

BEYOND THE STANDARD MODEL

LEC 1A: WHY NEW PHYSICS?

Flip Tanedo

UC Riverside Particle Theory



29 JULY 2019



PHYSICS &
ASTRONOMY



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Outline

1a. **The Hierarchy problem**, qualitative

1b. **Supersymmetry**, overview

2a. **Supersymmetry**, tools

2b. **Extra dimensions/compositeness**, qualitative

3a. **Extra dimensions**, tools

3b. **Compositeness**, tools

3c. **Naturalness and WIMP dark matter**

Suggestion: take notes of some sort

References

arXiv.org > hep-ph > arXiv:1602.04228

High Energy Physics – Phenomenology

Beyond the Standard Model

Csaba Csáki, Philip Tanedo

(Submitted on 12 Feb 2016 (v1), last revised 19 Dec 2016 (this version,



We introduce aspects of physics beyond the Standard Model focusing on supersymmetry, extra dimensions, and a composite Higgs as solutions to the Hierarchy problem. Lectures at the European School of High Energy Physics, Parádfürdő, Hungary, 5–18 June 2013.

Comments: 119 pages, 16 figures, minor revisions and corrections from published version. Proceedings of the 2013 European School of High–Energy Physics, Paradfurdo, Hungary, 5–18 June 2013, edited by M. Mulders and G. Perez, CERN–2015–004 (CERN, Geneva, 2015), ISBN: 9789290834205. v2: typos corrected, references updated

.... and references therein

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Why go *beyond* the Standard Model?

My top 5 (personal, not exhaustive)

1. **The Hierarchy Problem:** why is the Higgs light?
2. **Dark matter:** what is it? (*Graciela*)
3. **Baryogenesis:** why so much matter?
4. **Strong CP problem:** why no neutron EDM?
5. **Grand unification:** why is Y quantized?
Bonus: relation to neutrino mass? (*Joachim*)

Particle Physics, circa 1990s

(chiral)

fundamental forces

matter particles

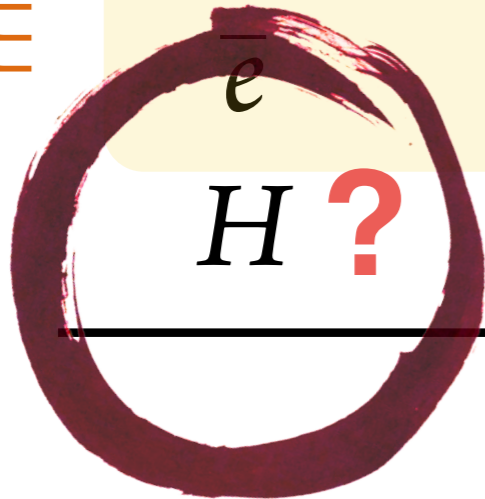
Field	Spin	$SU(3)_c$	$SU(2)_L$	$U(1)_Y$
Q	$1/2$	\square	\square	$1/6$
\bar{u}	$1/2$	$\bar{\square}$	$\mathbb{1}$	$-2/3$
\bar{d}	$1/2$	$\bar{\square}$	$\mathbb{1}$	$1/3$
L	$1/2$	$\mathbb{1}$	\square	$-1/2$
\bar{e}	$1/2$	$\mathbb{1}$	$\mathbb{1}$	-1
$H ?$	0	$\mathbb{1}$	\square	$1/2$

Particle Physics, circa 1990s

fundamental forces

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$H ?$	0	$\mathbb{1}$	\square	$1/2$

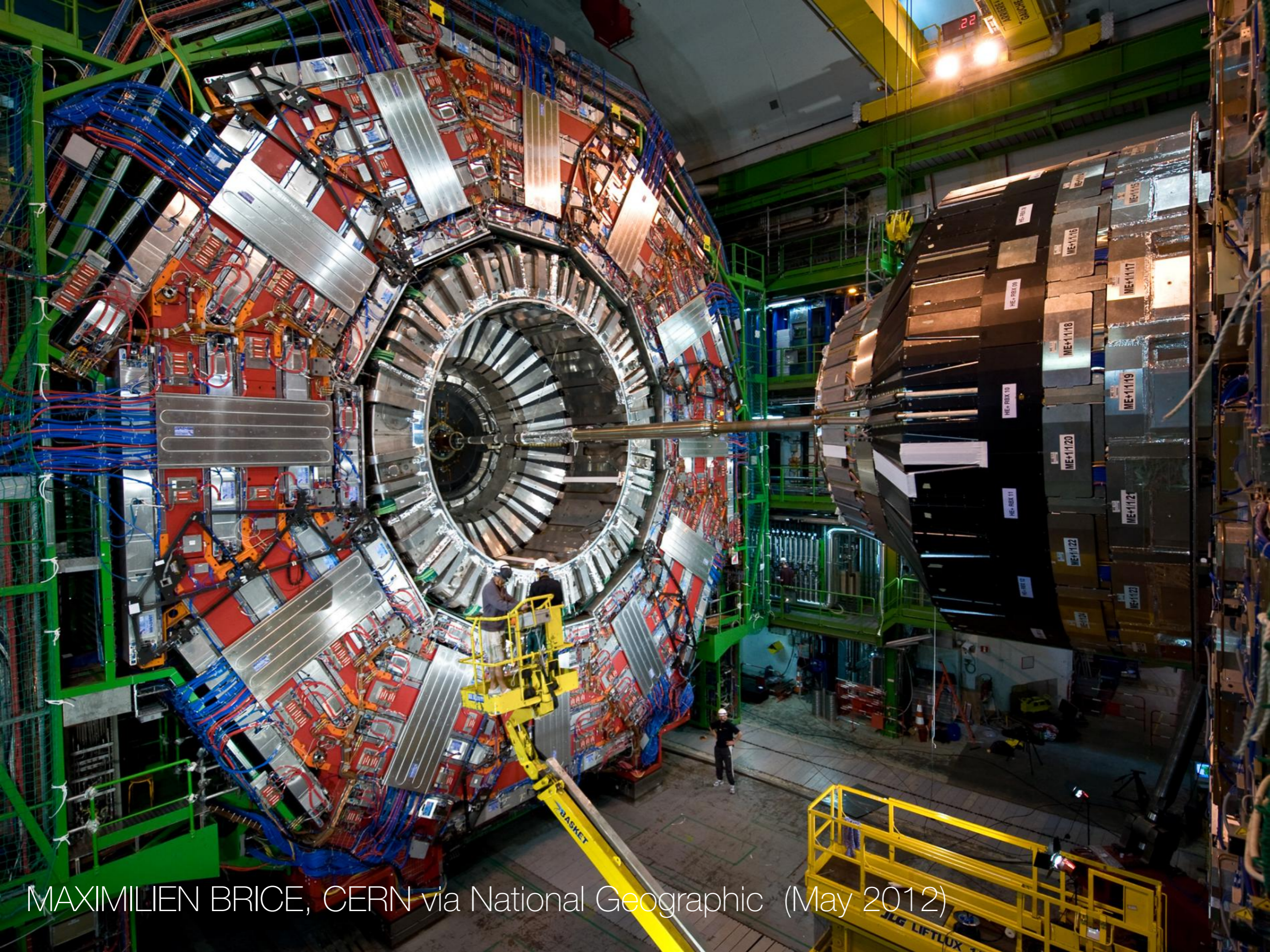
matter particles



or something to explain
unitarity of WW scattering

ME CALL IT A
"PARTICLE COLLIDER"





MAXIMILIEN BRICE, CERN via National Geographic (May 2012)

Physicists Find Elusive Particle Seen as Key to Universe

By DENNIS OVERBYE JULY 4, 2012



Scientists in Geneva on Wednesday applauded the discovery of a subatomic particle that looks like the Higgs boson.

Pool photo by Denis Balibouse

D. Overbye, New York Times, 4 July 2012

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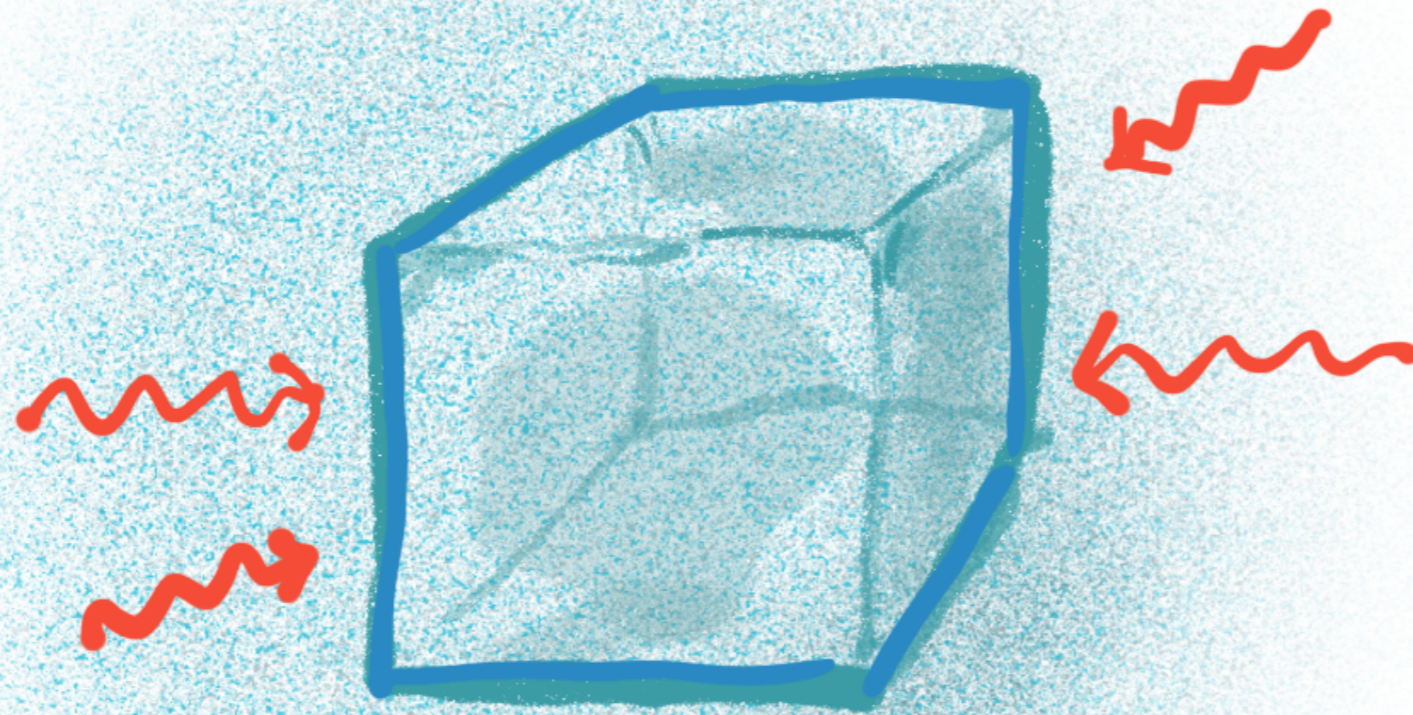
The Hierarchy Problem



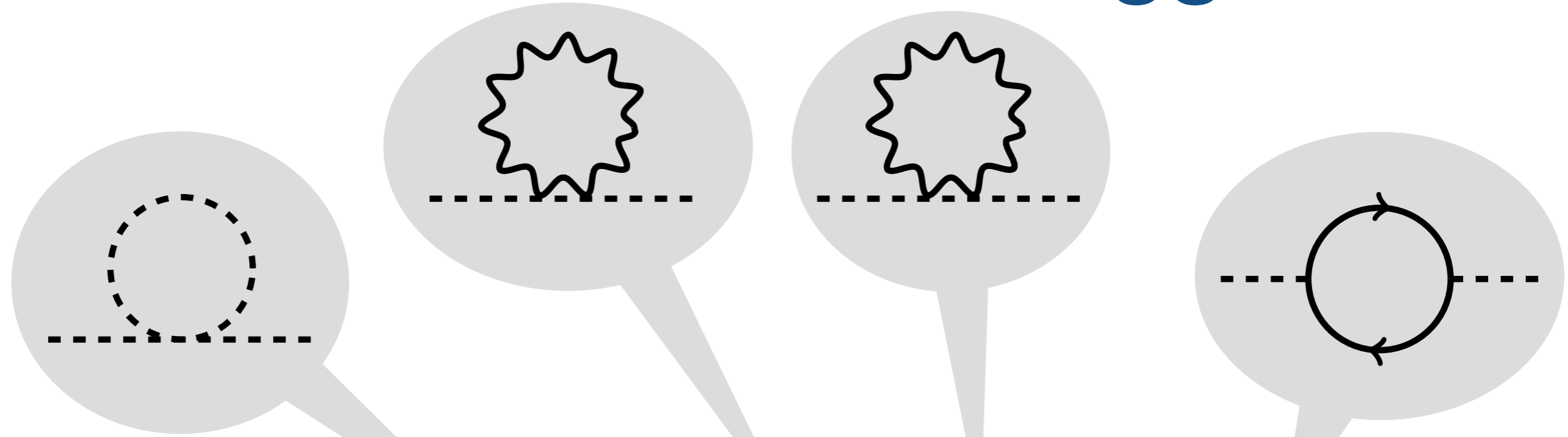
The Higgs has a *snowball's chance in hell* of being 125 GeV.

(and yet here we are)

Analogy: thermal randomness



quantum contributions to Higgs mass



$$\delta m_H^2 = \frac{\Lambda^2}{32\pi^2} \left[6\lambda + \frac{1}{4} (9g^2 + 3g'^2) - y_t^2 \right]$$

SCALARS

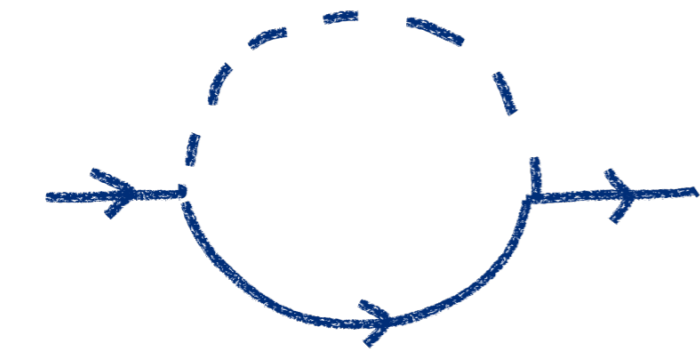
GAUGE

FERMIONS

Other “hierarchy problems” ?

Standard Model Fermions

$$\Delta m_e \sim m_e \ln \left(\frac{\Lambda}{m_e} \right)$$



Massive Gauge Bosons

$$\Delta M_W^2 \sim M_W^2 \ln \left(\frac{\Lambda}{M_W} \right)$$



...what about dim-reg?

$$\int \frac{d^d \ell}{(2\pi)^d} \frac{1}{(\ell^2 - \Delta)^n} = \frac{(-1)^n i}{(4\pi)^{d/2}} \frac{\Gamma(n - \frac{d}{2})}{\Gamma(n)} \left(\frac{1}{\Delta}\right)^{n - \frac{d}{2}}$$

$$\frac{\Gamma(2 - \frac{d}{2})}{(4\pi)^{d/2}} \left(\frac{1}{\Delta}\right)^{2 - \frac{d}{2}} = \frac{1}{(4\pi)^2} \left(\frac{2}{\epsilon} - \log \Delta - \gamma + \log(4\pi) + \mathcal{O}(\epsilon) \right)$$

$\epsilon = 4 - d.$

voila, no quadratic divergences!

