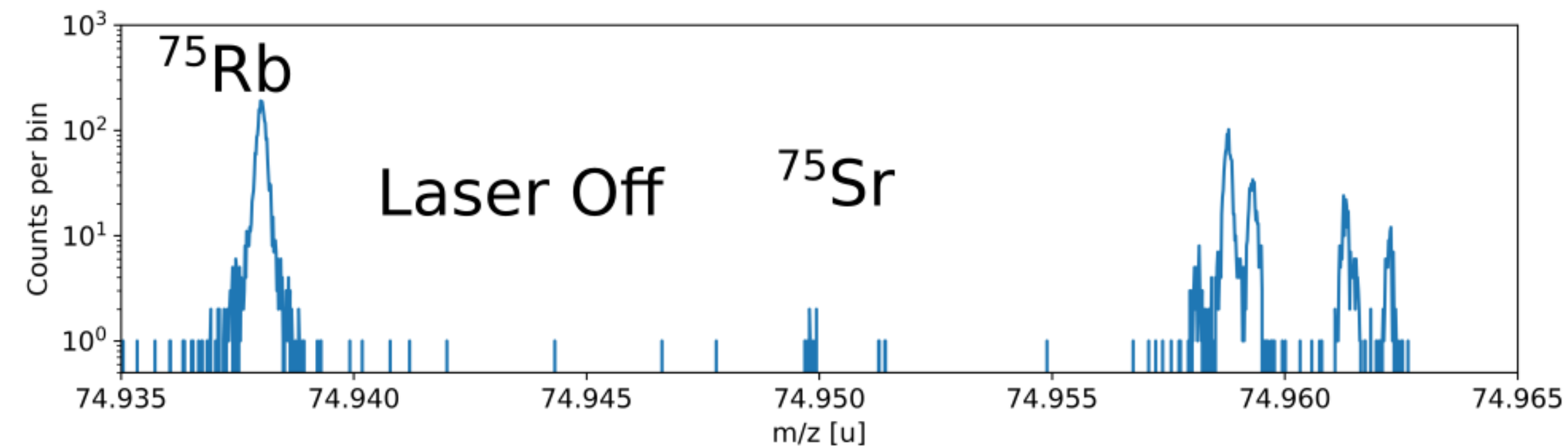
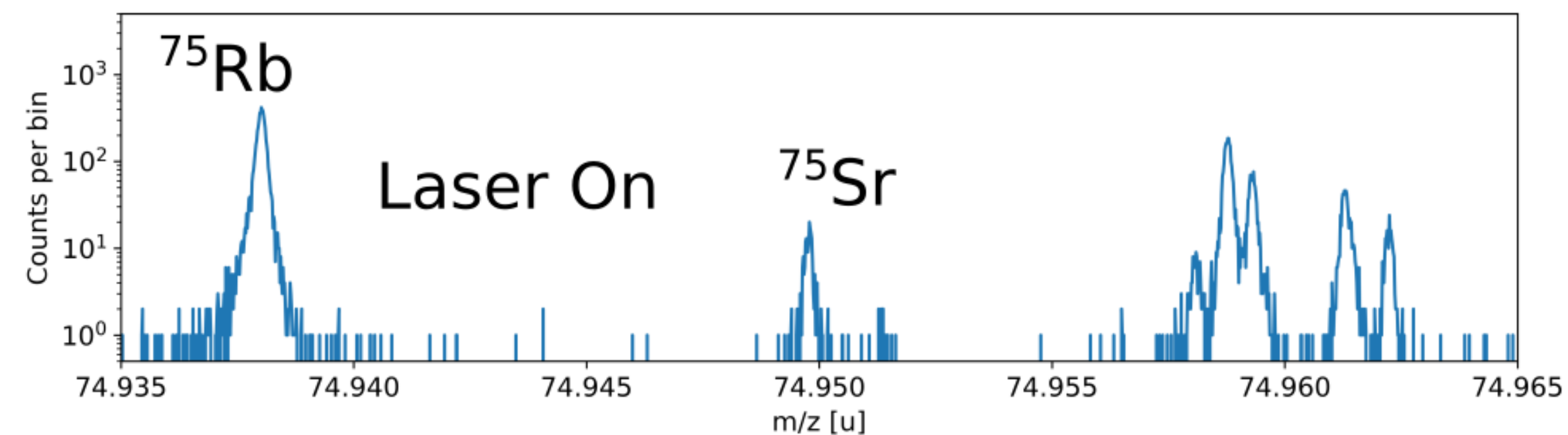
Strontium mass measurements



⁷⁵Sr Yield: 500 counts/hour



⁷⁴Sr Yield: 4.8 counts/hour



Verification of presence of strontium isotope using resonant ionization lasers

Precision mass measurements of neutron-deficient strontium

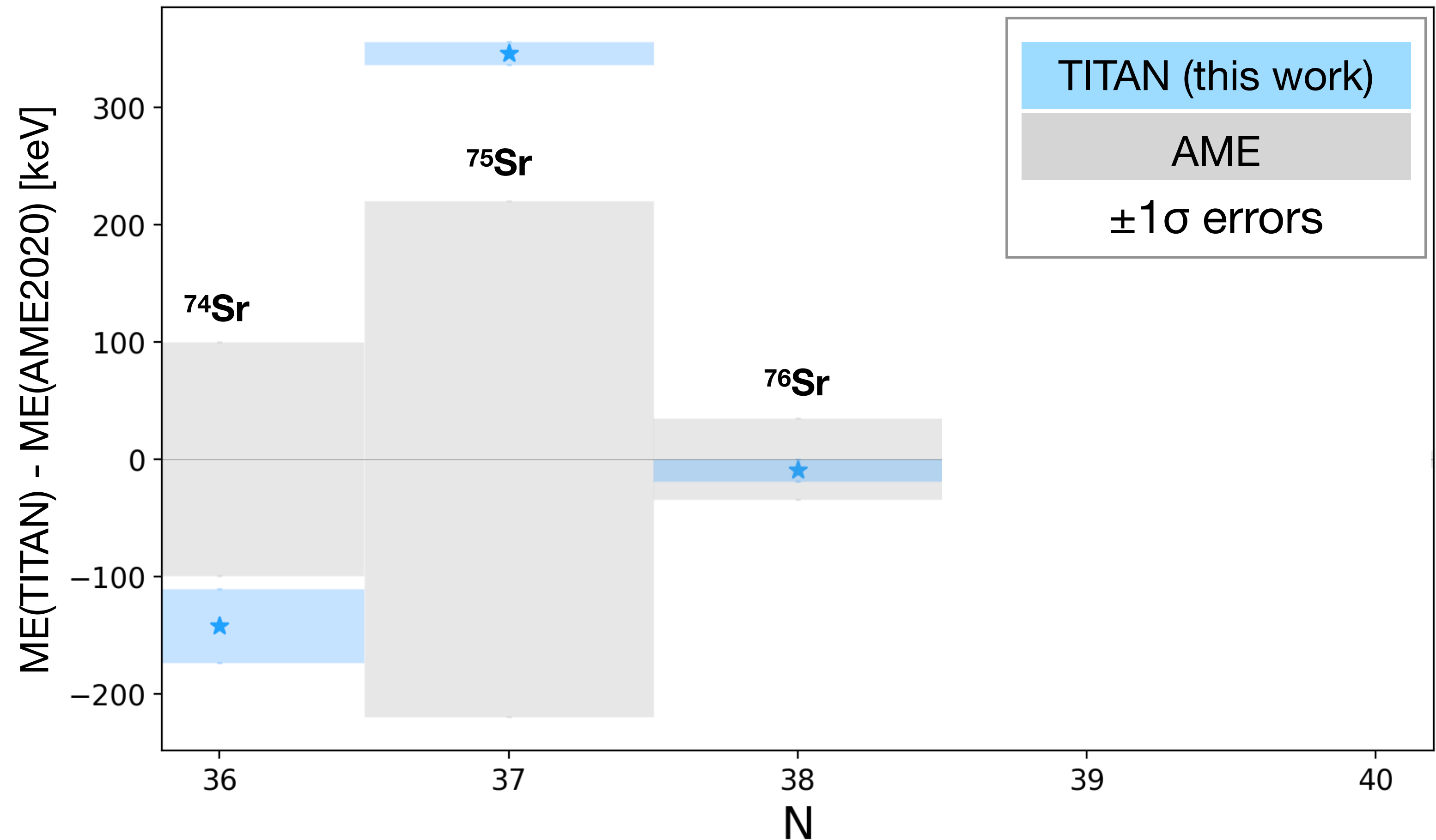
Comparison of TITAN values with 2020 Atomic Mass Evaluation (AME):

