

The SuperCDMS experiment at SNOLAB

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TRIUMF Science Week 2024

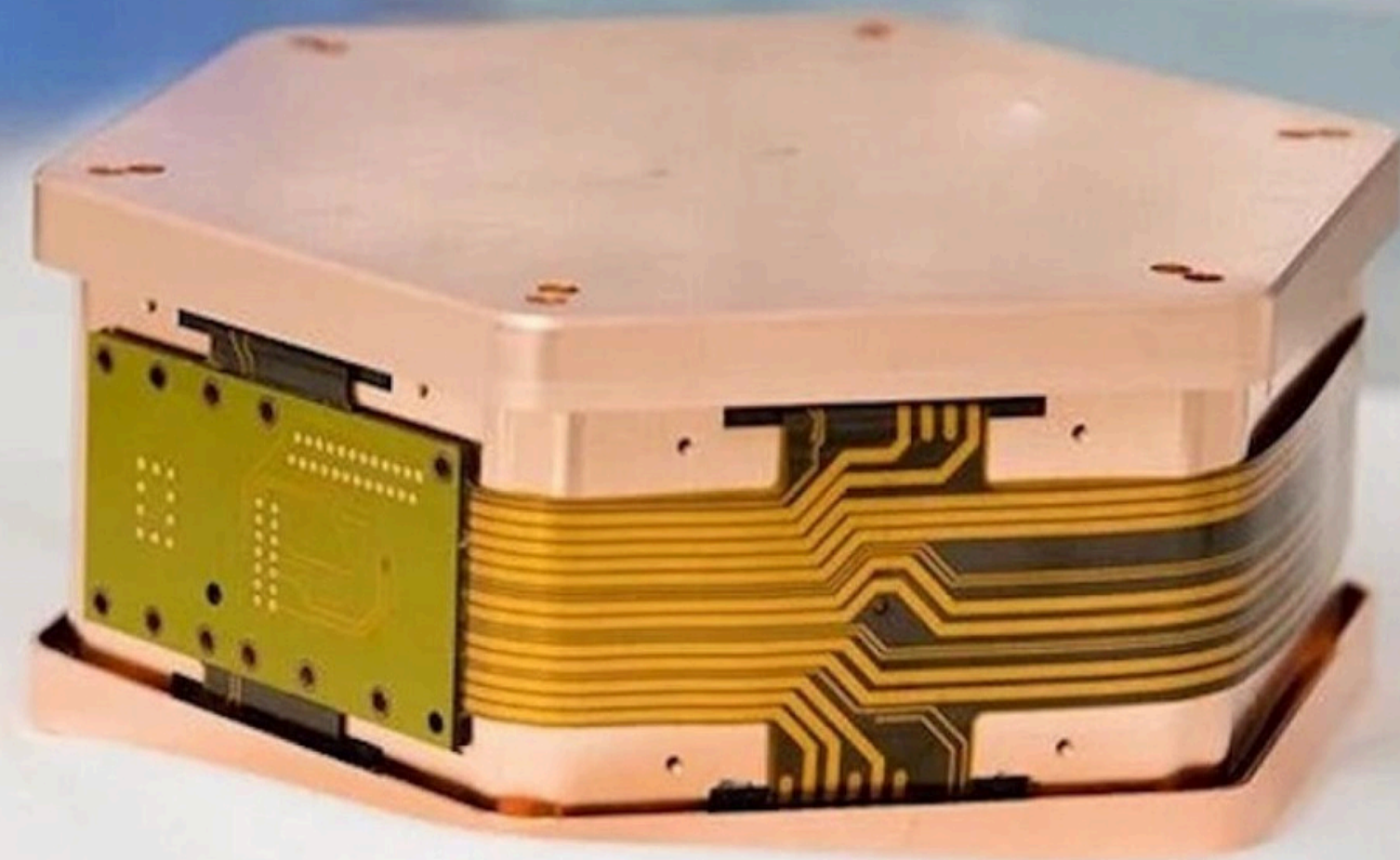
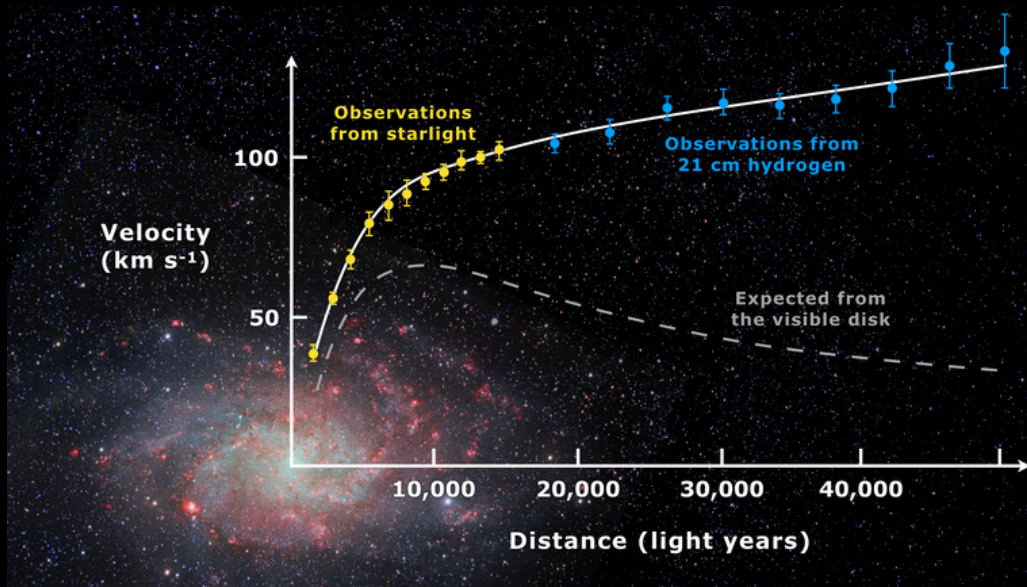


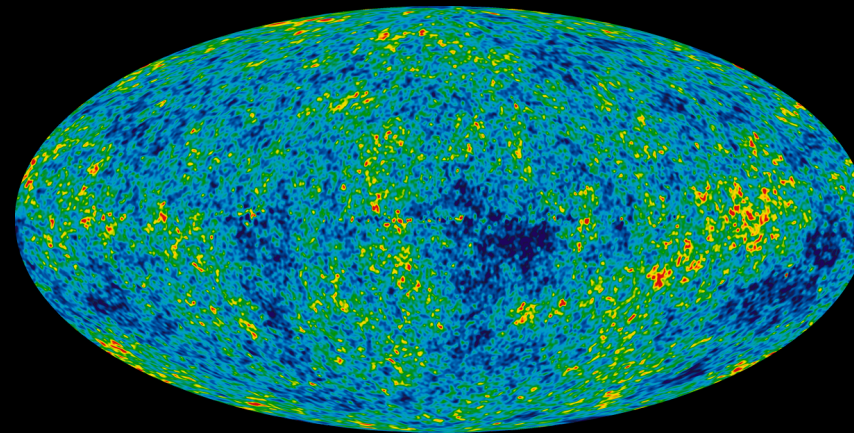
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- 2.** The SuperCDMS SNOLAB experiment
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Evidence for Dark Matter and its properties