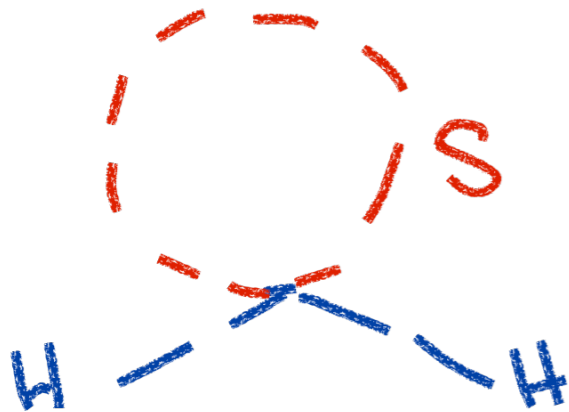
... but it was never about quadratic divergences

...what about dim-reg?

$$\Delta \mathcal{L} = \lambda_S |H|^2 S^2$$

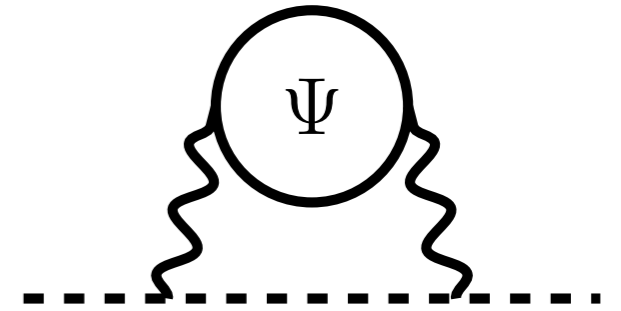
$$\delta m_H^2 = \frac{\lambda_S}{16\pi^2} \left[\cancel{\Lambda_{UV}^2} - 2m_S^2 \ln \left(\frac{\Lambda_{UV}}{m_S} \right) + (\text{finite}) \right]$$

μ_{RG} SCALE



Hierarchy problem is about **separation of scales**.

...what about dim-reg?



$$\delta m_H^2 \sim \left(\frac{g^2}{16\pi^2} \right)^2 \left[a\Lambda_{UV}^2 + 48m_F^2 \ln \frac{\Lambda_{UV}}{m_F} + (\text{finite}) \right]$$

... even if we sequester the new physics.

Frameworks for Hierarchy Problem

Supersymmetry: enforce a cancellation

Extra dimensions: effective UV scale is lower...
maybe because gravity is diluted in 5D?

Compositeness: effective UV scale is lower...
because you resolve Higgs substructure

Other? maybe that's just the way things are.

BEYOND THE STANDARD MODEL

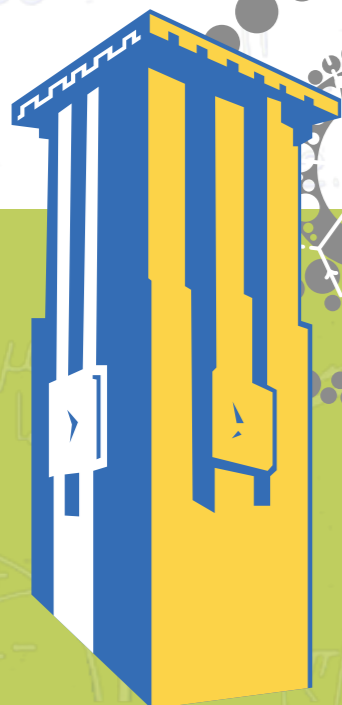
LEC1B: SUSY, QUALITATIVELY

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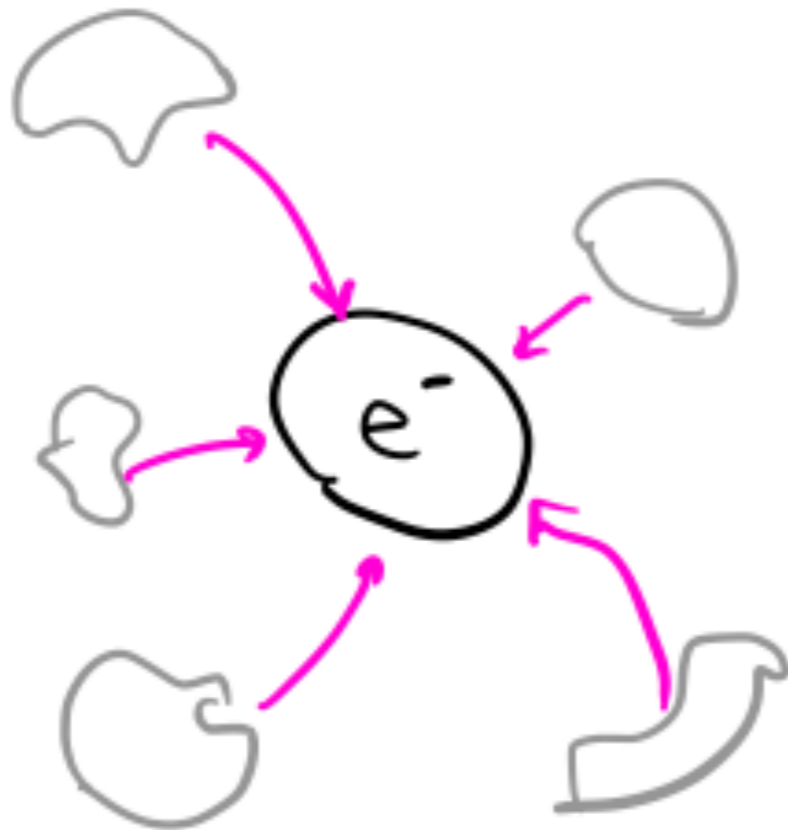
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The semi-classical electron

Why is the mass of electron small?



Imagine bringing together pieces of the electron from far away.

$$r_e \approx 10^{-17} \text{ cm}$$

The semi-classical electron

$$E = m_e c^2 \quad @ \text{ REST} \quad M = \frac{1}{2} \text{ MeV}$$
$$= M_0 c^2 + \Delta E_{\text{Coulomb}}$$

↑ "BARE MASS" not observable ↑ "SELF ENERGY" from self-repulsion

$$\Delta E_{\text{Coulomb}} = \frac{\alpha}{r_e}$$

← "radius" of e^-
expt: $r_e \approx 10^{-17}$
 $\Rightarrow \Delta E \approx 10 \text{ GeV}$

$$\text{GeV cm} = 5 \times 10^{13}$$

The semi-classical electron

$$M_e = M_0 + \Delta E_{\text{coulomb}}$$

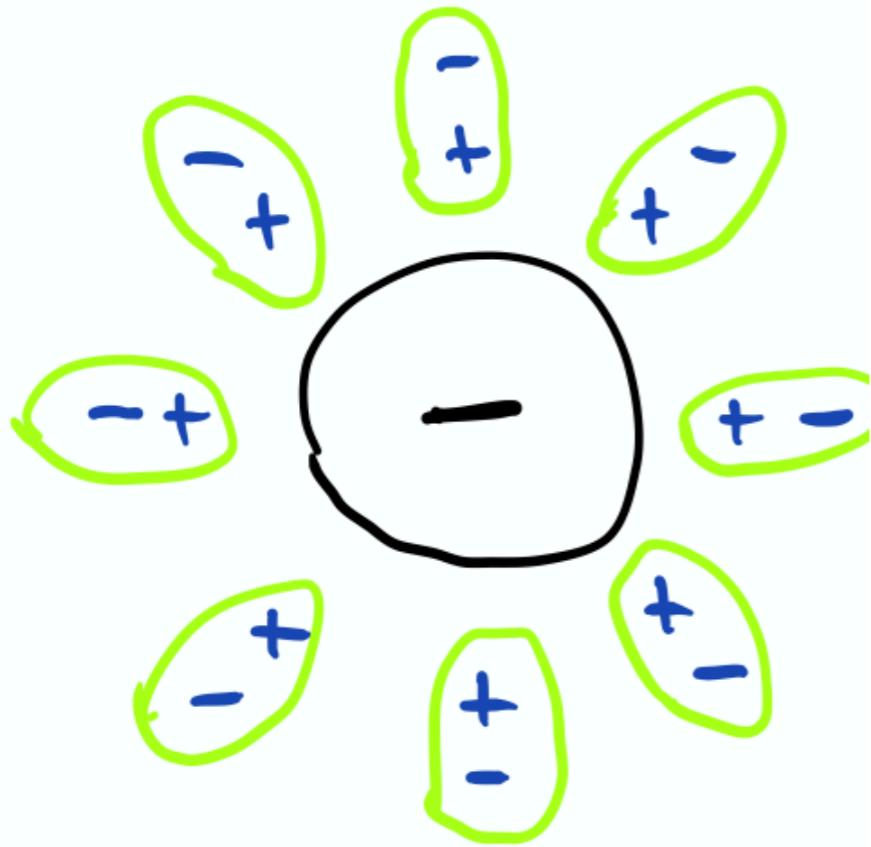
0.5 MeV 10,000 MeV

-9,999.5 MeV

$\approx 0.1\%$ TUNING!

The semi-classical electron

We know something interesting happens on length scales on the order of the electron mass.



vacuum polarization from virtual particle-antiparticle pairs

renormalization of charge, screening by virtual pairs

HAPPENS @ $r \sim (2m_e)^2$
(eg HEISENBERG $\Delta t \Delta E \sim 1$)

$1/\alpha \sim 100$ times sooner than naturalness requires!

$$\text{GeV cm} = 5 \times 10^{13}$$

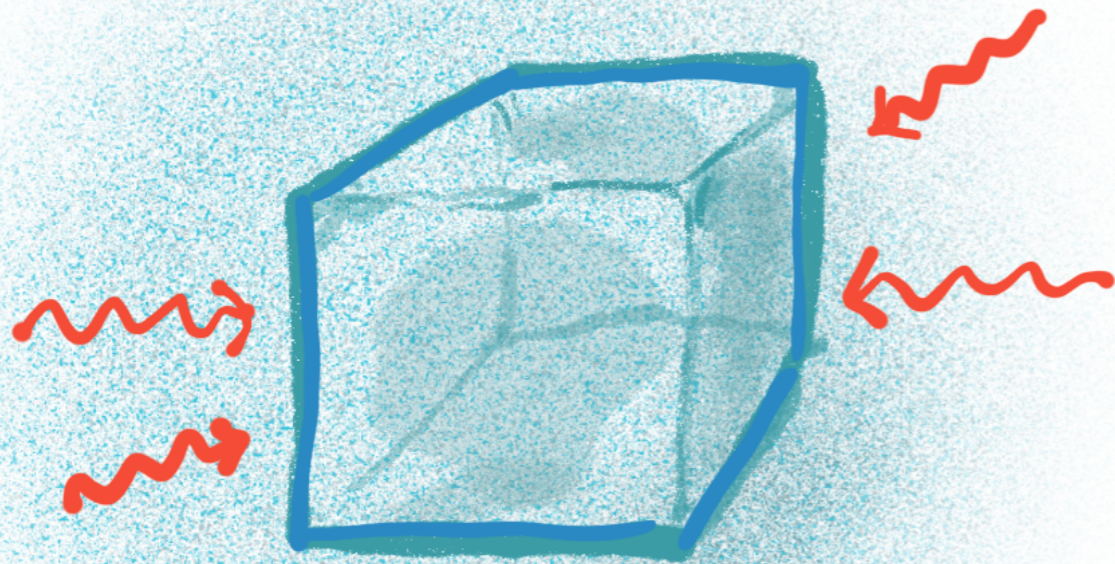
Supersymmetry

a qualitative introduction

S. Martin, *A SUSY Primer* hep-ph/9709356

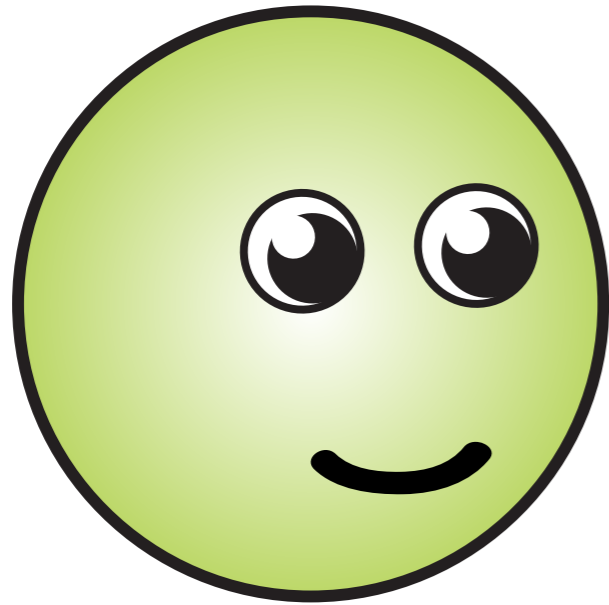
Quevedo et al. *Cambridge Lectures on SUSY* 1011.1491

Supersymmetry



Supersymmetry

VISIBLE STUFF



matter particle

force particle

NEW PARTICLES



force particle

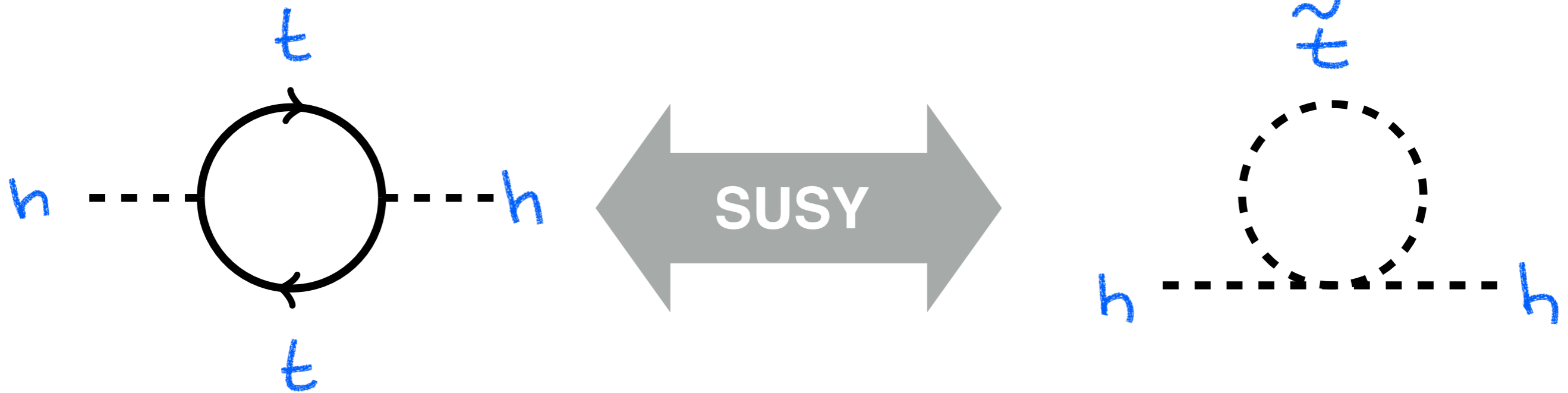
matter particle



Supersymmetry

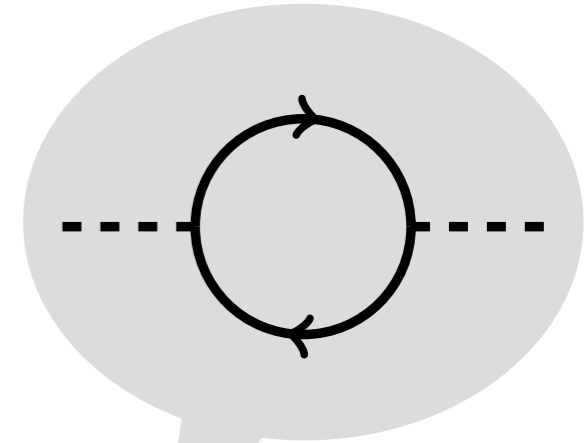
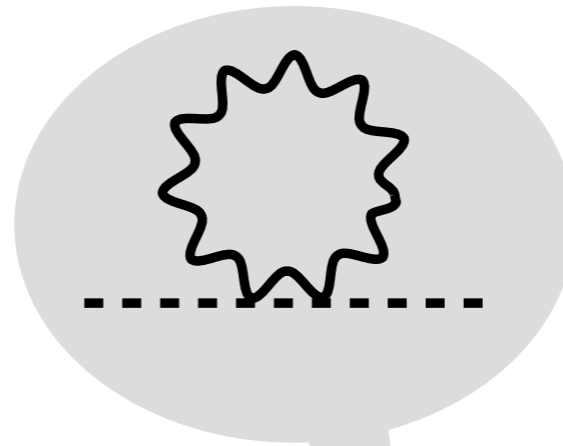
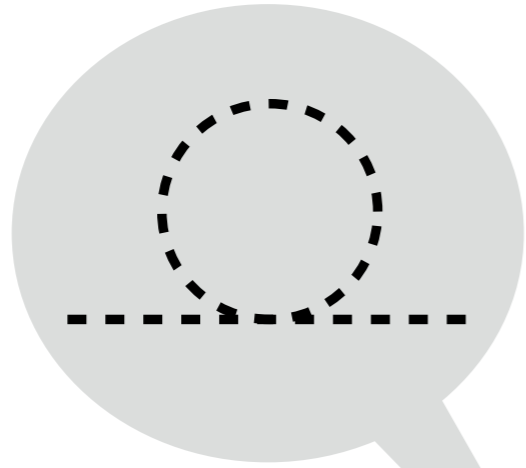
VISIBLE STUFF