- Mass Excess:

- $ME(Z, N) \equiv M(Z, N) - (Z + N)m_u$

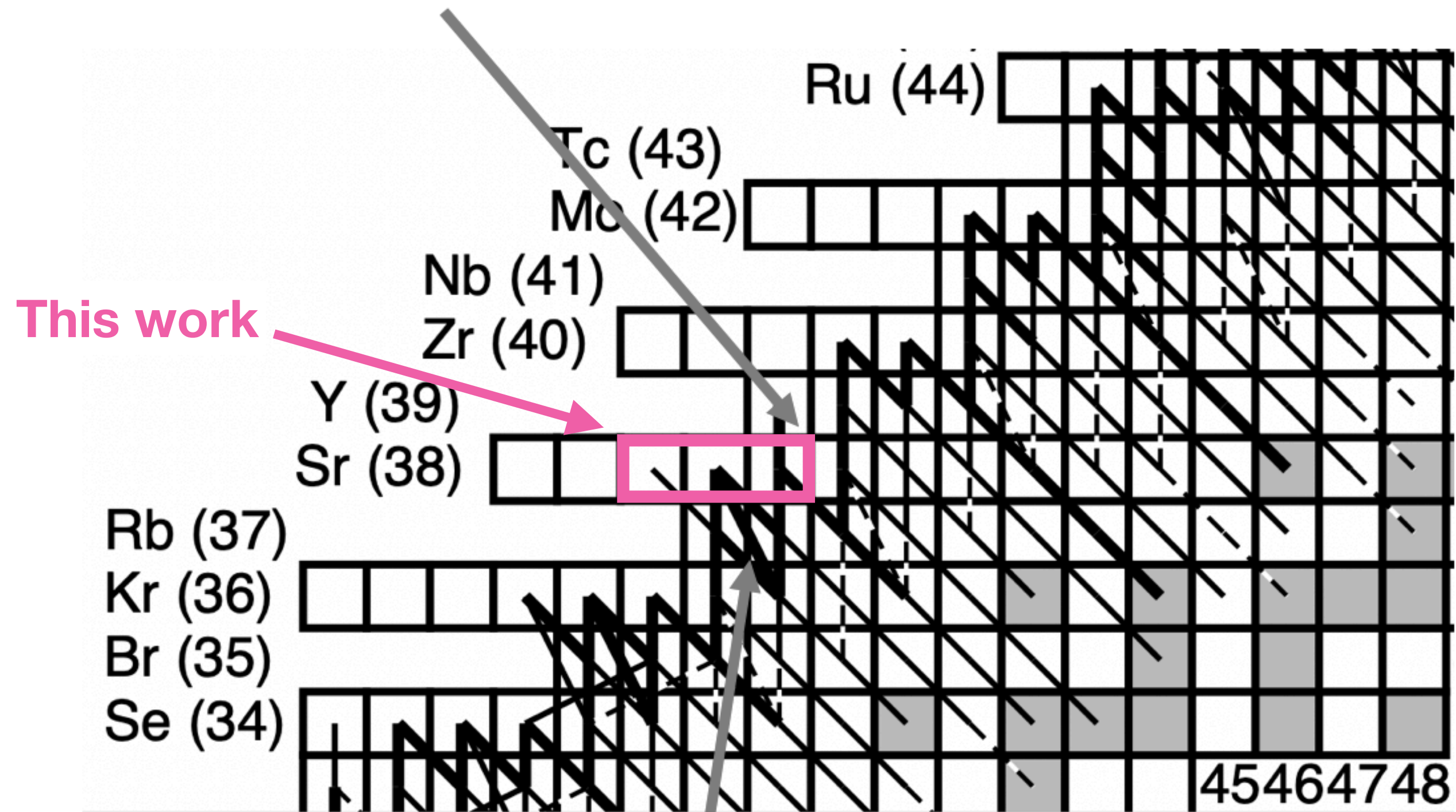
AME2020 values:

- ^{74}Sr : AME extrapolation [1]
- ^{75}Sr : indirect via β -decay [2]
- ^{76}Sr : direct ToF method [3]



1. Wang M *et. al.* Chinese Physics C. 2021 Mar 1;45(3):030003.
2. Huikari J *et. al.* The European Physical Journal A-Hadrons and Nuclei. 2003 Mar;16:359-63.
3. Lalleman AS *et. al.* Hyperfine Interactions. 2001 Jan;132:313-20.