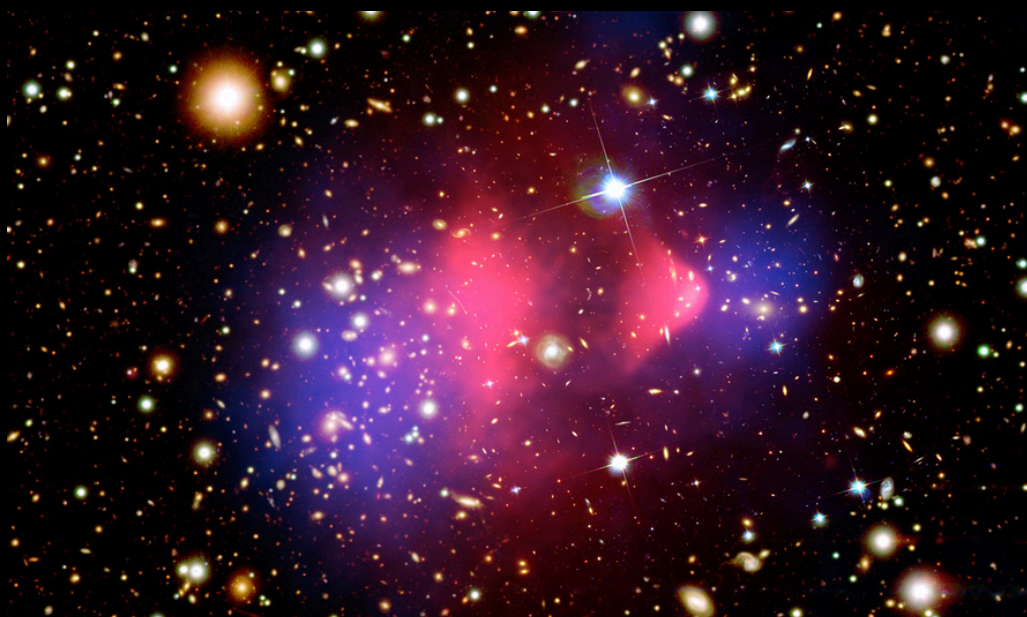


discrepancy in galactic rotational curve

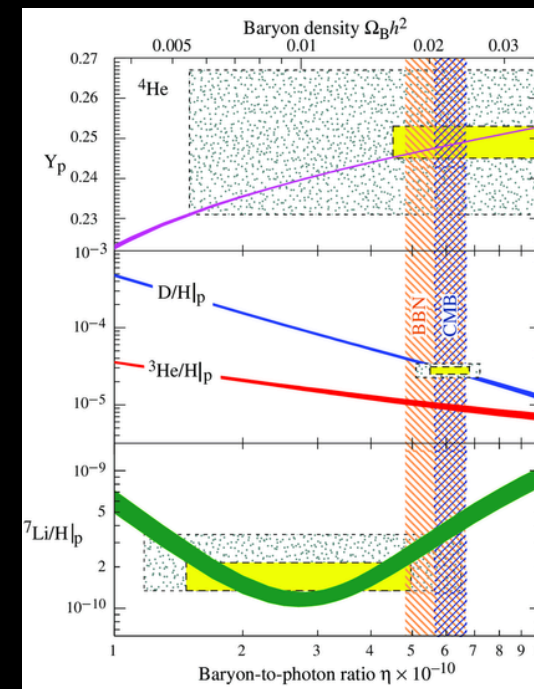


Cosmic Microwave Background

- Dark
- Stable
- Non-baryonic
- Not "hot"
- ~25% of the total universe mass



the Bullet cluster

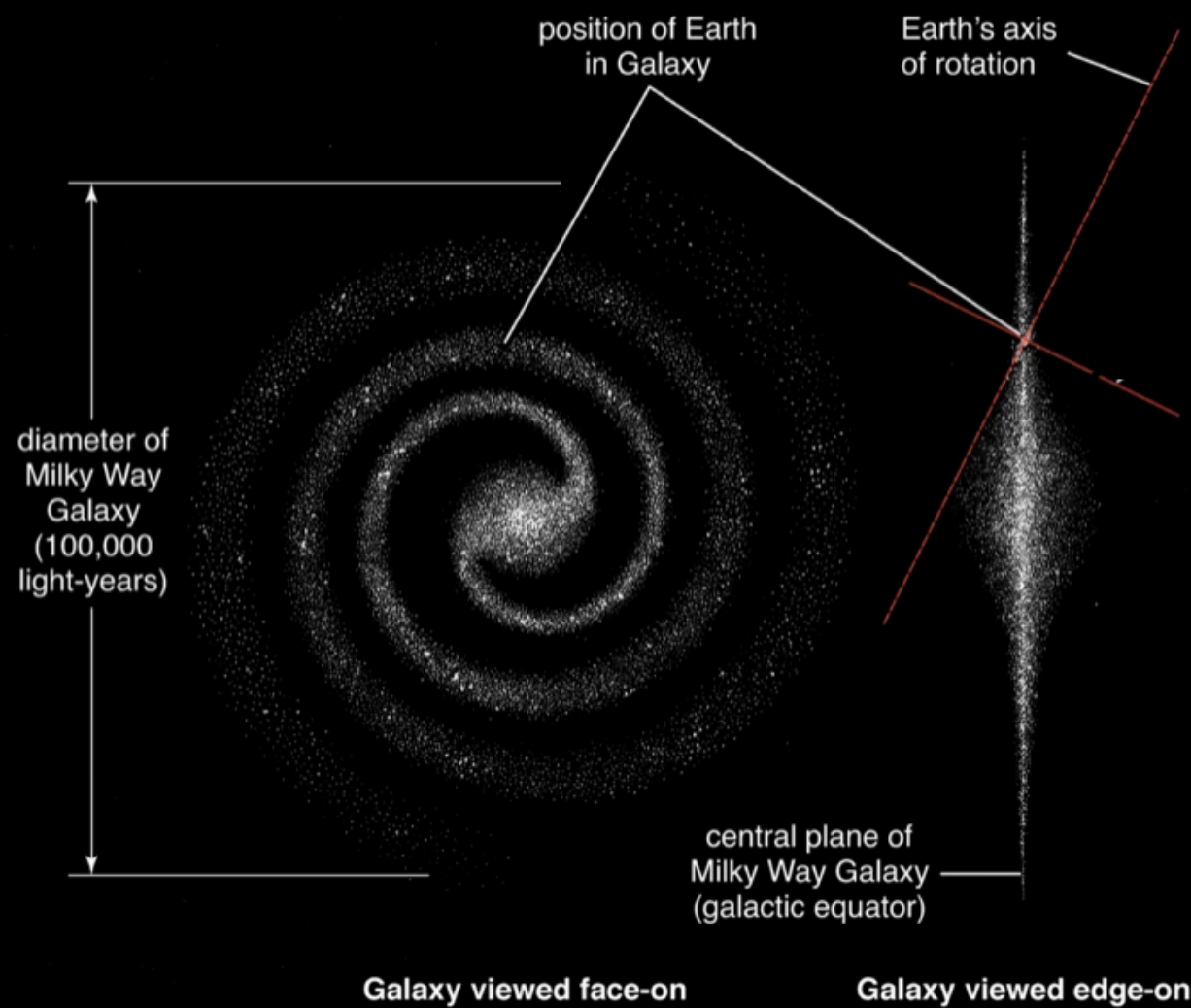


Big Bang Nucleosynthesis

In fact, we can assume some more...

Standard Halo Model

The Milky Way Galaxy



- Local velocity (velocity of Dark Matter at the Earth)

$$\sim 10^{-3} c$$

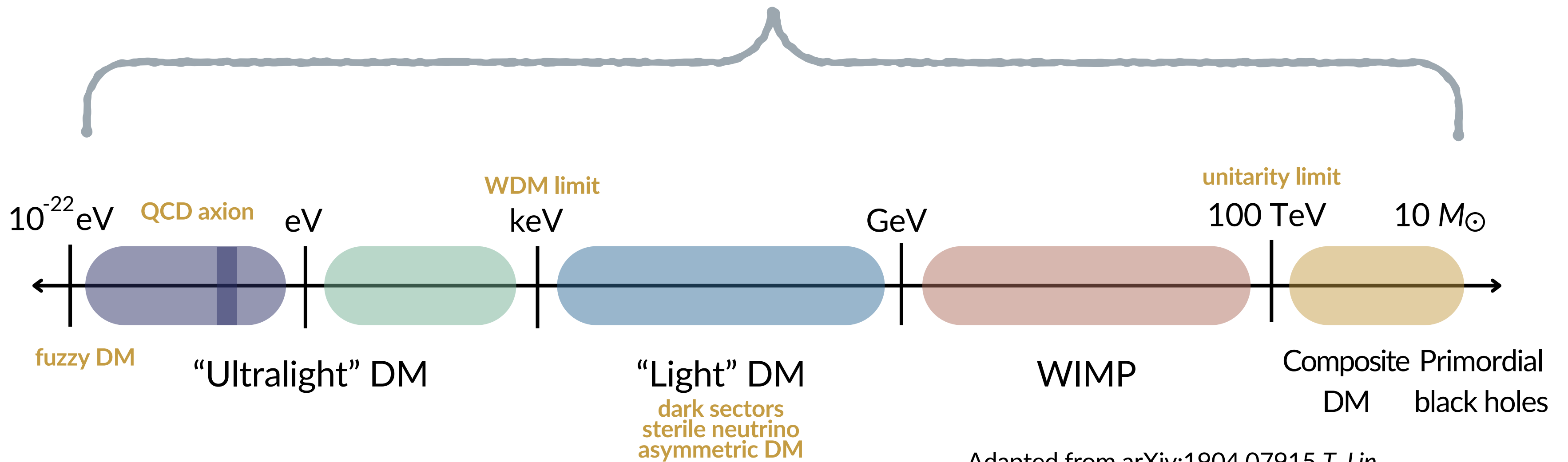
- Local density (density of Dark Matter at the Earth)