NEW PARTICLES



superpartners also contribute to Higgs mass

why this could work



$$\delta m_H^2 = \frac{\Lambda^2}{32\pi^2} \left[6\lambda + \frac{1}{4} (9g^2 + 3g'^2) - y_t^2 \right]$$

relative sign between boson
and fermion loops



... just need [super]symmetry to enforce
appropriate particles and couplings

the naming of squarks

The MSSM (s)particle content

... compare to “The Naming of Cats” by TS Eliot

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Multiplets to supermultiplets

also known as *superfields*

$$H = \begin{pmatrix} \phi^1 \\ \phi^2 \end{pmatrix} \quad \curvearrowright \quad \text{SU(2) ROTATION}$$

$$H = \phi^1 \mathbf{e}_1 + \phi^2 \mathbf{e}_2 \quad \leftarrow \text{BASIS OF GAUGE MULTIPLET}$$

FERMIONIC BASIS

$$\Phi(y, \theta) = \underbrace{\varphi(y)}_{\uparrow} + \sqrt{2}\theta\psi(y) + \theta^2 F(y)$$

$$y^\mu = x^\mu + i\theta\sigma^\mu\bar{\theta}$$

Rule of thumb: SUSY spectrum

$$\Phi(y, \theta) = \varphi(y) + \sqrt{2}\theta\psi(y) + \theta^2 F(y)$$

Chiral superfield: matter superfield

complex scalar, Weyl fermion, auxiliary field

$$V = -\theta\sigma^\mu\bar{\theta}V_\mu(x) + i\theta^2\bar{\theta}\bar{\lambda}(x) - i\bar{\theta}^2\theta\lambda(x) + \frac{1}{2}\theta^2\bar{\theta}^2 D(x)$$

Vector superfield: force superfield

spin-1, Majorana fermion, auxiliary field

auxiliary fields are analogous to gauge-redundant degrees of freedom

The superfields

vector superfields (forces)

χ^{SF}	$SU(3)_c$	$SU(2)_L$	$U(1)_Y$
Q	3	2	$1/6$
\bar{U}	$\bar{\mathbf{3}}$	1	$-2/3$
\bar{D}	$\bar{\mathbf{3}}$	1	$1/3$
L	1	2	$-1/2$
\bar{E}	1	1	-1
H_d	1	2	$-1/2$
H_u	1	2	$1/2$

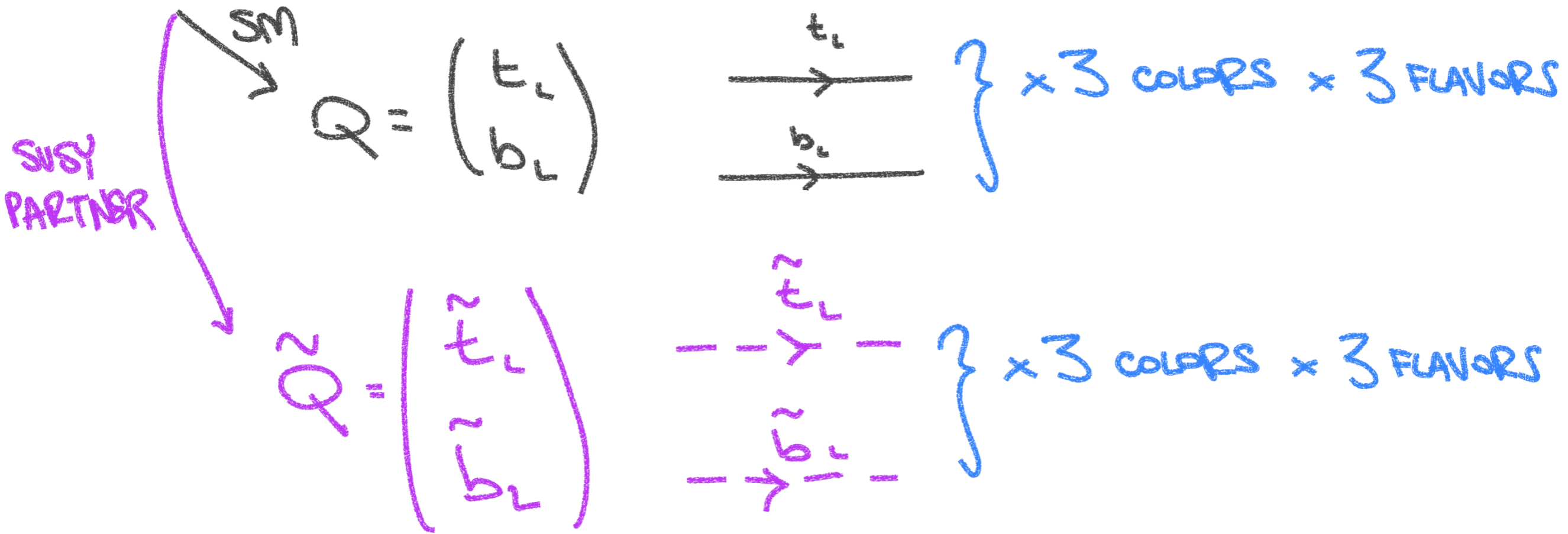
chiral (matter)
superfields

example: quark doublet

χ^{SF}	$SU(3)_c$	$SU(2)_L$	$U(1)_Y$
Q	$\mathbf{3}$	$\mathbf{2}$	$1/6$

example: quark doublet

χ^{SF}	$SU(3)_c$	$SU(2)_L$	$U(1)_Y$
Q	3	2	$1/6$



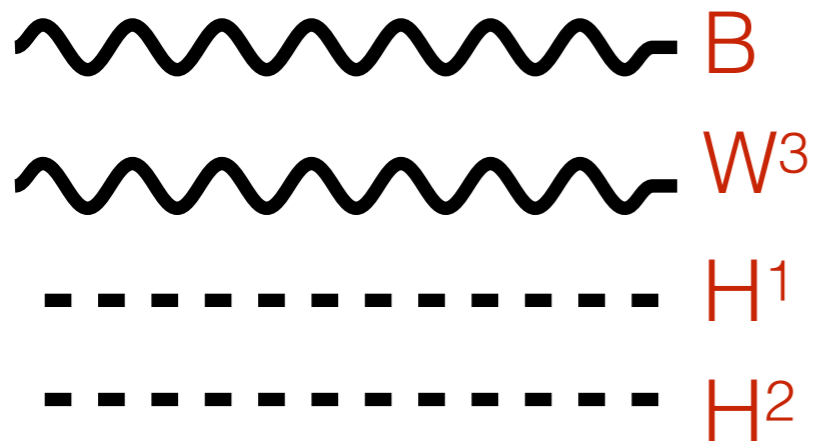
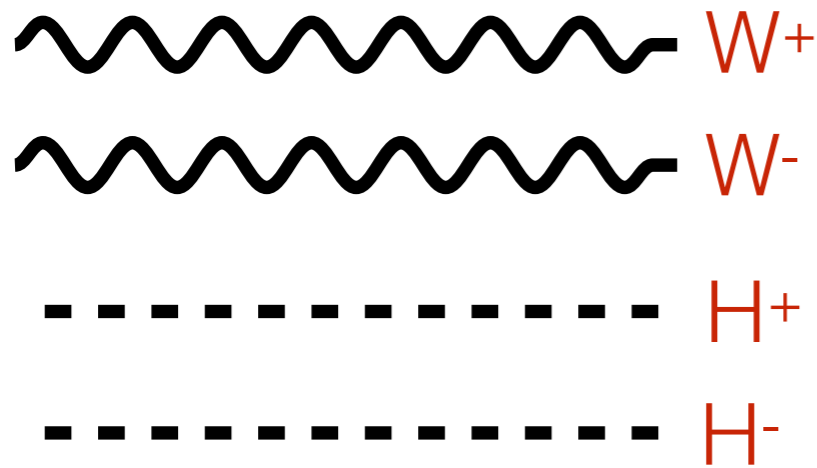
The superfields

SM BOSON \rightarrow -ino
 SM FERMION \rightarrow S-

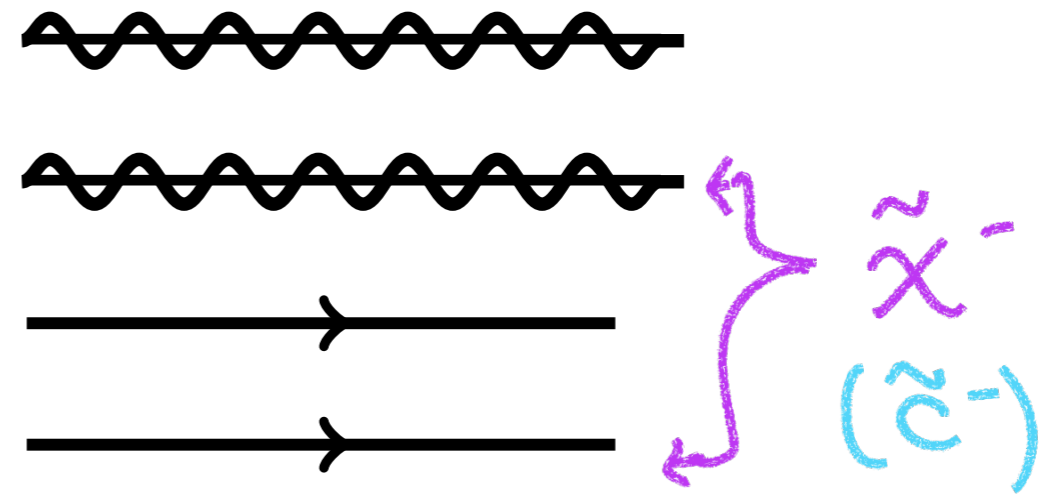
	χ^{SF}	g/gluinos $SU(3)_c$	W^a /winos $SU(2)_L$	B/bino $U(1)_Y$	
LH up					sup-L
LH down	Q	3	2	$1/6$	sdown-L
RH up*	\bar{U}	$\bar{\mathbf{3}}$	1	$-2/3$	sup-R*
RH down*	\bar{D}	$\bar{\mathbf{3}}$	1	$1/3$	sdown-R*
LH neutrino	L	1	2	$-1/2$	sneutrino
LH electron	\bar{E}	1	1	-1	selectron-L
RH electron*	H_d	1	2	$-1/2$	selectron-R
H^+, H^1	H_u	1	2	$1/2$	higgsinos
H^2, H^-					higgsinos

“sfermions” (squarks and sleptons) gauginos

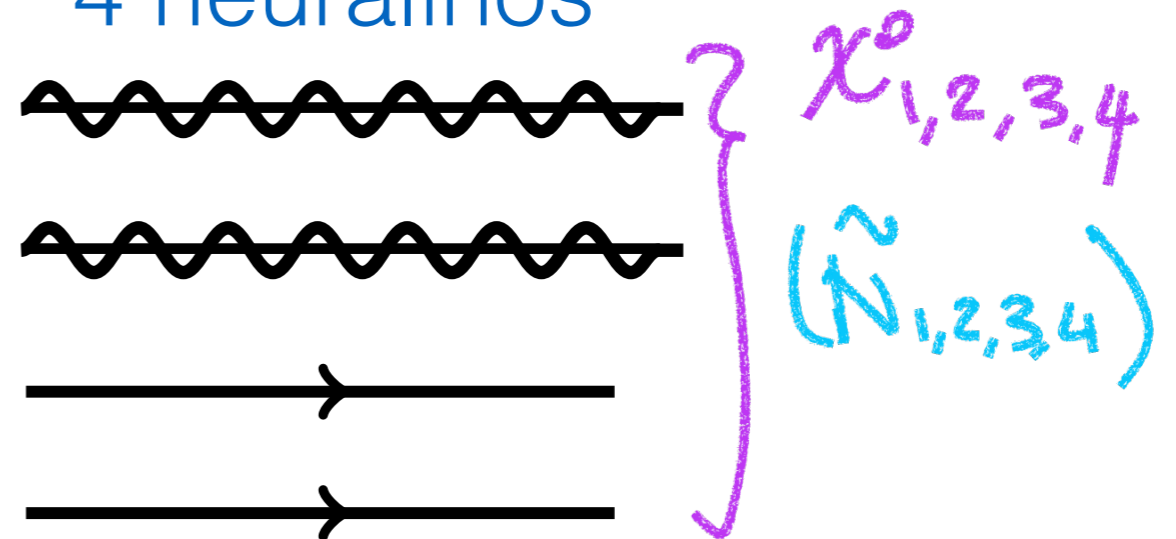
Gaugino mixing



2 charginos + h.c.



4 neutralinos



We never talk about photinos or zinos.