Effects of $^{74-76}\text{Sr}$ on rp-process



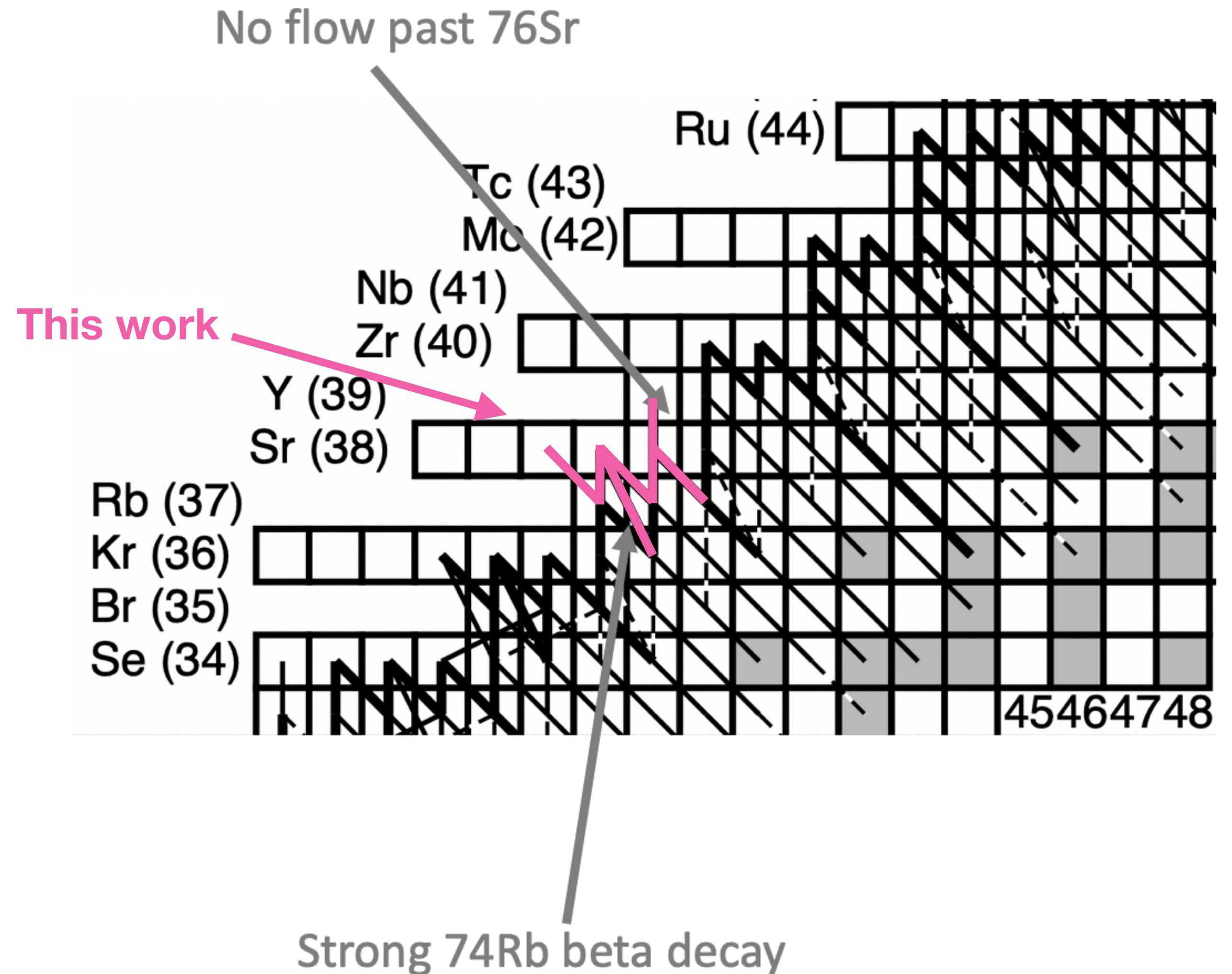
Reaction network plot

Effects of $^{74-76}\text{Sr}$ on rp-process

Network calculations for Type-I XRBs [1] performed by Hendrik Schatz:

Effect of $^{74-76}\text{Sr}$:

- Mass flow which passes ^{76}Sr reduced by a factor of ~ 3
- Uncertainty of $A=74$ ash production reduced from 16% to 0.7%
- Astrophysical observational differences are small, but as more masses are measured combined effect may be impactful



Isospin symmetry and nuclear structure

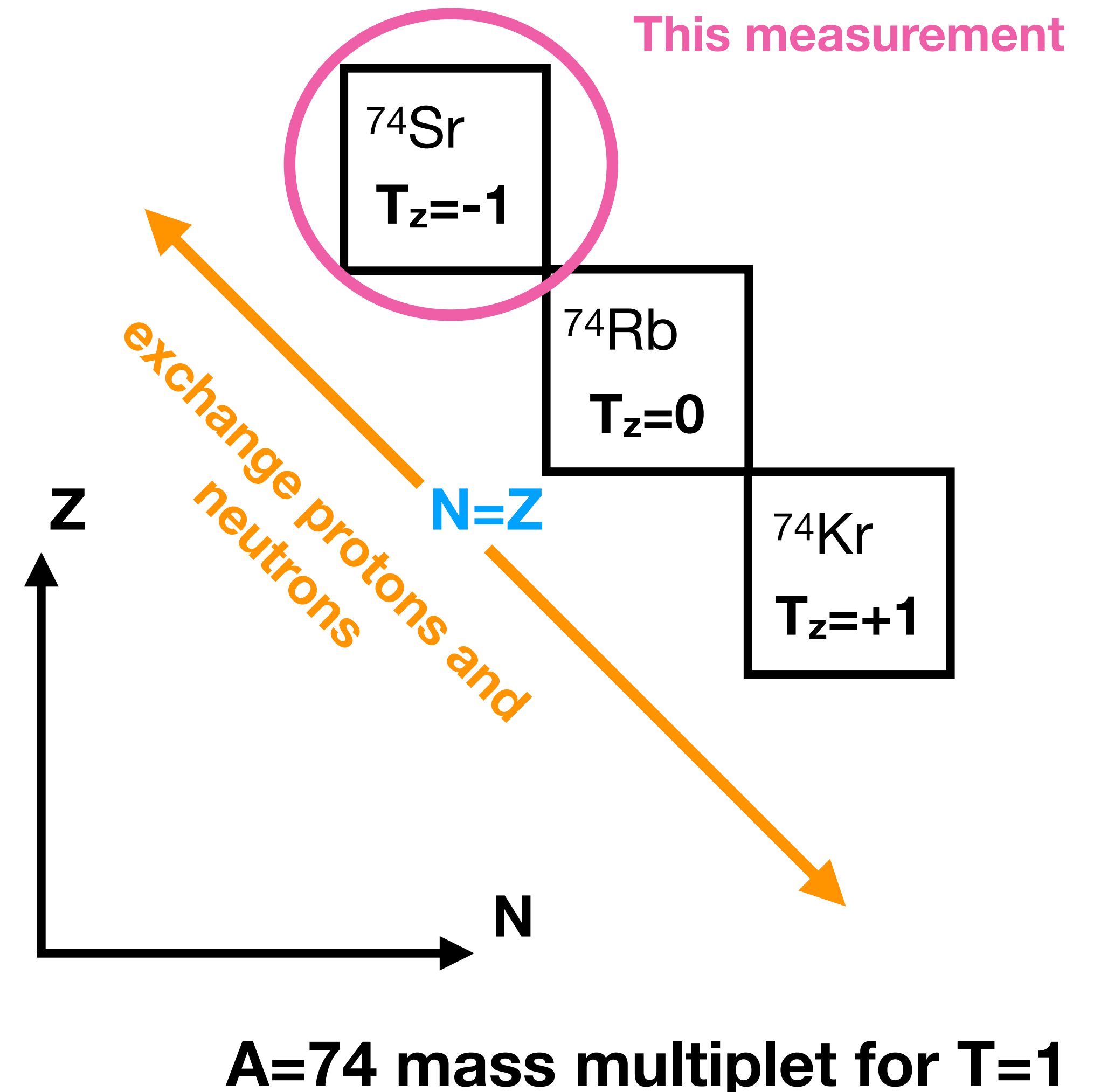
Isospin is **useful for examining proton-neutron asymmetry** in nuclei and how it impacts nuclear properties (e.g. mass, level structure)

