



Canada's National Laboratory for
Particle and Nuclear Physics

Science Goals and Priorities

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ARIEL Principal Scientist and TRIUMF Research Scientist

January 11th 2017

ARIEL Town Hall Meeting
Vancouver, BC, Canada

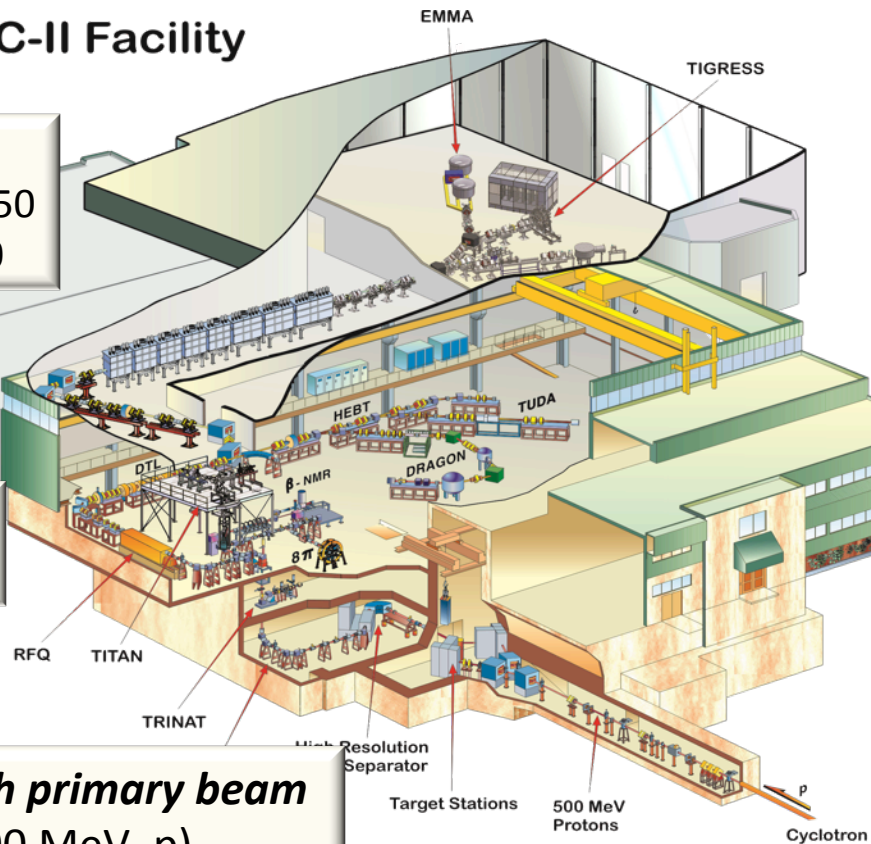
ISAC-I and ISAC-II Facility

ISAC II:

- 10 AMeV for $A < 150$
- 16 AMeV for $A < 30$

ISAC I:

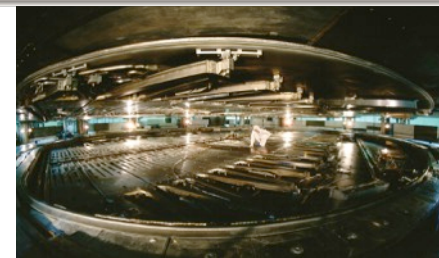
60 keV & 1.7 AMeV



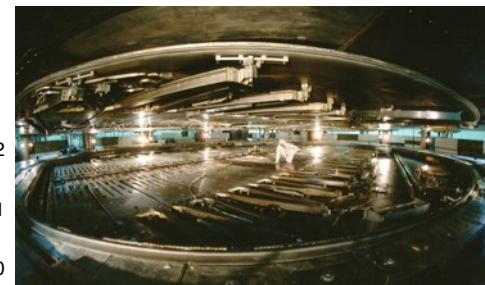
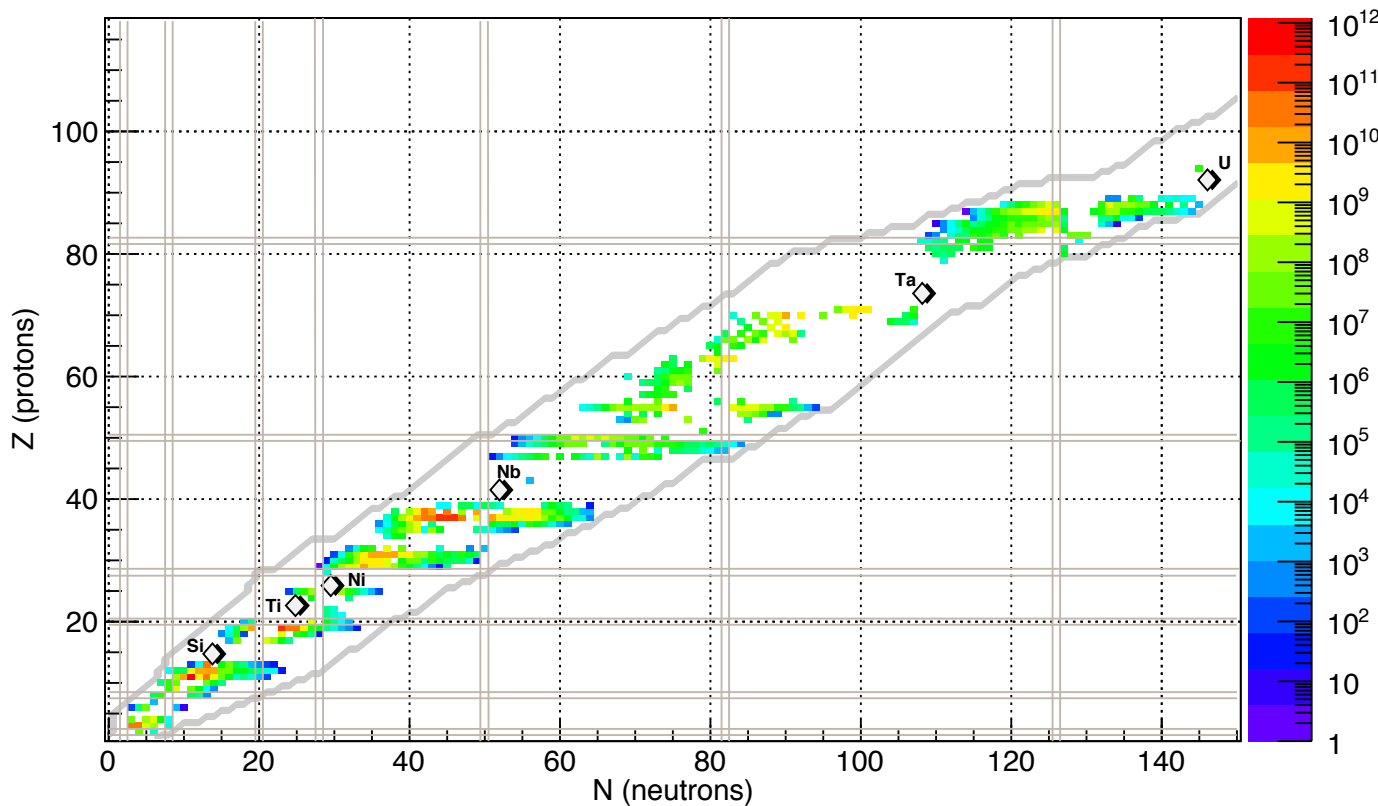
Programs in

- Nuclear Structure & Dynamics
- Nuclear Astrophysics
- Electroweak Interaction Studies
- Material Science
- 18 permanent experiments

ISOL facility with *high primary beam intensity* (100 μ A, 500 MeV, p)
Delivering RIBs since 1999.



Isotopes delivered at ISAC (P.Kunz, Updated June 2016)

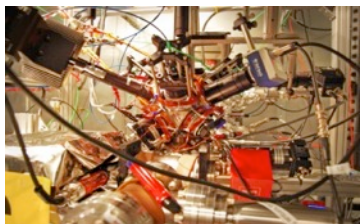


Target materials:
SiC, TiC, NiO, Nb,
ZrC, Ta, U

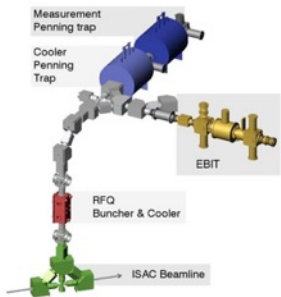
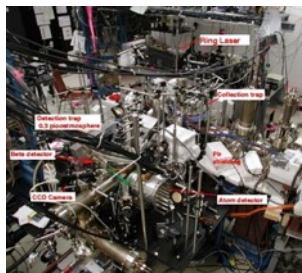
Ion sources:
Surface, FEBIAD,
IG-LIS

Low energy RIBs
< 60 keV

FRANCIUM MOT
(PNC, anapole moment)



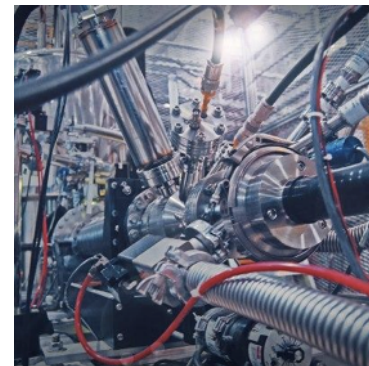
TRINAT
Neutral Atom Trap
($\beta\nu$ -neutrino correlations)



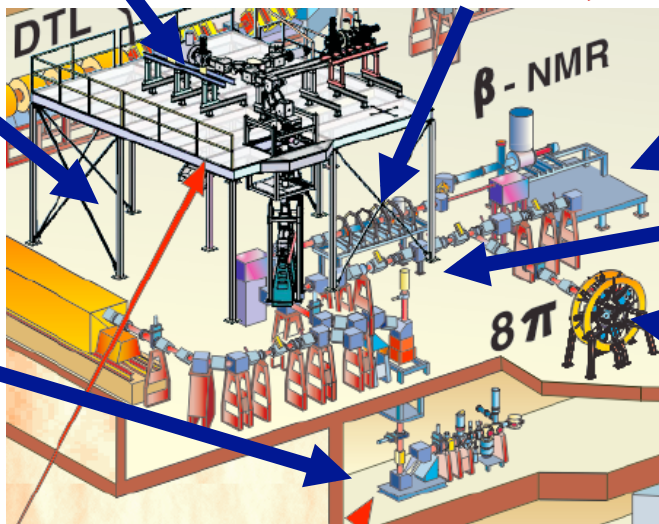
TITAN
Penning Traps
(masses,
in-trap decay)



Polarizer beamline
Laser spectroscopy, MTV
CPT test, betaNMR

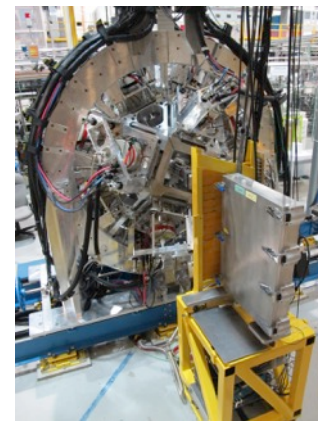


Beta-NMR
Material science



GPS
Beta counter
(superalloyed
decays)

GRIFFIN
Gamma & Electron
spectrometer
(decay spectroscopy,
superalloyed decays)

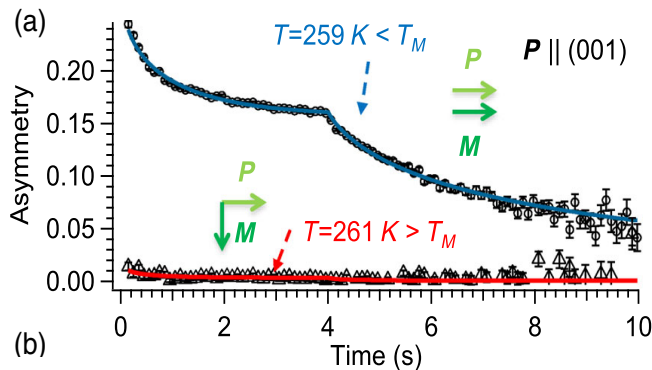


PRL **116**, 106103 (2016)

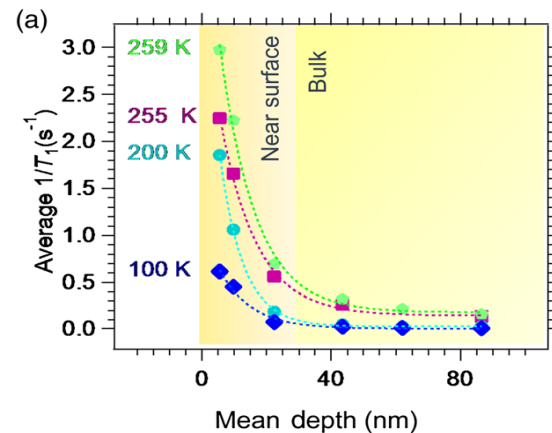
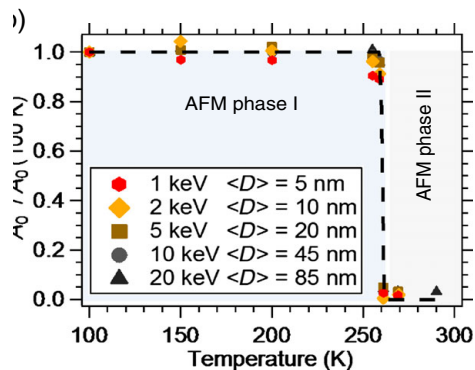
PHYSICAL REVIEW LETTERS

week ending
11 March 2016

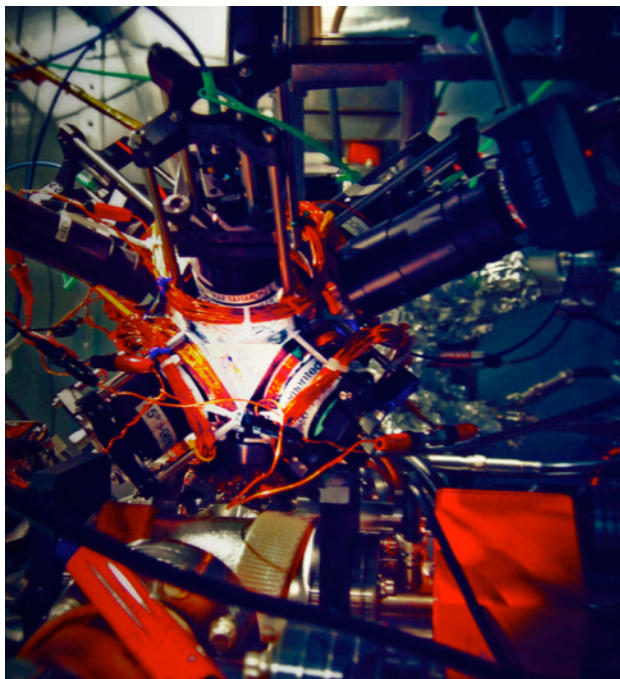
D. L. Cortie, T. Buck, M. H. Dehn, V. L. Karner, R. F. Kiefl, C. D. P. Levy, R. M. L. McFadden, G. D. Morris, I. McKenzie, M. R. Pearson, X. L. Wang, and W. A. MacFarlane



β-NMR was used to study the depth dependence of the Morin spin reorientation transition in α-Fe₂O₃ (hematite). The surface-localized dynamics decay towards the bulk with a characteristic length of 11 nm, indicating the presence of soft surface magnons.



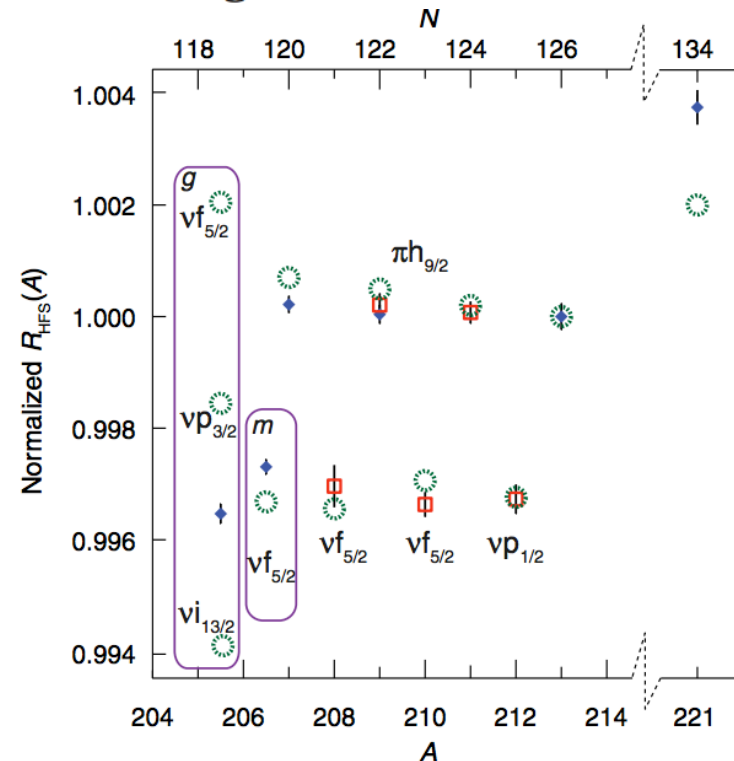
Hyperfine Anomalies in Fr: Boundaries of the Spherical Single Particle Model



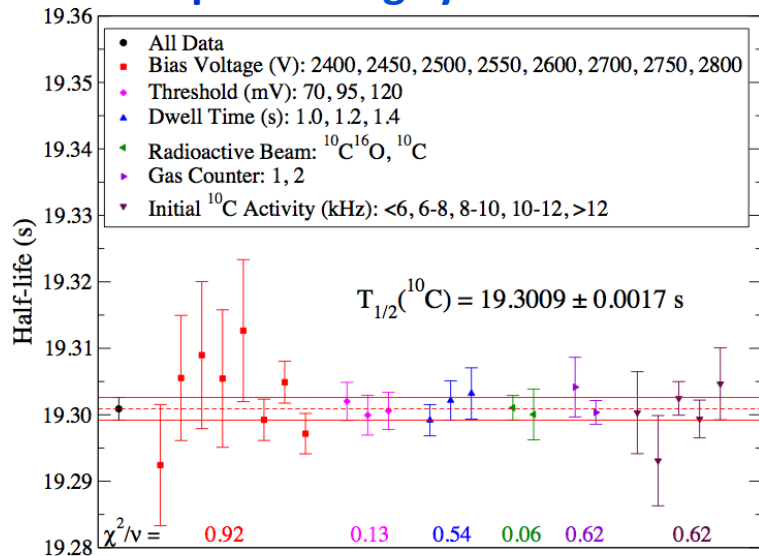
Goal is the measure parity-violating anapole moments.

Essential to understand all nuclear contributions, such as finite magnetic and charge distributions.

Hyperfine splitting of $7P_{1/2}$ state in Fr isotopes.