$$\sim 0.3 \text{ GeV}/c^2/\text{cm}^3$$

$$\text{roughly } 10^{-28} \text{ kg}/\text{cm}^3$$

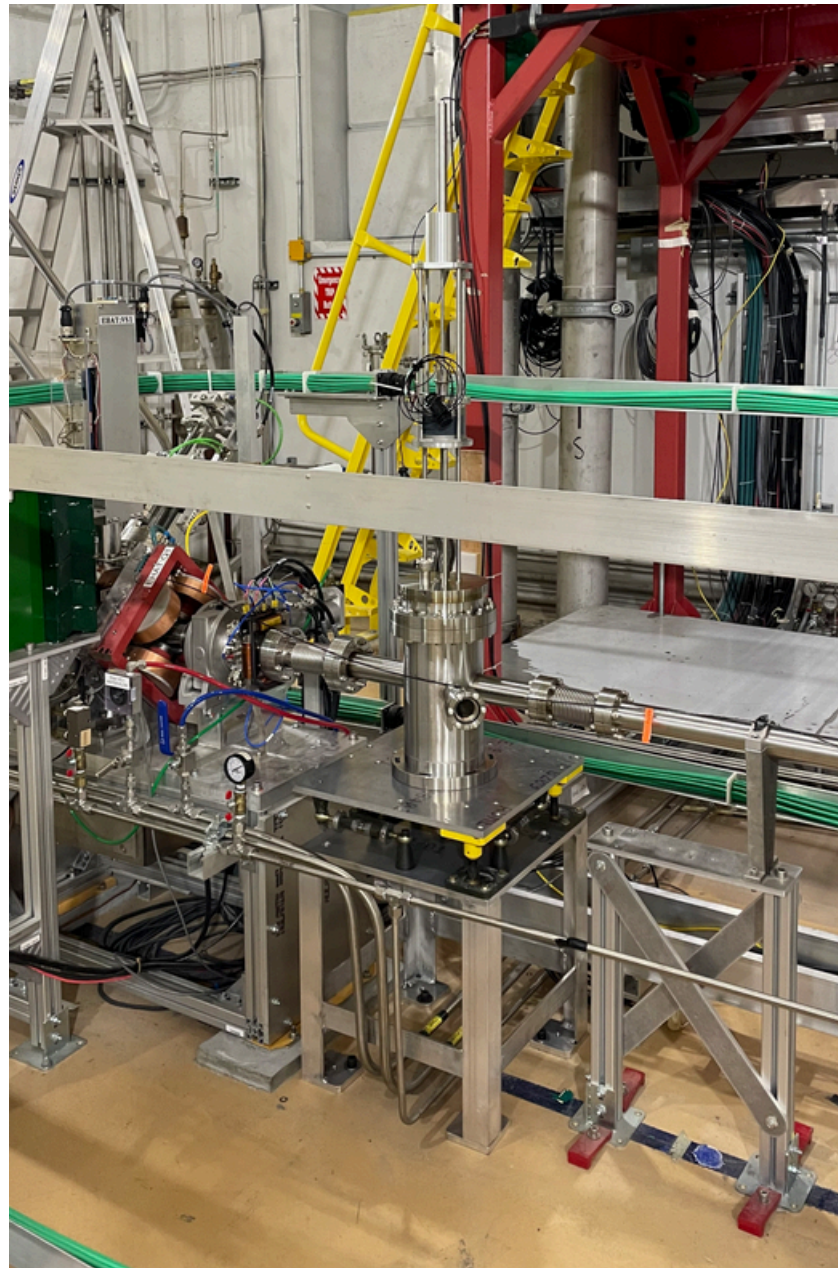
Dark Matter Mass

spanning across ~88 orders of magnitude

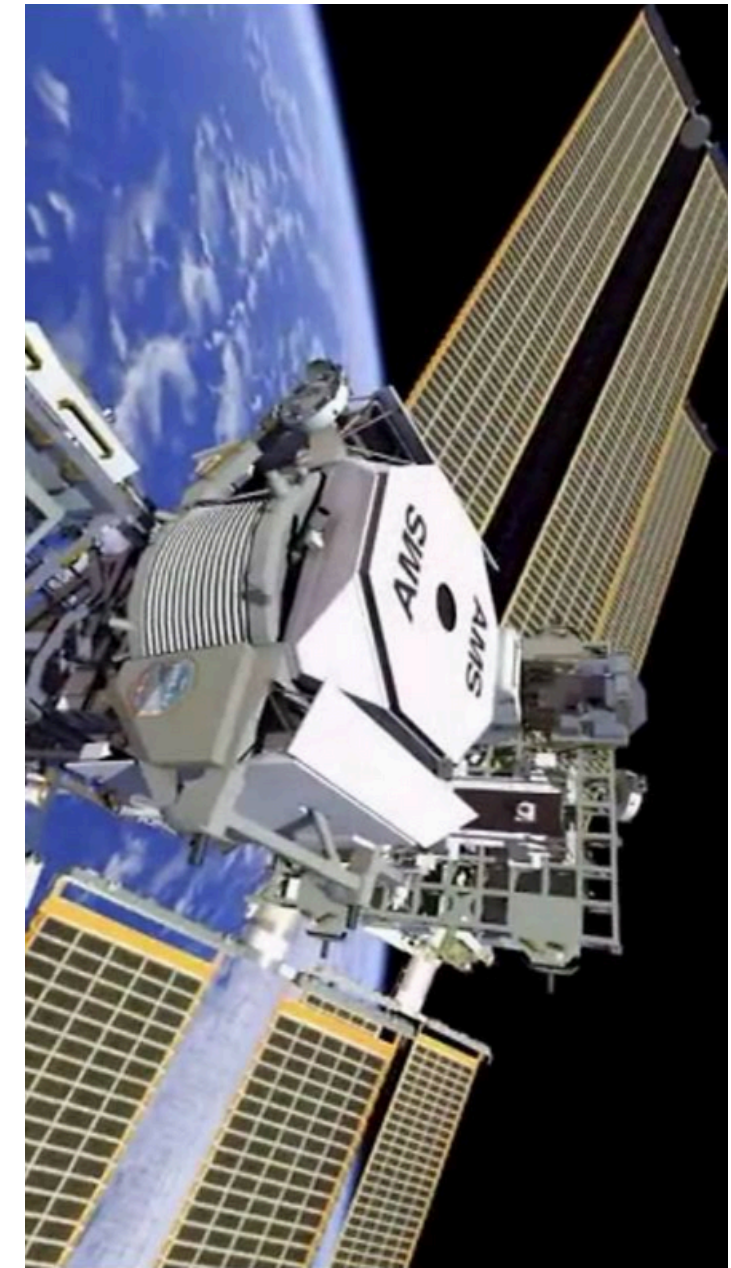
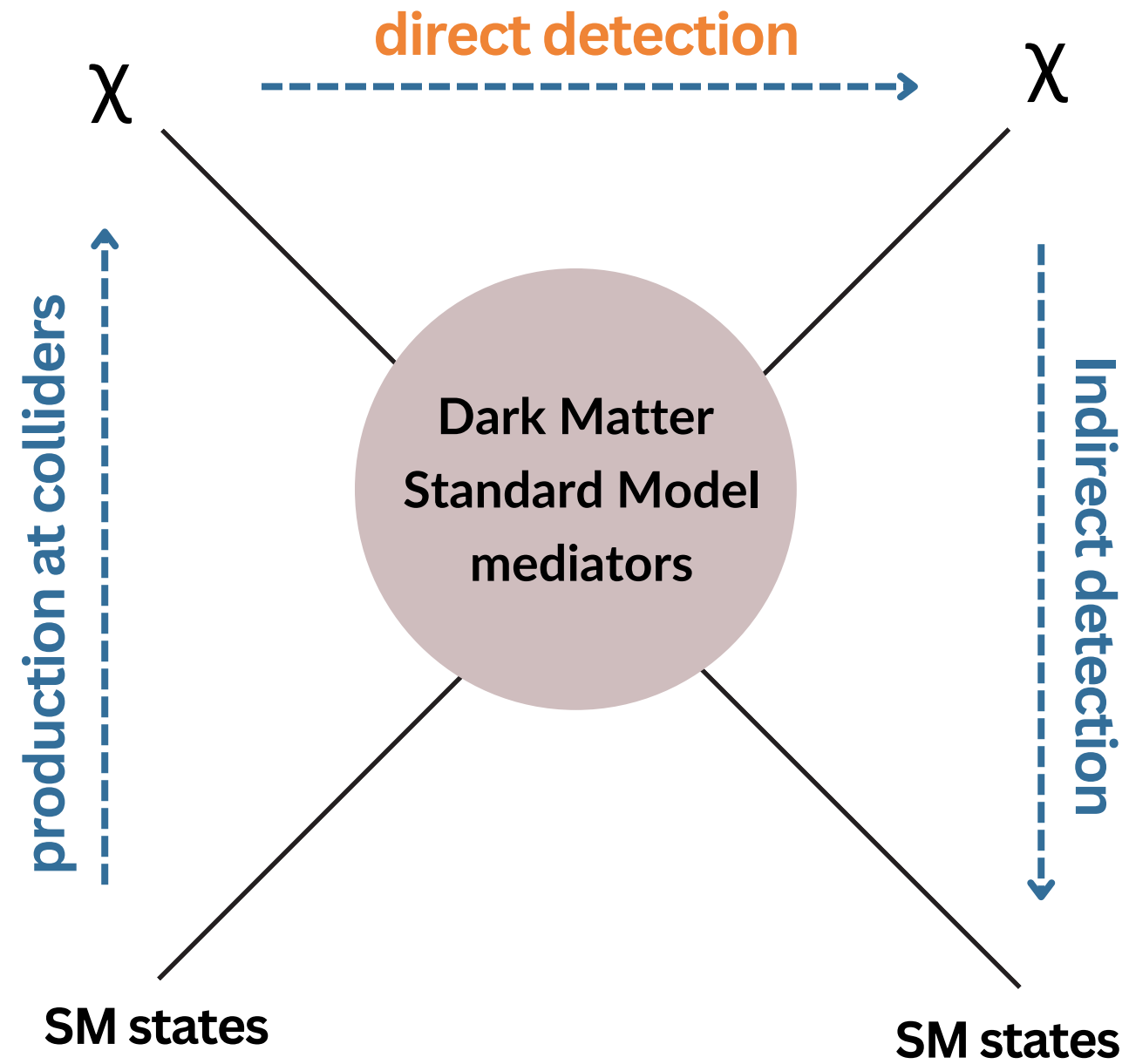