- Concept of isobaric mass multiplet
- **Mass Excess for members in an isobaric mass multiplet, assuming a quadratic form in T_z :**

Isobaric Multiplet Mass Equation (**IMME**):

$$ME(\alpha, T, T_z) = a + bT_z + cT_z^2$$

- Works very well for $A \approx 10-60$ [1]
- **Above $A=60$, IMME needs further testing [1]** (assumptions of IMME can break)



IMME c coefficient $\left(ME(\alpha, T, T_z) = a + bT_z + cT_z^2 \right)$

MacCormick M, Audi G. 2014 [1]:

- Extensive survey of IMME from A=10-60
- **Global parameterization of empirical c coefficient data** using simple assumption: nucleus is a homogeneously charged sphere

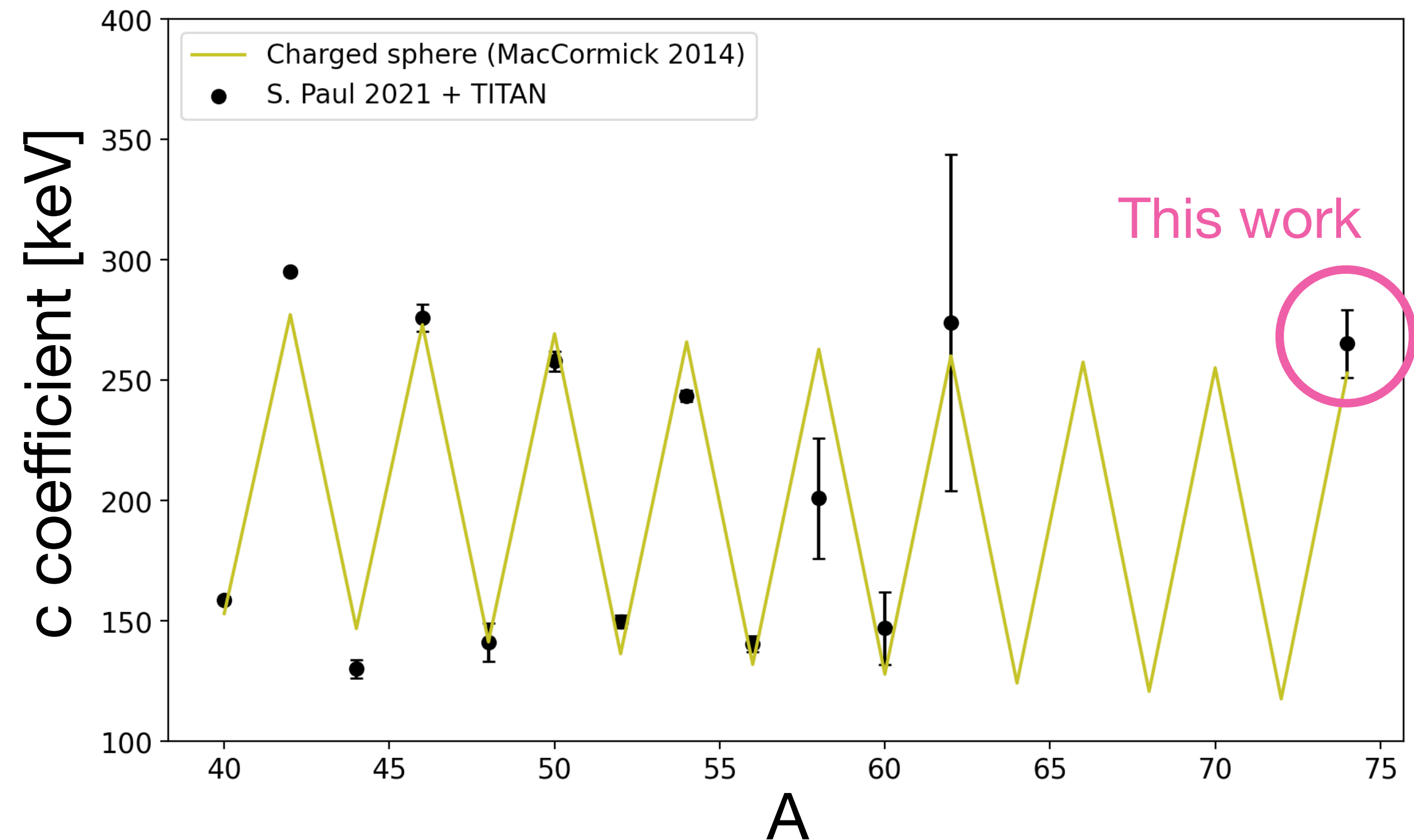
1. MacCormick M, Audi G. Nuclear Physics A. 2014 May 1;925:61-95.
2. Towner IS, Hardy JC. Physical Review C. 2008 Feb 7;77(2):025501.
3. Lam YH, Smirnova NA, Caurier E. Physical Review C. 2013 May 6;87(5):054304.
4. Martin MS *et. al.* Physical Review C. 2021 Jul 30;104(1):014324.

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We provide a new data point for IMME c coefficient (A=74 triplet (^{74}Sr , ^{74}Rb , ^{74}Rb))