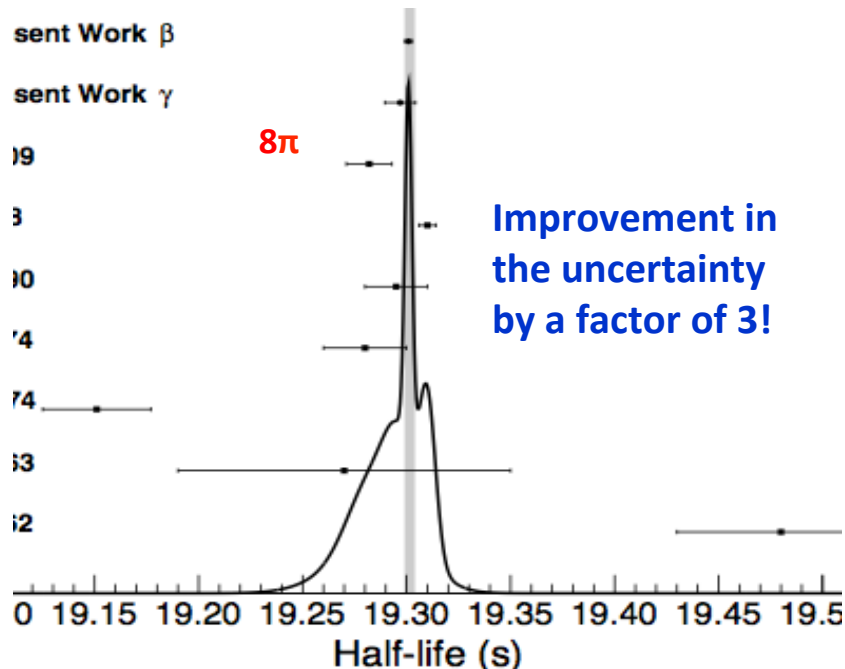


β -counting systematics



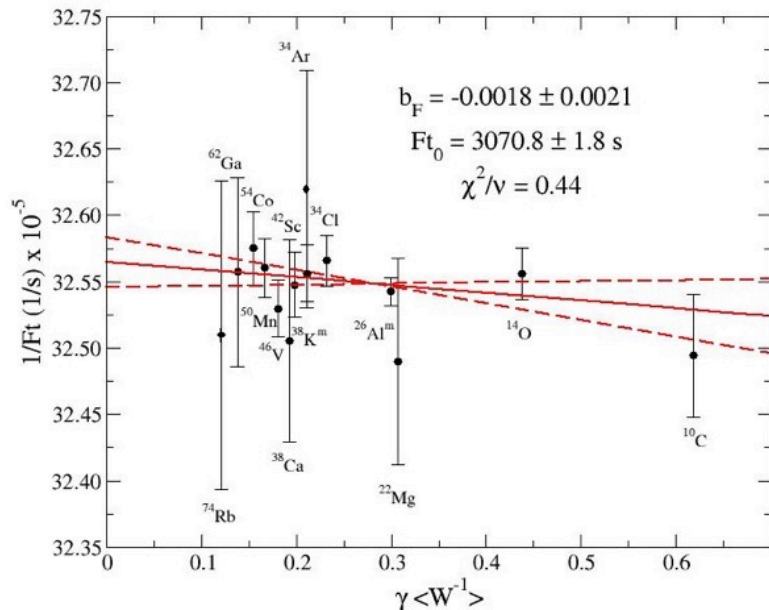
$$T_{1/2,\beta}(^{10}\text{C}) = 19.3009 \pm 0.0017 \text{ s}$$

The most precise (0.009%)
superalloyed $T_{1/2}$ reported to date!



M.R. Dunlop et al., Phys. Rev. Lett. 116, 172501 (2016)

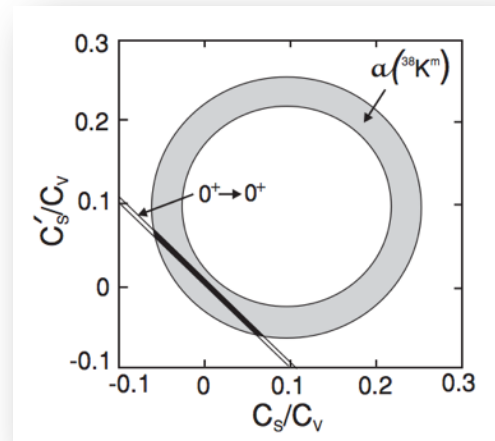
$$C_S/C_V = -b_F/2 = +0.0009 \pm 0.0011 \quad (\text{for } C_S = C_S')$$



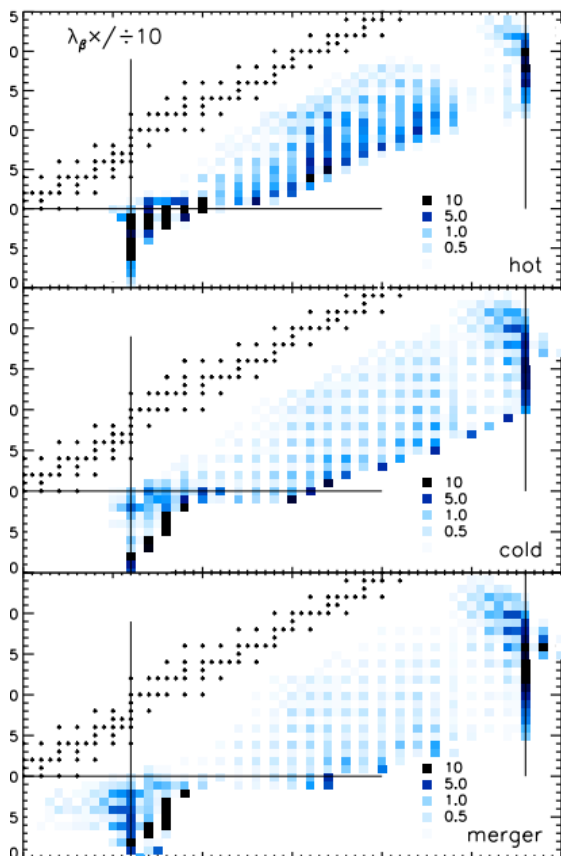
M.R. Dunlop et al., Phys. Rev. Lett. 116, 172501 (2016)

For a general C_S and C_S' , the β - ν angular correlation coefficient from ^{38m}K decay provides an independent constraint:

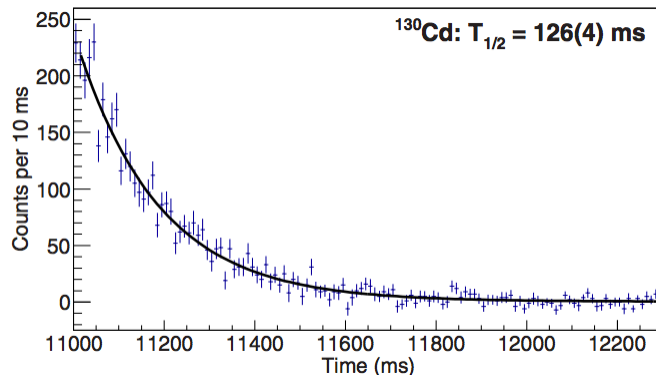
$$-0.065 \leq C_S/C_V, C_S'/C_V \leq 0.065$$



A. Gorelov et al., Phys. Rev. Lett. 94, 142501 (2005)

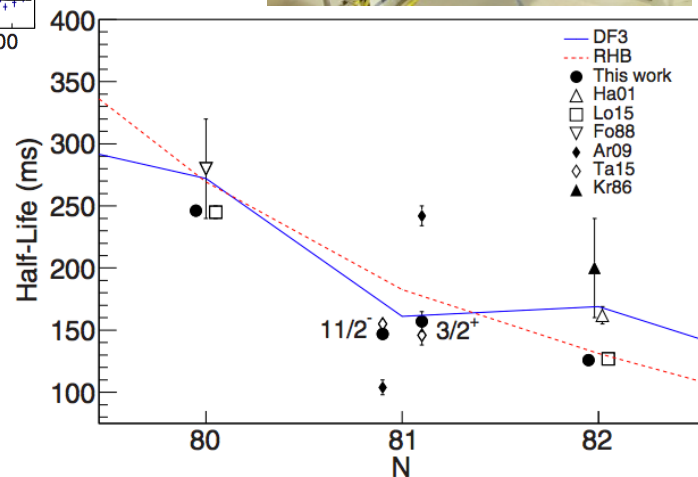

 PHYSICAL REVIEW C **93**, 062801(R) (2016)

Half-lives of neutron-rich $^{128-130}\text{Cd}$



Measurement of decay half lives and properties are important for astrophysical r-process calculations

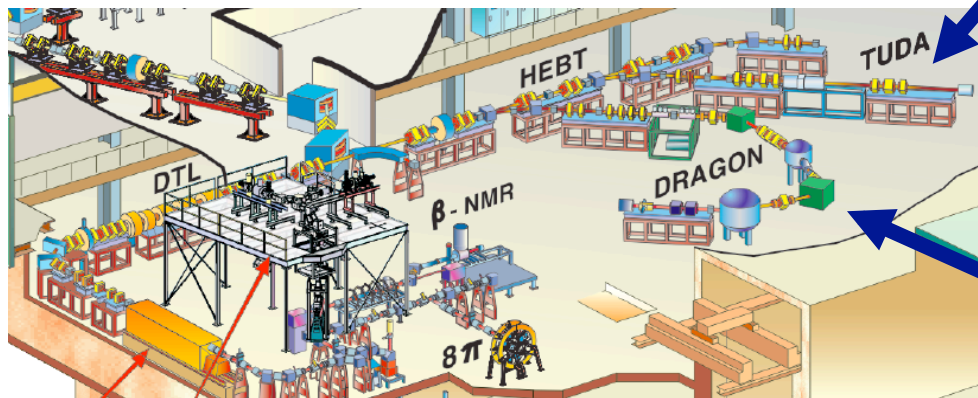
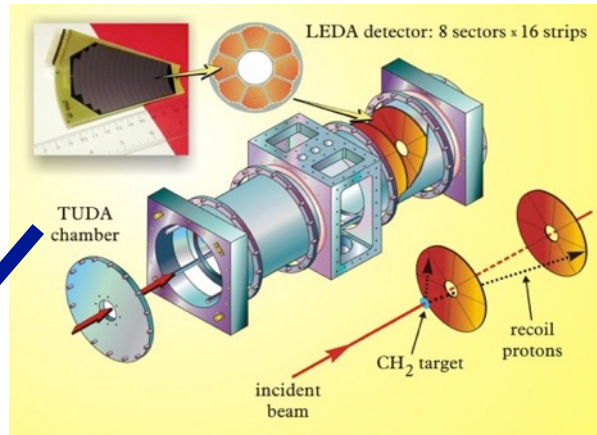
RAPID COMMUNICATIONS


 G. Lorusso et al. PRL **114** 192501 (2015)

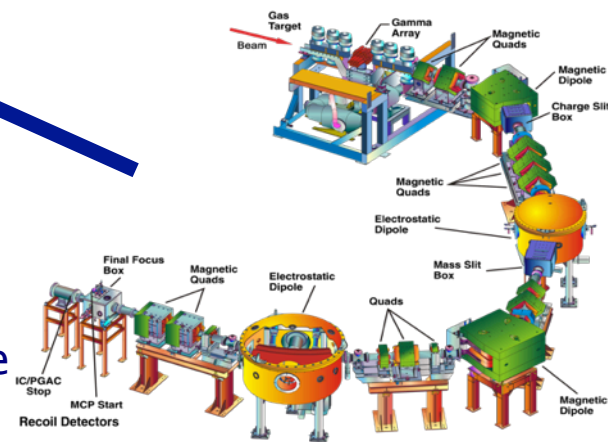
 M. Mumpower et al., Prog.Part.Nucl.Phys. **86**, 86 (2016)

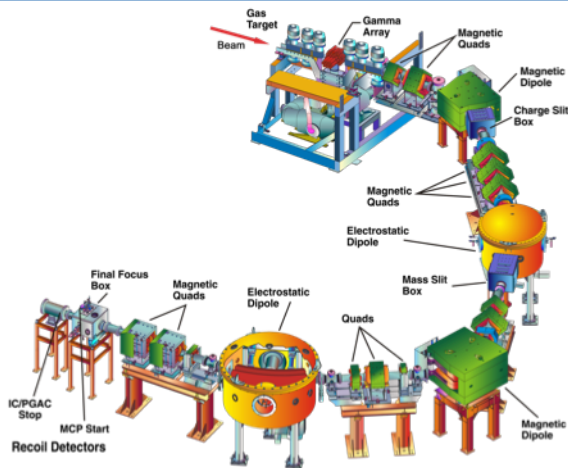
Medium energy RIBs
 ~ 0.15 - 1.7 AMeV

TUDA
 Astrophysical charged particle reactions



DRAGON
 Astrophysical capture reactions

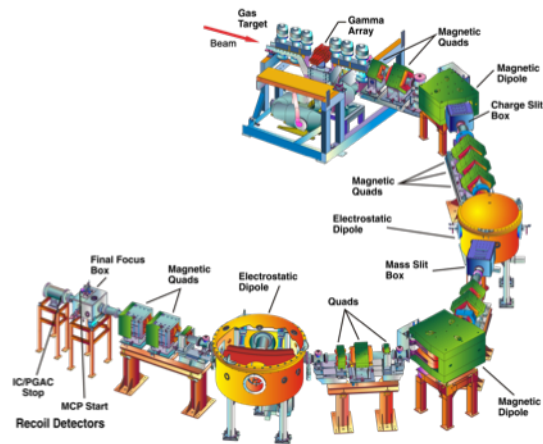




7 RIB
10 Stable beam

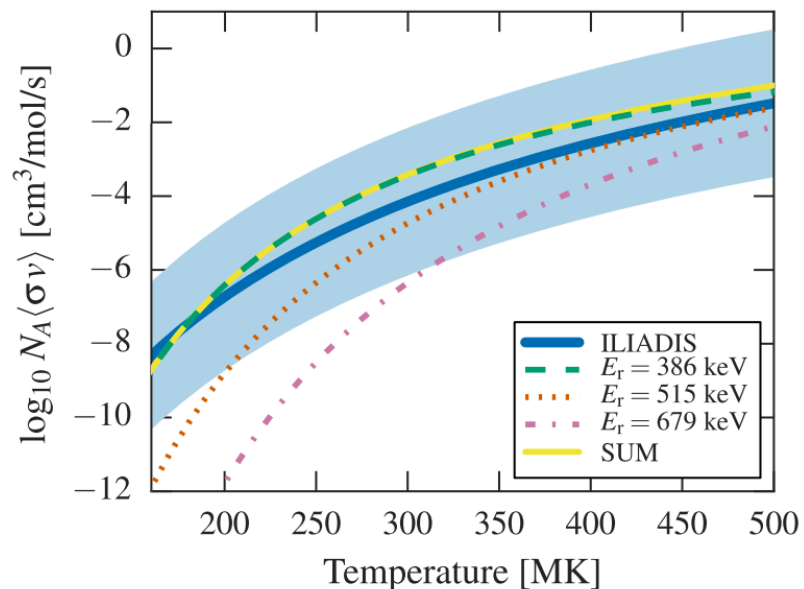
| Reaction | Motivation | Intensity (s ⁻¹) | Purity (beam:cont.) |
|---|--|------------------------------|---------------------|
| $^{21}\text{Na}(p,\gamma)^{22}\text{Mg}$ | 1.275 MeV line emission in ONe novae | 5×10^9 | 100% |
| $^{12}\text{C}(\alpha,\gamma)^{16}\text{O}$ | Helium burning in red giants | 6×10^{11} | |
| $^{26}\text{gAl}(p,\gamma)^{27}\text{Si}$ | Nova contribution to galactic ^{26}Al | 3×10^9 | 30,000:1 |
| $^{12}\text{C}(^{12}\text{C},\gamma)^{24}\text{Mg}$ | Nuclear cluster models | 3×10^{11} | |
| $^{40}\text{Ca}(\alpha,\gamma)^{44}\text{Ti}$ | Production of ^{44}Ti in SNIa | 3×10^{11} | 10,000:1 – 200:1 |
| $^{23}\text{Mg}(p,\gamma)^{24}\text{Al}$ | 1.275 MeV line emission in ONe novae | 5×10^7 | 1:20 – 1:1,000 |
| $^{17}\text{O}(\alpha,\gamma)^{21}\text{Ne}$ | Neutron poison in massive stars | 1×10^{12} | |
| $^{18}\text{F}(p,\gamma)^{19}\text{Ne}$ | 511 keV line emission in ONe novae | 2×10^6 | 100:1 |
| $^{33}\text{S}(p,\gamma)^{34}\text{Cl}$ | S isotopic ratios in nova grains | 1×10^{10} | |
| $^{16}\text{O}(\alpha,\gamma)^{20}\text{Ne}$ | Stellar helium burning | 1×10^{12} | |
| $^{17}\text{O}(p,\gamma)^{18}\text{F}$ | Explosive hydrogen burning in novae | 1×10^{12} | |
| $^3\text{He}(\alpha,\gamma)^7\text{Be}$ | Solar neutrino spectrum | 5×10^{11} | |
| $^{58}\text{Ni}(p,\gamma)^{59}\text{Cu}$ | High mass tests (p-process, XRB) | 6×10^9 | |
| $^{26}\text{mAl}(p,\gamma)^{27}\text{Si}$ | SNIa contribution to galactic ^{26}Al | 2×10^5 | 1:10,000 |
| $^{38}\text{K}(p,\gamma)^{39}\text{Ca}$ | Ca/K/Ar production in novae | 2×10^7 | 1:1 |
| $^{19}\text{Ne}(p,\gamma)^{20}\text{Na}$ | ^{19}F abundance in nova ejecta | 2×10^7 | 1:1 to 4:1 |
| $^{22}\text{Ne}(p,\gamma)^{23}\text{Na}$ | NeNa cycle; explosive H burning in classical novae | 2×10^{12} | |

Direct Measurement of the Astrophysical $^{38}\text{K}(p,\gamma)^{39}\text{Ca}$ Reaction and Its Influence on the Production of Nuclides toward the End Point of Nova Nucleosynthesis



First experimental measurement of this reaction rate.

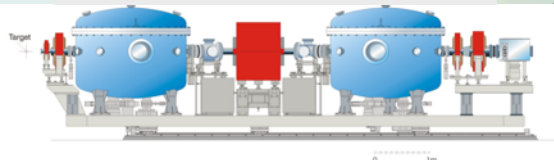
Constrains the rate of $^{38}\text{K}(p,\gamma)^{39}\text{Ca}$ in ONE nova and significantly reduces uncertainty in ^{38}Ar and ^{40}Ca abundances.



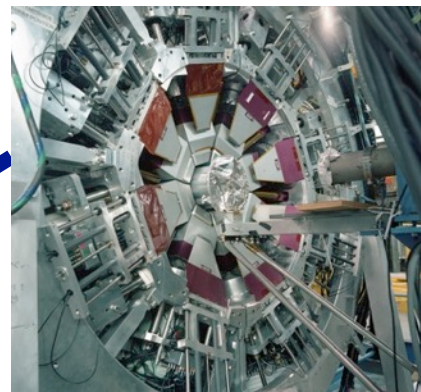
First charge-bred beam to DRAGON. Heaviest RIB direct radiative capture measurement.