Adapted from arXiv:1904.07915 T. Lin

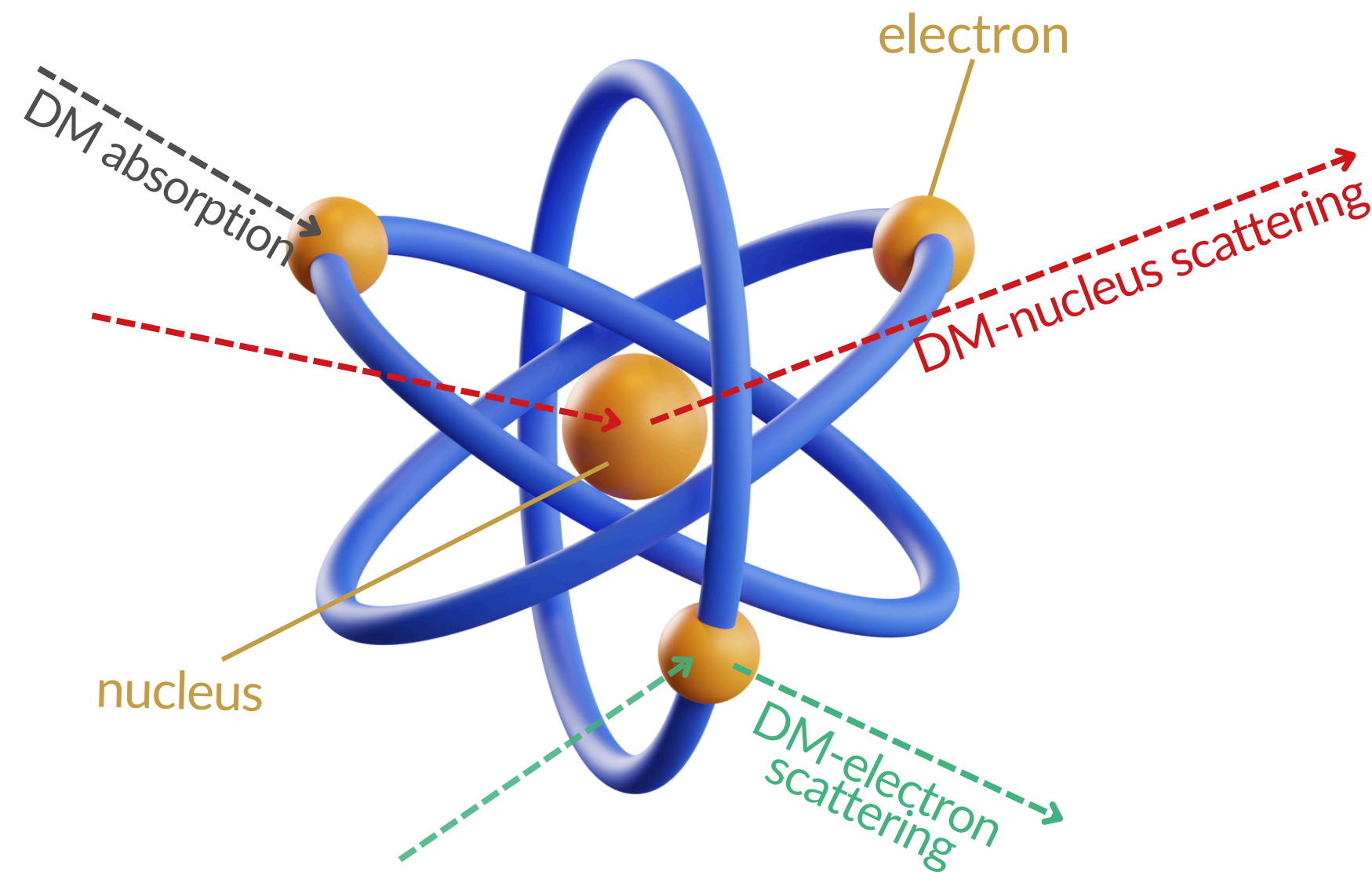
Make it, Break it, or Shake it



DarkLight@TRIUMF



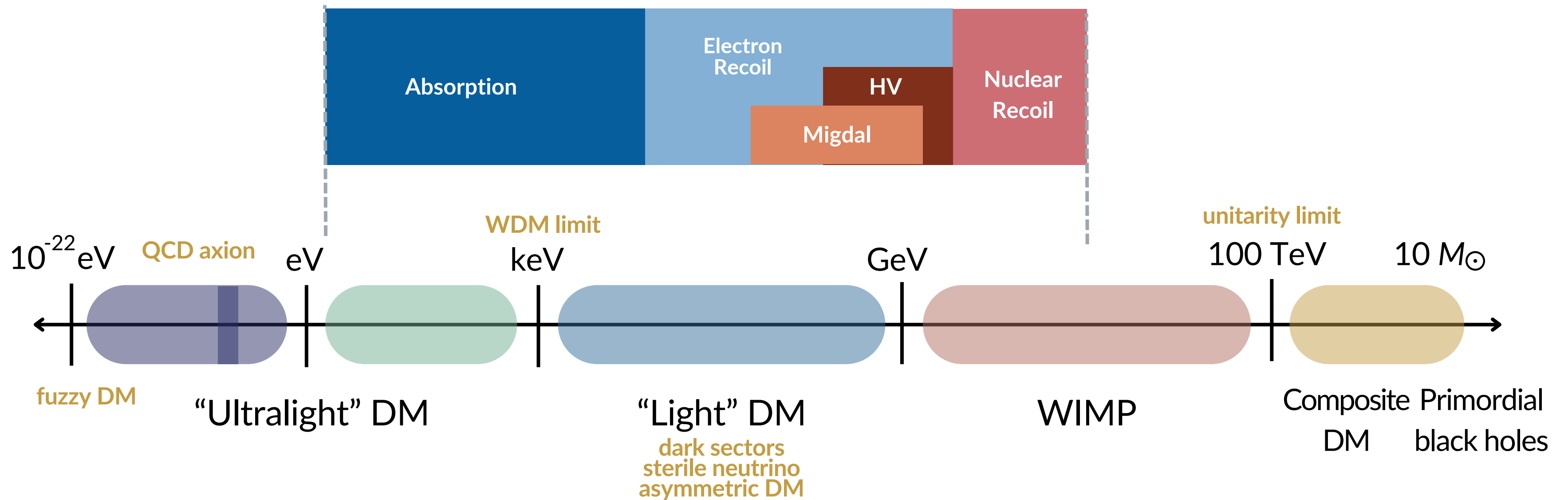
Typical Schemes for DM Direct Search



- **DM-nucleus scattering**
 - typically sensitive to $> \text{GeV}$ dark matter if assuming elastic interaction
- **DM-electron scattering**
 - inelastic process allowing full energy deposition
 - suitable for $< \text{GeV}$ dark matter search
- **DM absorption**
 - Clear spectral line feature
 - mass sensitivity limited by threshold