1. MacCormick M, Audi G. Nuclear Physics A. 2014 May 1;925:61-95.
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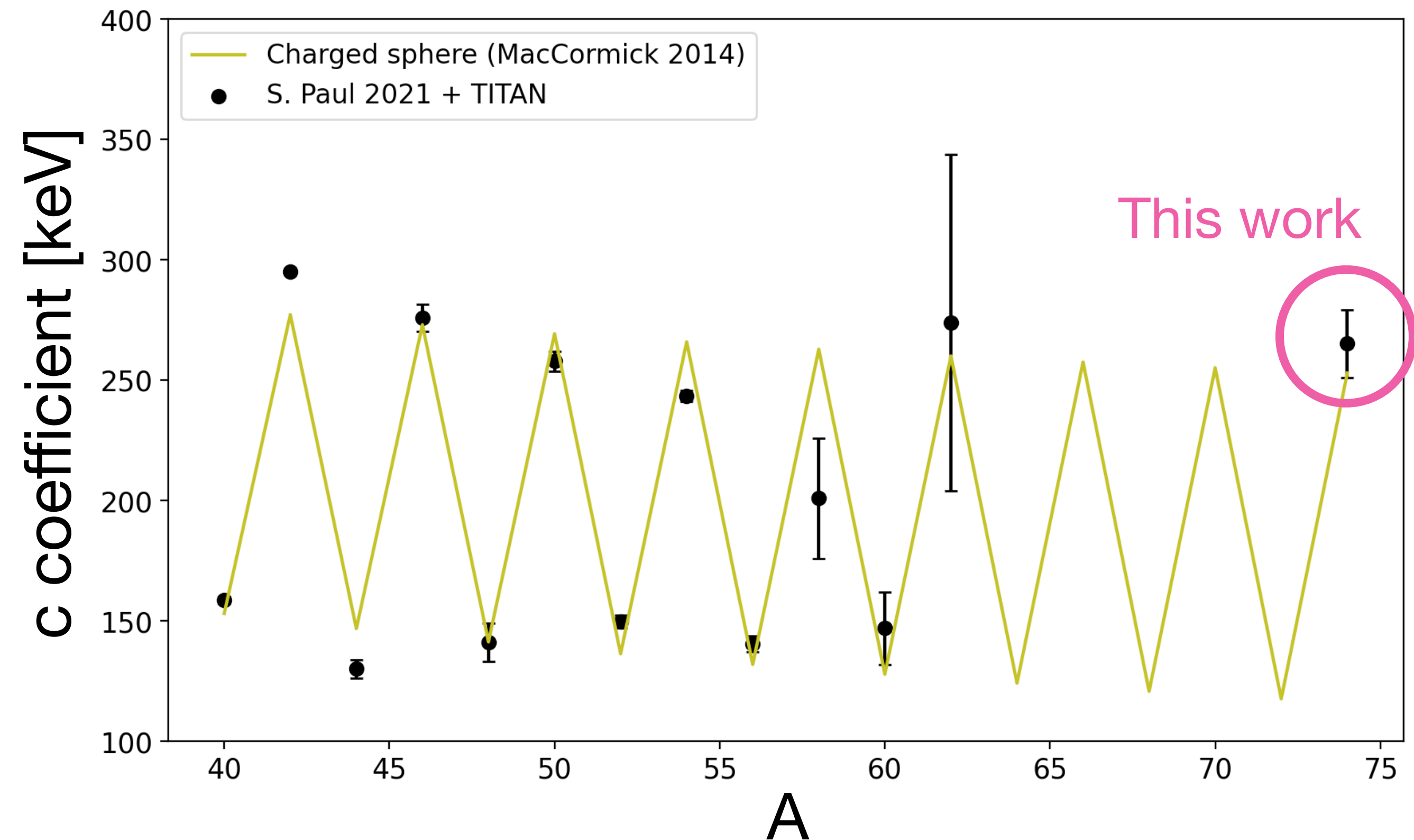
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IMME coefficients can be compared with theoretical models:

- Study contributions of isospin non-conserving forces [3]
- Tests of CKM Unitarity [2]
- Ab-initio calculations using VS-IMSRG [4] in-progress (Jason Holt and Baishan Hu)

1. MacCormick M, Audi G. Nuclear Physics A. 2014 May 1;925:61-95.
2. Towner IS, Hardy JC. Physical Review C. 2008 Feb 7;77(2):025501.
3. Lam YH, Smirnova NA, Caurier E. Physical Review C. 2013 May 6;87(5):054304.
4. Martin MS *et. al.* Physical Review C. 2021 Jul 30;104(1):014324.



Summary

- New precision mass data of $^{74-76}\text{Sr}$ using TITAN's MR-ToF MS
- Simulations of rp-process constrain burst ashes and mass flow
- New IMME data for studying isospin symmetry breaking and nuclear structure

TITAN Collaboration

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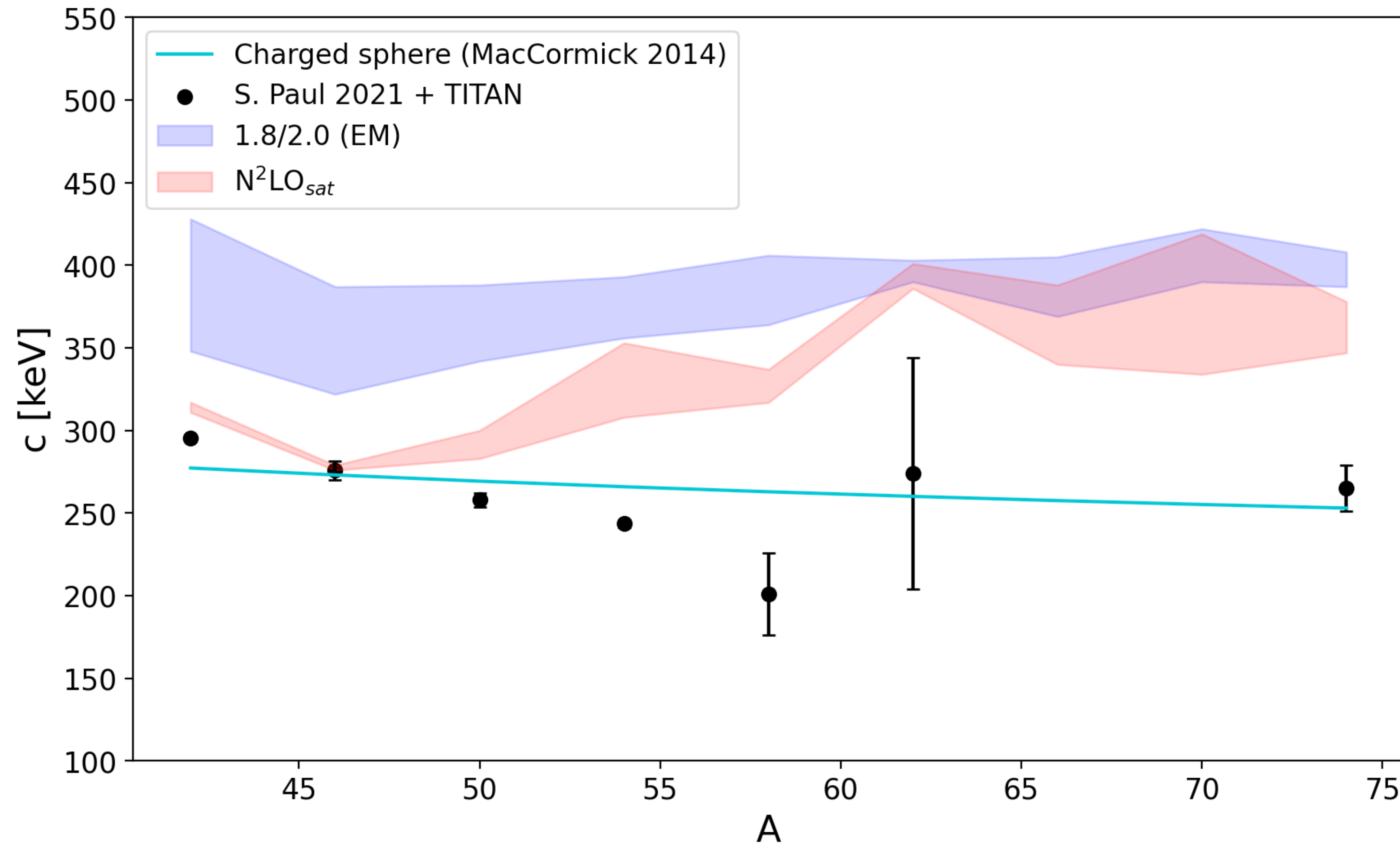