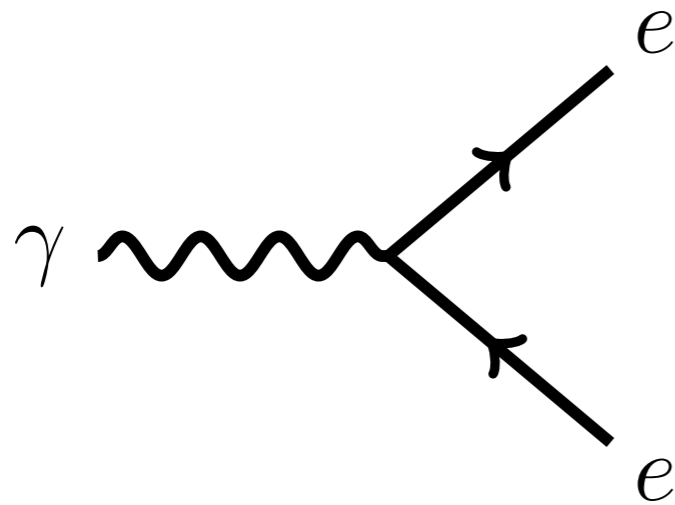
MSSM Feynman Rules

a cheater's guide (SUSY limit)

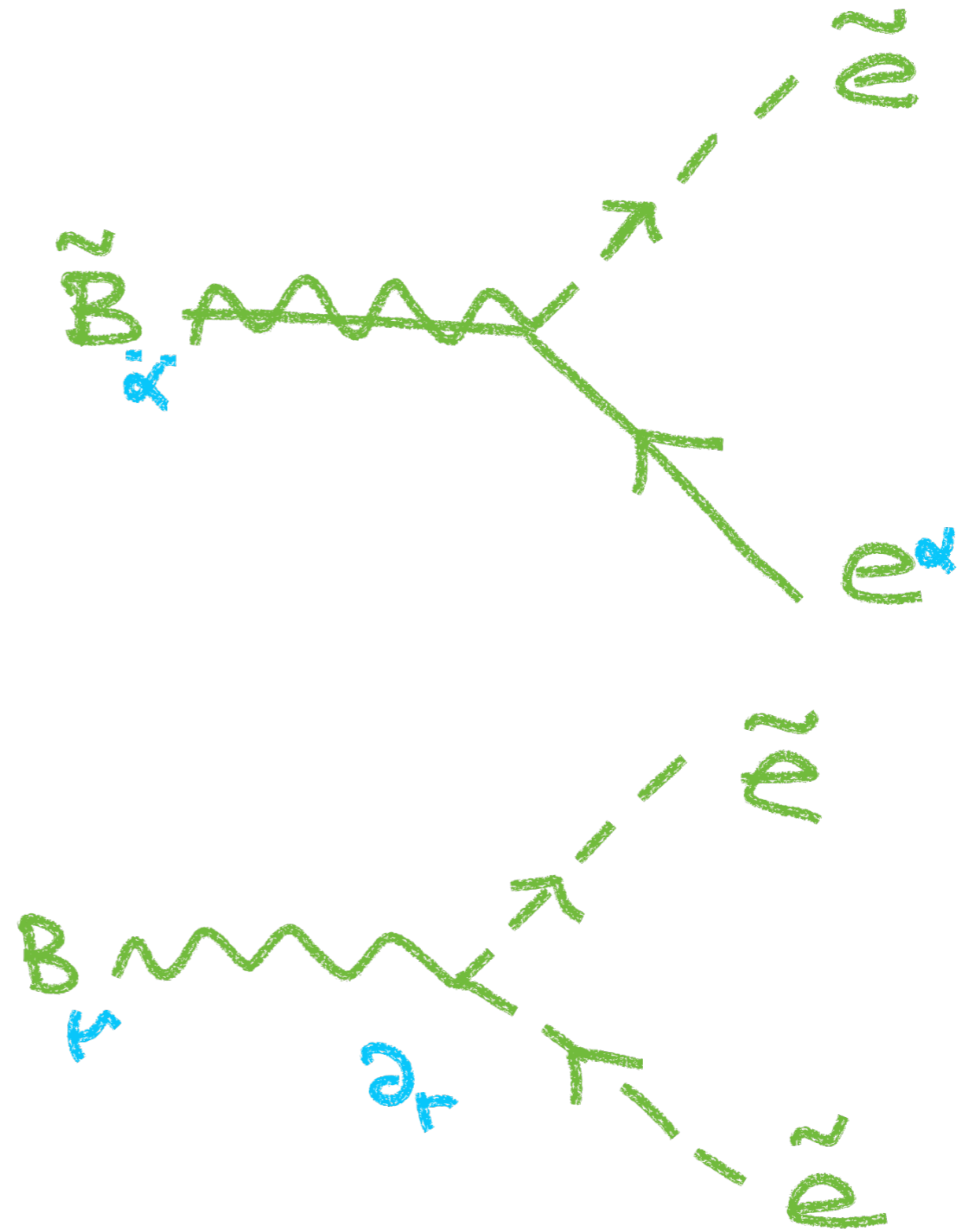
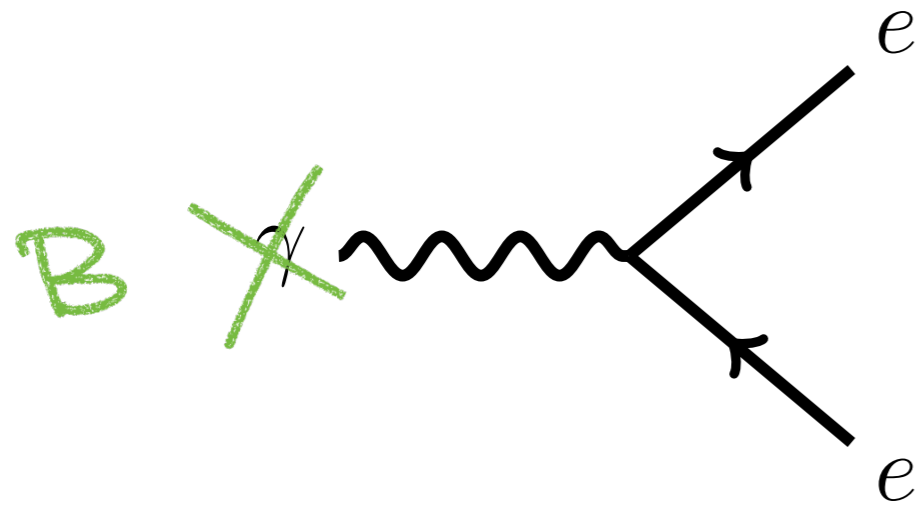
the dirty secret (approximate)

1. take a Standard Model vertex
2. replace two particles with SUSY partners
3. make sure indices contract (they will)

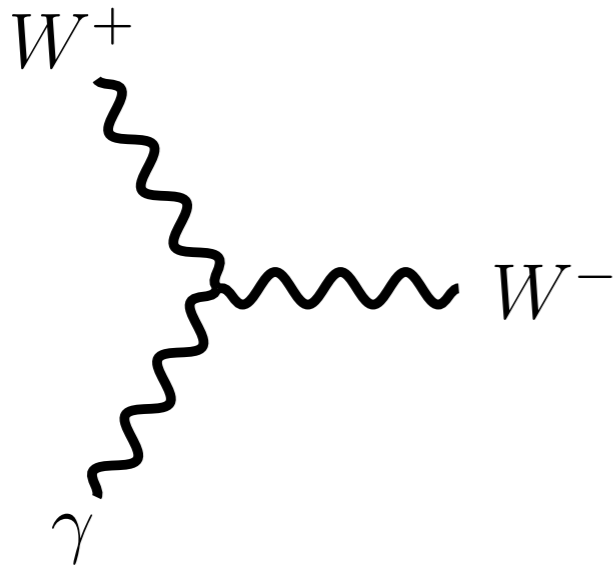
example



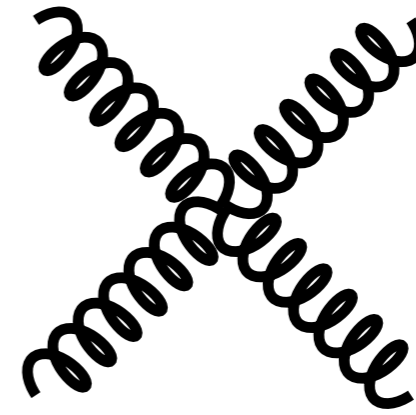
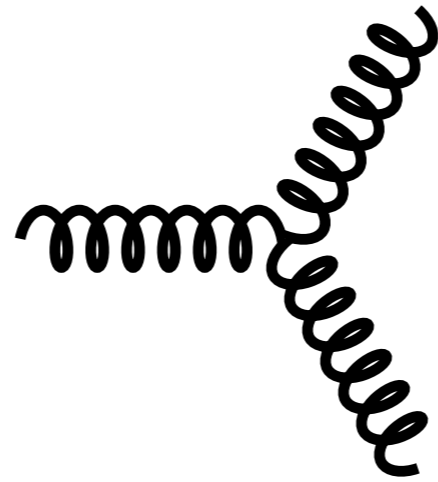
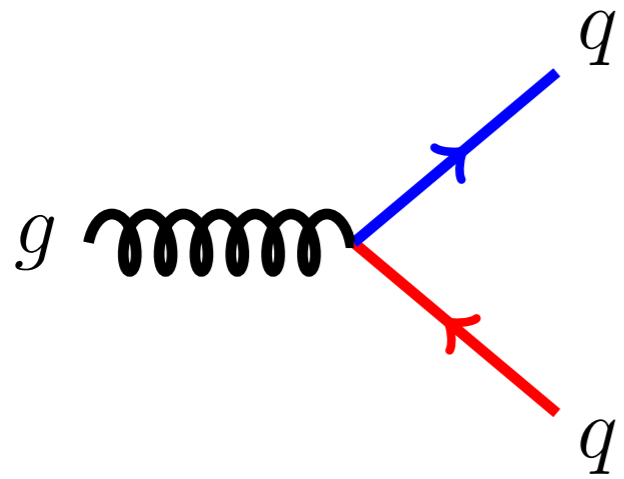
example



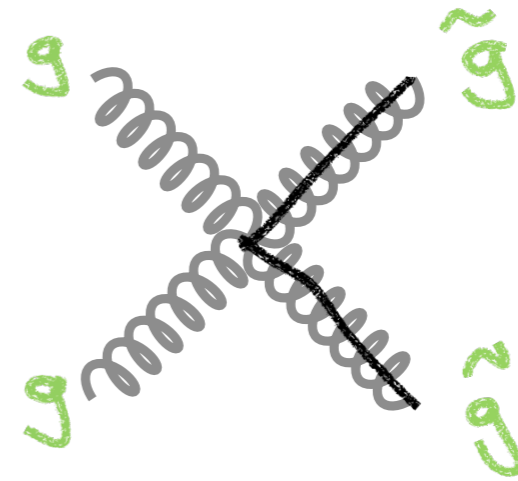
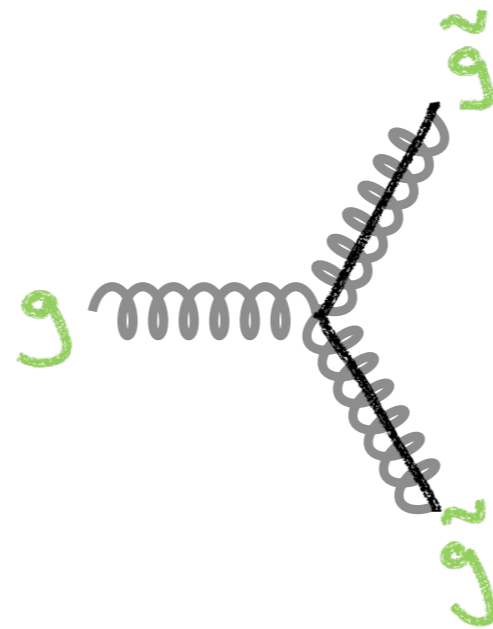
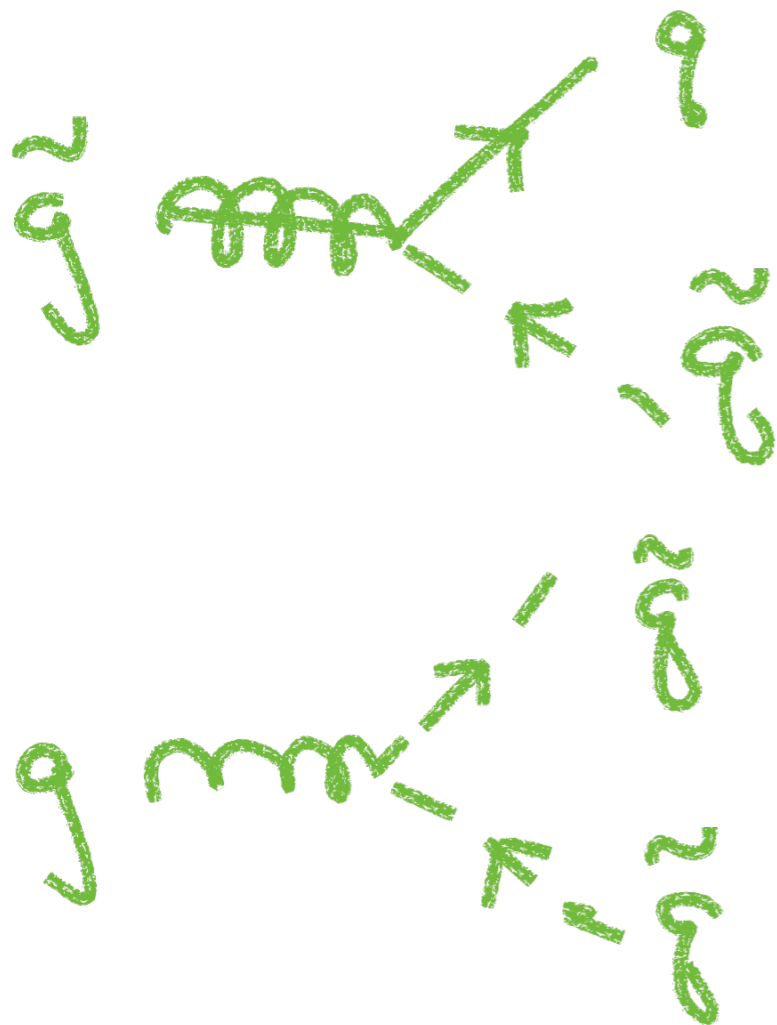
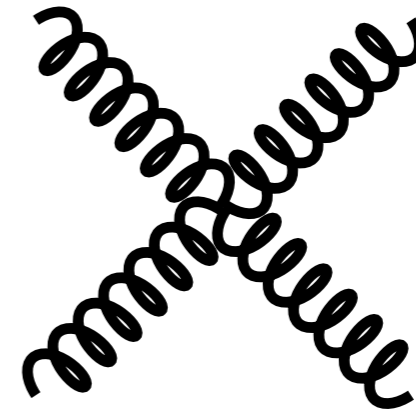
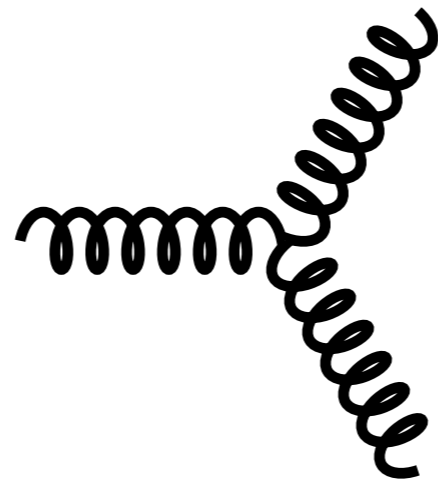
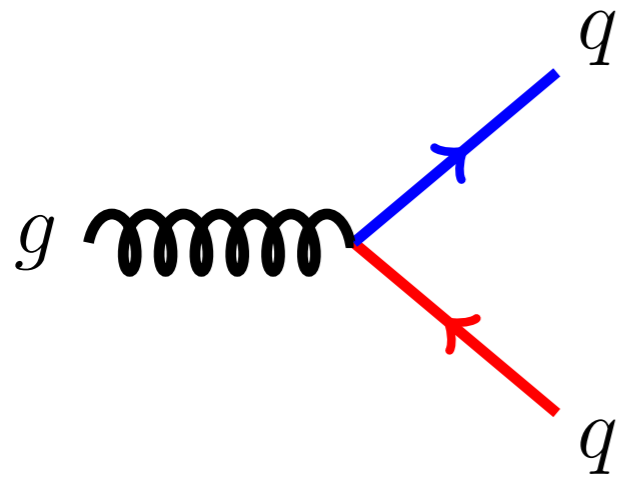
example