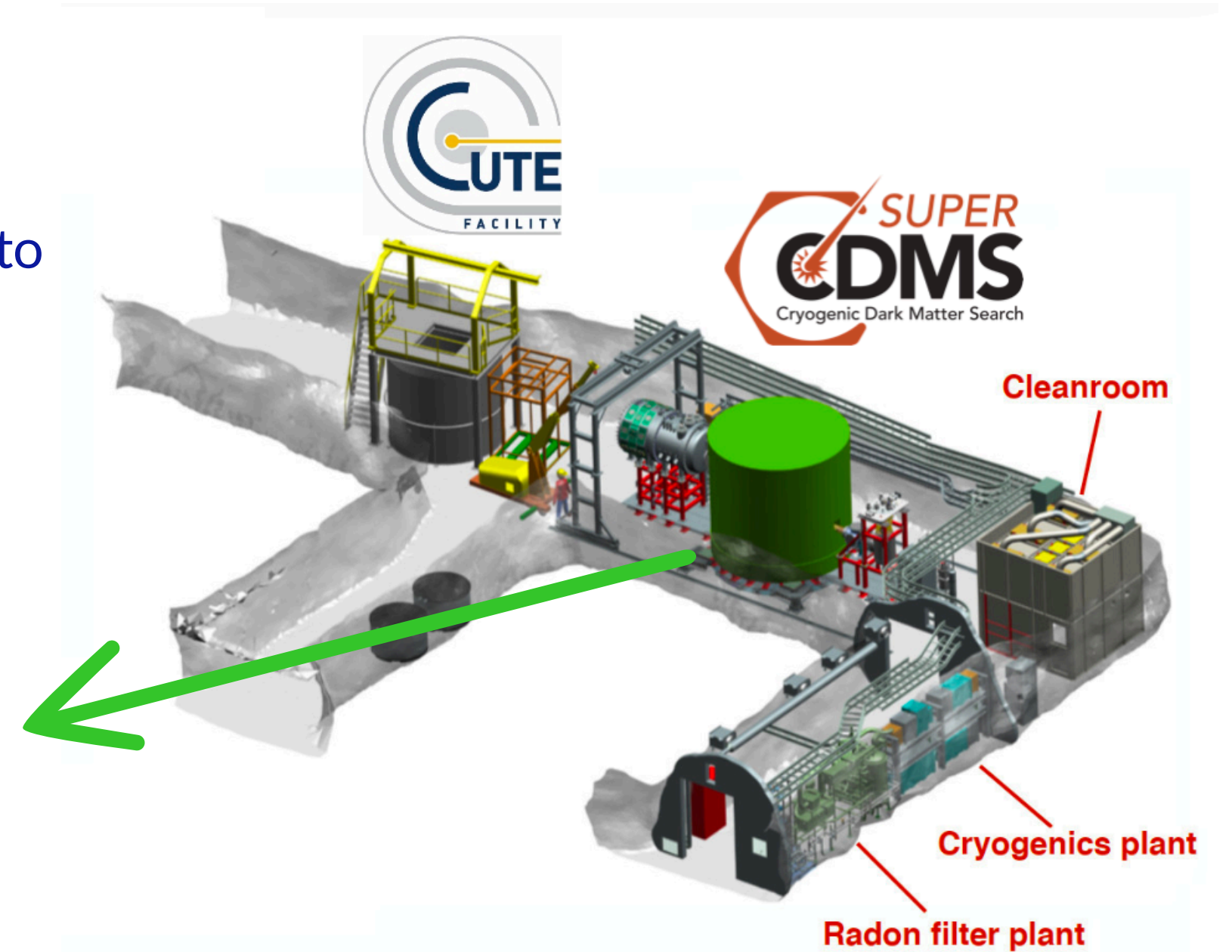
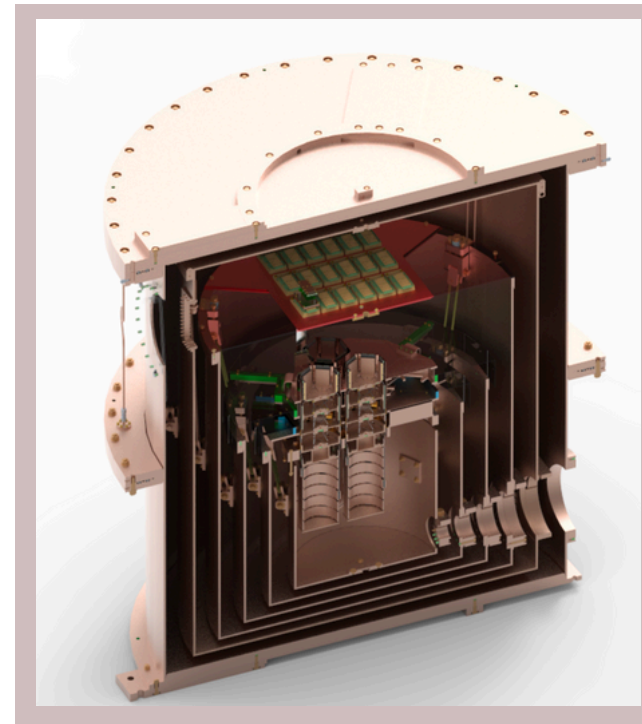
SuperCDMS as a Dark Matter Experiment

SuperCDMS SNOLAB is sensitive to various dark matter models across **10 orders of magnitude!**