Thank you
Merci



Extra

Ab-initio comparisons to c coefficient

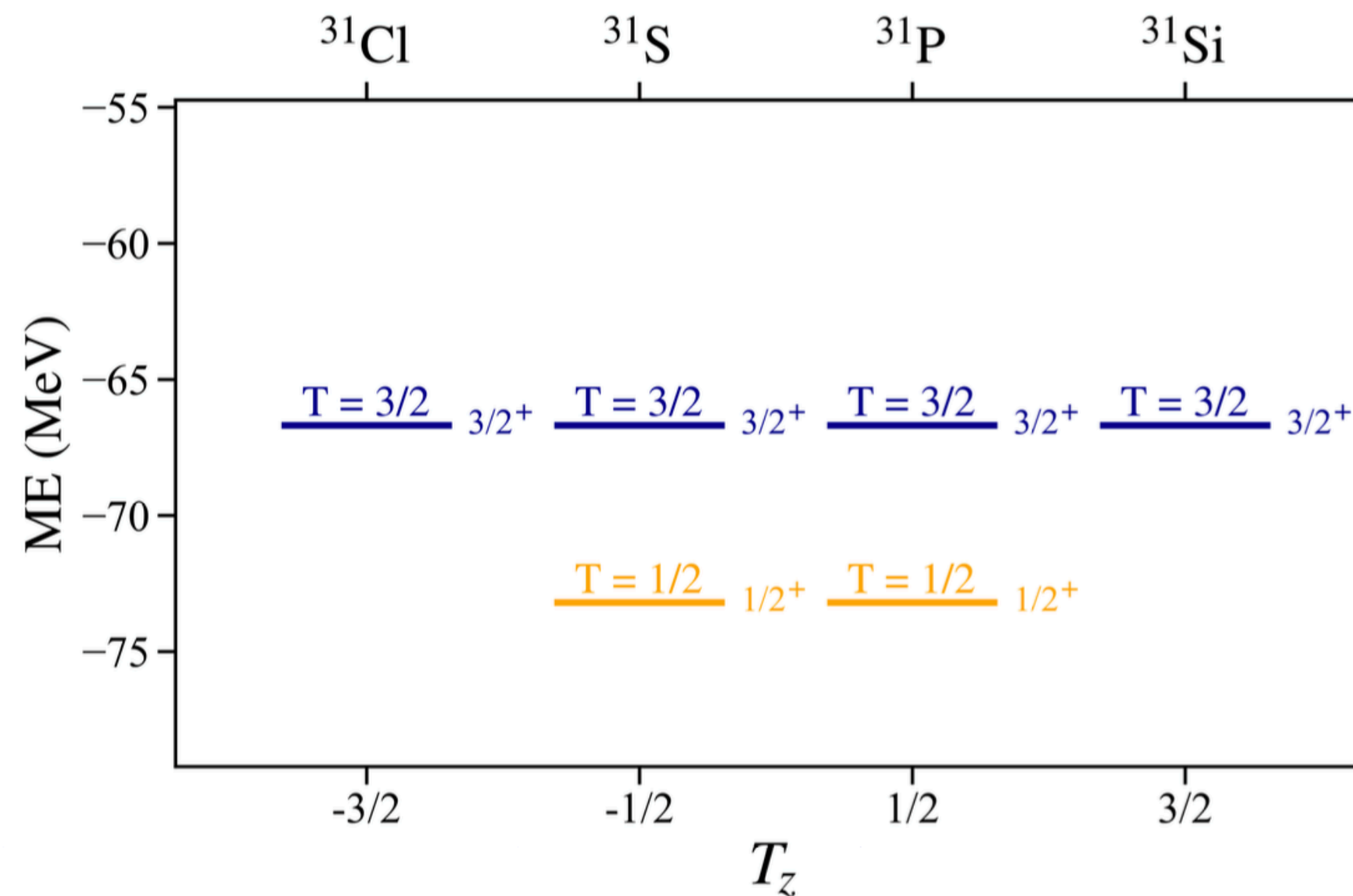


Isobaric mass multiplets

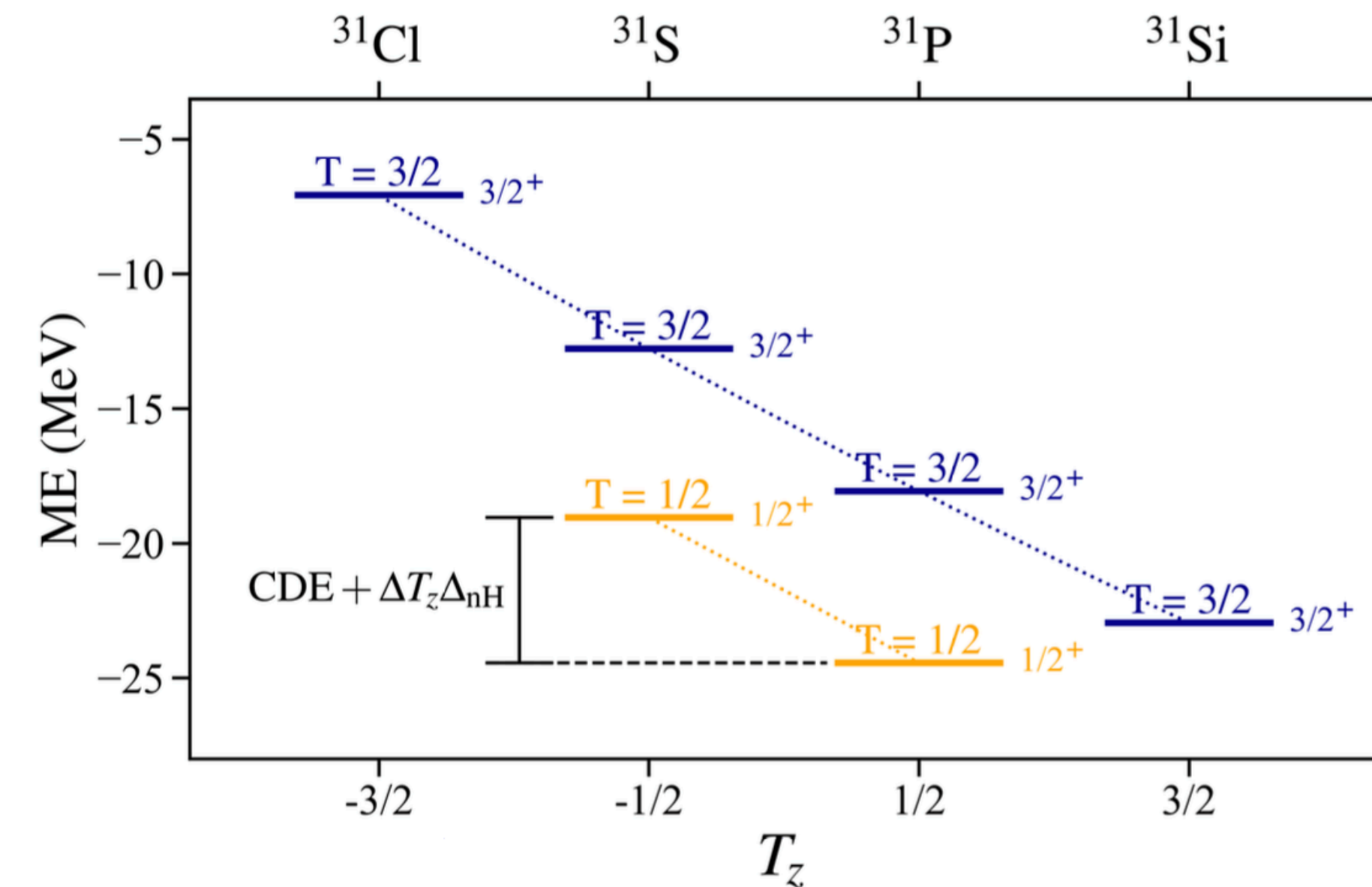
Different masses and electric charges of u and d quarks give **three distinct isospin-breaking effects**:

1. Coulomb interaction between protons
2. Mass difference between protons and neutrons
3. Charge-dependence of strong interaction

Hypothetical: neglecting isospin-symmetry breaking effects



Realistic situation:



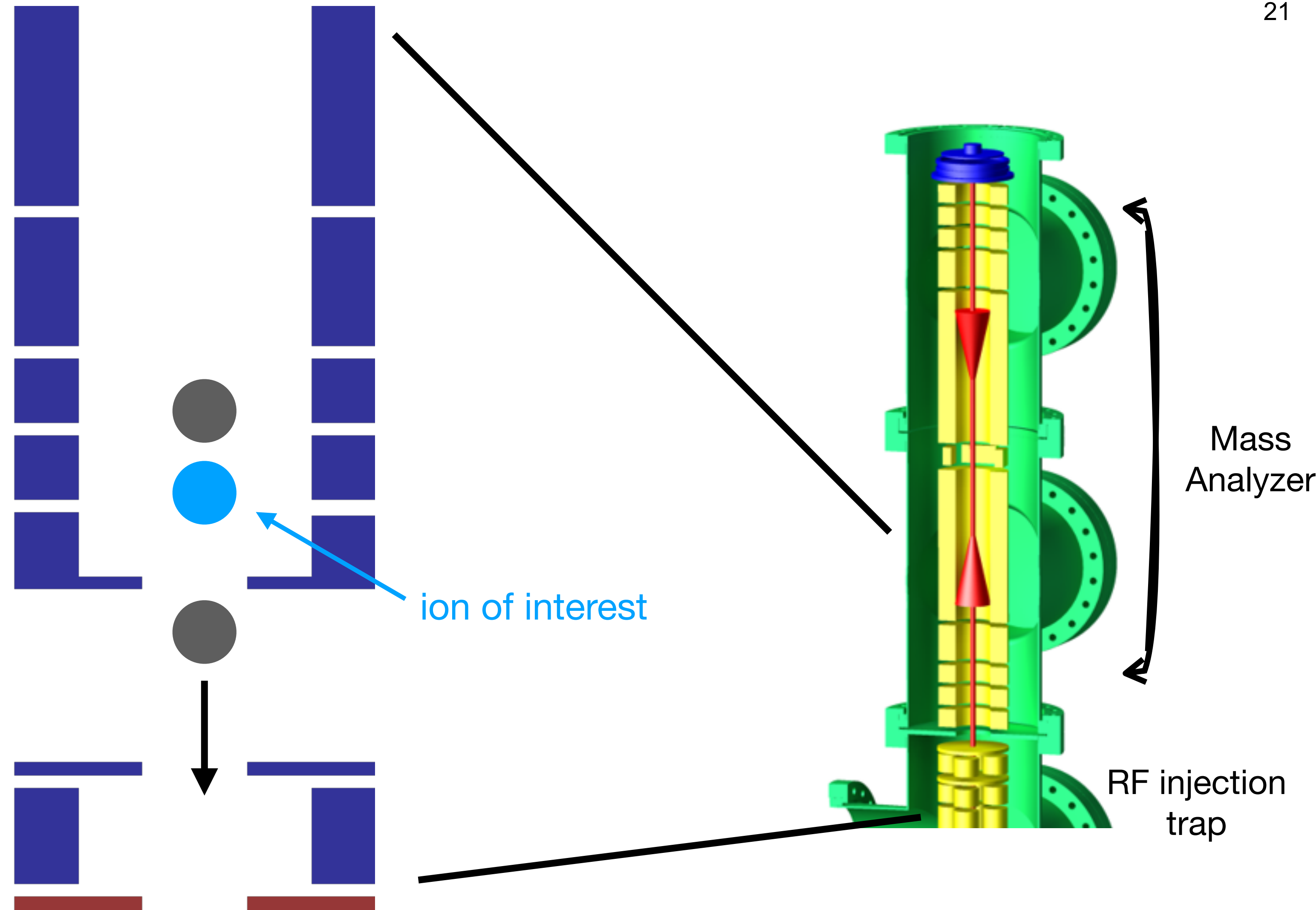
IMME above $A=60$

- Beyond $A=60$, large collection of protons in nucleus invalidates charge-independence
- In general, when IMME disagrees with experimental values, there is greater evidence for isospin-mixing (fragmentation of isospin over 2 or more nuclear states)
 - Because IMME supposes a nuclear state is assigned a single isospin (no fragmentation)
- Fragmentation readily observed over $A=40$ due to increasing density of levels (also observed at very low mass)