High-energy RIBs
> 6 AMeV

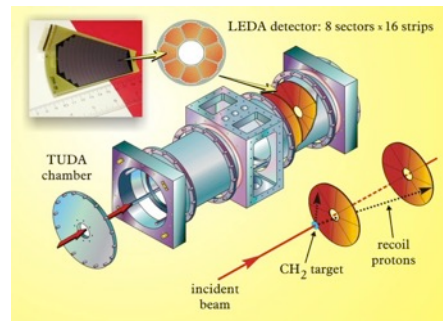
EMMA (2016)
Mass analyzer for
nuclear reactions



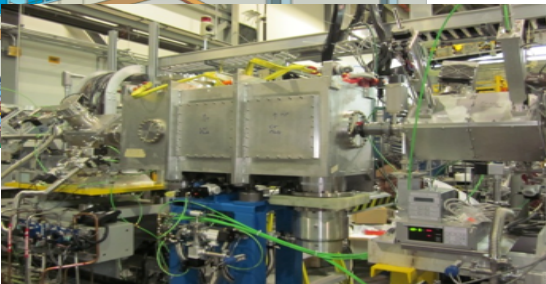
TIGRESS + auxiliary detectors
HPGe γ -ray spectrometer
in-beam spectroscopy of
nuclear reactions



TUDA
Scattering array
for direct reactions

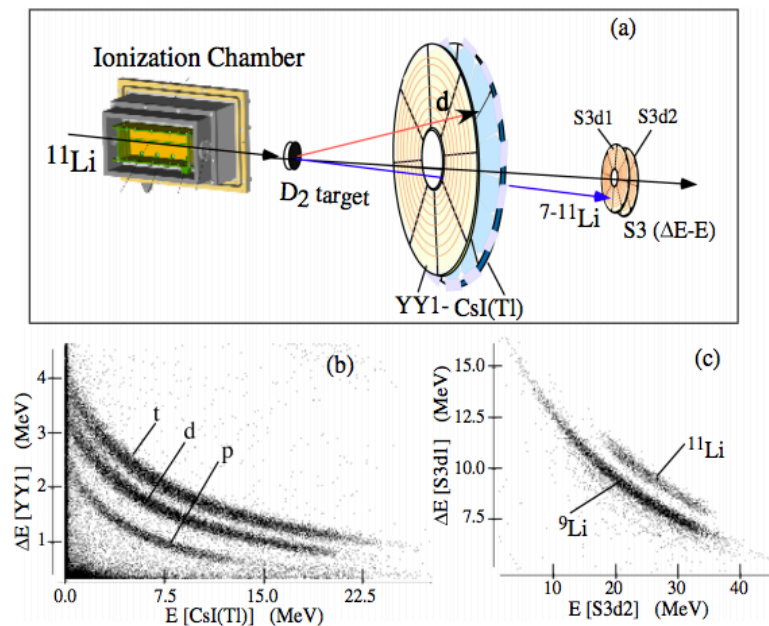


IRIS
Solid hydrogen target
for direct nuclear
reactions



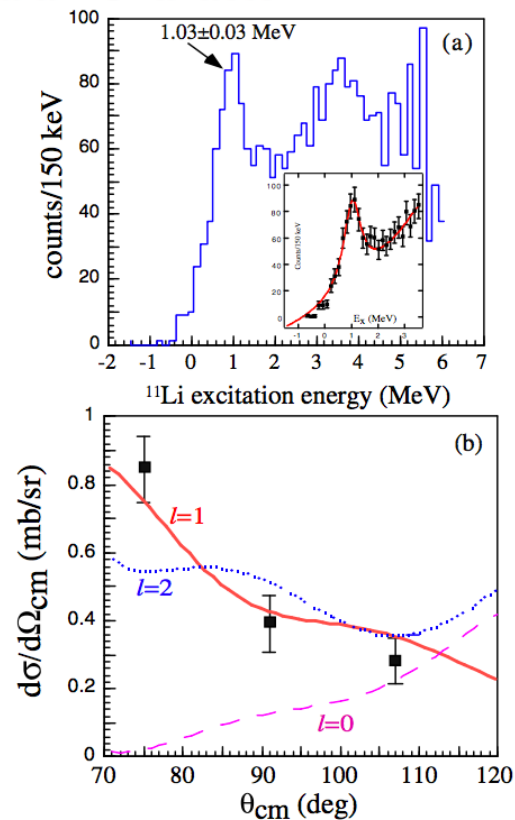
Evidence of Soft Dipole Resonance in ^{11}Li with Isoscalar Character

R. Kanungo, A. Sanetullaev *et al.*

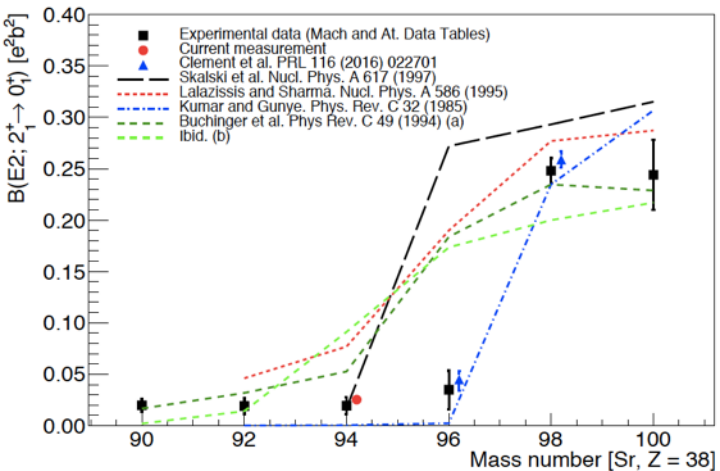
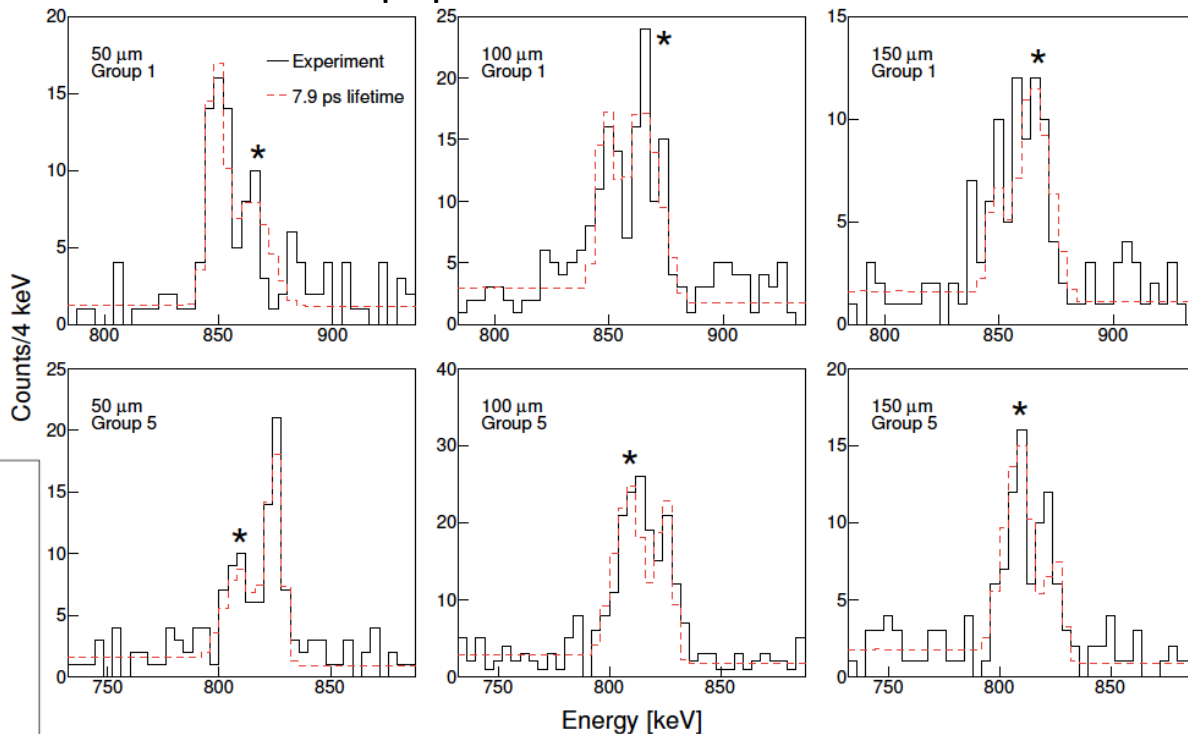
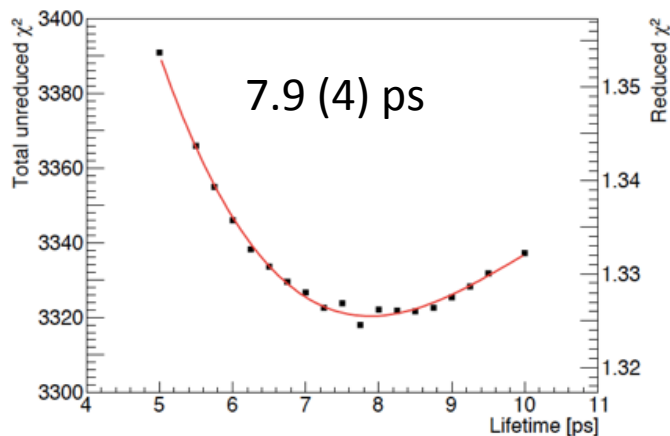


First evidence of a dipole resonance in ^{11}Li having an isoscalar character.

Provides stringent tests of *ab initio* theories and nuclear forces.

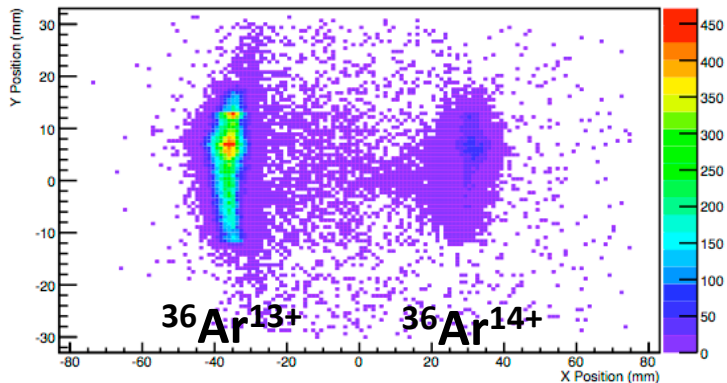


First 2^+ state populated in unsafe Coulex of ^{94}Sr beam

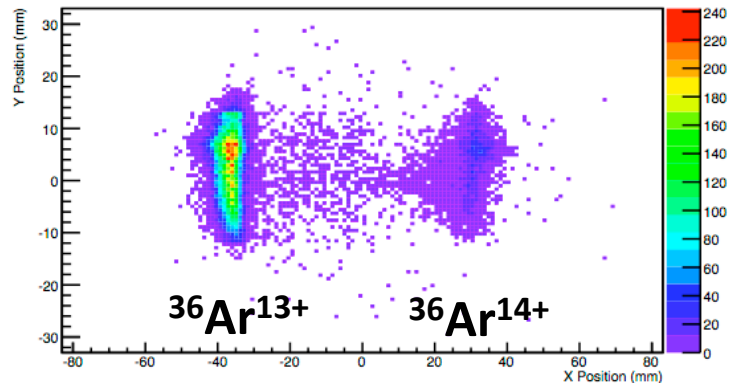


- Improves uncertainty in the $B(E2)$ value from 40% to $<6\%$.
- Confirms sudden onset of deformation in ground-state structure. *Submitted to PRL.*

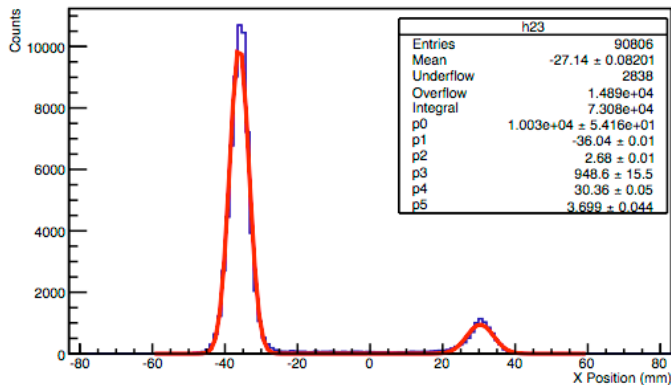
Position Spectrum (No Cuts)



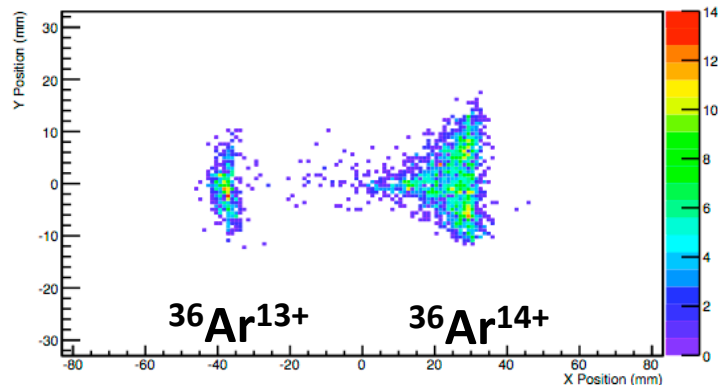
Position Spectrum (w/ Silicon Trigger)



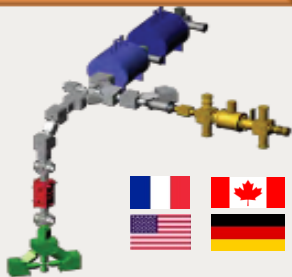
X Profile (No Cuts)



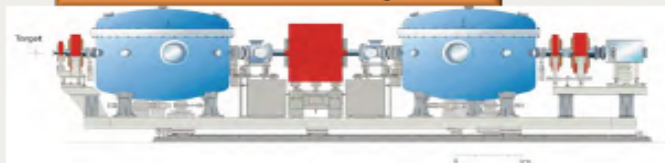
Position Spectrum (w/ High Silicon Trigger Threshold)



TITAN Penning Trap facility



EMMA recoil mass analyzer

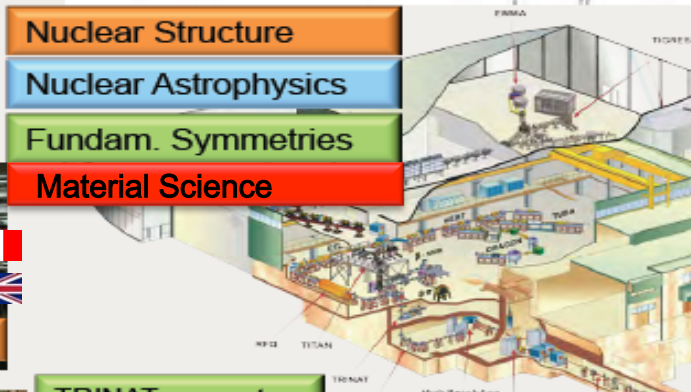


Nuclear Structure

Nuclear Astrophysics

Fundam. Symmetries

Material Science



MTV Mott scattering drift chamber



TIGRESS in-beam gamma-ray spectrometer



Laser polarizer line



IRIS solid hydrogen reaction set-up



Francium trapping facility

TRINAT magneto optical trap



DESCANT



GRIFFIN



TUDA reaction setup



DRAGON recoil separator



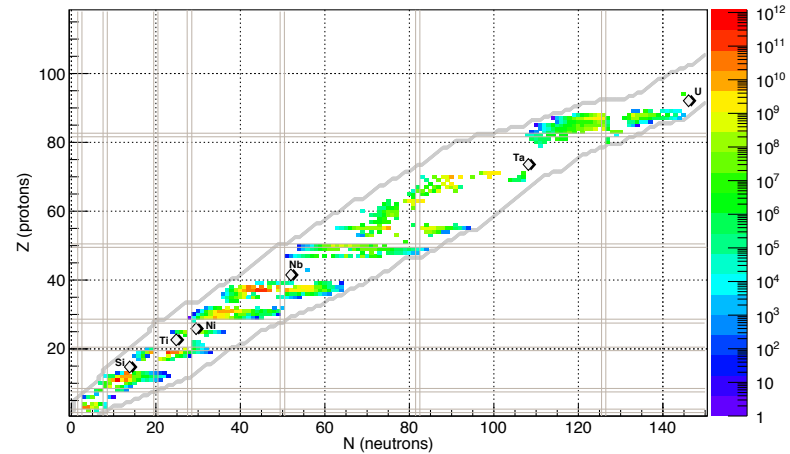
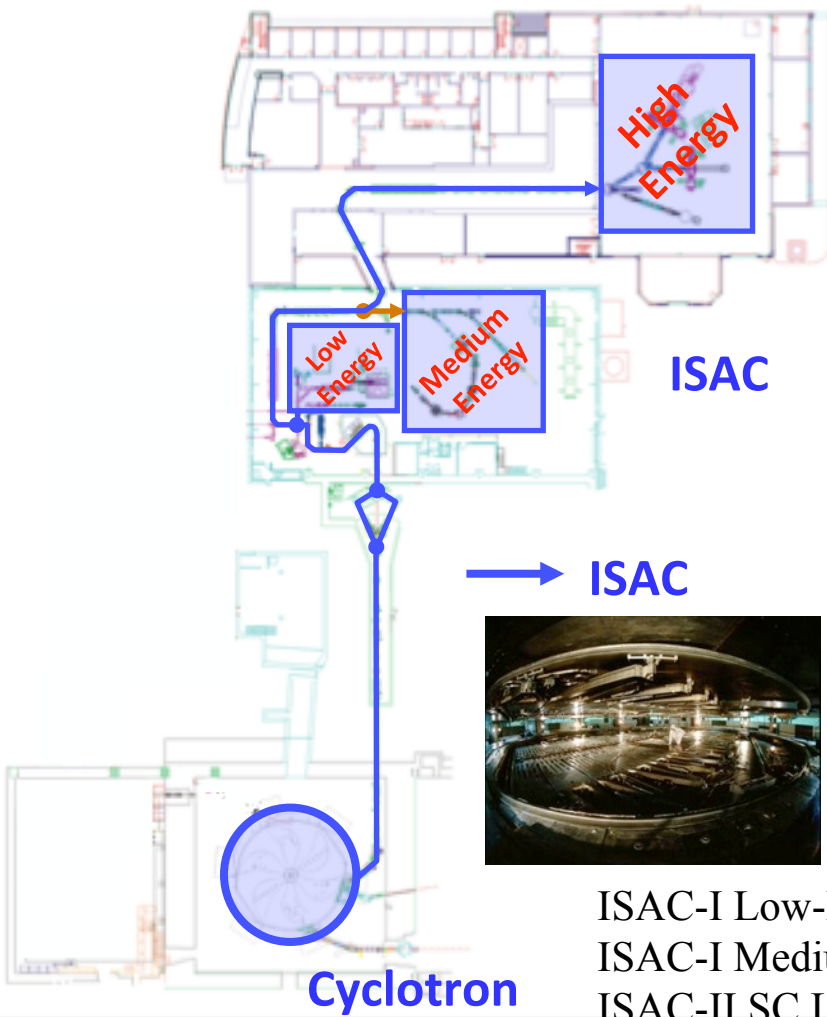
TRIUMF-ISAC

Isotope Separator and ACcelerator

1 RIB delivery to experiments

500MeV p^+ at 100 μ A on ISOL target

SiC, NiO, Nb, ZrC, Ta, UC_x Targets
Surface, FEBIAD, IG-LIS ion sources



ISAC-I Low-Energy <60keV

ISAC-I Medium E <1.5MeV/u

ISAC-II SC LINAC <10MeV/u

Ground state + decay, material science

Astrophysics

Nuclear reactions and structure

TRIUMF-ARIEL

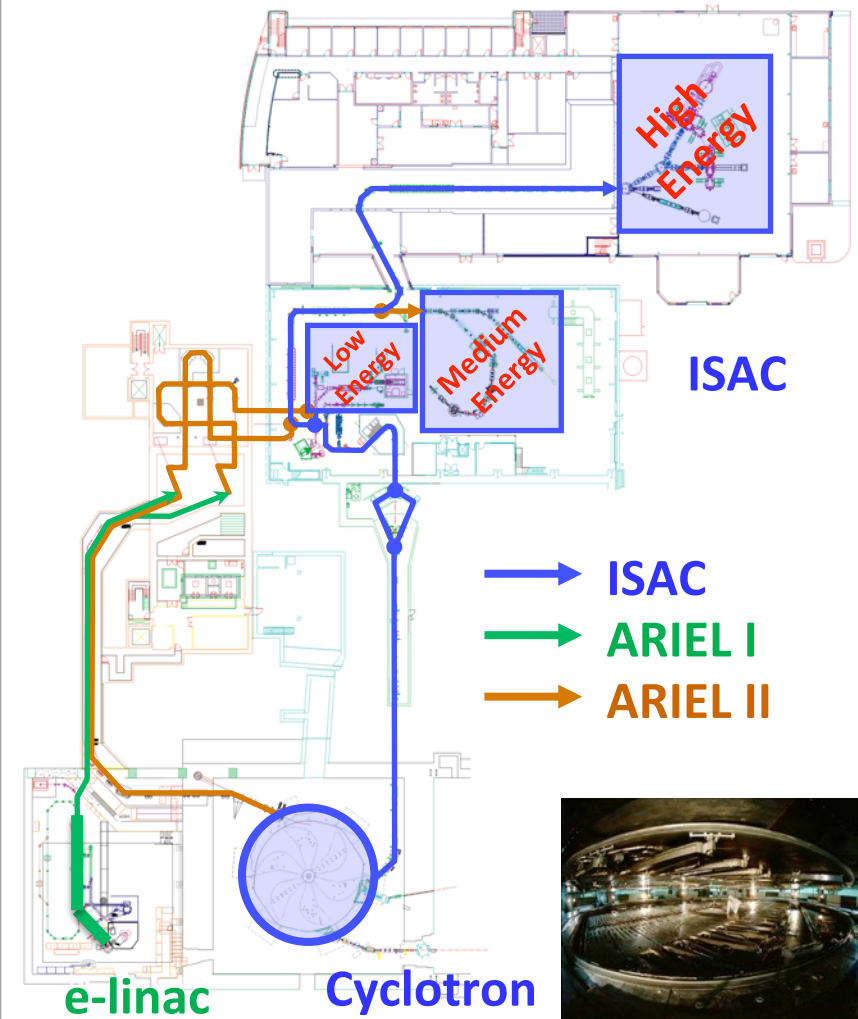
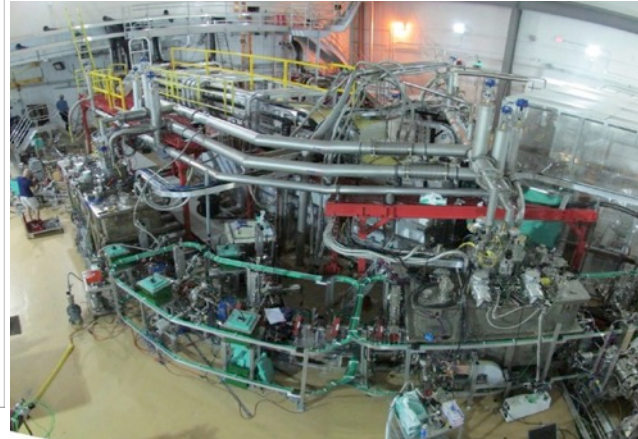
Advanced Rare-Isotope Laboratory