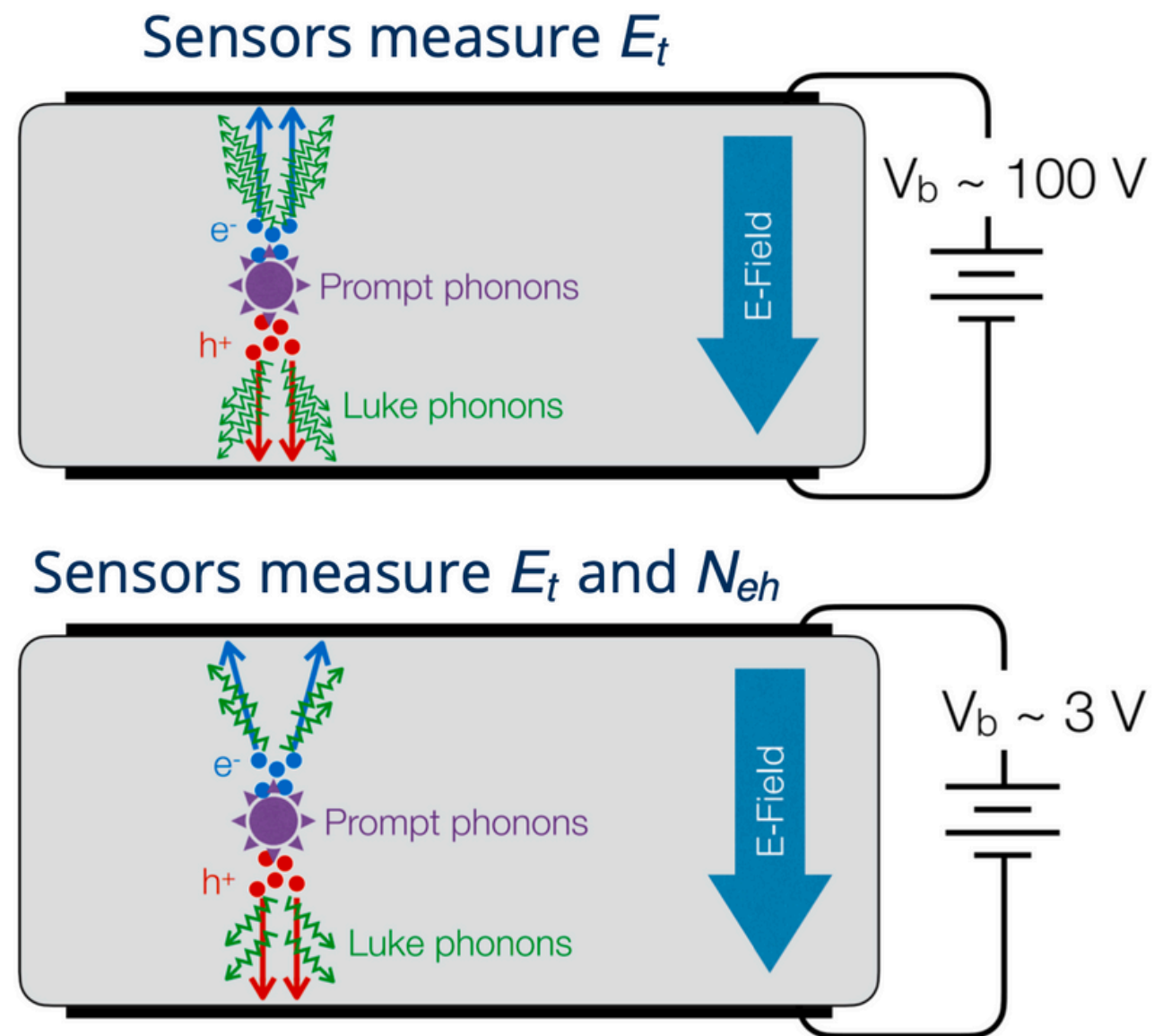
SuperCDMS SNOLAB Experiment

- US Generation 2 Dark Matter Experiment
- 2km underground in Sudbury, Canada
- Germanium and Silicon crystals cooled down to ~ 10 mK to be sensitive to minute dark matter signals



SuperCDMS SNOLAB Detector Technology

Two different detector configurations



- **HV detectors -> low threshold**
 - NTL effect to reach lower threshold
 - highest science reach

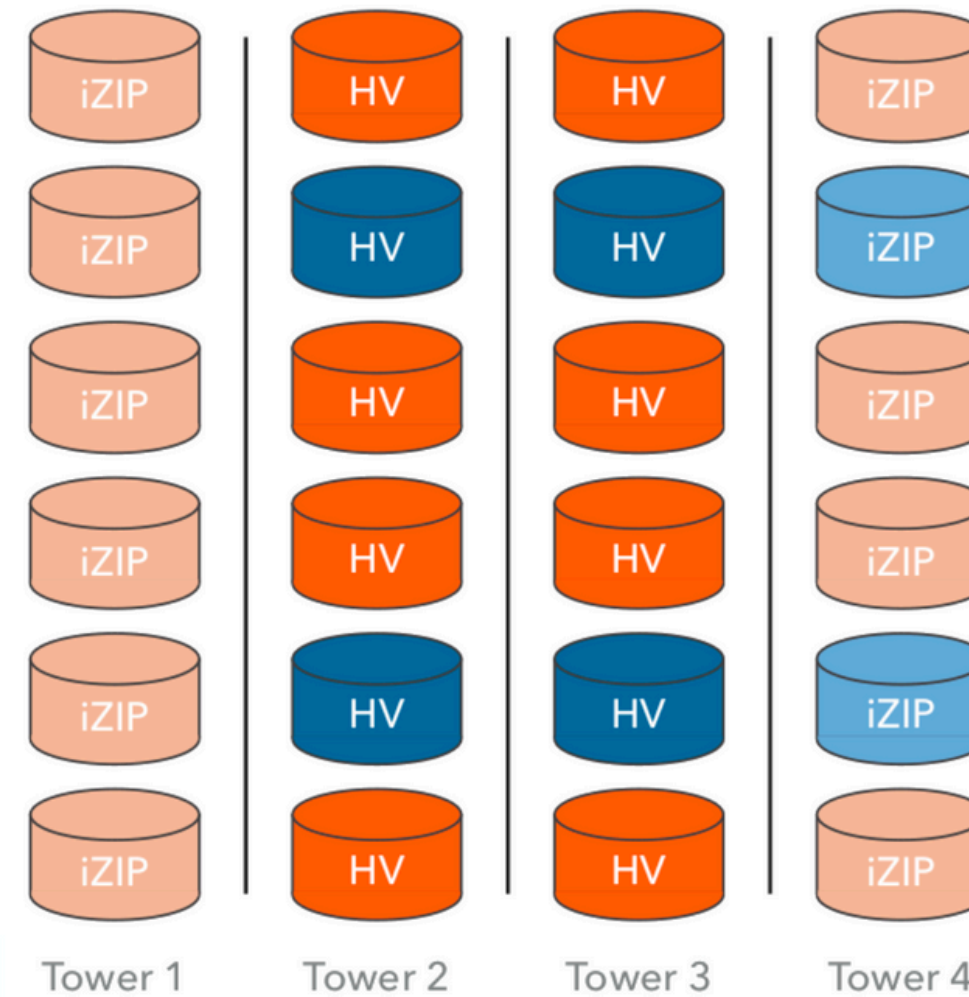
$$E_t = E_r + N_{eh}eV_b$$

E_t : total phonon energy
 E_r : primary recoil energy
 $N_{eh}eV_b$: Luke phonon energy

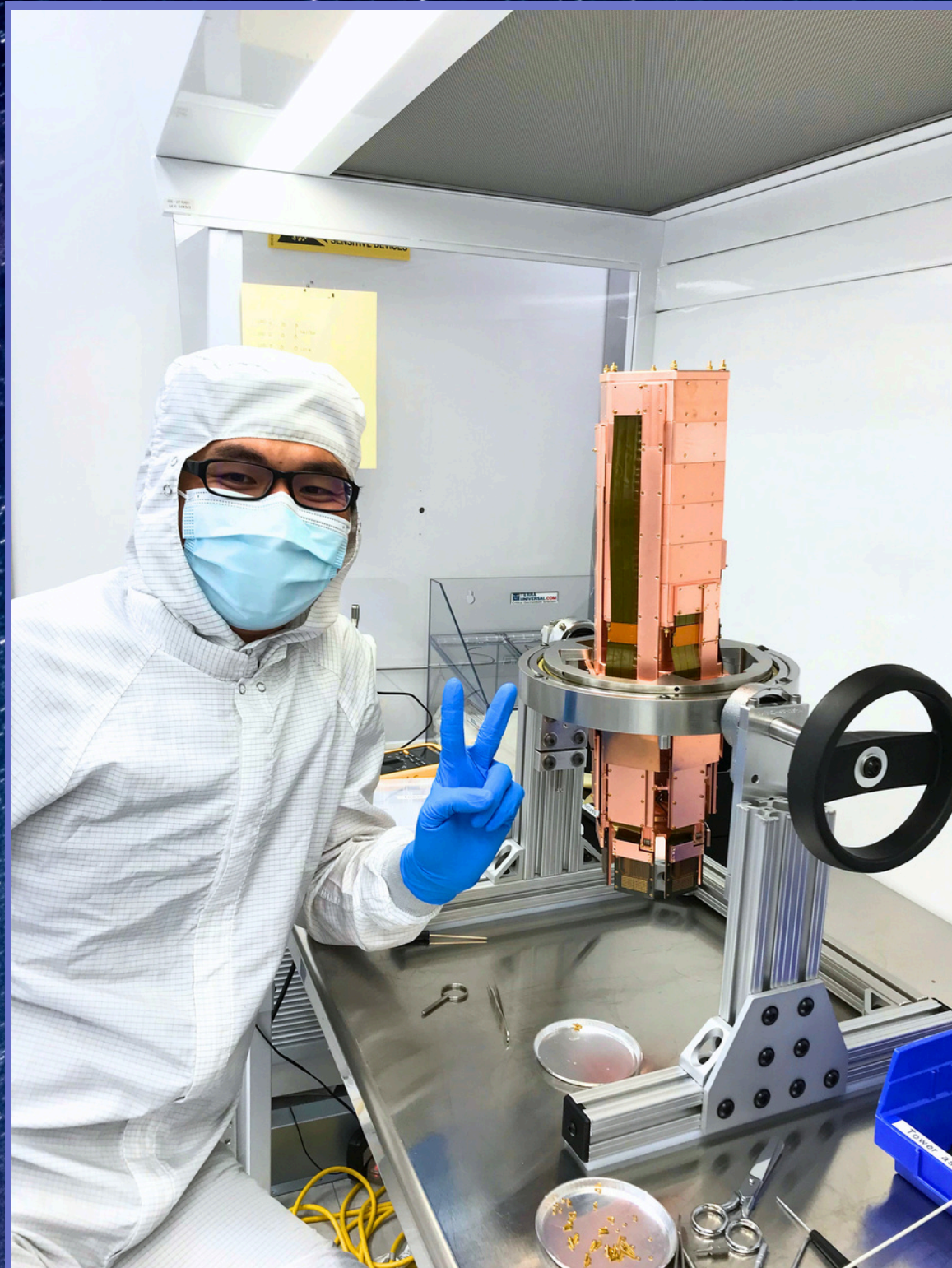
- **iZIP detectors -> low background**
 - ionization-based NR/ER background discrimination
 - higher dynamic range
 - radioactive background measurements

SuperCDMS SNOLAB Detector Technology

24 detectors were assembled in 4 towers at SLAC, and arrived at SNOLAB underground in 2023.