Mass-Selective Re-Trapping

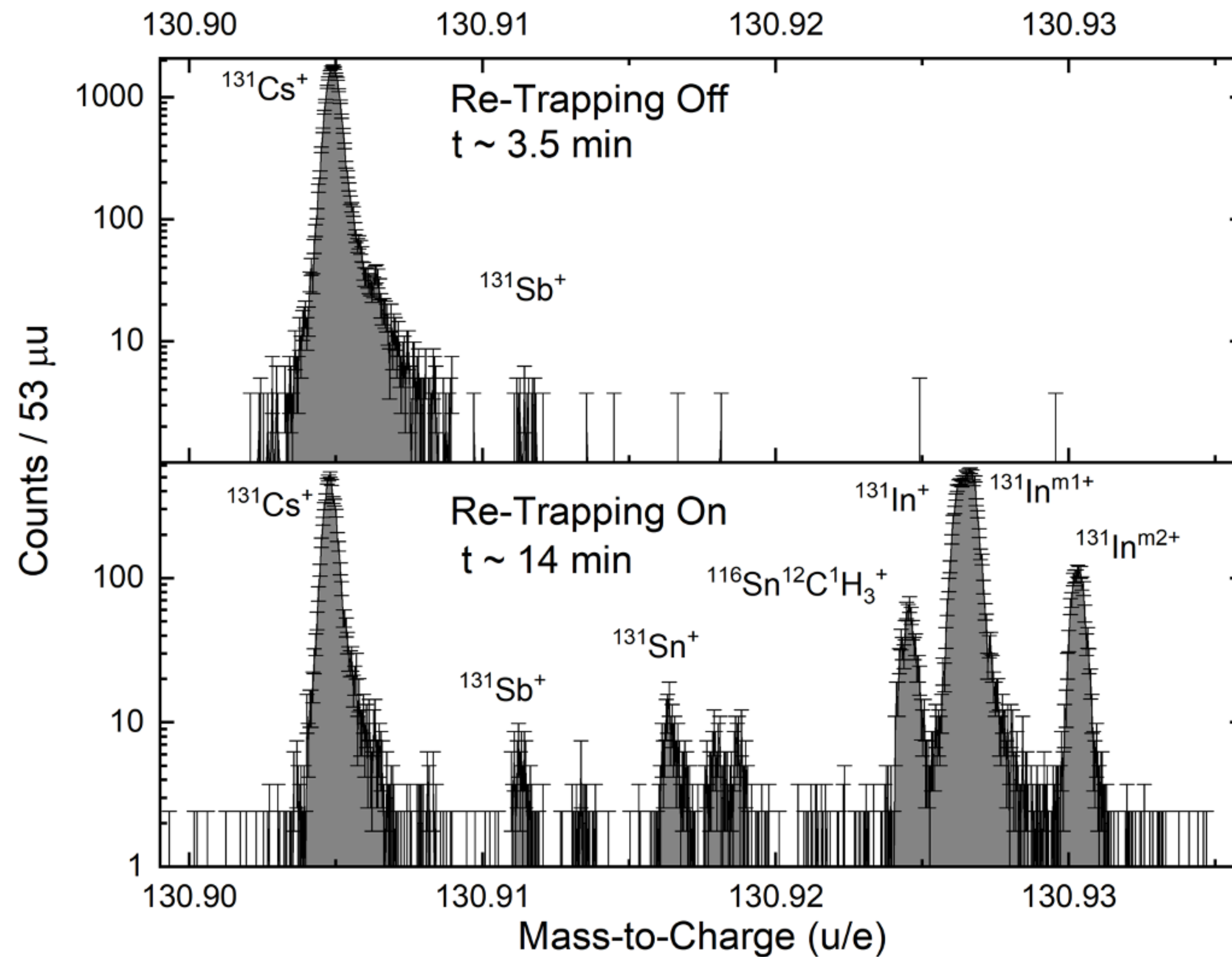
Re-trapping Mode:

1. Ions injected for an initial mass separating cycle (self-cleaning)
2. Extraction back into RF injection trap
3. Selective trapping of ion of interest
4. Injection back into Mass Analyzer for ToF analysis



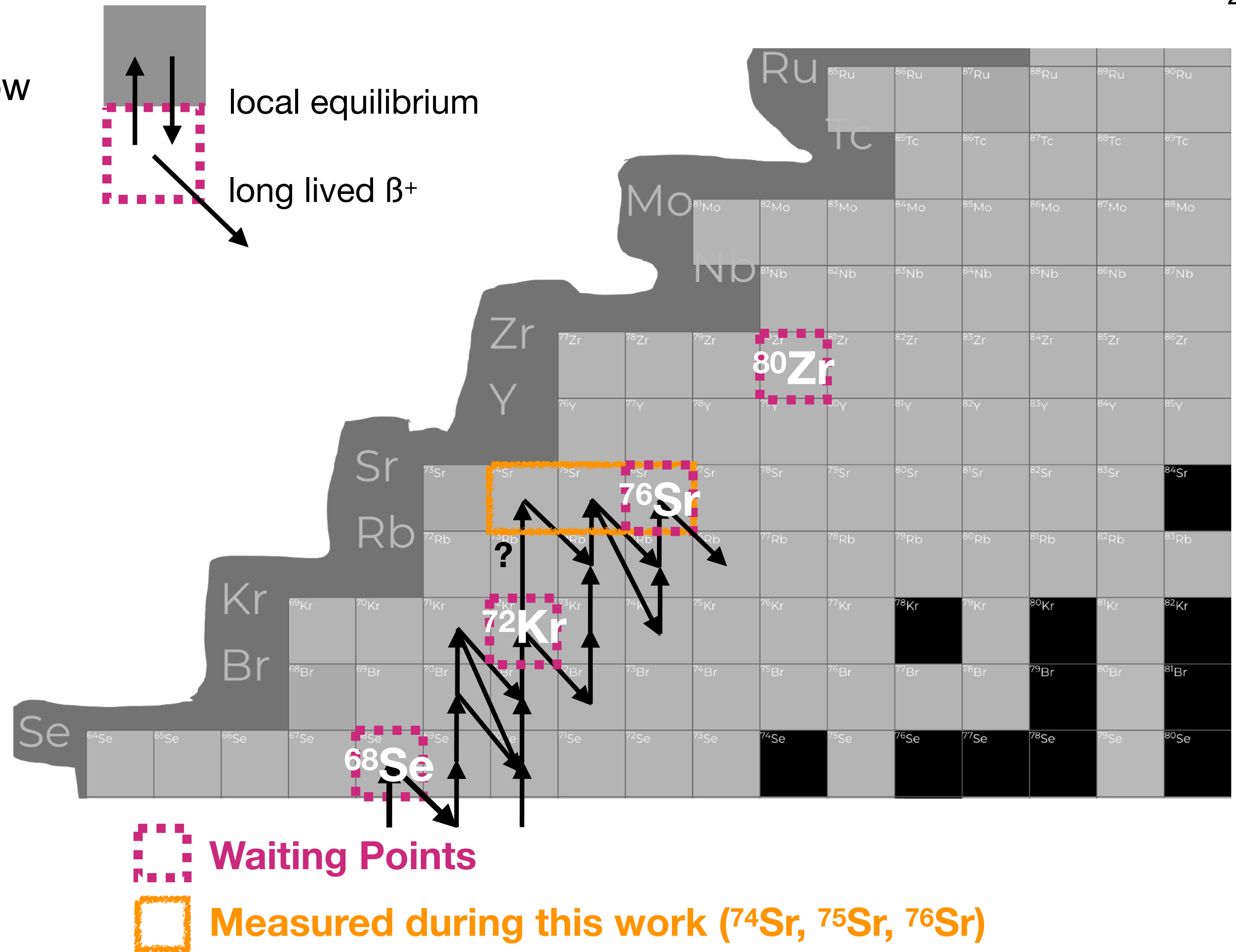
Mass-Selective Re-Trapping

4 orders of magnitude
background suppression:



Sequential 2p capture on Waiting Points

- **Waiting point nuclei:** a local equilibrium between neighboring isotones force mass flow onto a long lived β^+ branch.
 - Strong consequence on reaction flow and burst ashes
- **^{72}Kr ($t_{1/2}=17\text{s}$):** a well-known waiting point nucleus
 - $^{72}\text{Kr}(p,\gamma)^{73}\text{Rb}$ unfavored because ^{73}Rb is proton-unbound
 - Bypass with sequential 2p capture ($^{72}\text{Kr}(p,\gamma)^{73}\text{Rb}(p,\gamma)^{74}\text{Sr}$)?
 - Sequential 2p capture on ^{38}Ca is an important waiting point bypass¹



1. Görres J, Wiescher M, Thielemann FK. Bridging the waiting points: The role of two-proton capture reactions in the rp process. Physical Review C. 1995 Jan 1;51(1):392.