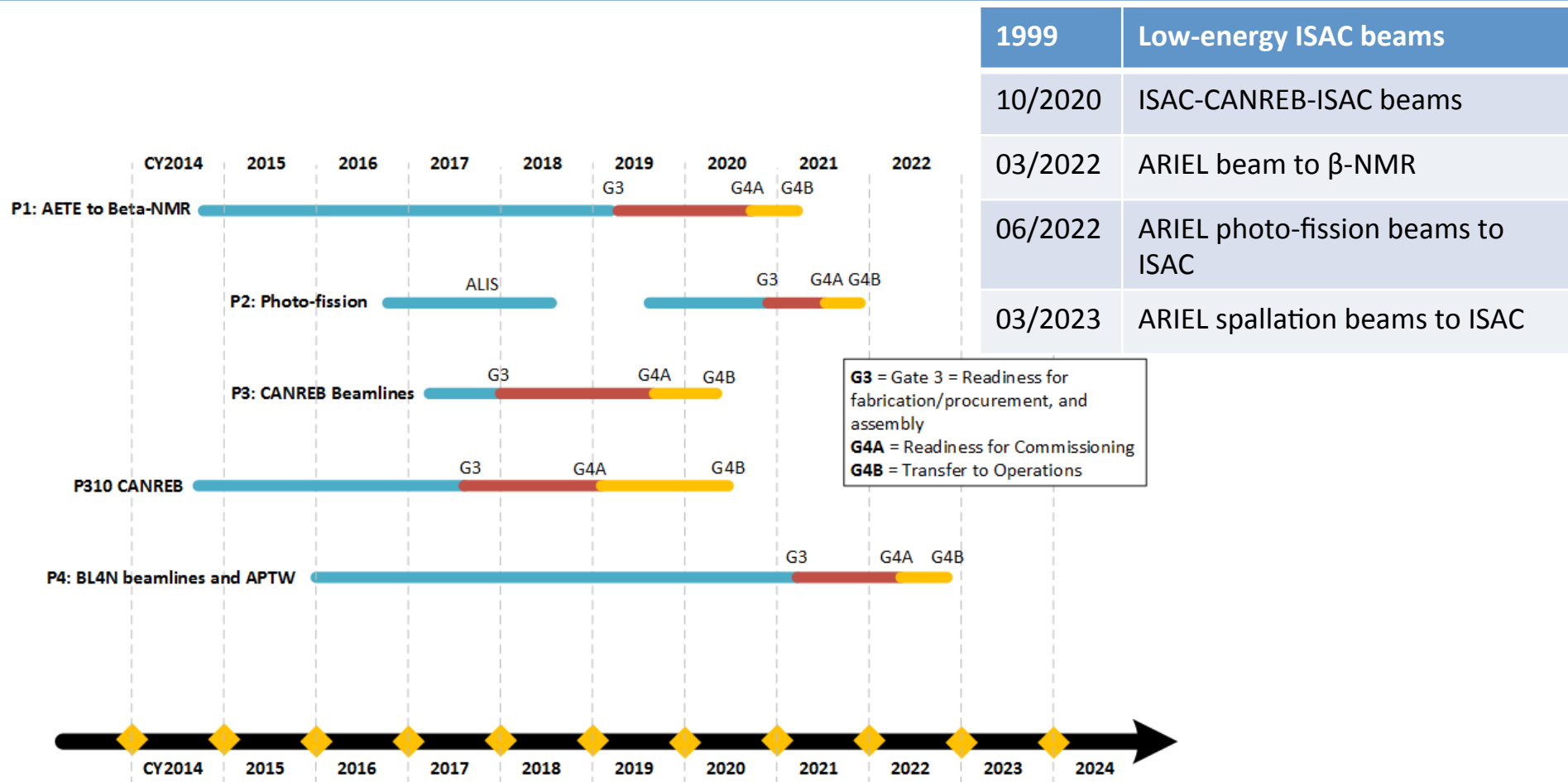
1 RIB → 3 simultaneous RIBs

ARIEL Project:

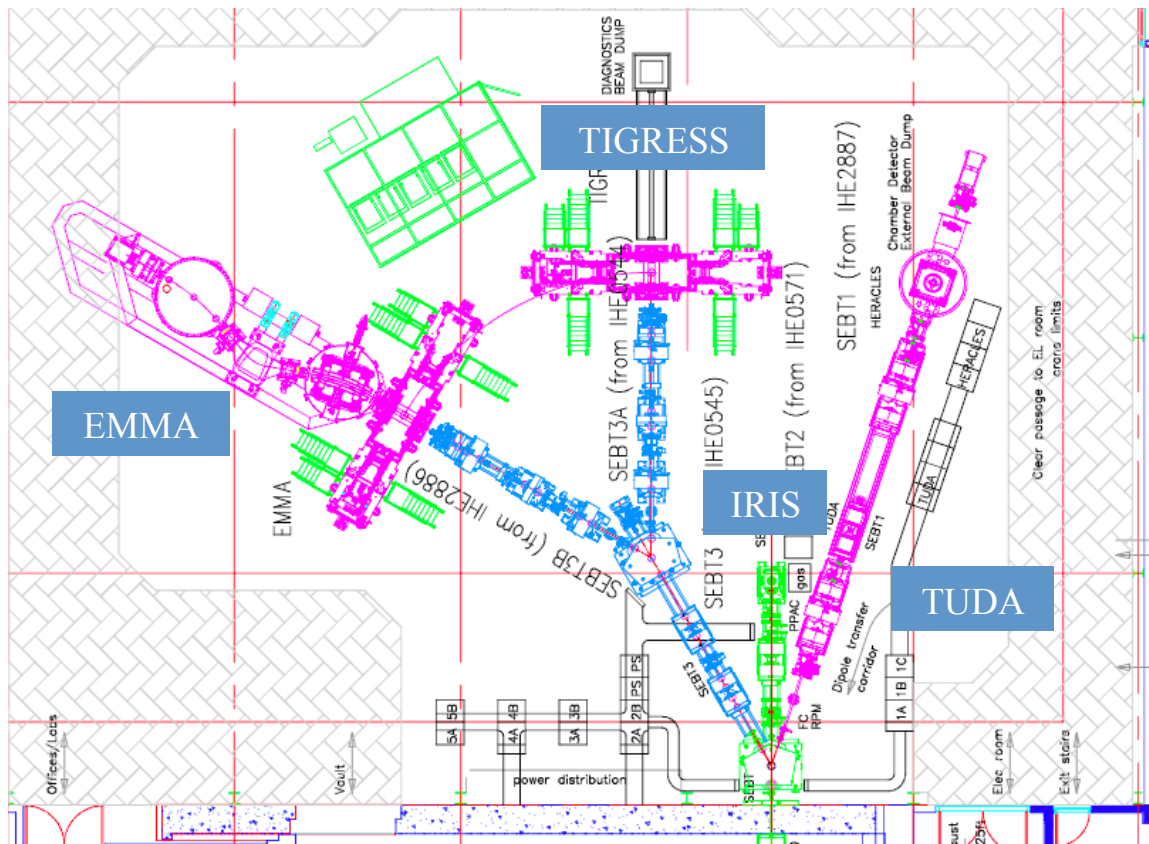
- new electron linac driver for photo-fission
- new proton beamline
- new target stations and front end

E-linac and electron beamline
Sept. 2014



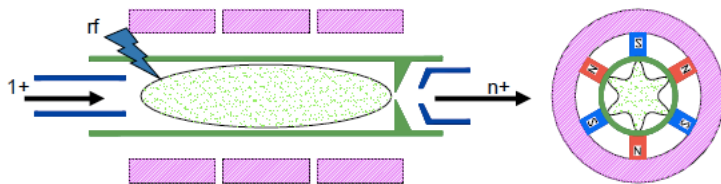


- Impact on accelerated beam program, primarily ISAC-II



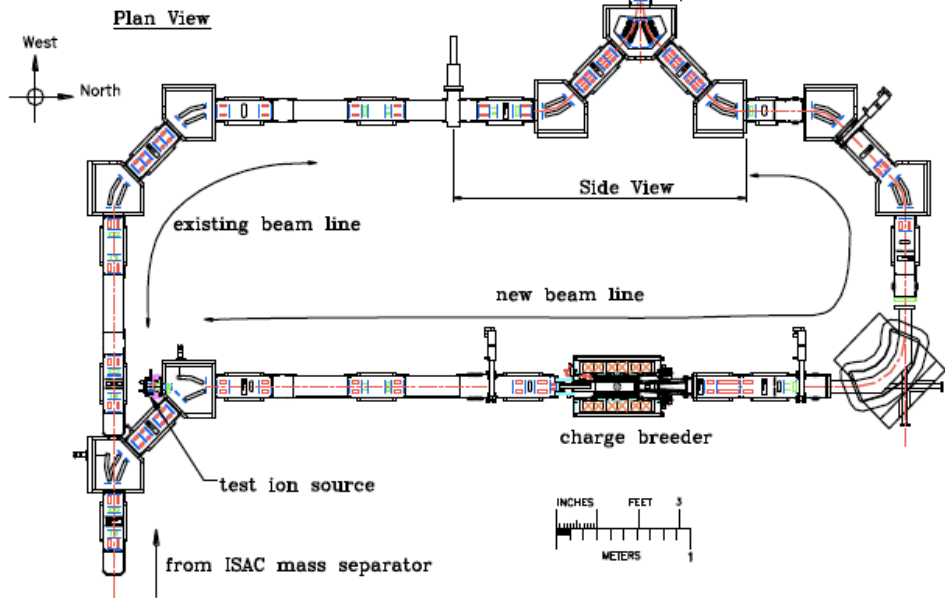
EMMA
IRIS
TIGRESS
TUDA-II

DRAGON
TUDA-I



$$\omega_{rf} = \omega_c = \frac{q}{m(E)} B$$

to experimental hall



Existing ISAC ECR CSB:

Modified 14.5 GHz PHOENIX ECR ion source from Pantechnik

Inject 1+ ions, extract n+ ions

Reduces A/q from <238 to <30 (7) for acceptance into RFQ (MEBT)

Advantages:

- Continuous output (DC beam)
- High intensity capability
- No pre-bunching/cooling required

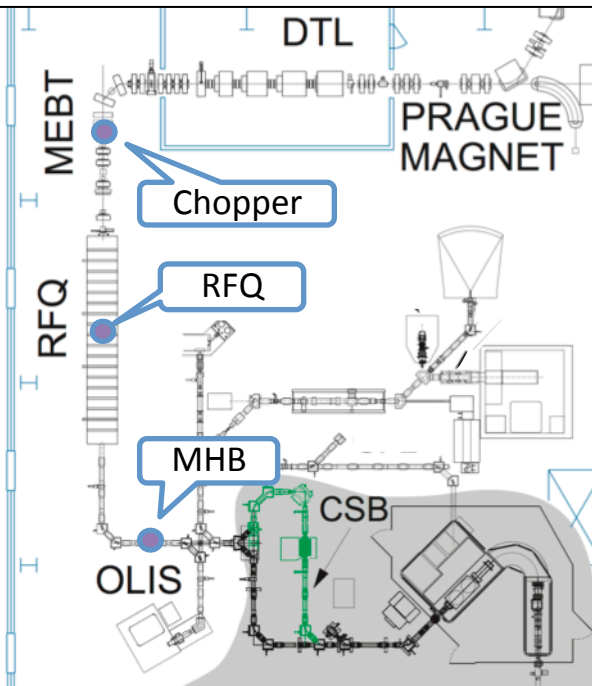
Issues:

- Efficiency <5%
- Stable backgrounds at all A/q

Stage 1:

CSB-LEBT-RFQ-MEBT-DTL

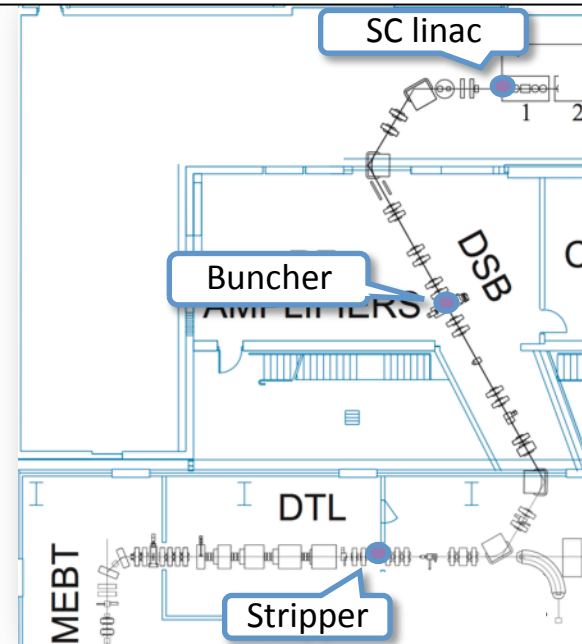
- Time-of-flight separation in LEBT
- Pre-buncher phase used to tune for selection
- Prague Diagnostic station used for setup
- Theoretical: 1/1000 resolution in A/q



Stage 2 (NOT ALWAYS REQUIRED):

DSB-SCLINAC-SEBT-Experiment

- Stripping foil at 1.5MeV/u
- Change in A/q and differential TOF
- DSB slits used for selection
- TBragg detector used for setup
- Theoretical: 1/800 resolution in A/q



Future CANREB EBIS CSB:

Electron-Beam Ion Source built at Max-Planck Institute for Nuclear Physics, Heidelberg

Inject cooled, bunched $1+$ ions, extract bunched $n+$ ions

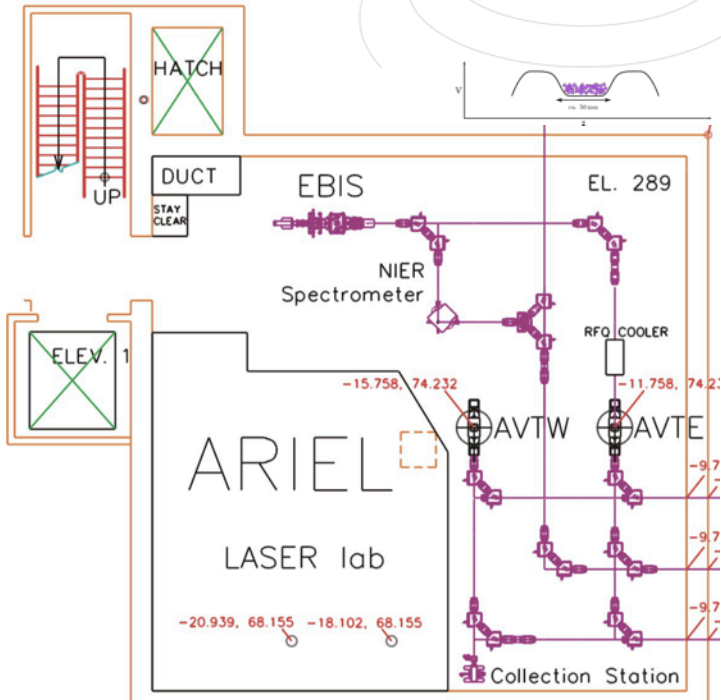
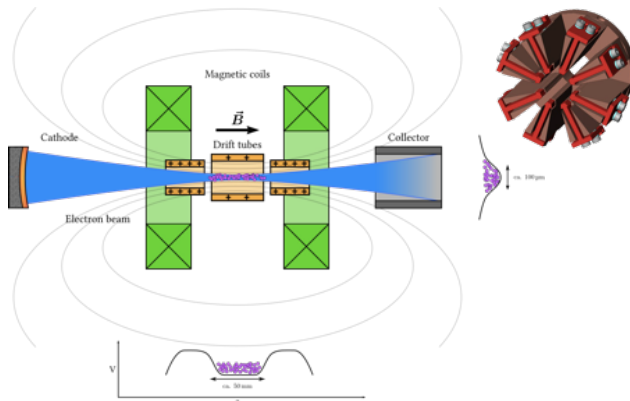
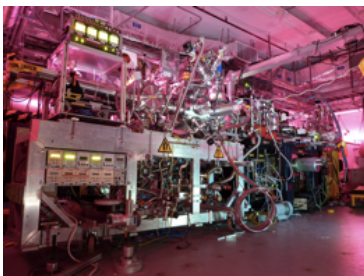
Reduces A/q from <238 to <30 (7) for acceptance into RFQ (MEBT)

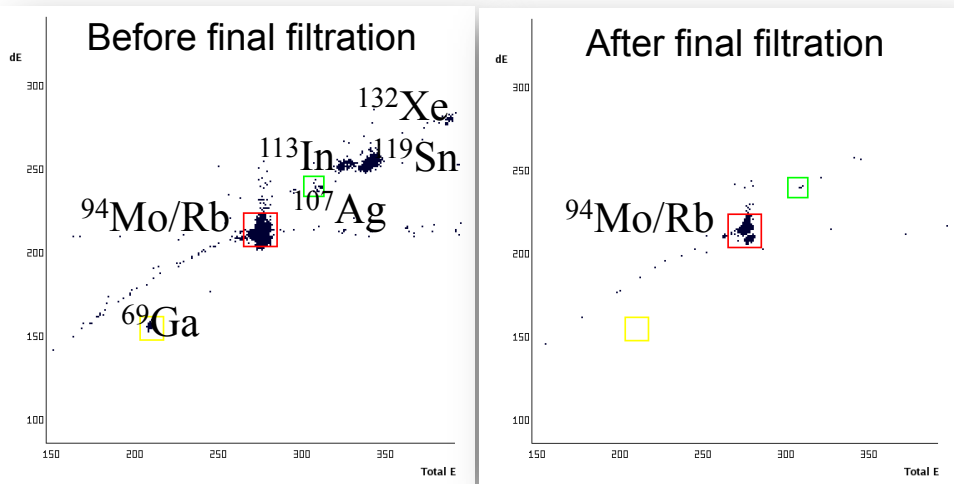
Advantages:

- High Efficiency 10-20%
- High-purity beams

Issues:

- Need cooled and bunched injection
- Pulsed beam extraction





Factor ~100 increase in charge-bred beam intensities.