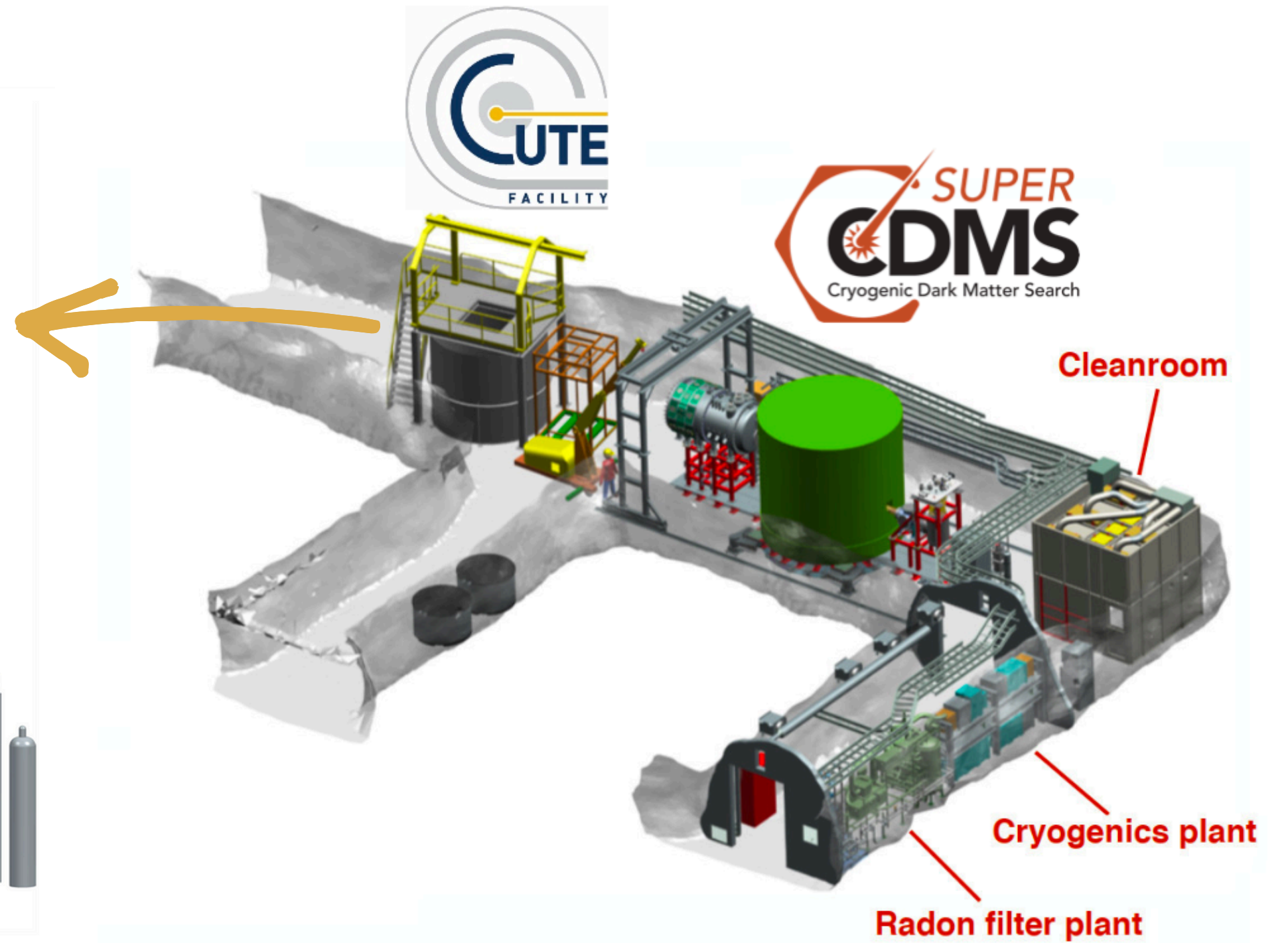
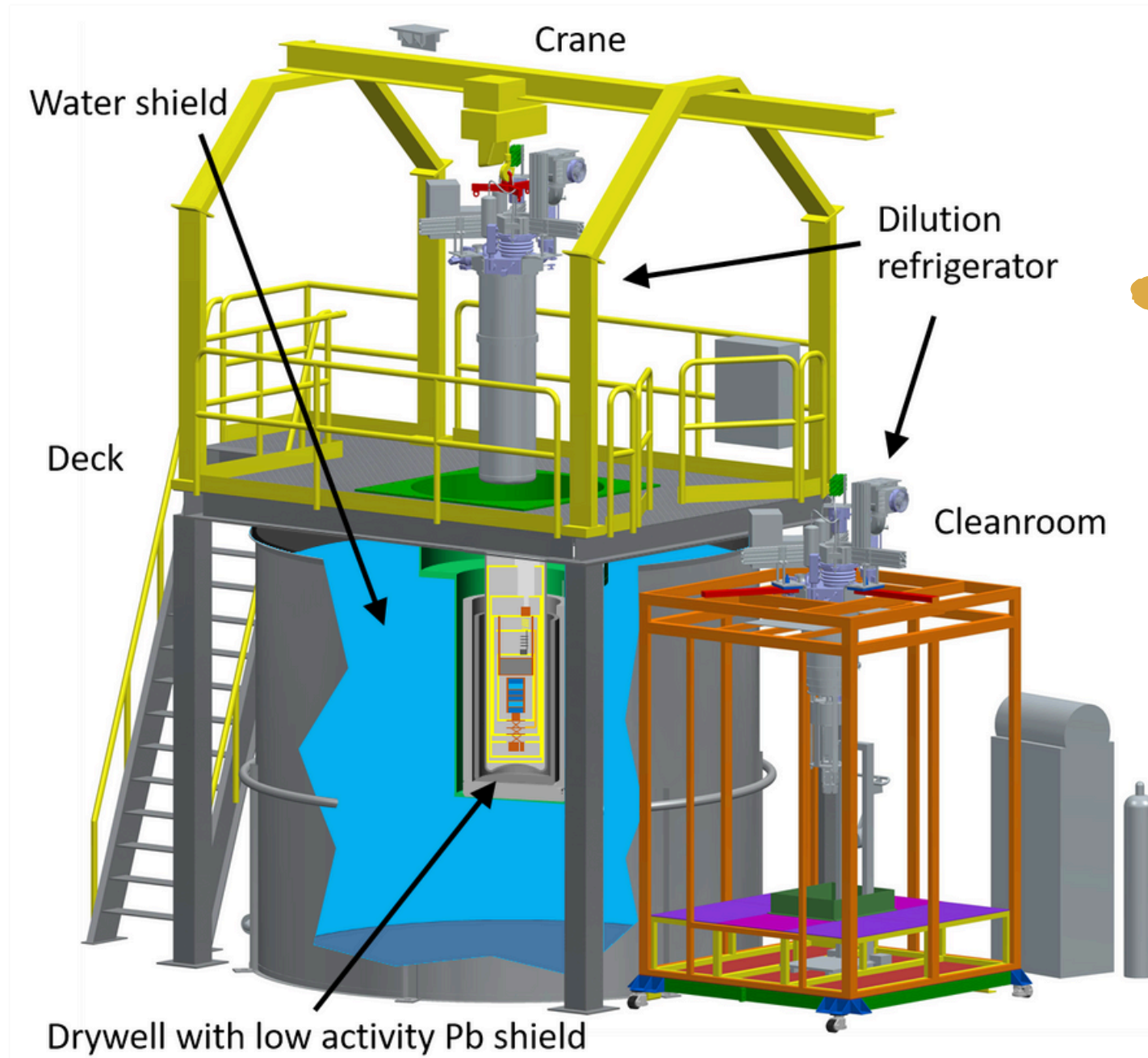