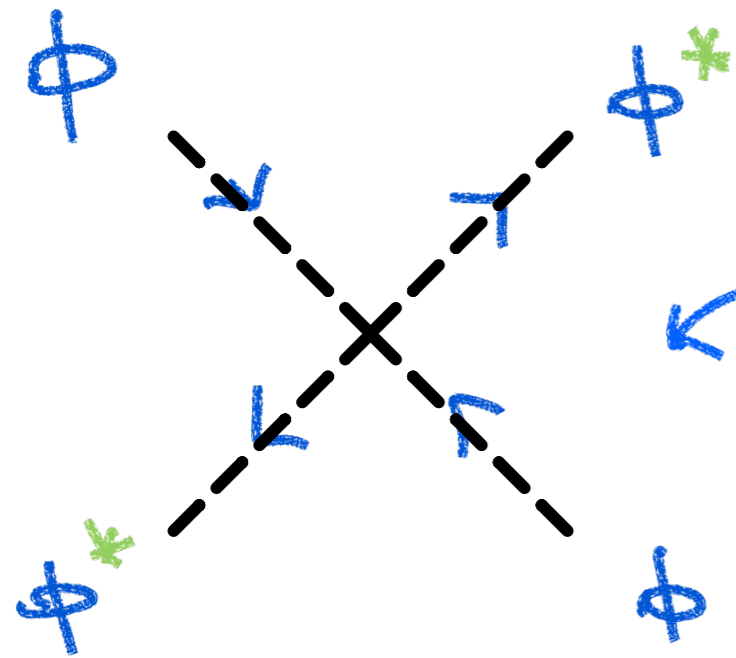
example



example



what you miss



$$\mathcal{L} = \frac{1}{2} g^2 \sum_a (\phi^* T^a \phi)^2$$

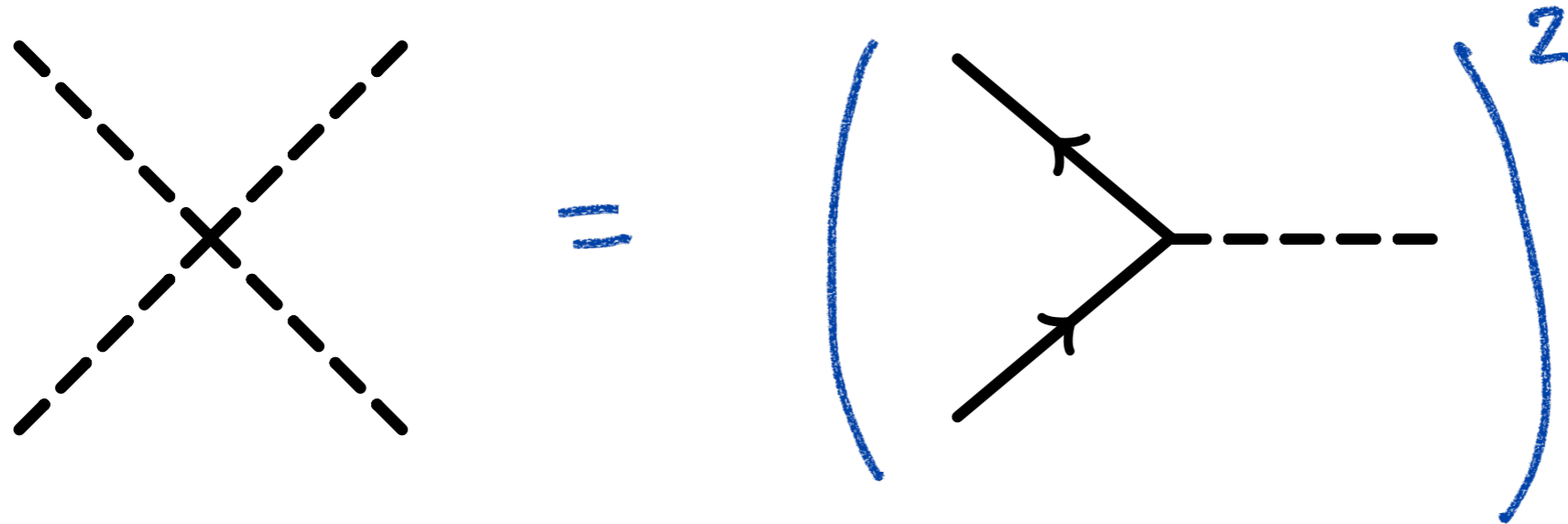
Scalars charged
under gauge group

sum over gauge
generators

Some quartic terms come from kinetic terms,
proportional to gauge coupling.

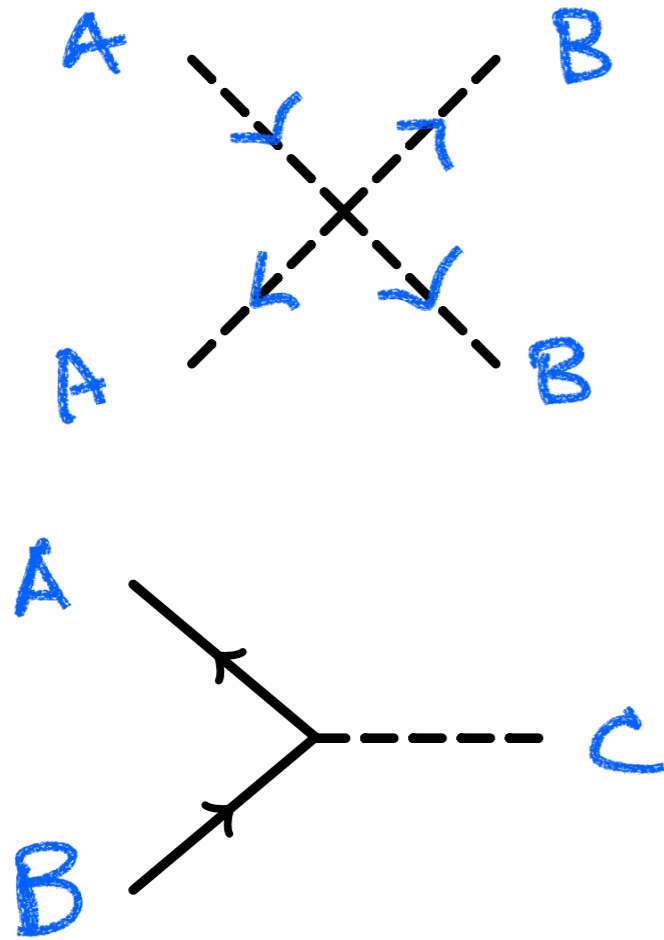
... may make you worry about Higgs sector!

what you miss



other quartic terms come from Yukawa terms
... this is really important for cancellation of loops

Other Interactions

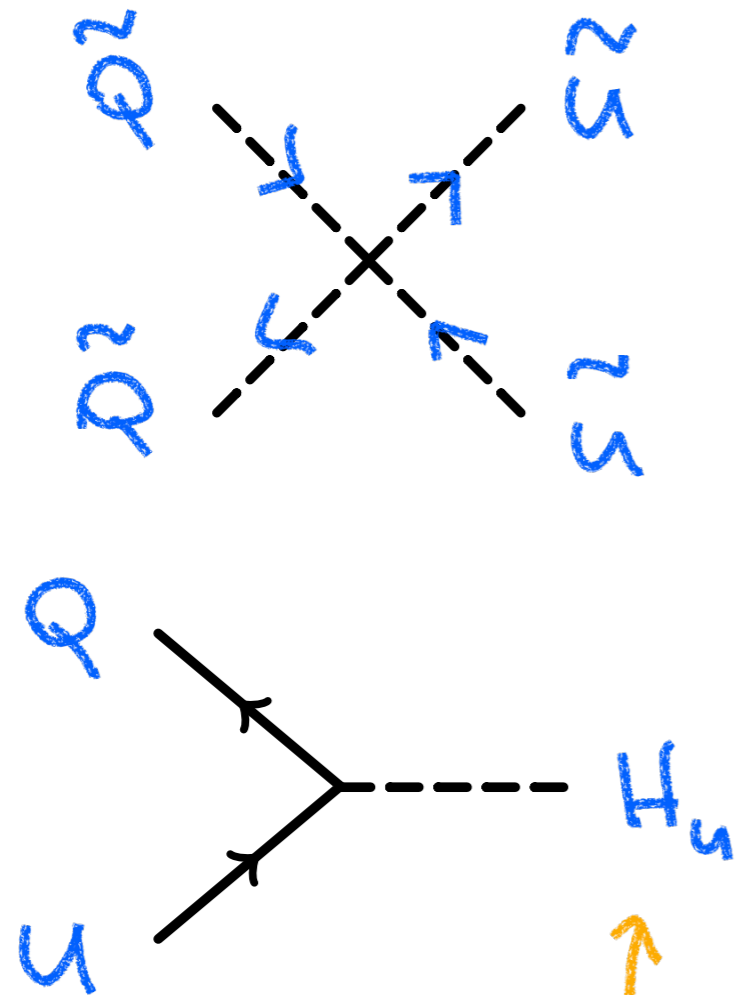


χ^{SF}	$SU(3)_c$	$SU(2)_L$	$U(1)_Y$
Q	3	2	$1/6$
\bar{U}	$\bar{3}$	1	$-2/3$
\bar{D}	$\bar{3}$	1	$1/3$
L	1	2	$-1/2$
\bar{E}	1	1	-1
H_d	1	2	$-1/2$
H_u	1	2	$1/2$

More fields, more ways to put them together.

Pick gauge invariant combination of 3 XSF, A, B, C

Other Interactions



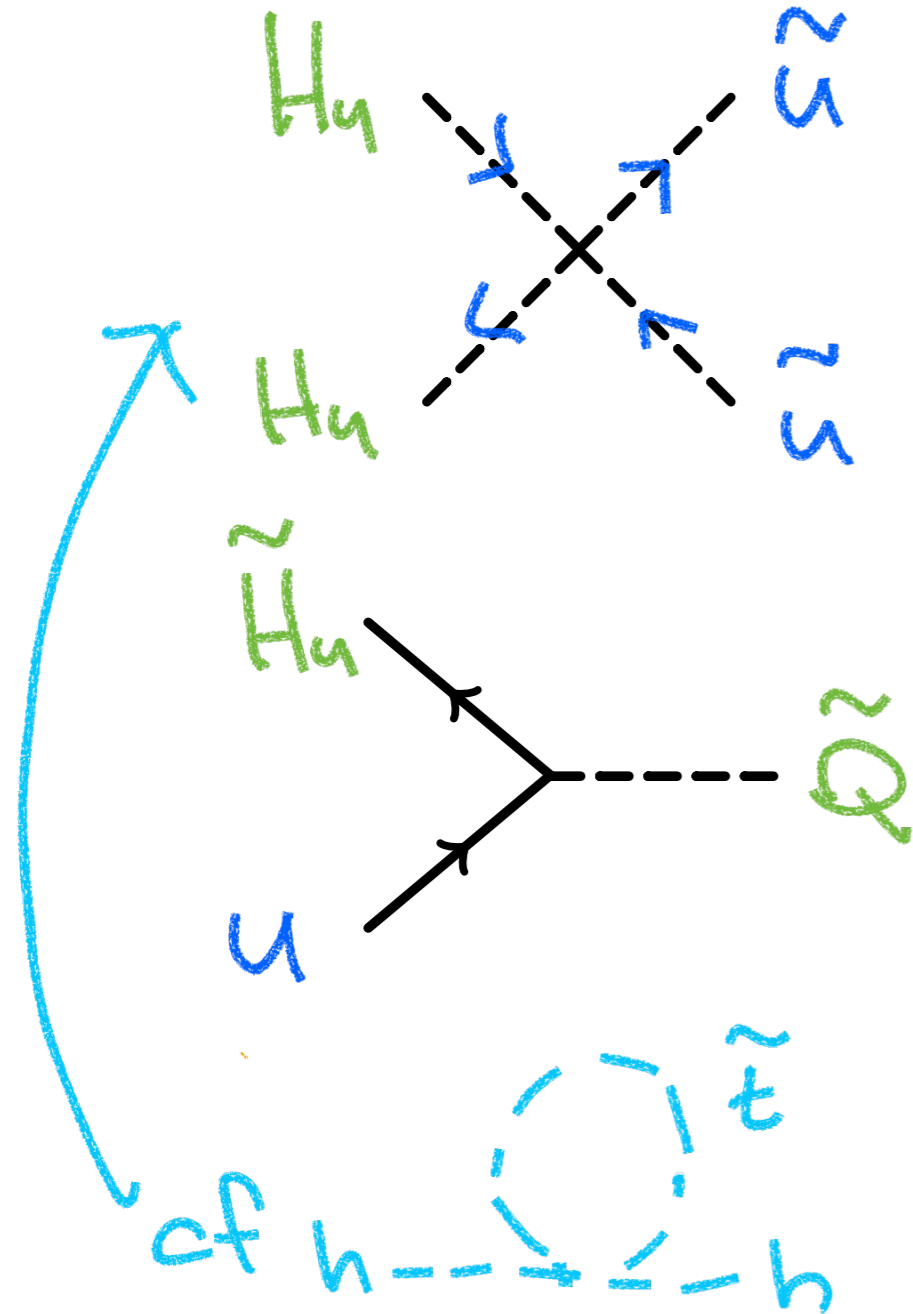
χ_{SF}	$SU(3)_c$	$SU(2)_L$	$U(1)_Y$
Q	3	2	$1/6$
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L	1	2	$-1/2$
\bar{E}	1	1	-1
H_d	1	2	$-1/2$
H_u	1	2	$1/2$

$$\begin{pmatrix} H_u + \\ 4 \\ 4 \end{pmatrix}$$

USUAL UP-TYPE YUKAWA

Pick gauge invariant combination of 3 XSF, A, B, C

Other Interactions



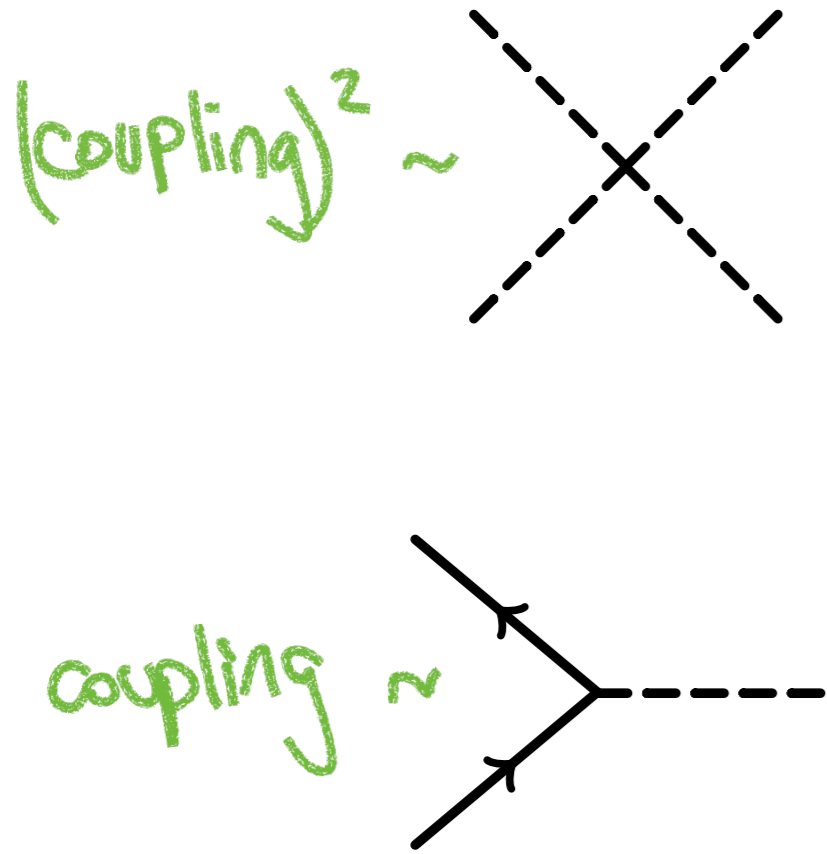
(u, d)

χ_{SF}	$SU(3)_c$	$SU(2)_L$	$U(1)_Y$	
Q	3	2	$1/6$	A C
\bar{U}	$\bar{3}$	1	$-2/3$	B
\bar{D}	$\bar{3}$	1	$1/3$	
L	1	2	$-1/2$	
\bar{E}	1	1	-1	
H_d	1	2	$-1/2$	
H_u	1	2	$1/2$	C A

USUAL UP-TYPE YUKAWA
+ SUSY "PARTNER" VERTICES

Pick gauge invariant combination of 3 XSF, A, B, C

... first look at a superpotential



χ_{SF}	$\text{SU}(3)_c$	$\text{SU}(2)_L$	$\text{U}(1)_Y$
Q	3	2	$1/6$
\bar{U}	$\bar{\mathbf{3}}$	1	$-2/3$
\bar{D}	$\bar{\mathbf{3}}$	1	$1/3$
L	1	2	$-1/2$
\bar{E}	1	1	-1
H_d	1	2	$-1/2$
H_u	1	2	$1/2$

HOW TO READ: [coupling] ABC χ_{SF}

$$W^{(\text{good})} = y_u^{ij} Q^i H_u \bar{U}^j + y_d^{ij} Q^i H_d \bar{D}^j + y_e^{ij} L^i H_d \bar{E}^j + \mu H_u H_d$$

$$W^{(\text{bad})} = \lambda_1^{ijk} Q^i L^j \bar{D}^k + \lambda_2^{ijk} L^i L^j \bar{E}^k + \lambda_3^i L^i H_u + \lambda_4^{ijk} \bar{D}^i \bar{D}^j \bar{U}^k$$

