

Chiral Symmetry in Nuclear Medium Observed in Pionic Atoms

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Chiral symmet	ry restoration at high matter
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- T.Nishi, KI et al., N. Phys. 19, 788 (2023)
 Article DOI: 10.1038/s41567-023-02001-x
- Nature Physics (2023/3/23)
 News and Views "Modified in Medium"



Chiral Symmetry in Nuclear Medium Observed in Pionic Atoms

- Dominant symmetry of the vacuum in low-energy region.
- Spontaneous breakdown due to non-perturbative strong interaction.
- Non-trivial structure of the QCD vacuum.

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on at high matter

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Chiral condensate qq on Tp plane



Lattice QCD calculated T dependence of <qq>



Remark: sign problem makes it difficult for lattice to approach non-zero ρ region

PRD105(2022)034504

ρ dependence of < $\bar{q}q$ > known so far



ρ dependence of < $\bar{q}q$ > known so far



Need high-quality experimental information to quantify <qq> reduction and confirm theoretical scenario of vacuum evolution



Precision Spectroscopy of Pionic Atoms



$$U_{s}(\mathbf{r}) = b_{0} \rho + \mathbf{b}_{1} (\rho_{n} - \rho_{p}) + B_{0} \rho^{2}$$
$$U_{p}(r) = \frac{2\pi}{\mu} \vec{\nabla} \cdot [c(r) + \varepsilon_{2}^{-1} C_{0} \rho^{2}(r)] L(r) \vec{\nabla}$$



Pion-nucleus strong interaction

Overlap between pion w.f. and nucleus → π works as a probe at ρ_e~0.58ρ_s

π-nucleus **interaction is changed** in nuclear medium for wavefunction renormalization effect

> Ericson-Ericson potential $U_{opt}(r) = U_{s}(r) + U_{p}(r),$ $U_{s}(r) = b_{0} \rho + b_{1} (\rho_{n} - \rho_{p}) + B_{0} \rho^{2}$ $U_{p}(r) = \frac{2\pi}{\mu} \vec{\nabla} \cdot [c(r) + \varepsilon_{2}^{-1} C_{0} \rho^{2}(r)] L(r) \vec{\nabla}$



Strong interaction and chiral condensate

Overlap between pion w.f. and nucleus → π works as a probe at ρ_e~0.58ρ_s

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Ericson-Ericson potential

 $U_{\text{opt}}(r) = U_s(r) + U_p(r),$ $U_s(r) = b_0 \rho + b_1 (\rho_n - \rho_p) + B_0 \rho^2$ $U_p(r) = \frac{2\pi}{\mu} \vec{\nabla} \cdot [c(r) + \varepsilon_2^{-1} C_0 \rho^2(r)] L(r) \vec{\nabla}$

In-medium Glashow-Weinberg relation



γ=0.184±0.003

Jido, Hatsuda, Kunihiro, PLB670, 109 (2008)

Strong interaction and chiral condensate







Deduction of pion-nucleus interaction in medium by

Spectroscopy of pionic atoms in (*d*,³He) reactions

Based on energy-momentum conservation law:

Excitation energy $\sim T_d - T_{3He}$



(d,³He) Reaction Spectroscopy in RIBF



Kenta Itahashi, RIKEN





First pionic atom in RIBF (2010) **Pionic** ¹²¹**Sn atom**

> First observation of θ dependence of π atom cross section



T. Nishi KI et al., PRL120, 152505 (2018)



16

1s and 2p pionic atom cross sections in (d,³He)



T. Nishi KI et al., PRL120, 152505 (2018)

Theory calculates 5x larger cross section for 1s



High Precision Spectrum of ¹²²Sn(*d*,³He) in RIBF-54



Resolution 287 keV (FWHM)







Before determining b₁, we included updated theories, nuclear parameters etc.