Opportunities for the accelerated beam program

- Coulex: B(E2) → Quadrupole moment
- Transfer reactions become feasible in some cases where they were not before
- Transfer: state identification → precision

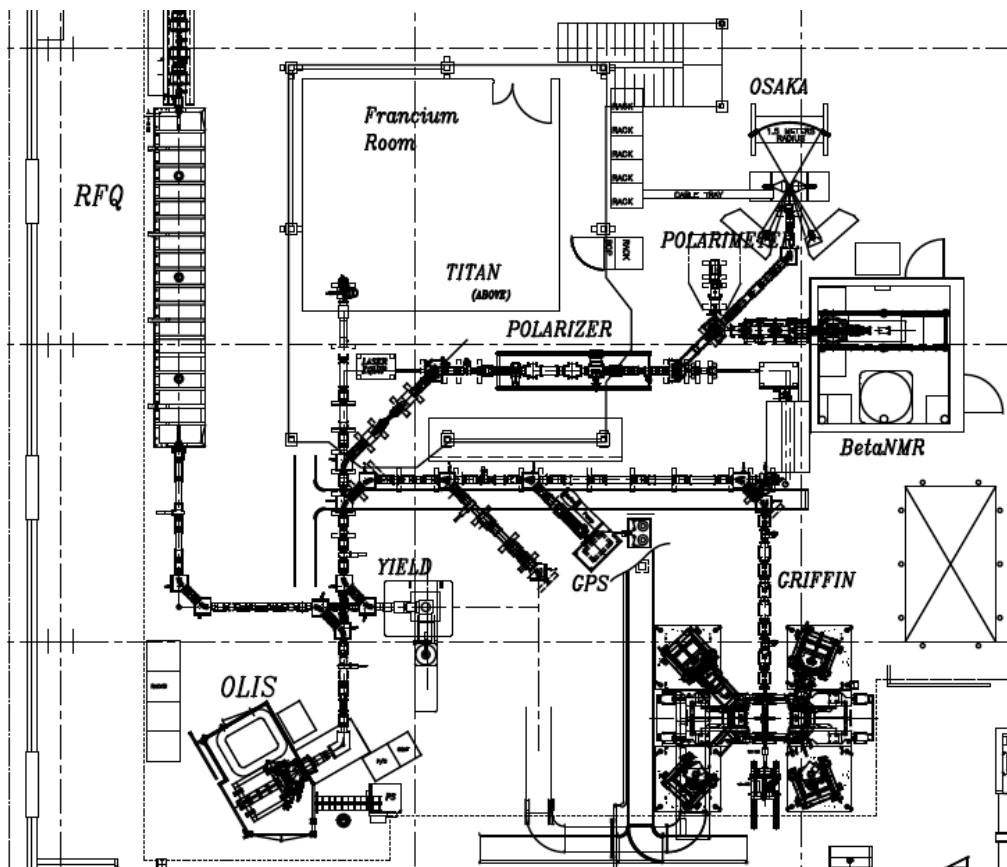
Call to Users:

What are the first experiments, science goals and priorities with CANREB charge-bred beams?

In-target production rates [$1 \text{ kW}^{-1} \cdot \text{s}^{-1}$]:

from BeO: ${}^8\text{Li}$: $5 \cdot 10^9$ pps

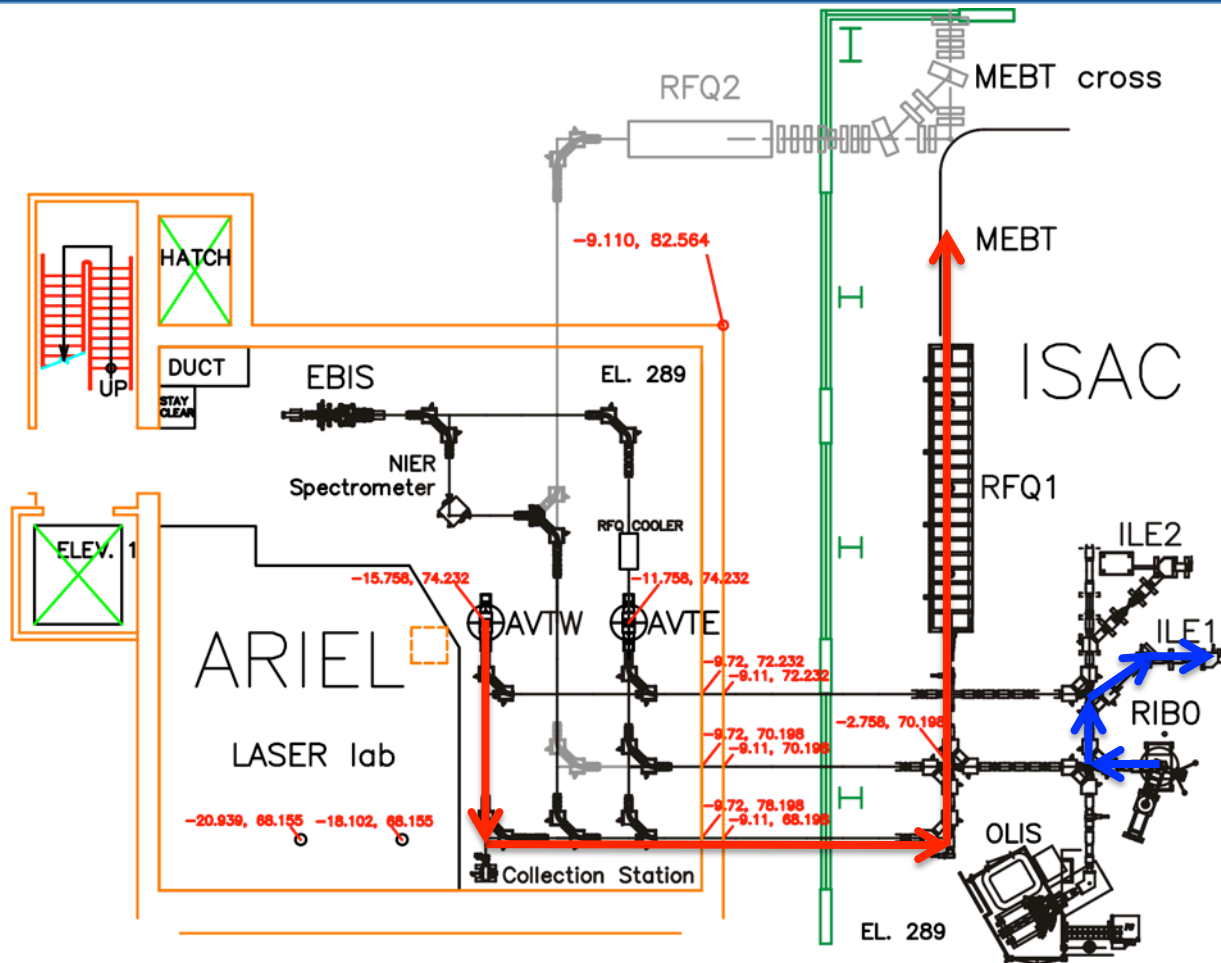
- β -NMR currently ~ 6 weeks of ISAC schedule \rightarrow ~ 3 months of ARIEL beam
- This will not happen instantly. Access still required in ARIEL target hall for installation work
- Low-energy area is poorly laid out. When delivering to β -NMR from ARIEL:
 - cannot deliver to TITAN, Co-liner Laser Spec., OSAKA, MTV
- Increase in beam availability to β -NMR, GRIFFIN, GPS, Francium, DRAGON, ISAC-II



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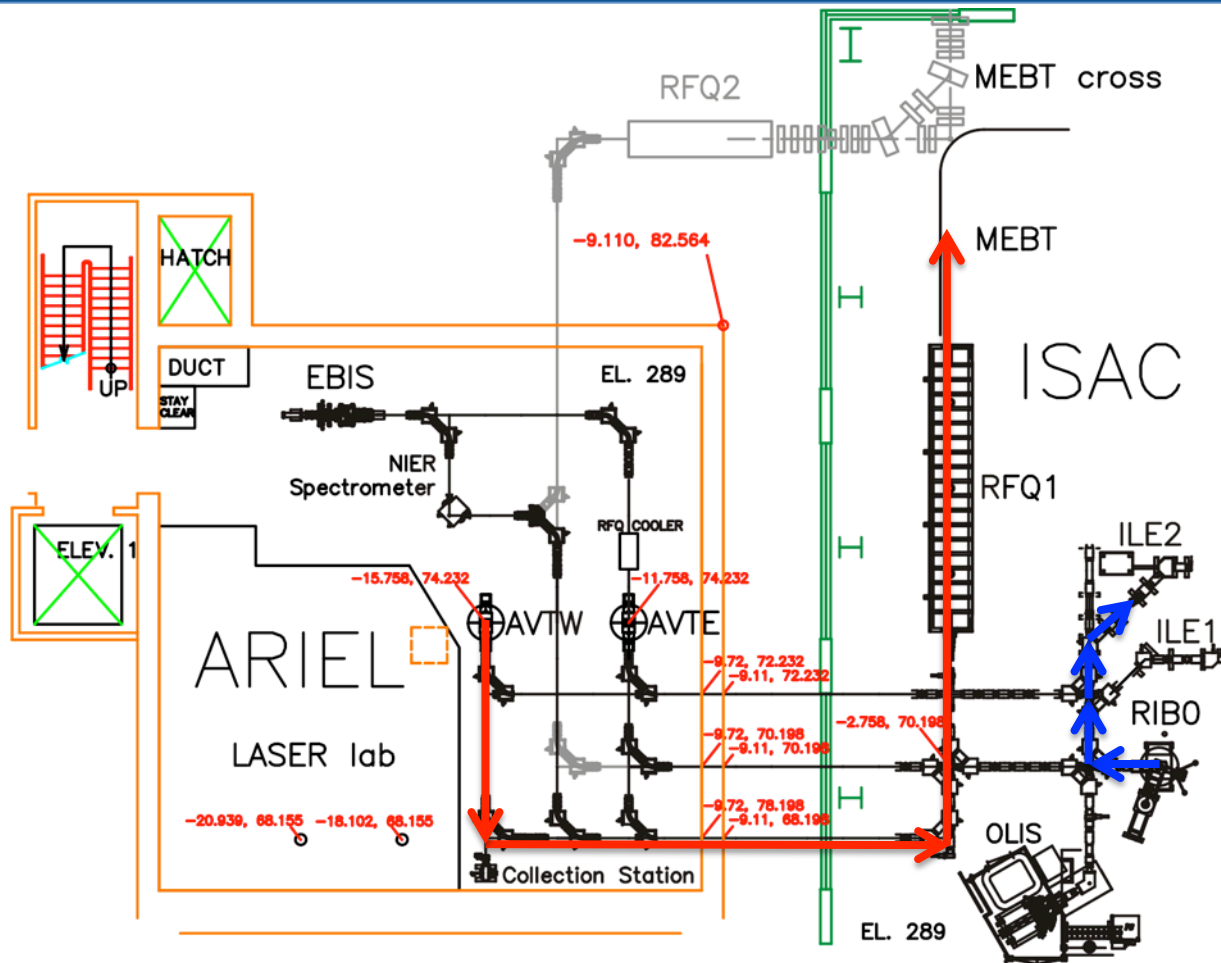
When delivering to β -NMR from ARIEL:

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ISAC – LEBT
GPS, GRIFFIN

ARIEL – MEBT/SEBT
DRAGON, TUDA-I
EMMA, IRIS,
TIGRESS, TUDA-II

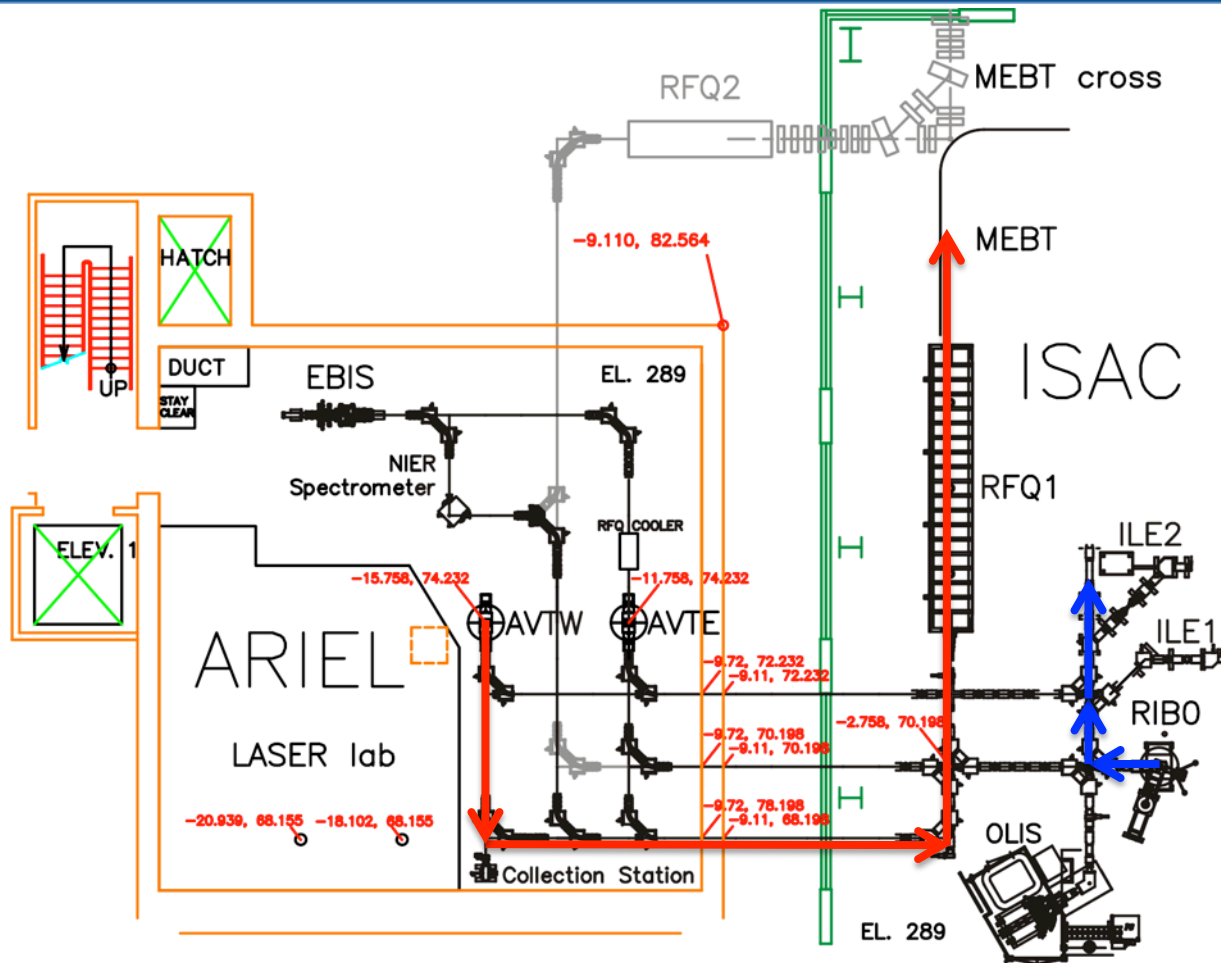


ISAC – LEBT

TITAN, Laser Spec.,
beta-NMR, OSAKA, MTV

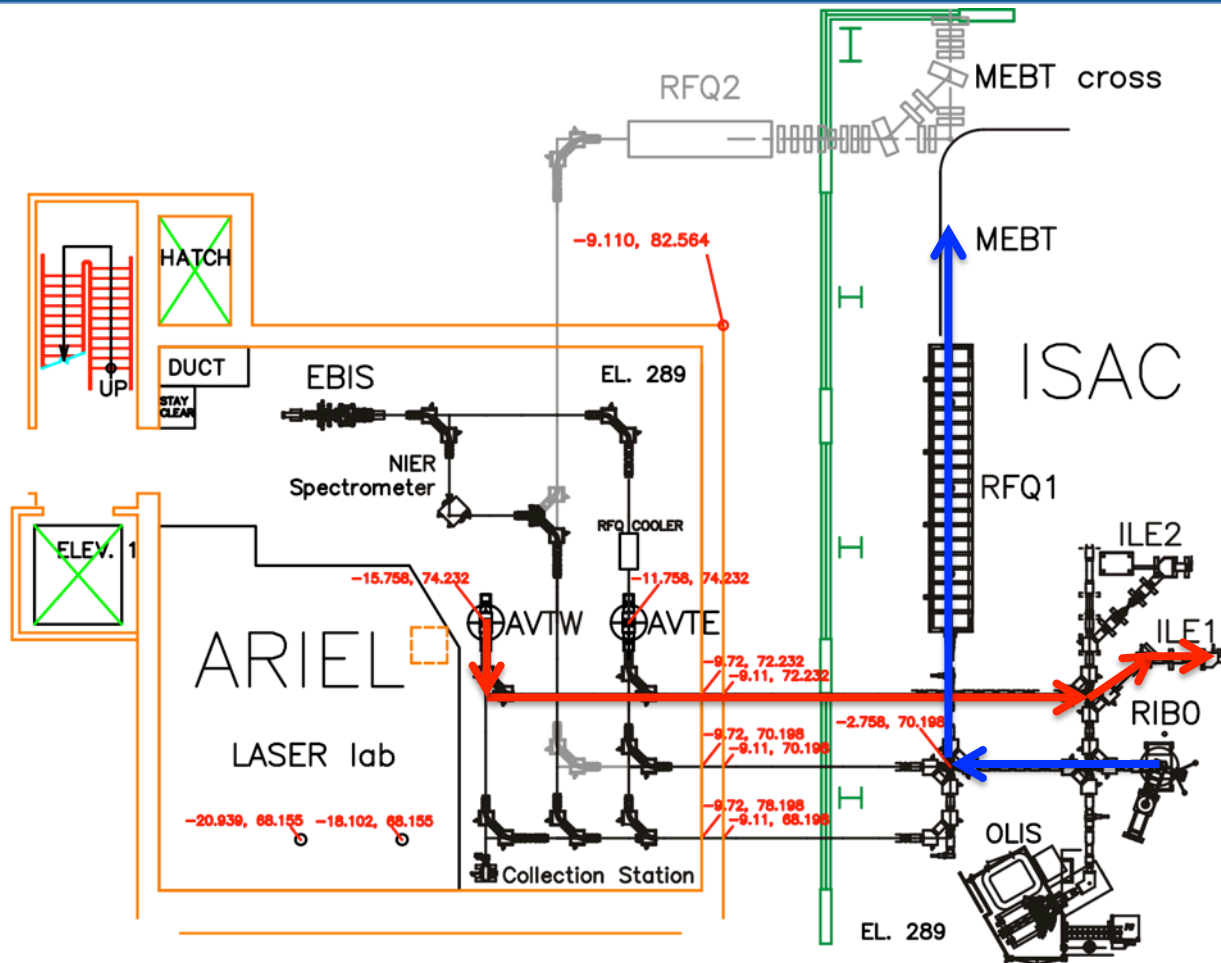
ARIEL – MEBT/SEBT

DRAGON, TUDA-I
EMMA, IRIS,
TIGRESS, TUDA-II



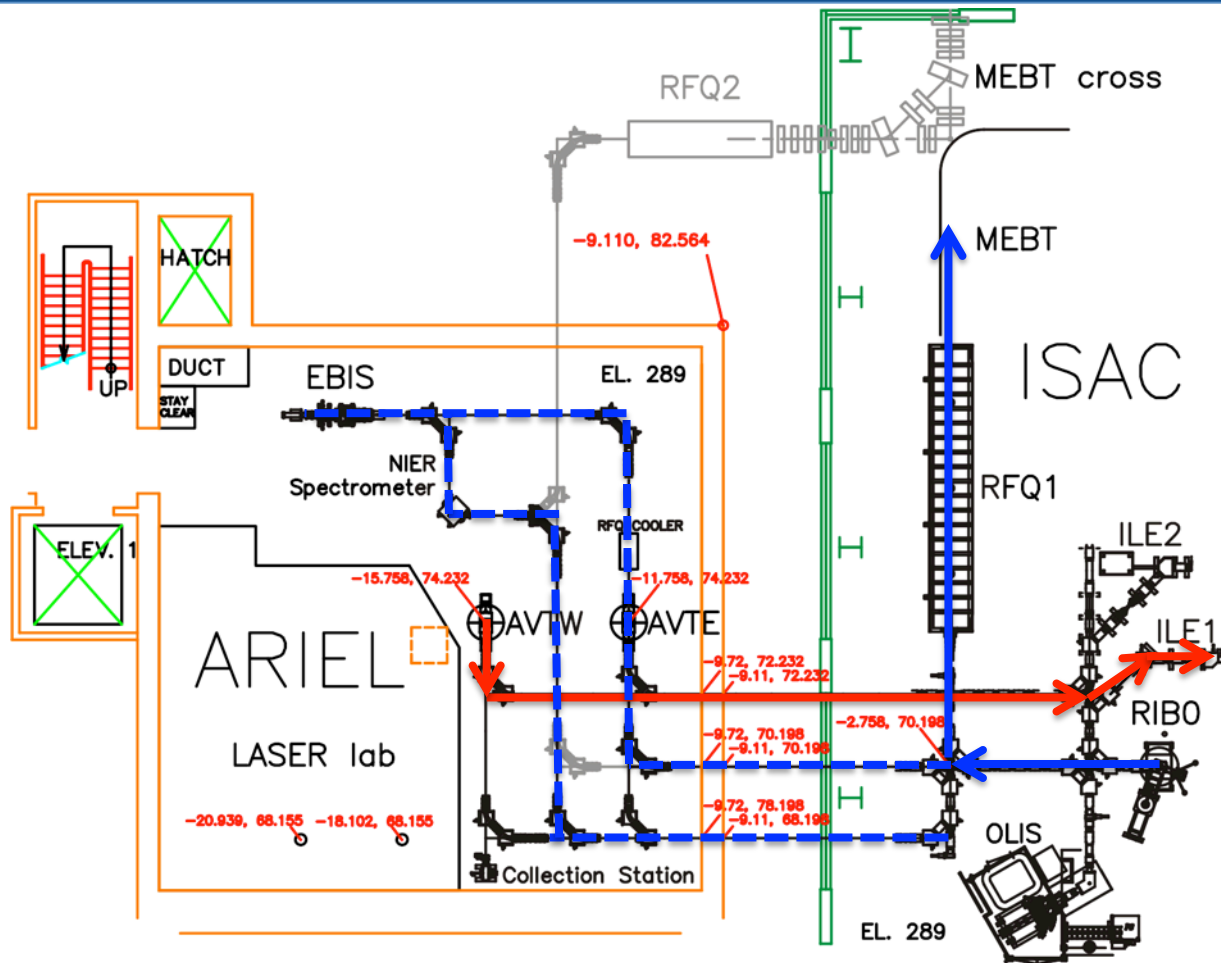
ISAC – LEBT
Francium

ARIEL – MEBT/SEBT
DRAGON, TUDA-I
EMMA, IRIS,
TIGRESS, TUDA-II



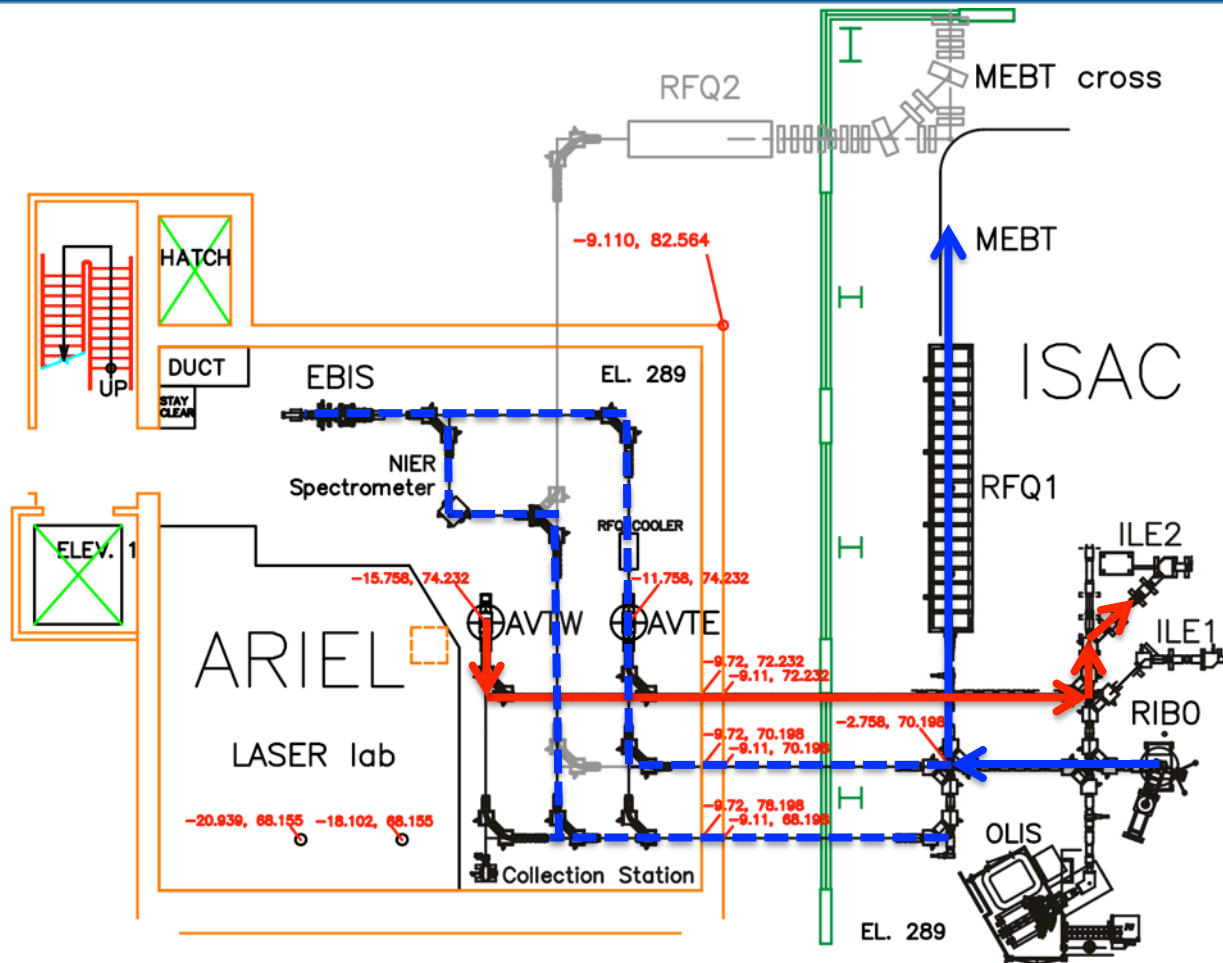
ISAC – LEBT
 DRAGON, TUDA-I
 EMMA, IRIS,
 TIGRESS, TUDA-II

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 GPS, GRIFFIN



ISAC – LEBT
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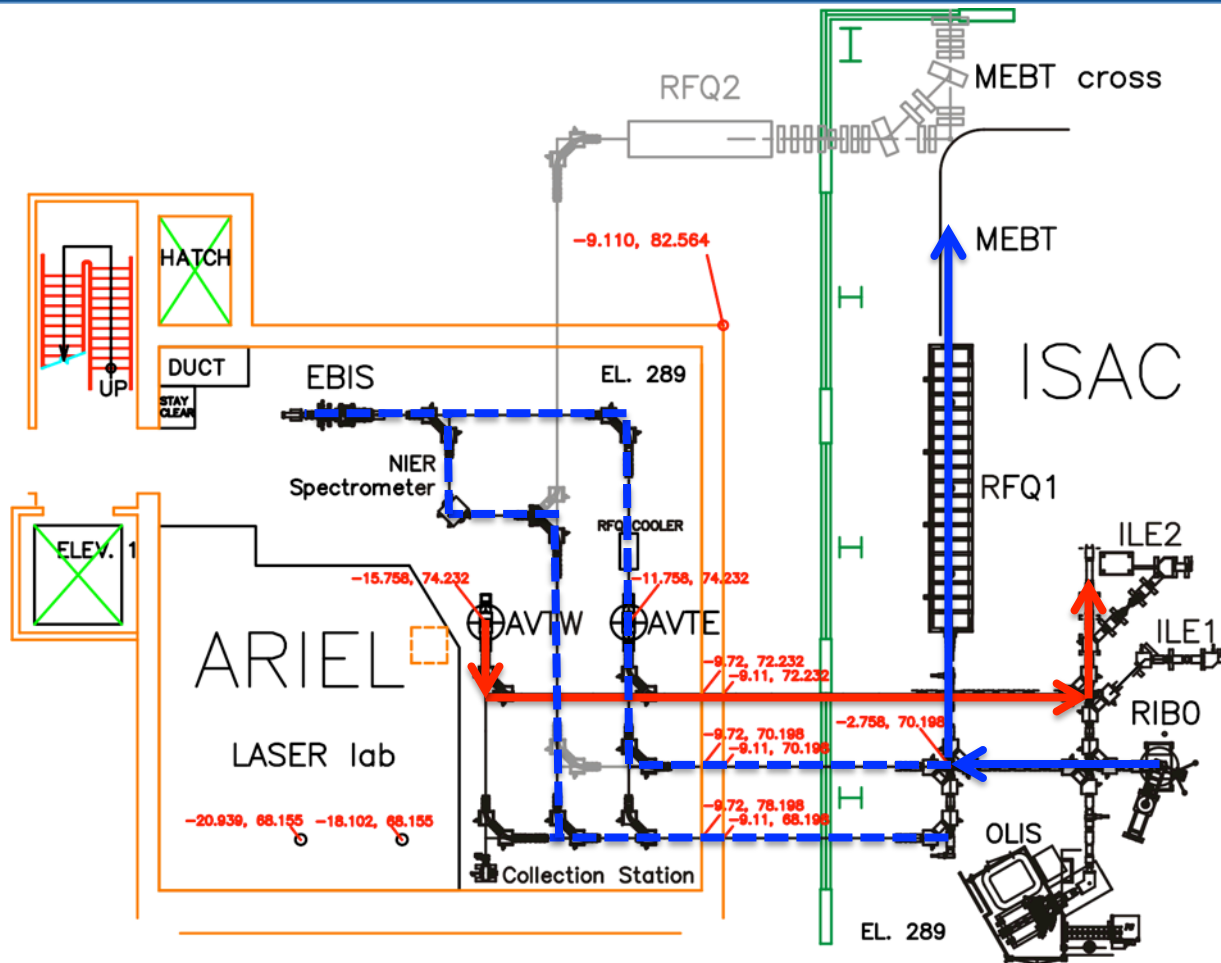


ISAC – LEBT

DRAGON, TUDA-I
EMMA, IRIS,
TIGRESS, TUDA-II

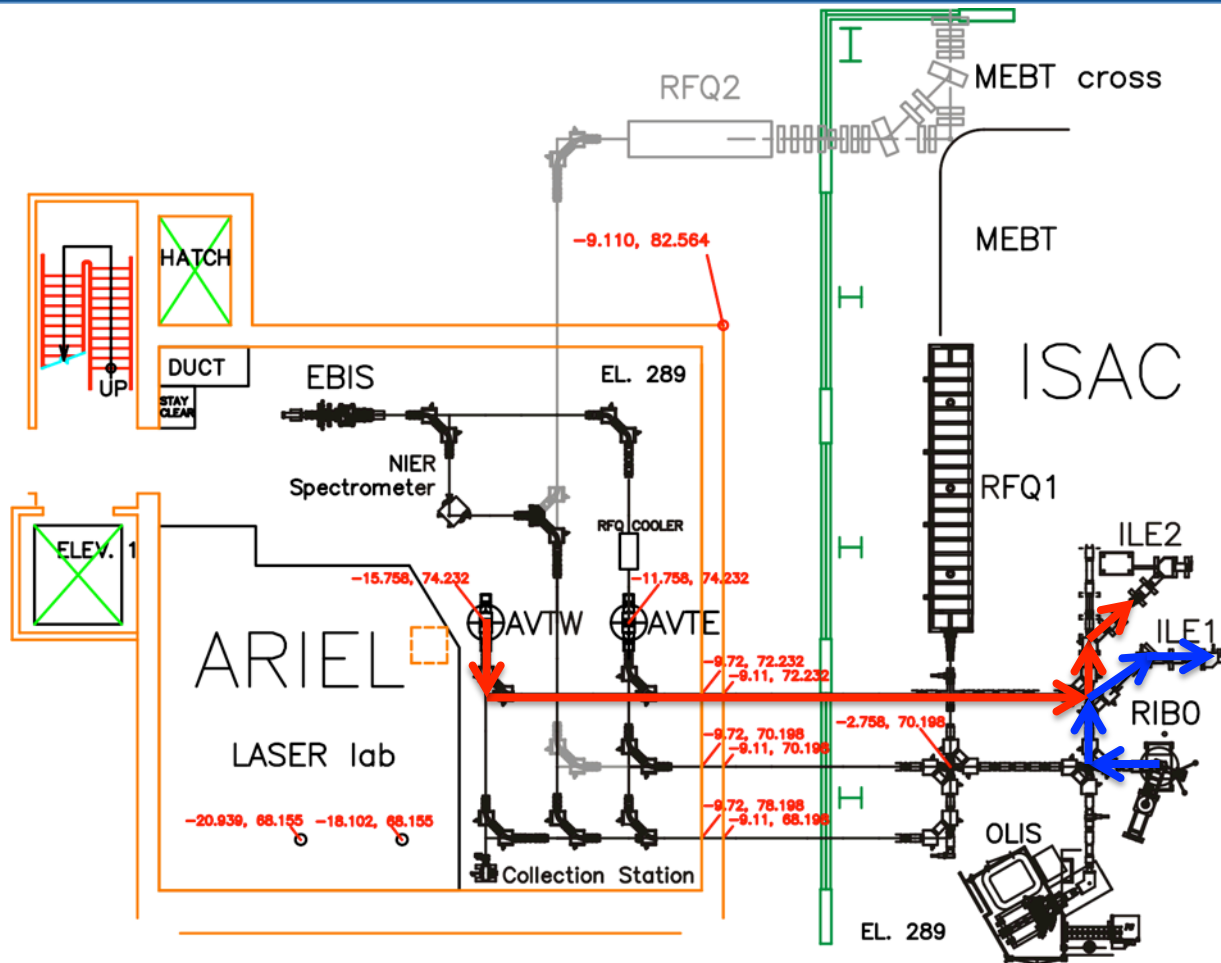
ARIEL – LEBT

TITAN, Laser Spec.,
beta-NMR, OSAKA, MTV



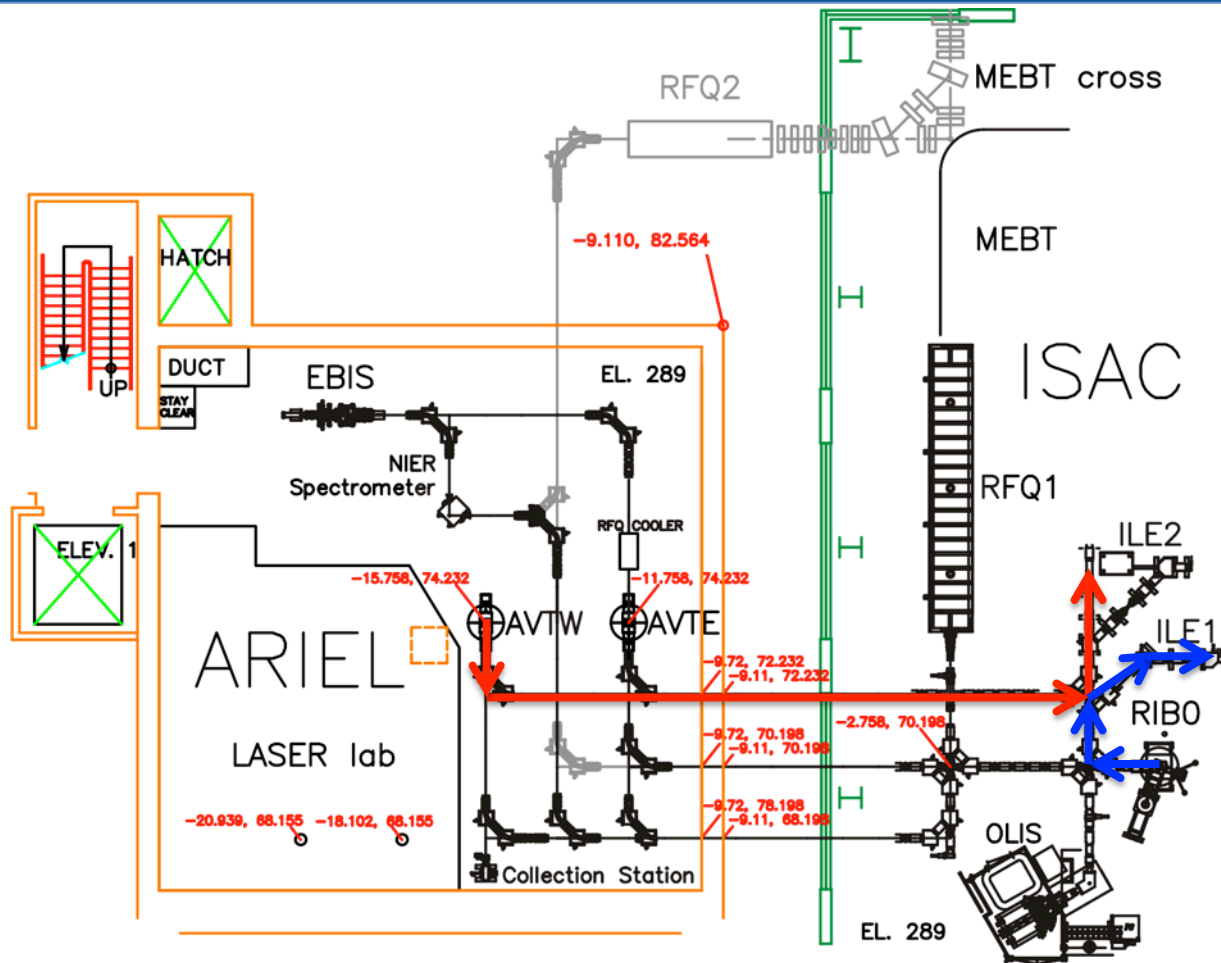
ISAC – LEBT
 DRAGON, TUDA-I
 EMMA, IRIS,
 TIGRESS, TUDA-II