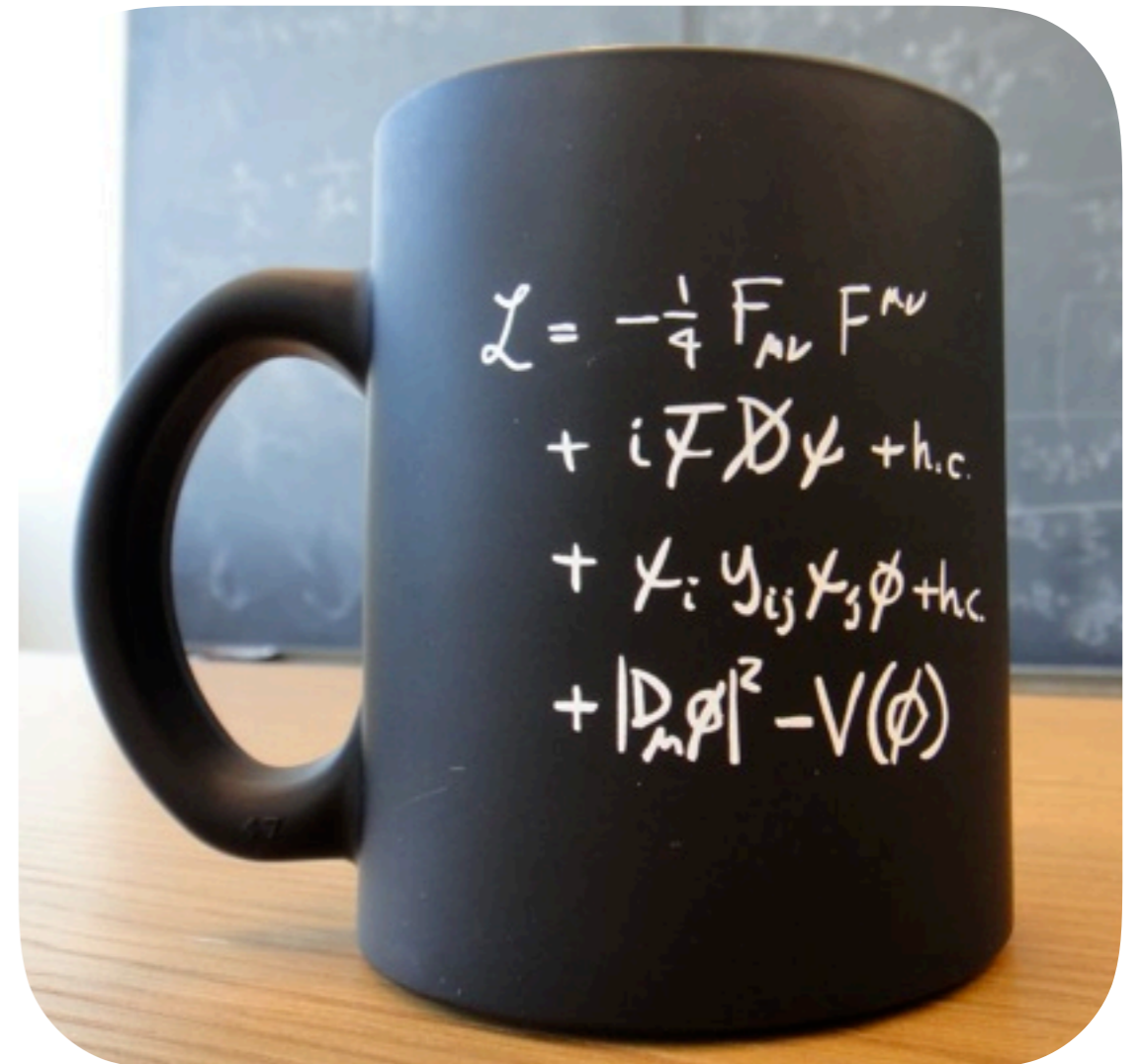
↑ PROBLEMATIC

The MSSM... approximately

Exercise 1: what are the particles of the **minimal supersymmetric Standard Model**?

Exercise 2: what are the SUSY interactions?

ignore quartics,
hint on the right



Supersymmetry

super-shortcomings

there are no sparticles

The electron has an antiparticle with the same properties except CP

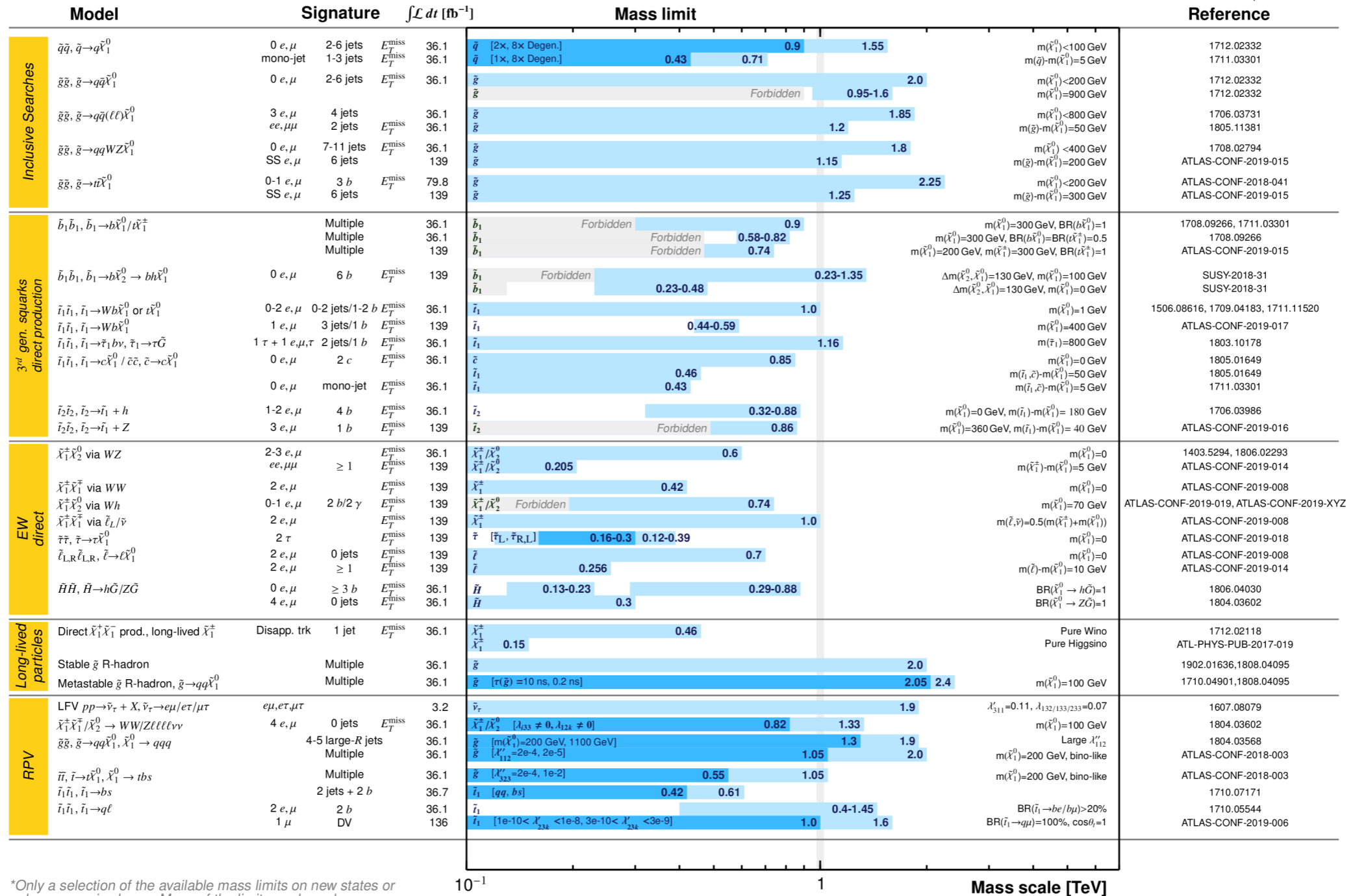
... but it definitely **does not have a super-partner** with the same properties except

We would have discovered it a long time ago.

there are no sparticles

ATLAS SUSY Searches* - 95% CL Lower Limits
July 2019

ATLAS Preliminary
 $\sqrt{s} = 13$ TeV



*Only a selection of the available mass limits on new states or phenomena is shown. Many of the limits are based on simplified models, c.f. refs. for the assumptions made.

there are no sparticles

ATLAS SUSY Searches* - 95% CL Lower Limits
July 2019

ATLAS Preliminary
 $\sqrt{s} = 13$ TeV

Model	Signature	$\int \mathcal{L} dt$ [fb $^{-1}$]	Mass limit	Reference	
Inclusive Searches	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$	0 e, μ	2-6 jets E_T^{miss} 36.1	\tilde{q} [2x, 8x Degener.] 0.9 1.55	$m(\tilde{\chi}_1^0) < 100$ GeV 1712.02332
		mono-jet	1-3 jets E_T^{miss} 36.1	\tilde{q} [1x, 8x Degener.] 0.43 0.71	$m(\tilde{q}) - m(\tilde{\chi}_1^0) = 5$ GeV 1711.03301
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	0 e, μ	2-6 jets E_T^{miss} 36.1	Forbidden 0.95-1.6 2.0	$m(\tilde{\chi}_1^0) < 200$ GeV 1712.02332
					$m(\tilde{\chi}_1^0) = 900$ GeV 1712.02332
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(\ell\ell)\tilde{\chi}_1^0$	3 e, μ	4 jets E_T^{miss} 36.1		$m(\tilde{\chi}_1^0) < 800$ GeV 1706.03731
		$ee, \mu\mu$	2 jets E_T^{miss} 36.1		$m(\tilde{g}) - m(\tilde{\chi}_1^0) = 50$ GeV 1805.11381
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qqWZ\tilde{\chi}_1^0$	0 e, μ	7-11 jets E_T^{miss} 36.1		$m(\tilde{\chi}_1^0) < 400$ GeV 1708.02794
		SS e, μ	6 jets E_T^{miss} 36.1		$m(\tilde{g}) - m(\tilde{\chi}_1^0) = 200$ GeV ATLAS-CONF-2019-015
		0-1 e, μ	3 b E_T^{miss} 73.8		$m(\tilde{\chi}_1^0) < 200$ GeV ATLAS-CONF-2018-041
		SS e, μ	6 jets E_T^{miss} 139		$m(\tilde{g}) - m(\tilde{\chi}_1^0) = 300$ GeV ATLAS-CONF-2019-015
RPV		Multiple	36.1	\tilde{b}_1 0.9	$m(\tilde{\chi}_1^0) = 300$ GeV, $BR(\tilde{b}\tilde{\chi}_1^0) = 1$ 1708.09266, 1711.03301
		Multiple	36.1	\tilde{b}_1 0.58-0.82	$m(\tilde{\chi}_1^0) = 300$ GeV, $BR(\tilde{b}\tilde{\chi}_1^0) = BR(\tilde{t}\tilde{\chi}_1^0) = 0.5$ 1708.09266
		Multiple	139	\tilde{b}_1 0.74	$m(\tilde{\chi}_1^0) = 200$ GeV, $m(\tilde{\chi}_2^0) = 300$ GeV, $BR(\tilde{\chi}_1^0) = 1$ ATLAS-CONF-2019-015
	$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$	6 b	E_T^{miss} 139		$\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 130$ GeV, $m(\tilde{\chi}_1^0) = 100$ GeV SUSY-2018-31
		1-2 b	E_T^{miss} 36.1		$\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 130$ GeV, $m(\tilde{\chi}_1^0) = 0$ GeV SUSY-2018-31
		b	E_T^{miss} 139		$m(\tilde{\chi}_1^0) = 1$ GeV 1506.08616, 1709.04183, 1711.11520
		b	E_T^{miss} 36.1		$m(\tilde{\chi}_1^0) = 400$ GeV ATLAS-CONF-2019-017
		b	E_T^{miss} 36.1		$m(\tilde{\tau}_1) = 800$ GeV 1803.10178
			E_T^{miss} 36.1		$m(\tilde{\chi}_1^0) = 0$ GeV 1805.01649
			E_T^{miss} 139		$m(\tilde{\tau}_1, \tilde{\nu}) - m(\tilde{\chi}_1^0) = 50$ GeV 1805.01649
		E_T^{miss} 139		$m(\tilde{\tau}_1, \tilde{\nu}) - m(\tilde{\chi}_1^0) = 5$ GeV 1711.03301	
		E_T^{miss} 36.1		$m(\tilde{\chi}_1^0) = 0$ GeV, $m(\tilde{\tau}_1) - m(\tilde{\chi}_1^0) = 180$ GeV 1706.03986	
		E_T^{miss} 139		$m(\tilde{\chi}_1^0) = 360$ GeV, $m(\tilde{\tau}_1) - m(\tilde{\chi}_1^0) = 40$ GeV ATLAS-CONF-2019-016	
		E_T^{miss} 36.1		$m(\tilde{\chi}_1^0) = 0$ 1403.5294, 1806.02293	
		E_T^{miss} 139		$m(\tilde{\chi}_1^0) - m(\tilde{\chi}_2^0) = 5$ GeV ATLAS-CONF-2019-014	
		E_T^{miss} 139		$m(\tilde{\chi}_1^0) = 0$ ATLAS-CONF-2019-008	
		E_T^{miss} 139		$m(\tilde{\chi}_1^0) = 70$ GeV ATLAS-CONF-2019-019, ATLAS-CONF-2019-XYZ	
		E_T^{miss} 139		$m(\tilde{\ell}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^0) + m(\tilde{\chi}_2^0))$ ATLAS-CONF-2019-008	
		E_T^{miss} 139		$m(\tilde{\chi}_1^0) = 0$ ATLAS-CONF-2019-018	
		E_T^{miss} 139		$m(\tilde{\chi}_1^0) = 0$ ATLAS-CONF-2019-008	
		E_T^{miss} 139		$m(\tilde{\ell}) - m(\tilde{\chi}_1^0) = 10$ GeV ATLAS-CONF-2019-014	
	3 b	E_T^{miss} 36.1		$BR(\tilde{\chi}_1^0 \rightarrow h\tilde{G}) = 1$ 1806.04030	
	0 jets	E_T^{miss} 36.1		$BR(\tilde{\chi}_1^0 \rightarrow Z\tilde{G}) = 1$ 1804.03602	
	1 jet	E_T^{miss} 36.1		Pure Wino 1712.02118	
		E_T^{miss} 36.1		Pure Higgsino ATL-PHYS-PUB-2017-019	
	Multiple	36.1		\tilde{g} 2.0 1902.01636, 1808.04095	
	Multiple	36.1		\tilde{g} [$\tau(\tilde{g}) = 10$ ns, 0.2 ns] 2.05 2.4 $m(\tilde{\chi}_1^0) = 100$ GeV 1710.04901, 1808.04095	
	$e\mu, e\tau, \mu\tau$	E_T^{miss} 3.2		$\tilde{\nu}_\tau$ 1.9 $\lambda'_{311} = 0.11, \lambda'_{132/133/233} = 0.07$ 1607.08079	
	4 e, μ	0 jets E_T^{miss} 36.1		$\tilde{\chi}_1^\pm / \tilde{\chi}_2^0$ [$\lambda_{333} \neq 0, \lambda_{124} \neq 0$] 0.82 1.33 $m(\tilde{\chi}_1^0) = 100$ GeV 1804.03602	
	4-5 large- R jets	E_T^{miss} 36.1		$\tilde{g}\tilde{g}$ [$m(\tilde{\chi}_1^0) = 200$ GeV, 1100 GeV] 1.3 1.9 Large λ'_{112} 1804.03568	
	Multiple	E_T^{miss} 36.1		$\tilde{g}\tilde{g}$ [$\lambda'_{112} = 2e-4$] 1.05 2.0 $m(\tilde{\chi}_1^0) = 200$ GeV, bino-like ATLAS-CONF-2018-003	
	Multiple	E_T^{miss} 36.1		$\tilde{g}\tilde{g}$ [$\lambda'_{112} = 2e-4$] 1.05 $m(\tilde{\chi}_1^0) = 200$ GeV, bino-like ATLAS-CONF-2018-003	
	2 jets + 2 b	E_T^{miss} 36.7		$\tilde{g}\tilde{g}$ [$\lambda'_{112} = 2e-4$] 1.05 $m(\tilde{\chi}_1^0) = 200$ GeV, bino-like ATLAS-CONF-2018-003	
	2 e, μ	2 b E_T^{miss} 36.1		$\tilde{g}\tilde{g}$ [$\lambda'_{112} = 2e-4$] 1.05 $m(\tilde{\chi}_1^0) = 200$ GeV, bino-like ATLAS-CONF-2018-003	
	1 μ	DV E_T^{miss} 36.1		$\tilde{g}\tilde{g}$ [$\lambda'_{112} = 2e-4$] 1.05 $m(\tilde{\chi}_1^0) = 200$ GeV, bino-like ATLAS-CONF-2018-003	
				$BR(\tilde{\tau}_1 \rightarrow b\mu) > 20\%$ 1710.07171	
				$BR(\tilde{\tau}_1 \rightarrow q\mu) = 100\%$, $\cos\theta = 1$ 1710.05544	
				ATLAS-CONF-2019-006	