	ImB _o [n	0.0475	
	[keV]	Statistical	Systematic
$B_{\pi}(1s)$	3831	± 3	+78 - 76
$B_{\pi}(2p)$	2276	± 3	+84 - 83
$B_{\pi}(1s) - B_{\pi}(2p)$	1555	± 4	± 12
$\Gamma_{\pi}(1s)$	316	± 12	+36 - 39
$\Gamma_{\pi}(2p)$	164	± 17	+41 - 32
$\Gamma_{\pi}(1s) - \Gamma_{\pi}(2p)$	152	± 20	+28 - 36



 $b_{1} [m_{\pi}^{-1}]$

Nuclear p distribution measured in Sn(p,p')

at RCNP, Osaka



23

Neutron spectroscopic factors in Sn isotopes



T. Nishi KI et al., PRL120, 152505 (2018)

Neutron spectroscopic factors in Sn isotopes

(*d*,³He) requires *n*-spectroscopic factor information

Spectroscopic factors are measured in (*p,d*), (*d,p*)... nuclear reactions confirming vacancy+occupancy~2*j*+1

S.V. Szwec et al., PRC104, 054308 (2021)





Deduced b₁ with updated parameters



Result: deduced chiral condensate



Result: deduced chiral condensate



Result: deduced chiral condensate



Support existence of non-trivial structure in the vacuum



To step further forward, RIBF-135 for systematic study of Sn isotopes

NPI512-RIBF135 Density Dependence of Chiral Condensate

Q. what can be achieved by measuring isotopes? why not single isotope? How far can we discuss?

Ans.: **p derivative of <qq>= d<qq>/dp** can be studied based on pionic Sn isotopes

Densities probed by pionic Sn with wide range of A

Important for σ_{πN} for investigation of origin of matter mass



Pionic atoms are known to probe $\sim 0.6 \rho_s$

Systematic spectroscopy of pionic Sn isotopes RIBF-135

113	114	115	116	117	118	119	120	121	122	123	124	125	126	127	128
Ι	I	I	I	I	I	I	I	I	I	I	I	I	I	I	I
112	113	114	115	116	117	118	119	120	121	122	123	124	125	126	127
Те															
111	112	113	114	115	116	117	118	119	120	121	122	123	124	125	126
Sb															
110	111	112	113	114	115	116	117	118	119	120	121	122	123	124	125
Sn															
109	110	111	112	113	114	115	116	117	118	119	120	121	122	123	124
In															
108	109	110	111	112	113	114	115	116	117	118	119	120	121	122	123
Cd															
107	108	109	110	111	112	113	114	115	116	117	118	119	120	121	122
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Online spectra in RIBF-135 (2021)



D-candidate: S.Y. Matsumoto

We are preparing for Inverse kinematics (RIBF-214)

For kinematical reasons, ambiguities in the incident beam energy do not affect the results. **The resolution will be even improved.**

Proposing D(136 Xe, 3 He) reaction at T = 250 MeV/u at RIBF



72 hours with 10¹⁰/s ¹³⁶Xe beam

NP2212-RIBF214

Experimental setup



S. Purushothaman et al., APR **53**, 134 (2019)

Summary

- Chiral condensate at ρ_e is evaluated to be reduced by 77±2%, which is linearly extrapolated to 60±3% at the nuclear saturation density [N. Phys. 19 (2023) 788].
- The binding energies and widths of the pionic 1s and 2p states in Sn121 are determined with very high precision. Difference between the 1s and 2p values drastically reduces the systematic errors.
- Recent theoretical progress was adopted to the <qbar q> deduction, which directly relates the chiral condensate and the pion-nucleus interaction.
- We included various updates for the first time. The updated parameters made substantial effects leading to much higher accuracy.
- For future, we are analyzing data of systematic study of pionic Sn isotopes to deduce ρ derivative of chiral condensate. We also plan measurement with "inverse kinematics" reactions for pionic xenon, which leads to future experiments for pionic unstable nuclei.

Hadron 2025 (Osaka, Mar. 27-31) Hadron in Nucleus 2025 (Kyoto, Apr. 2-4) for yielding different V(0) values. If we allow Re B_0 to be varied, we have to change the V(0) value accordingly. However the two parameters, b_0 and Re B_0 , are interrelated as in the Seki-Masutani relation obtained by reading from Fig. 1 in Ref. [26],

$$b_0 \rho(0) + 0.50 \times \operatorname{Re} B_0 \rho^2(0) = 0.062 \text{ fm}^{-2}.$$
 (4.8)

This relation can be derived by asserting that the binding energies are determined essentially by the local potential strength at the nuclear radius $(r=R_0)$. Since $\rho(R_0)$



KI et al., PRC62,025202