ARIEL – LEBT
 Francium



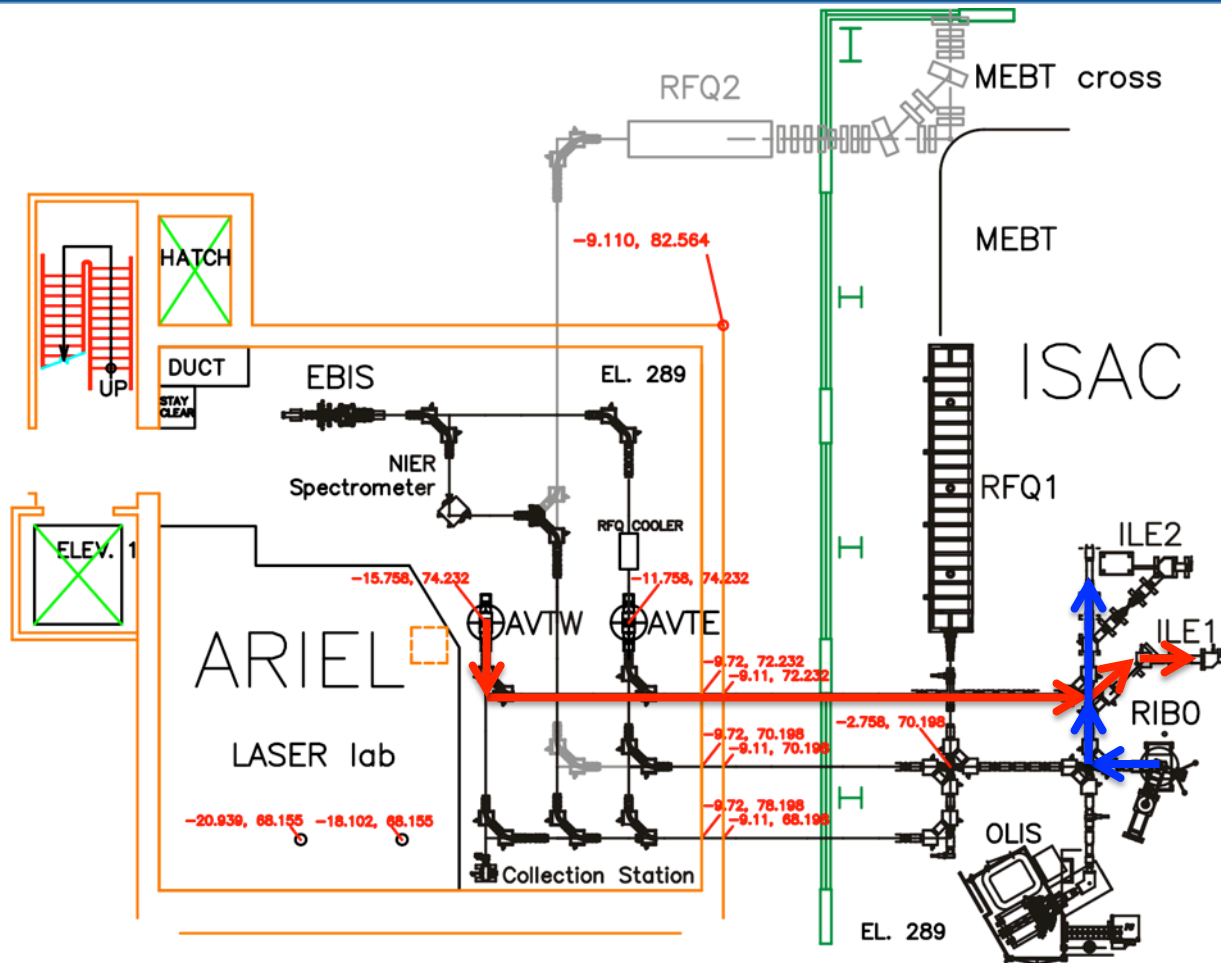
ISAC – LEBT
GPS, GRIFFIN

ARIEL – LEBT
TITAN, Laser Spec.,
beta-NMR, OSAKA, MTV



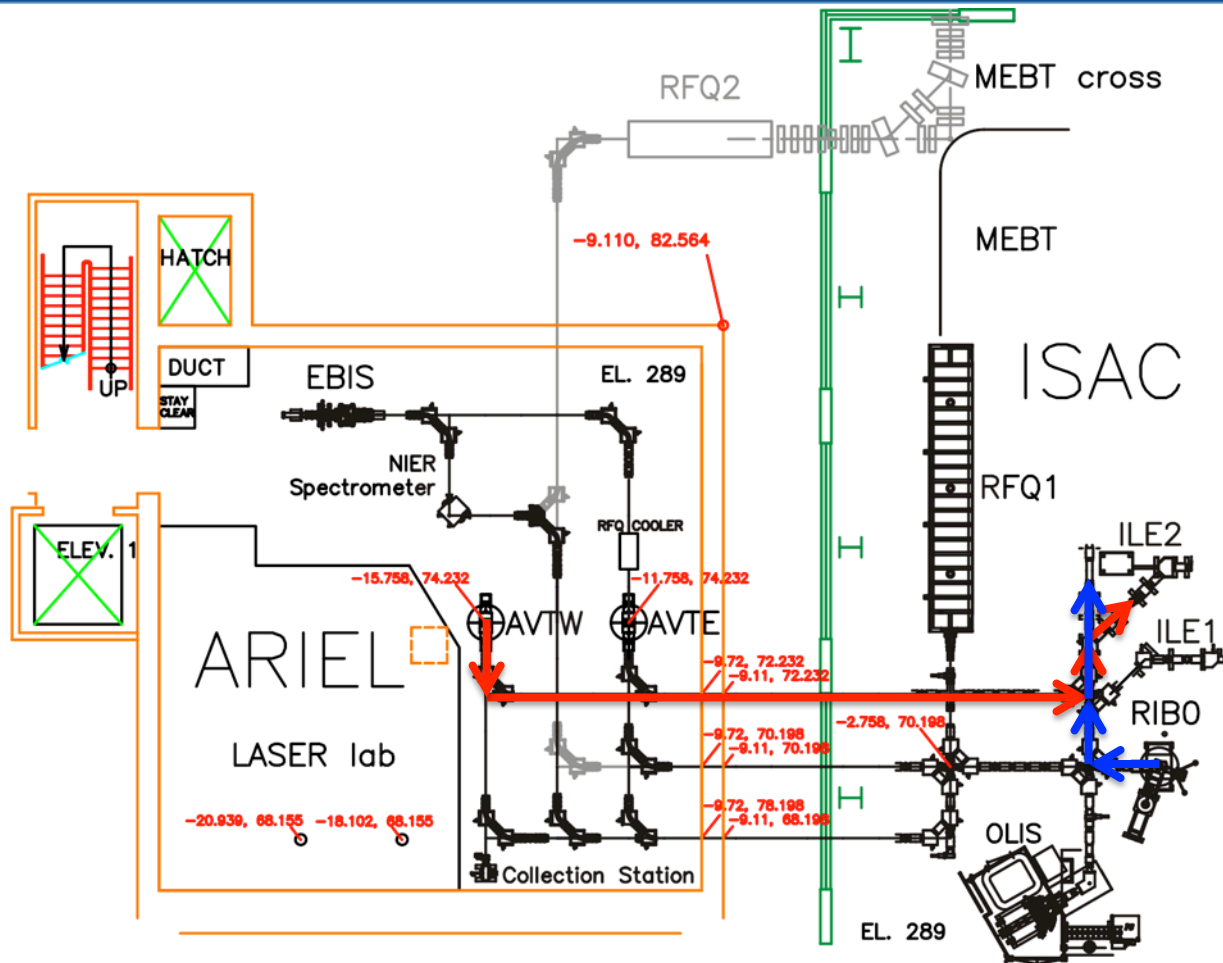
ISAC – LEBT
GPS, GRIFFIN

ARIEL – LEBT
Francium



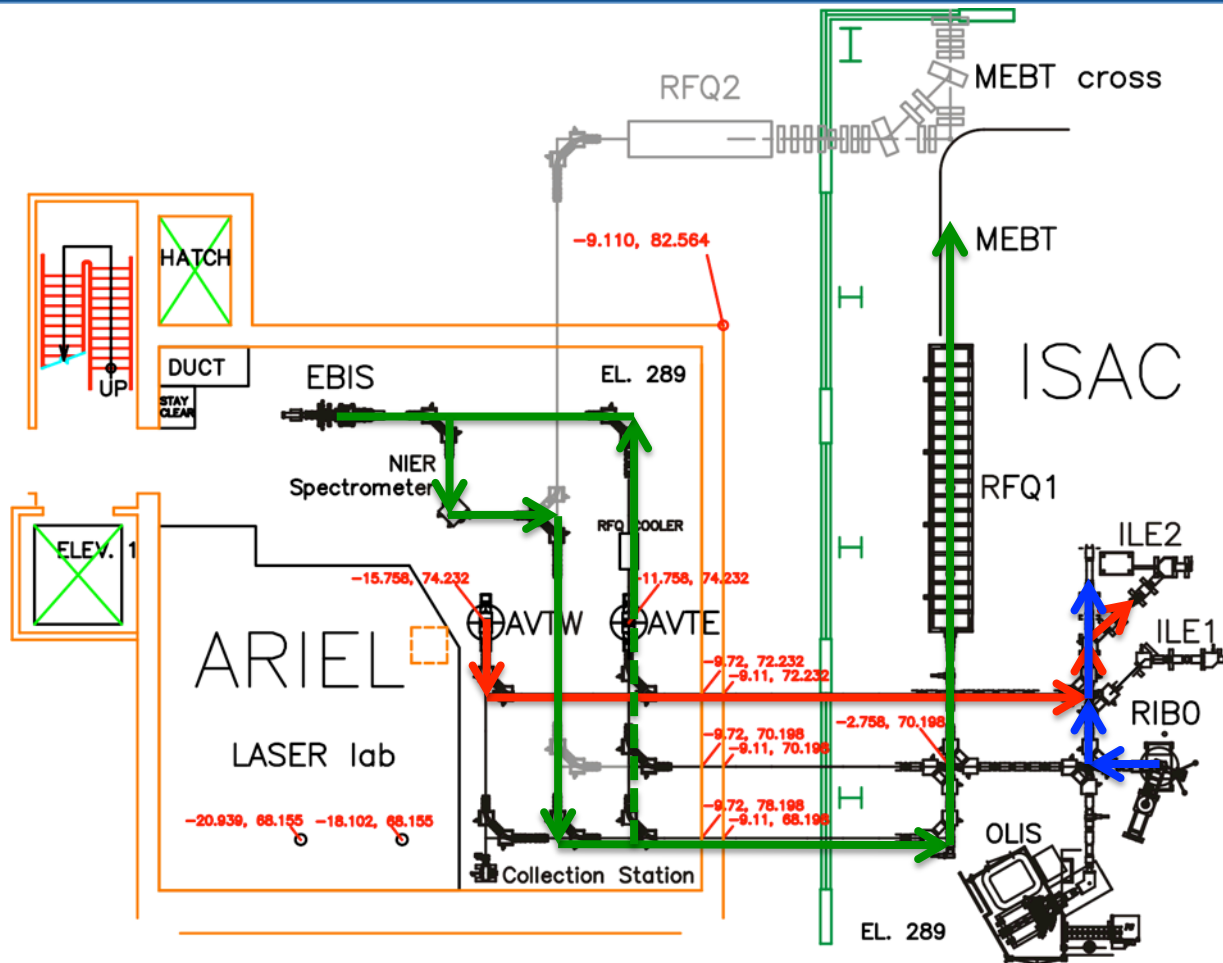
ISAC – LEBT
Francium

ARIEL – LEBT
GPS, GRIFFIN



ISAC – LEBT
Francium

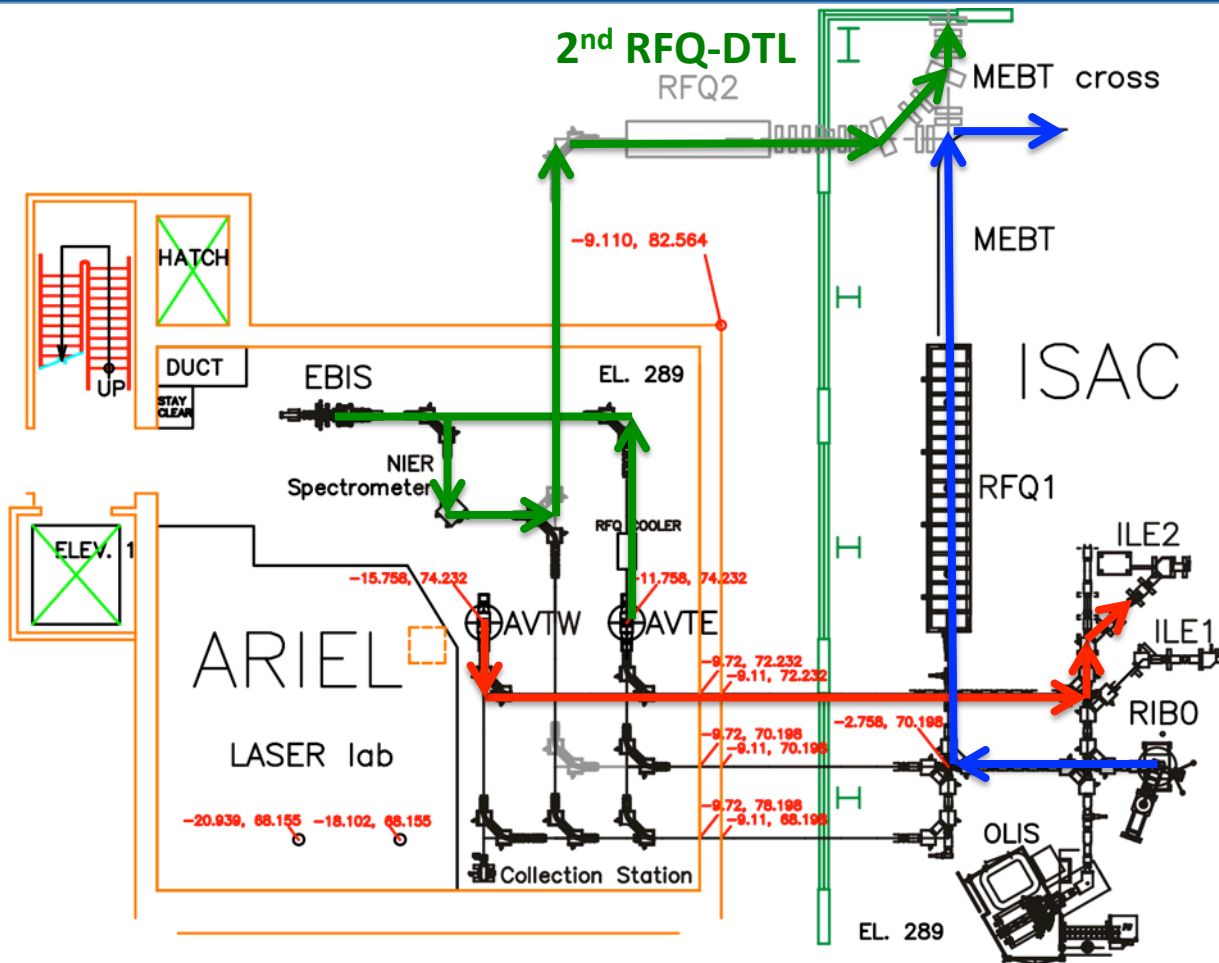
ARIEL – LEBT
TITAN, Laser Spec.,
beta-NMR, OSAKA, MTV



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ISAC – MEBT
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ARIEL – LEBT
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 EMMA, IRIS,
 TIGRESS, TUDA-II

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Call to Users:

**What are the first experiments, science goals and priorities
for material science with β -NMR?**

Rare-isotope beams will be produced by from proton and electron driver beams.

Proton energy above 350 MeV

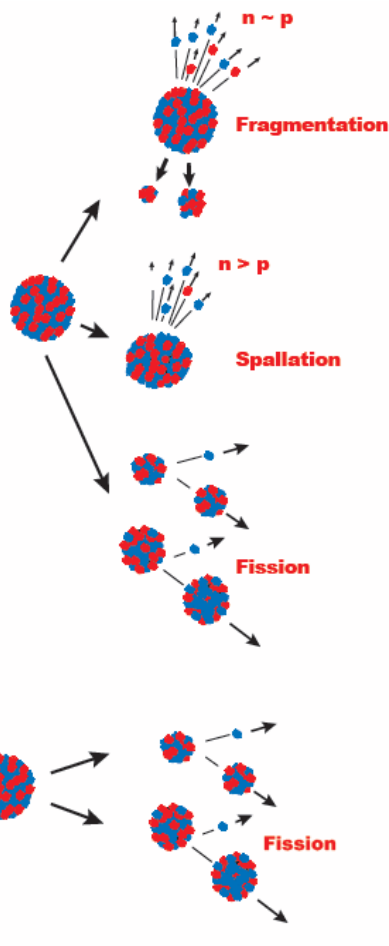
Protons

n
p

Electron energy ~ 50 MeV

High energy gamma

n
p



500MeV Protons on UCx

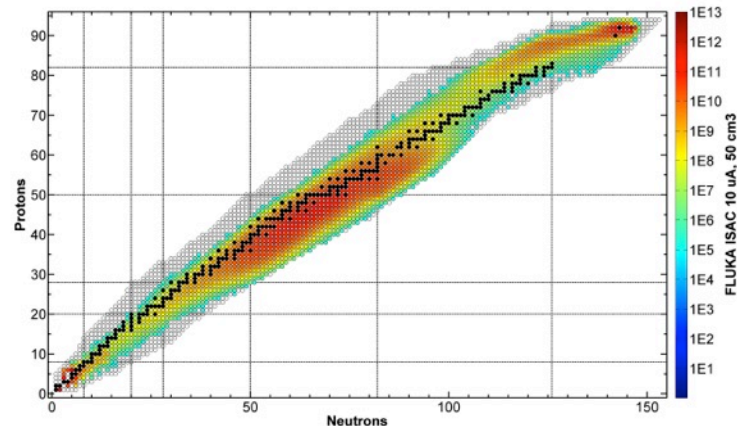
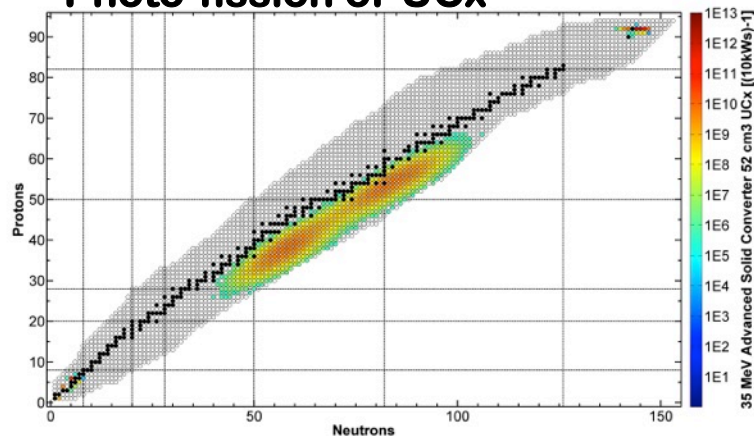
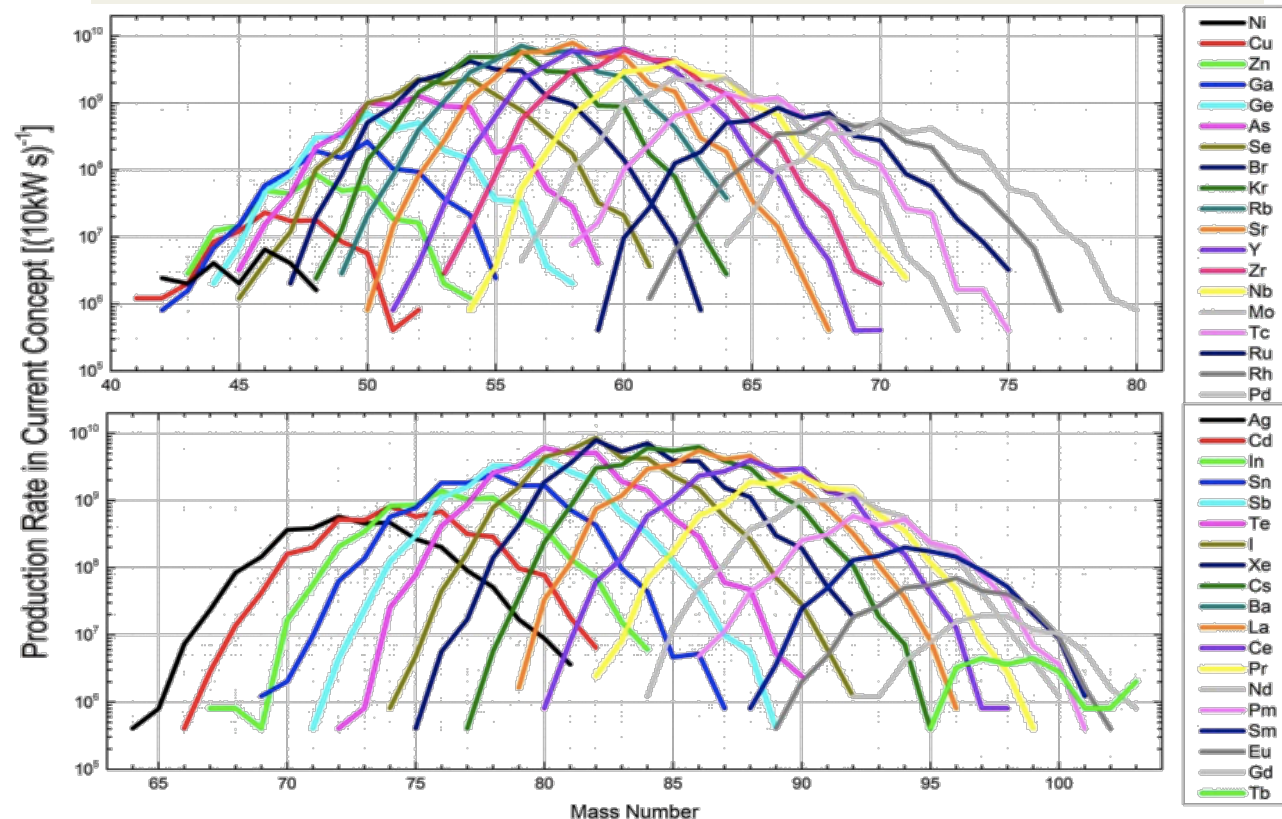