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Inclusive Searches

$\tilde{q}\tilde{q}, \tilde{q} \rightarrow q\tilde{\chi}_1^0$
 $\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$
 $\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(\ell\ell)\tilde{\chi}_1^0$
 $\tilde{g}\tilde{g}, \tilde{g} \rightarrow qqWZ\tilde{\chi}_1^0$

Mass scale [TeV]

Mass scale [TeV]

Supersymmetry is broken

$$m_{\text{fermion}} \neq m_{\text{boson}}$$

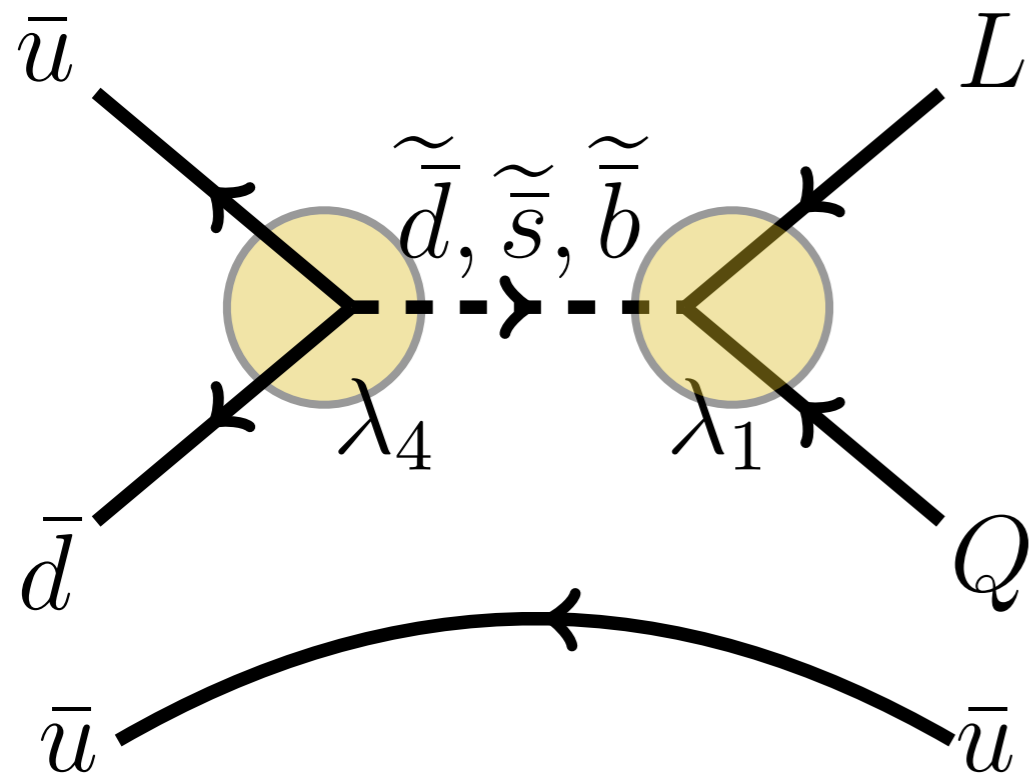
SUSY is not a good symmetry of nature.
c.f. electroweak symmetry

Can it solve the Hierarchy Problem?

Soft SUSY breaking

$$\Delta m_H^2 = m_{\text{soft}}^2 \left[\frac{\lambda}{16\pi^2} \ln(\Lambda_{\text{UV}}/m_{\text{soft}}) + \dots \right]$$

proton decay



$$P_R = (-)^{3(B-L)+2s}$$

$$P_R[\text{ordinary matter}] = +$$

$$P_R[\text{superpartner}] = -$$

$$W^{(\text{good})} = y_u^{ij} Q^i H_u \bar{U}^j + y_d^{ij} Q^i H_d \bar{D} + y_e^{ij} L^i H_d \bar{E}^j + \mu H_u H_d$$

$$W^{(\text{bad})} = \lambda_1^{ijk} Q^i L^j \bar{D}^k + \lambda_2^{ijk} L^i L^j \bar{E}^k + \lambda_3^i L^i H_u + \lambda_4^{ijk} \bar{D}^i \bar{D}^j \bar{U}^k$$

Added bonus:
lightest superpartner is stable.



Image: *We Have No Idea*, Whiteson & Cham

electroweak symmetry breaking

The Higgs potential is a challenge in SUSY.

$$V_H = \frac{1}{8}(g^2 + g'^2) (|H_u^0|^2 - |H_d^0|^2)^2 + \sum_{i=u,d} (|\mu|^2 + m_{H_i}^2) |H_i^0|^2 - 2B_\mu \text{Re}(H_u^0 H_d^0)$$

Two Higgses, both get vevs.

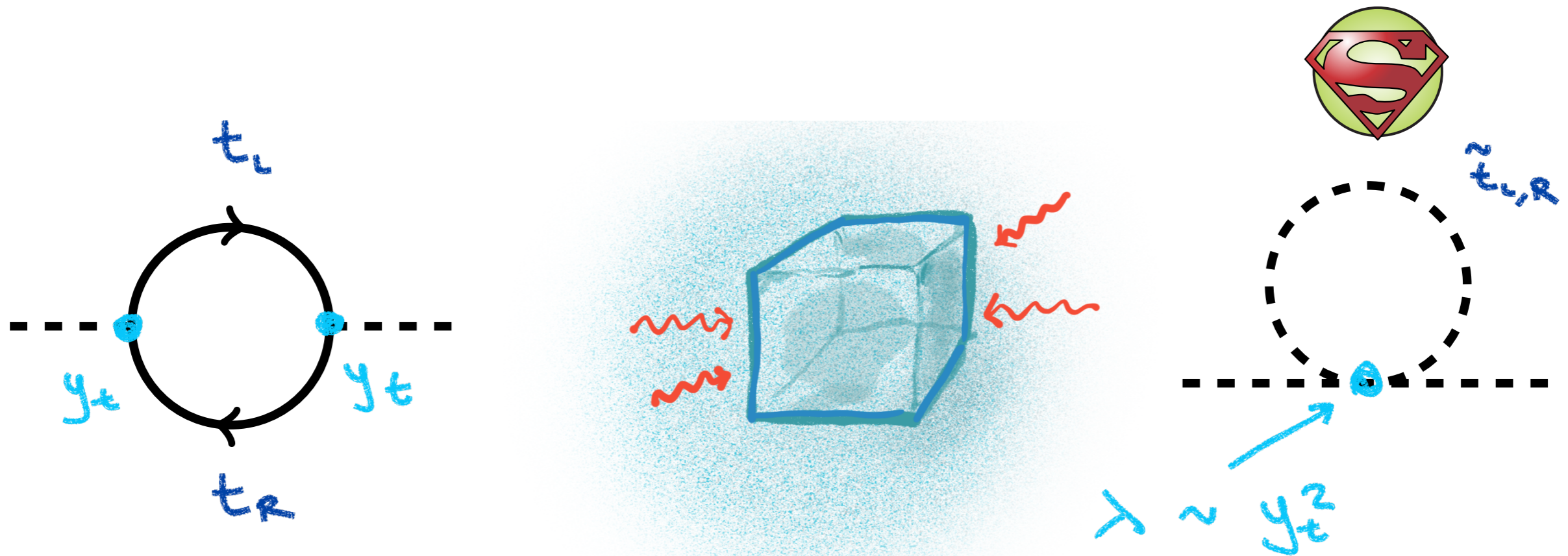
Only one quartic direction (other is D-flat)

Relies on apparent conspiracy between supersymmetric and SUSY-breaking terms.

Supersymmetry

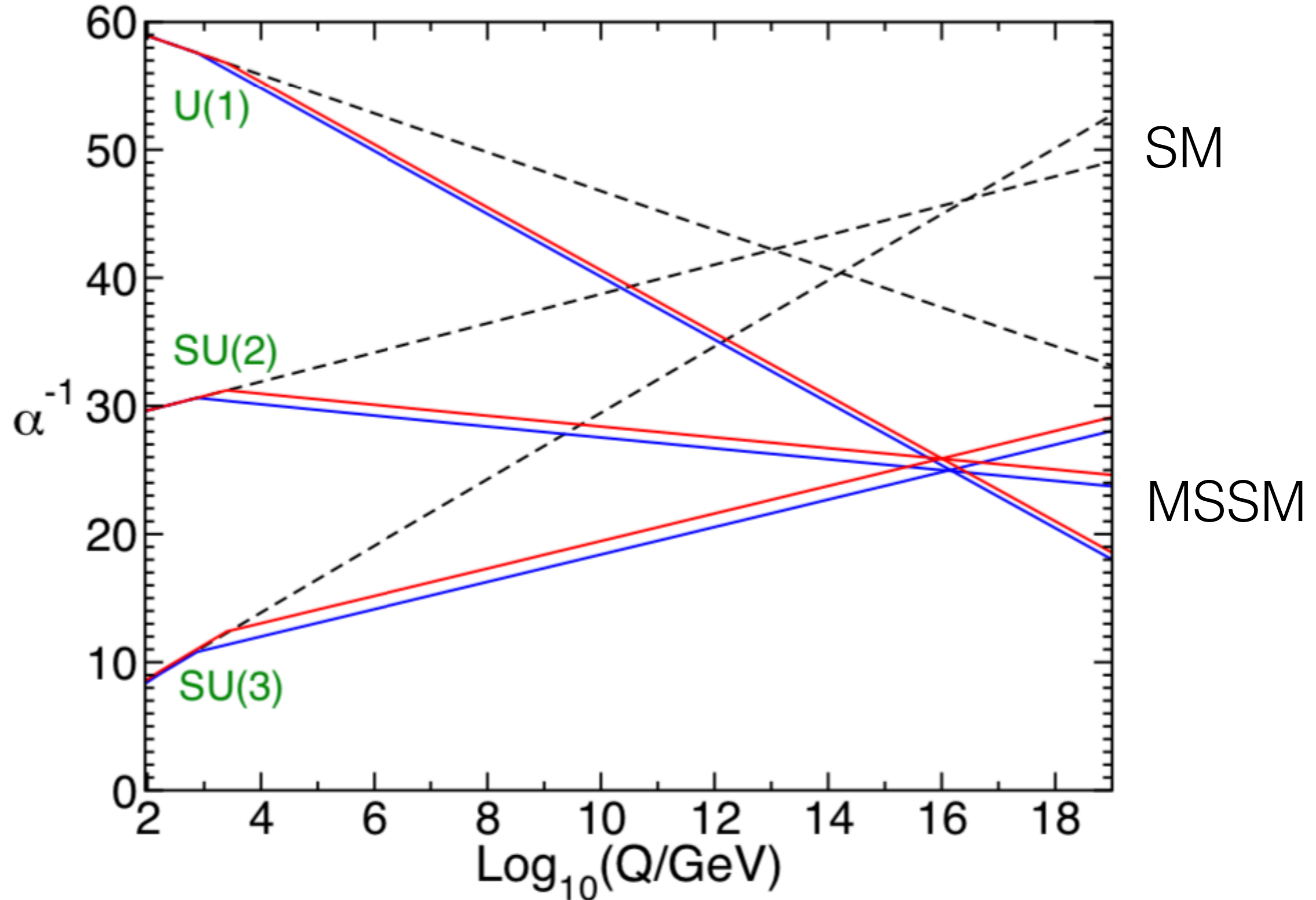
why we really like(d) it

Solves Hierarchy Problem



Symmetry principle that protects Higgs mass from quantum corrections in the UV.

gauge unification for free



S. Martin hep-ph/9709356 Fig 6.8

flip.tanedo@ucr.edu

TRISEP SUMMER SCHOOL 2019

62/70

nice dark matter candidate

Requirements: stable, uncharged.

What's a good dark matter candidate?

caveat: there's a lot we're sweeping under the rug
... R-parity and all that (for now)

nice dark matter candidate

Requirements: stable, uncharged.

What's a good dark matter candidate?

caveat: there's a lot we're sweeping under the rug
... R-parity and all that (for now)

the \tilde{N}_1 ← LIGHTEST NEUTRALINO
(COMBINATION OF $\tilde{H}_u^0, \tilde{H}_d^0, \tilde{B}, \tilde{W}_3$)

sometimes: **gravitino** is lightest such state

The MSSM is...

simple enough to understand **how and why** it works (at least in the SUSY limit)

complex enough to generate many kinds of phenomenological signatures

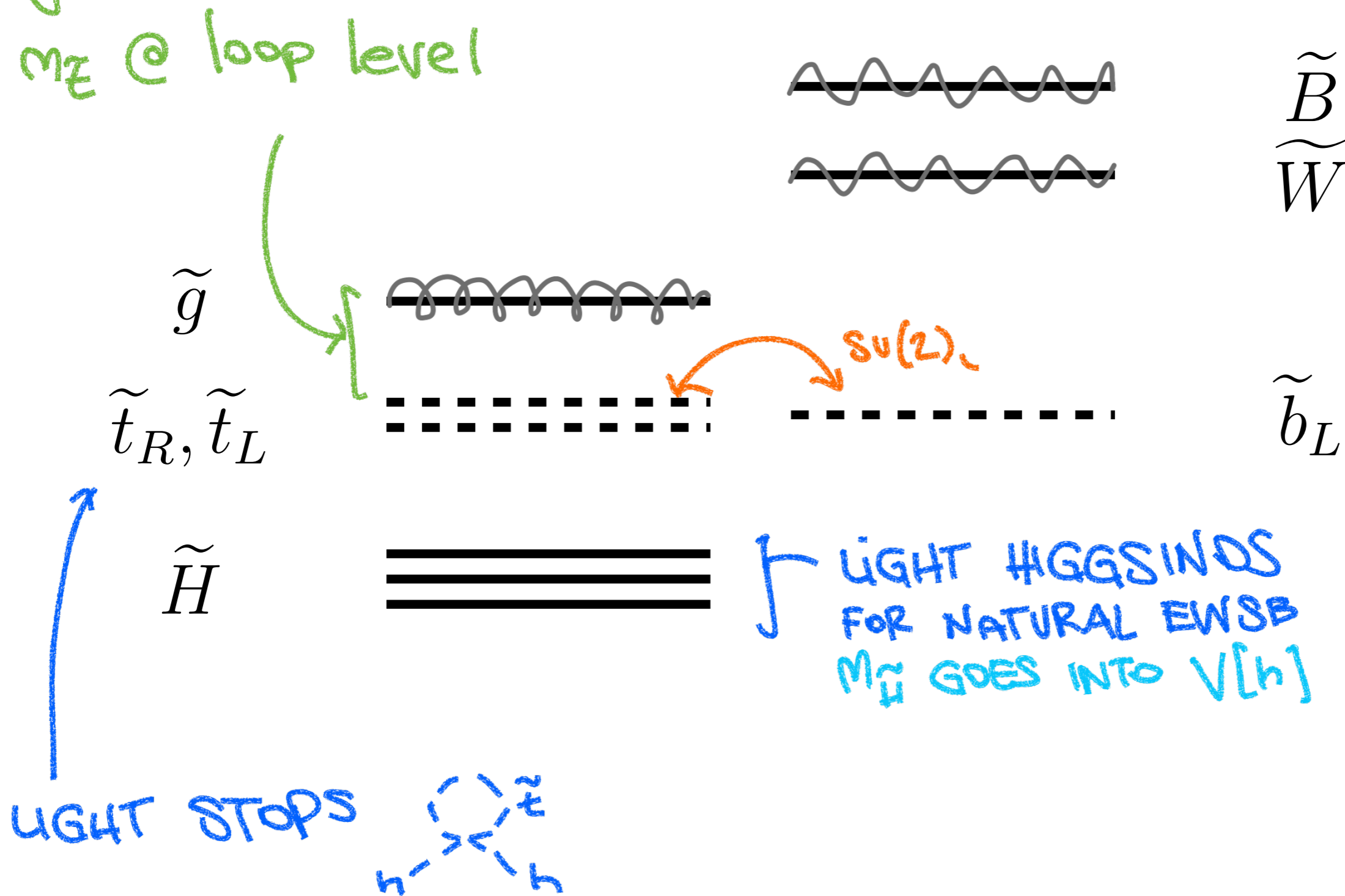
MSSM

basic pheno

natural SUSY spectrum

\tilde{g} contribute to m_Z @ loop level

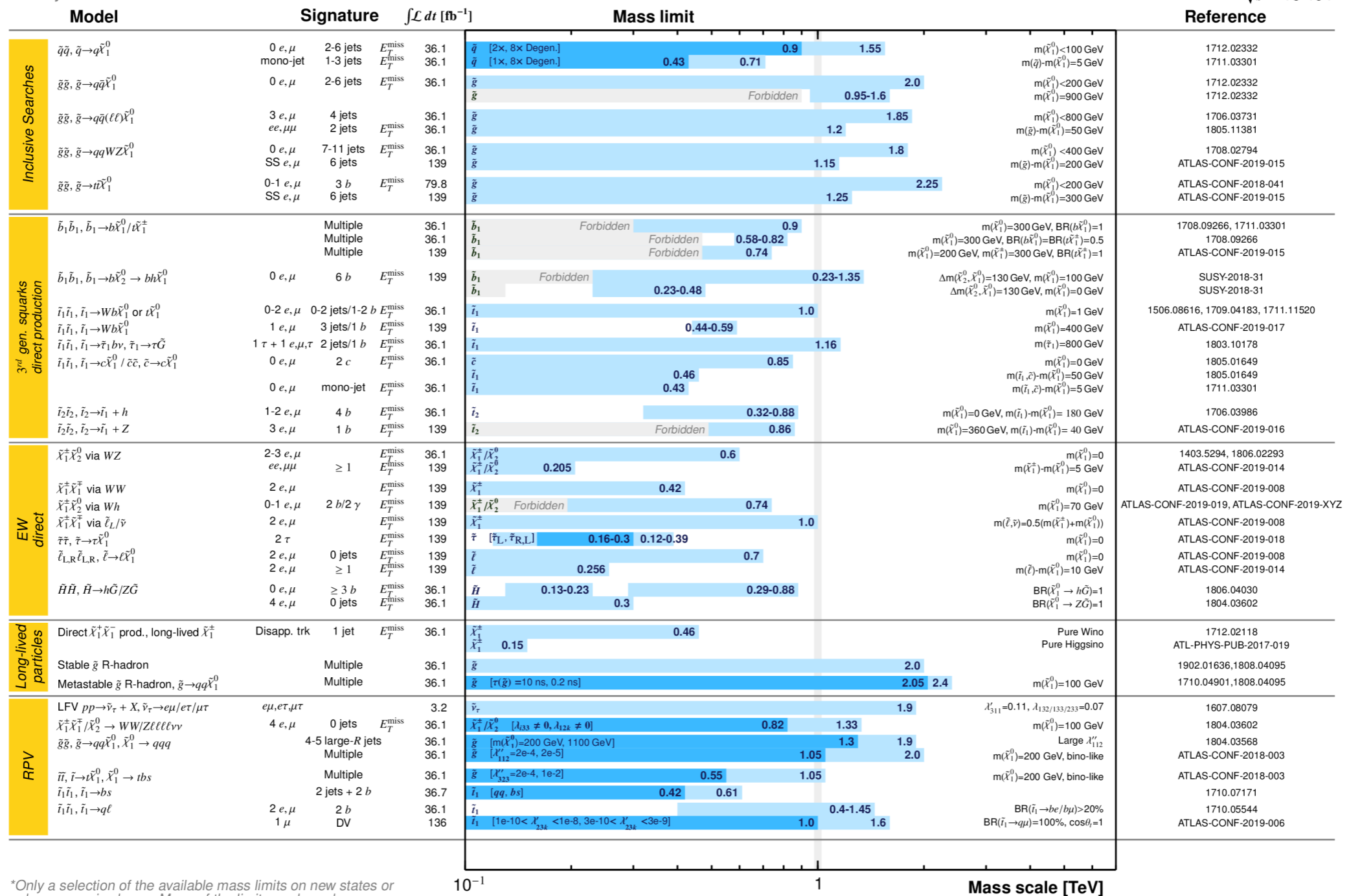
EW-inos
not too heavy
(for GUT)



there are no sparticles

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Model	Signature	$\int \mathcal{L} dt [\text{fb}^{-1}]$	Mass limit	Reference	
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		mono-jet	1-3 jets E_T^{miss} 36.1	\tilde{q} [1x, 8x Degener.] 0.43 0.71	$m(\tilde{q}) - m(\tilde{\chi}_1^0) = 5 \text{ GeV}$ 1711.03301
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}\tilde{\chi}_1^0$	0 e, μ	2-6 jets E_T^{miss} 36.1	Forbidden 0.95-1.6 2.0	$m(\tilde{\chi}_1^0) < 200 \text{ GeV}$ 1712.02332
					$m(\tilde{\chi}_1^0) = 900 \text{ GeV}$ 1712.02332
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\tilde{q}(\ell\ell)\tilde{\chi}_1^0$	3 e, μ	4 jets E_T^{miss} 36.1		$m(\tilde{\chi}_1^0) < 800 \text{ GeV}$ 1706.03731
		$ee, \mu\mu$	2 jets E_T^{miss} 36.1		$m(\tilde{g}) - m(\tilde{\chi}_1^0) = 50 \text{ GeV}$ 1805.11381
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qqWZ\tilde{\chi}_1^0$	0 e, μ	7-11 jets E_T^{miss} 36.1		$m(\tilde{\chi}_1^0) < 400 \text{ GeV}$ 1708.02794
		SS e, μ	6 jets E_T^{miss} 36.1		$m(\tilde{g}) - m(\tilde{\chi}_1^0) = 200 \text{ GeV}$ ATLAS-CONF-2019-015
		0-1 e, μ	3 b E_T^{miss} 73.8		$m(\tilde{\chi}_1^0) < 200 \text{ GeV}$ ATLAS-CONF-2018-041
		SS e, μ	6 jets E_T^{miss} 139		$m(\tilde{g}) - m(\tilde{\chi}_1^0) = 300 \text{ GeV}$ ATLAS-CONF-2019-015
Model	Multiple		36.1	\tilde{b}_1 0.9	$m(\tilde{\chi}_1^0) = 300 \text{ GeV}, \text{BR}(\tilde{b}\tilde{\chi}_1^0) = 1$ 1708.09266, 1711.03301
	Multiple		36.1	\tilde{b}_1 0.58-0.82	$m(\tilde{\chi}_1^0) = 300 \text{ GeV}, \text{BR}(\tilde{b}\tilde{\chi}_1^0) = \text{BR}(\tilde{t}\tilde{\chi}_1^0) = 0.5$ 1708.09266
	Multiple		139	\tilde{b}_1 0.74	$m(\tilde{\chi}_1^0) = 200 \text{ GeV}, m(\tilde{\chi}_2^0) = 300 \text{ GeV}, \text{BR}(\tilde{\chi}_2^0) = 1$ ATLAS-CONF-2019-015
	6 b	E_T^{miss}	139		$\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 130 \text{ GeV}, m(\tilde{\chi}_1^0) = 100 \text{ GeV}$ SUSY-2018-31
	1-2 b	E_T^{miss}	36.1		$\Delta m(\tilde{\chi}_2^0, \tilde{\chi}_1^0) = 130 \text{ GeV}, m(\tilde{\chi}_1^0) = 0 \text{ GeV}$ SUSY-2018-31
	b	E_T^{miss}	139		$m(\tilde{\chi}_1^0) = 1 \text{ GeV}$ 1506.08616, 1709.04183, 1711.11520
	b	E_T^{miss}	36.1		$m(\tilde{\chi}_1^0) = 400 \text{ GeV}$ ATLAS-CONF-2019-017
	b	E_T^{miss}	36.1		$m(\tilde{\tau}_1) = 800 \text{ GeV}$ 1803.10178
	b	E_T^{miss}	36.1		$m(\tilde{\chi}_1^0) = 0 \text{ GeV}$ 1805.01649
		E_T^{miss}	139		$m(\tilde{\tau}_1, \tilde{\nu}) - m(\tilde{\chi}_1^0) = 50 \text{ GeV}$ 1805.01649
	E_T^{miss}	139		$m(\tilde{\tau}_1, \tilde{\nu}) - m(\tilde{\chi}_1^0) = 5 \text{ GeV}$ 1711.03301	
	E_T^{miss}	36.1		$m(\tilde{\chi}_1^0) = 0 \text{ GeV}, m(\tilde{\tau}_1) - m(\tilde{\chi}_1^0) = 180 \text{ GeV}$ 1706.03986	
	E_T^{miss}	139		$m(\tilde{\chi}_1^0) = 360 \text{ GeV}, m(\tilde{\tau}_1) - m(\tilde{\chi}_1^0) = 40 \text{ GeV}$ ATLAS-CONF-2019-016	
	E_T^{miss}	36.1		$m(\tilde{\chi}_1^0) = 0$ 1403.5294, 1806.02293	
	E_T^{miss}	139		$m(\tilde{\chi}_1^0) - m(\tilde{\chi}_2^0) = 5 \text{ GeV}$ ATLAS-CONF-2019-014	
	E_T^{miss}	139		$m(\tilde{\chi}_1^0) = 0$ ATLAS-CONF-2019-008	
	E_T^{miss}	139		$m(\tilde{\chi}_1^0) = 70 \text{ GeV}$ ATLAS-CONF-2019-019, ATLAS-CONF-2019-XYZ	
	E_T^{miss}	139		$m(\tilde{\ell}, \tilde{\nu}) = 0.5(m(\tilde{\chi}_1^0) + m(\tilde{\chi}_2^0))$ ATLAS-CONF-2019-008	
	E_T^{miss}	139		$m(\tilde{\chi}_1^0) = 0$ ATLAS-CONF-2019-018	
	E_T^{miss}	139		$m(\tilde{\chi}_1^0) = 0$ ATLAS-CONF-2019-008	
	E_T^{miss}	139		$m(\tilde{\ell}) - m(\tilde{\chi}_1^0) = 10 \text{ GeV}$ ATLAS-CONF-2019-014	
	3 b	E_T^{miss}	36.1	\tilde{H} 0.13-0.23 0.29-0.88	$\text{BR}(\tilde{\chi}_1^0 \rightarrow h\tilde{G}) = 1$ 1806.04030
	0 jets	E_T^{miss}	36.1	\tilde{H} 0.3	$\text{BR}(\tilde{\chi}_1^0 \rightarrow Z\tilde{G}) = 1$ 1804.03602
	1 jet	E_T^{miss}	36.1	$\tilde{\chi}_1^\pm$ 0.46	Pure Wino 1712.02118
		E_T^{miss}	36.1	$\tilde{\chi}_1^\pm$ 0.15	Pure Higgsino ATL-PHYS-PUB-2017-019
	Multiple		36.1	\tilde{g} 2.0	1902.01636, 1808.04095
	Multiple		36.1	\tilde{g} [$\tau(\tilde{g}) = 10 \text{ ns}, 0.2 \text{ ns}$] 2.05 2.4	$m(\tilde{\chi}_1^0) = 100 \text{ GeV}$ 1710.04901, 1808.04095
RPV	$\tilde{\chi}_1^\pm \tilde{\chi}_1^0 \rightarrow WWZ\ell\ell\nu\nu$	$e\mu, e\tau, \mu\tau$	3.2	$\tilde{\nu}_\tau$ 1.9	$\lambda'_{311} = 0.11, \lambda'_{132/133/233} = 0.07$ 1607.08079
	$\tilde{g}\tilde{g}, \tilde{g} \rightarrow qq\tilde{\chi}_1^0, \tilde{\chi}_1^0 \rightarrow qq\tilde{\chi}_1^0$	4 e, μ	0 jets E_T^{miss} 36.1	$\tilde{\chi}_1^\pm \tilde{\chi}_2^0$ [$\lambda_{333} \neq 0, \lambda_{124} \neq 0$] 0.82 1.33	$m(\tilde{\chi}_1^0) = 100 \text{ GeV}$ 1804.03602
			36.1	$\tilde{g}\tilde{g}$ [$m(\tilde{\chi}_1^0) = 200 \text{ GeV}, 1100 \text{ GeV}$] 1.3 1.9	Large λ'_{112} 1804.03568
			36.1	$\tilde{g}\tilde{g}$ [$\lambda'_{112} = 2e-4$] 1.05 2.0	$m(\tilde{\chi}_1^0) = 200 \text{ GeV}, \text{bino-like}$ ATLAS-CONF-2018-003
			36.1	$\tilde{g}\tilde{g}$ [$\lambda'_{112} \rightarrow q\mu$] = 100%, 0.5 1.05	$m(\tilde{\chi}_1^0) = 200 \text{ GeV}, \text{bino-like}$ ATLAS-CONF-2018-003
			36.7	$\tilde{g}\tilde{g}$ [$\lambda'_{112} \rightarrow q\mu$] = 100%, 0.5 1.05	$m(\tilde{\chi}_1^0) = 200 \text{ GeV}, \text{bino-like}$ ATLAS-CONF-2018-003
			36.7	$\tilde{g}\tilde{g}$ [$\lambda'_{112} \rightarrow q\mu$] = 100%, 0.5 1.05	$m(\tilde{\chi}_1^0) = 200 \text{ GeV}, \text{bino-like}$ ATLAS-CONF-2018-003
		2 e, μ	2 b E_T^{miss} 36.1		$\text{BR}(\tilde{\tau}_1 \rightarrow b\mu) > 20\%$ 1710.07171
		1 μ	DV E_T^{miss} 36.1		$\text{BR}(\tilde{\tau}_1 \rightarrow q\mu) = 100\%, \cos\theta = 1$ 1710.05544
					ATLAS-CONF-2019-006

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Mass scale [TeV]

Mass scale [TeV]

there are no sparticles

many jets

Model

$$\tilde{g}\tilde{g}, \tilde{g} \rightarrow q\bar{q}\tilde{\chi}_1^0$$

Signature

$$\int \mathcal{L} dt [\text{fb}^{-1}]$$

0 e, μ

2-6 jets

E_T^{miss}

36.1

no leptons

missing E.

250	$m(\tilde{\chi}_1^0) < 200 \text{ GeV}$		2.0
150	$m(\tilde{\chi}_1^0) = 900 \text{ GeV}$	Forbidden	0.95-1.6