Photo-fission of UCx



ARIEL Current Concept Design In-Target Production Yields [$10 \text{ kW}^{-1} \cdot \text{s}^{-1}$]



In-target production rates
[$10 \text{ kW}^{-1} \cdot \text{s}^{-1}$]:

from BeO:

$${}^8\text{Li}: 5 \cdot 10^{10}$$

from UC_x :

$${}^{78}\text{Ni}: 1 \cdot 10^5$$

$${}^{98}\text{Kr}: 8 \cdot 10^7$$

$${}^{100}\text{Rb}: 1 \cdot 10^8$$

$${}^{98}\text{Sr}: 5 \cdot 10^9$$

$${}^{132}\text{Sn}: 5 \cdot 10^8$$

$${}^{146}\text{Xe}: 2 \cdot 10^7$$

$${}^{144}\text{Ba}: 5 \cdot 10^9$$

$${}^{150}\text{Cs}: 4 \cdot 10^5$$

**500MeV
Protons on
 UC_x :**

$${}^{78}\text{Ni}: 2 \cdot 10^6$$

$${}^{98}\text{Kr}: 1 \cdot 10^8$$

$${}^{100}\text{Rb}: 9 \cdot 10^7$$

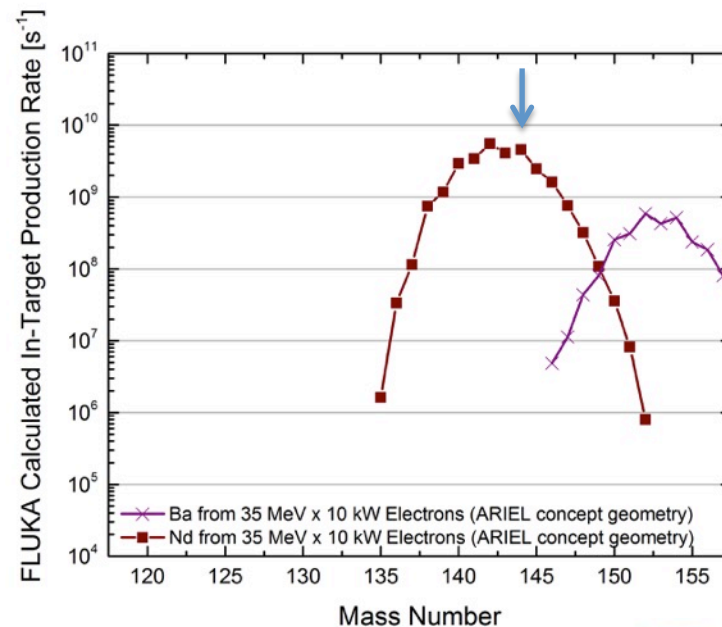
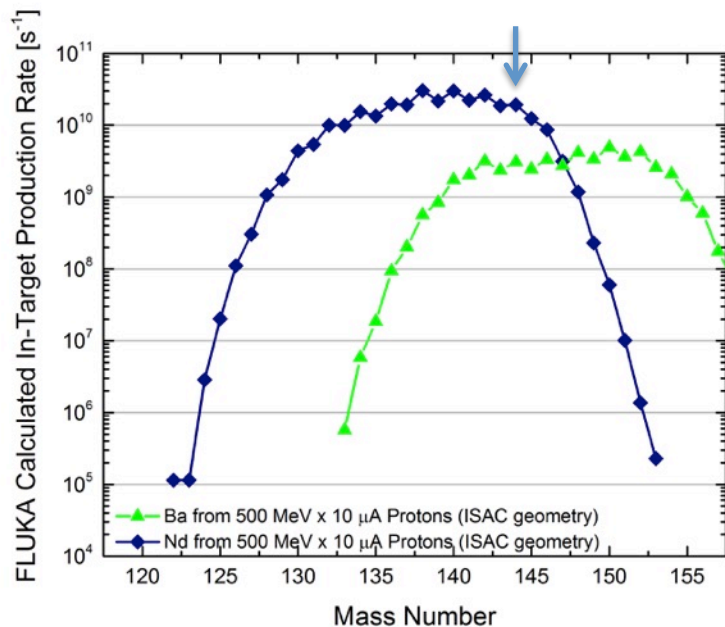
$${}^{98}\text{Sr}: 1 \cdot 10^{10}$$

$${}^{132}\text{Sn}: 5 \cdot 10^9$$

$${}^{146}\text{Xe}: 1 \cdot 10^7$$

$${}^{144}\text{Ba}: 2 \cdot 10^{10}$$

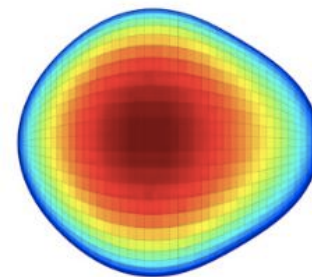
$${}^{150}\text{Cs}: 5 \cdot 10^5$$



^{144}Ba is doubly-magic for octupole deformation; $Z=56$, $N=88$.

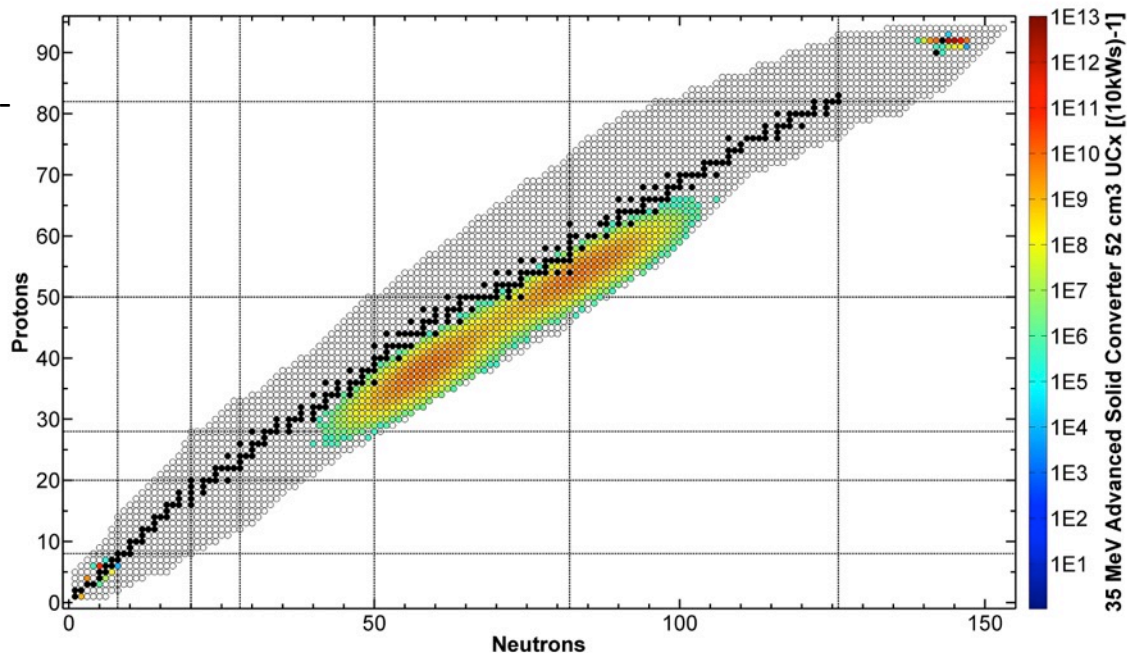
500MeV protons: 2×10^{10} with 3×10^9 Nd (and Ce, Pr, Pm, Sm, Eu, Gd etc)

10kW electrons: 5×10^9 with zero Nd



- Beams from electron-induced photo-fission of UC_x
- Intense, clean neutron-rich beams
- Lots of interesting physics:
r-process, shell structure, quadrupole deformation, collectivity, Octupole deformation, phase transitions

Photo-fission of UC_x



Call to Users:

What are the first experiments, science goals and priorities with beams from photo-fission?

- More beamtime availability with 3 simultaneous beams.
- More time available for beam development activities.
- Astrophysics – longer beamtimes means more precision and more complete studies
- Fundamental symmetries – longer beamtimes enable precision measurements
- Opportunities limited without a Medium Resolution Separator (MRS), in most cases only 2 beams can be utilized simultaneously. IG-LIS beams may be ok.
- Opportunities limited by only one RFQ-DTL, can only deliver to either medium or high-energy area.

- Longer beamtimes means more precision and more complete studies

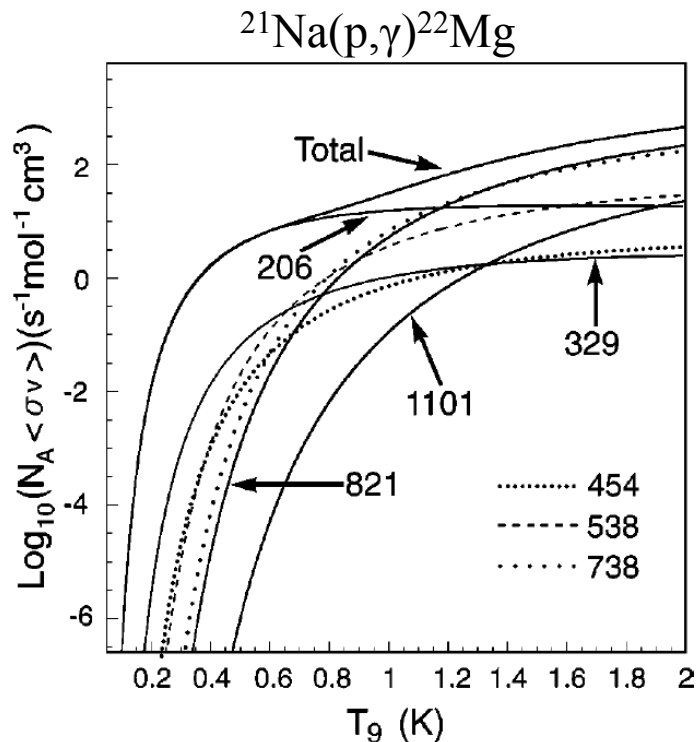
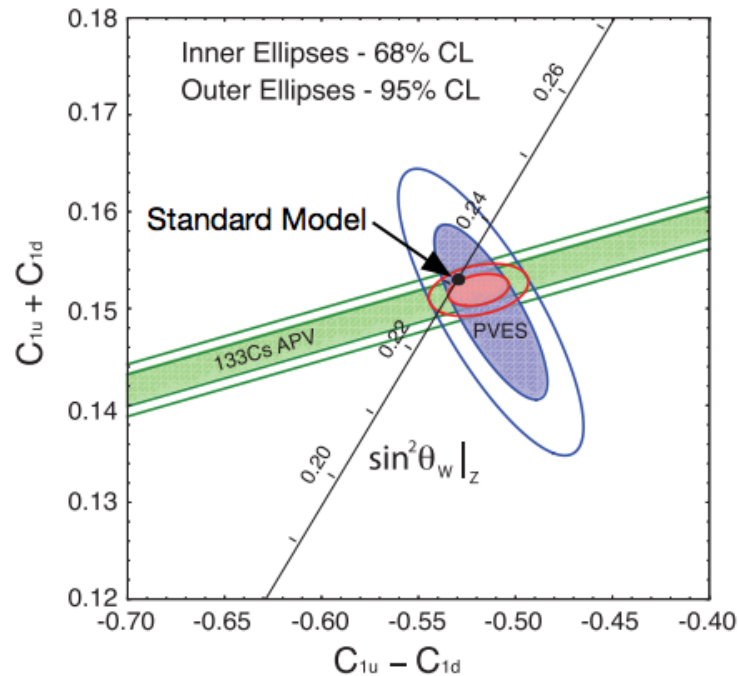
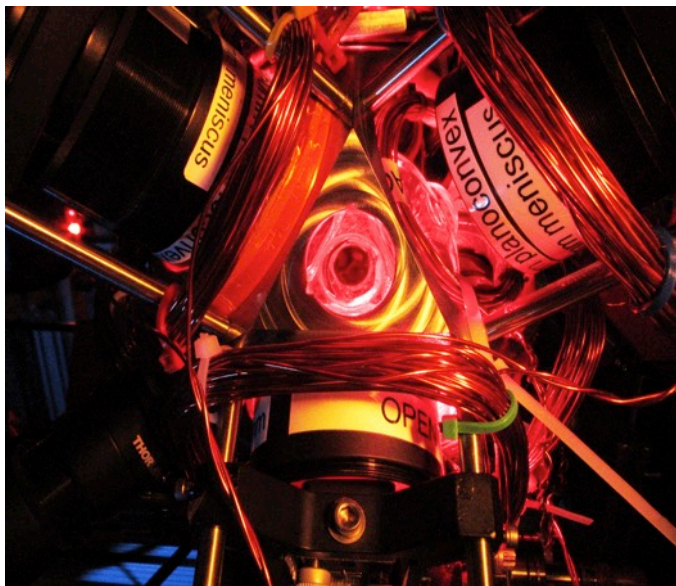


TABLE I. $^{21}\text{Na}(p,\gamma)^{22}\text{Mg}$ resonance strengths and energies.

| E_x (MeV) | $E_{c.m.}$ (keV) | Γ (keV) | $\omega\gamma$ (meV) |
|-------------|------------------|----------------|----------------------|
| 5.714 | 205.7 ± 0.5 | | 1.03 ± 0.21 |
| 5.837 | 329 | | ≤ 0.29 |
| 5.962 | 454 ± 5 | | 0.86 ± 0.29 |
| 6.046 | 538 ± 13 | | 11.5 ± 1.36 |
| 6.246 | 738.4 ± 1.0 | | 219 ± 25 |
| 6.329 | 821.3 ± 0.9 | 16.1 ± 2.8 | 556 ± 77 |
| 6.609 | 1101.1 ± 2.5 | 30.1 ± 6.5 | 368 ± 62 |

Limited beamtime usually results in only the strongest resonances being studied introducing significant uncertainty to the final reaction rate calculations.

- Longer beamtimes enables time-consuming precision measurements



Atomic Parity Violation measurements in Fr to put constraints on weak electron-quark couplings.

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- More time available for beam development activities.
- Astrophysics – longer beamtimes means more precision and more complete studies
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- Opportunities limited without a MRS, in most cases only 2 beams can be utilized simultaneously unless IG-LIS is used.
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Call to Users:

**What are the first experiments, science goals and priorities
when 3 simultaneous beams are available?**

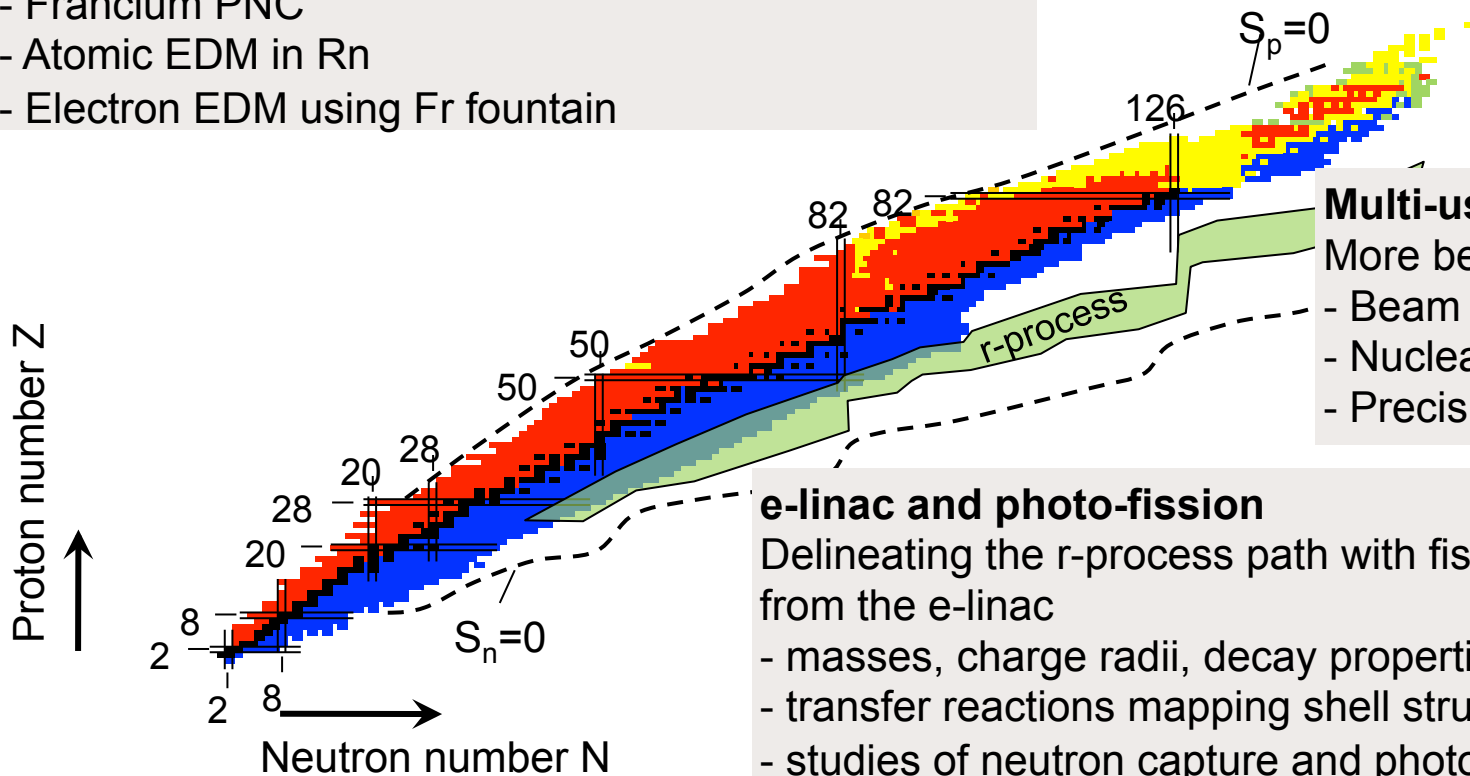
- Medium-Resolution Separator
 - 2nd RFQ-DTL accelerator path
 - New ISAC frontend
 - 50MeV eLINAC
 - 500kW converter development
 - Re-circulation ring
 - ...
- Storage ring
 - HELIOS-type device
 - Total-Absorption Spectrometer
 - ...

**User consultation in the summer as part of next TRIUMF 5-year planning process.
Starting thinking about this.**

Actinide proton beam-line:

High intensity, clean beams for electroweak precision experiments using hundreds of days of beam per year

- Francium PNC
- Atomic EDM in Rn
- Electron EDM using Fr fountain



Multi-user operations:

- More beam time for
- Beam development
 - Nuclear astrophysics
 - Precision experiments

e-linac and photo-fission

Delineating the r-process path with fission fragment beams from the e-linac

- masses, charge radii, decay properties
- transfer reactions mapping shell structure
- studies of neutron capture and photo dissociation rates

Call to Users:

Please help define the science goals and priorities for the different stages of ARIEL:

- More intense, cleaner accelerated beams from ISAC-CANREB-ISAC,
- Increase in polarized Li beam availability to beta-NMR, OSAKA, MTV
- Beams from electron-induced Photo-fission of UC_x ,
- when 3 simultaneous beams are available.

What are the priorities for future upgrades and installations?

Any other suggestions please.



Canada's national laboratory for
particle and nuclear physics

Laboratoire national canadien
pour la recherche en physique
nucléaire et en physique des
particules

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Victoria | Western | Winnipeg | York

Thank you!
Merci!

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