Ge: 1.4 kg
Si: 0.6 kg

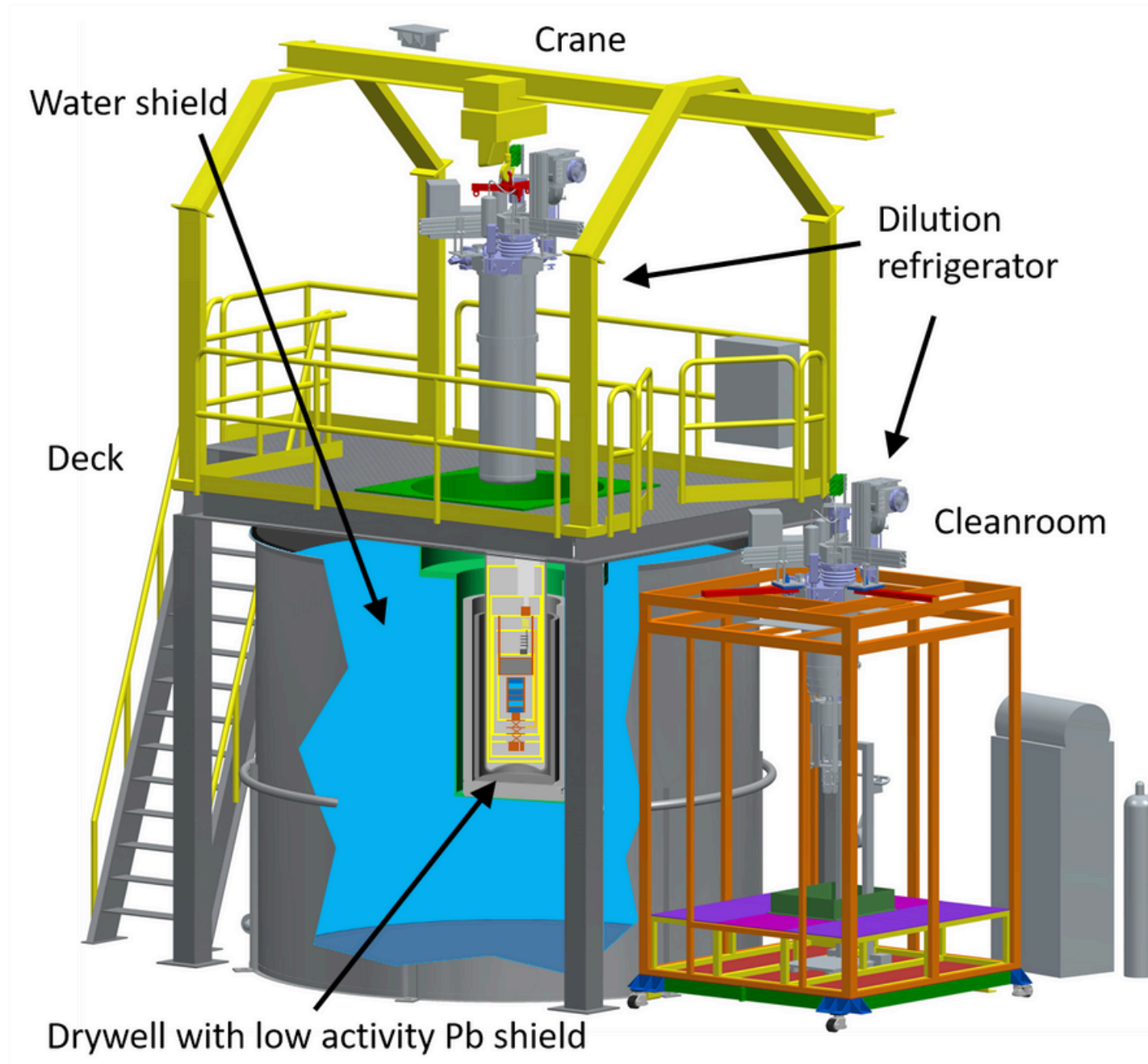


Assembling the last tower at SLAC
summer 2022

Cryogenic Underground Test Facility



Cryogenic Underground Test Facility



- CUTE is a unique deep underground, low background cryogenic test facility
 - Background level ~10 DRU
 - Suspension system to reduce vibration
- Efficient turnaround of cryogenic device testing

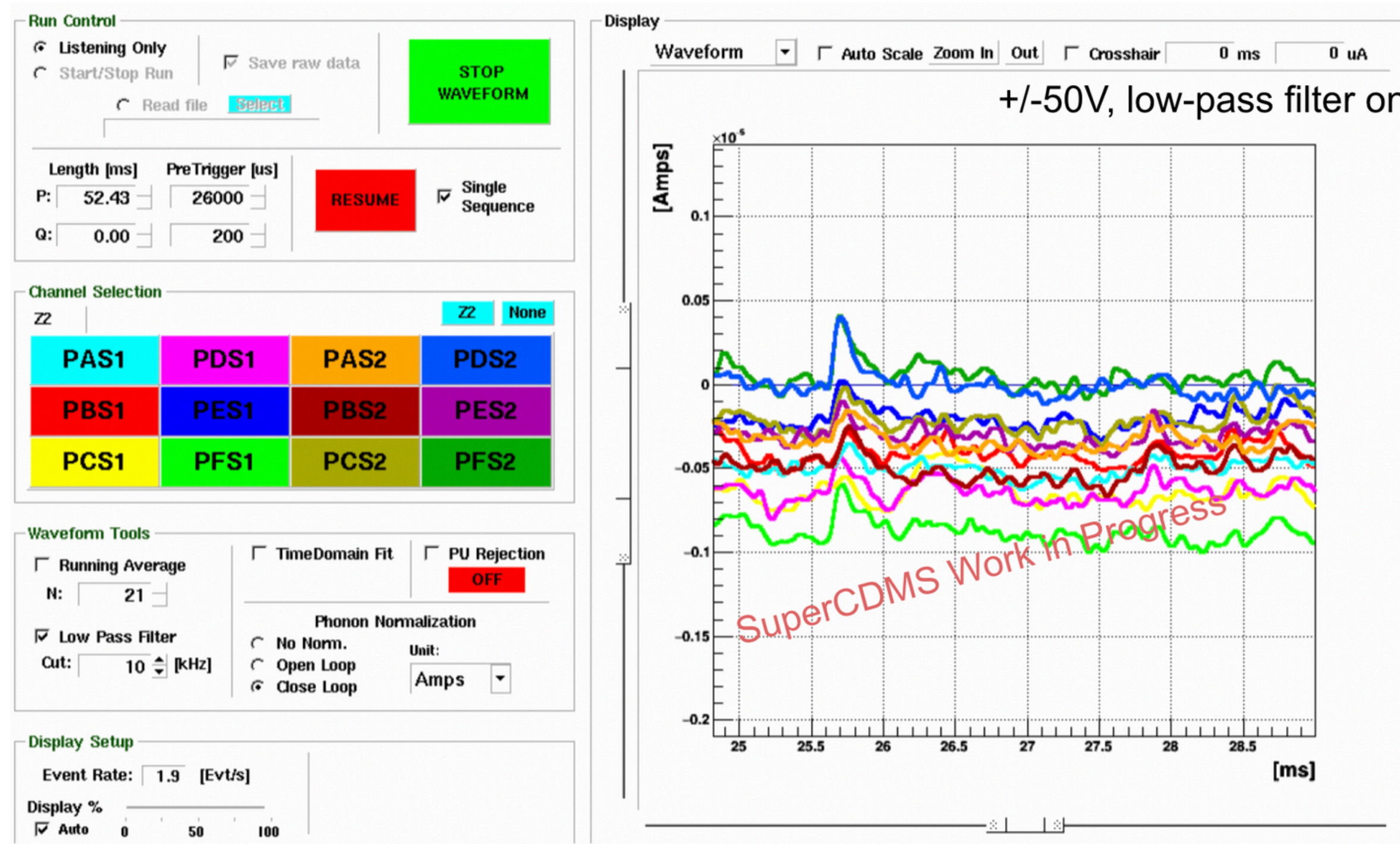
Tower testing at CUTE

- A significant milestone for the experiment
 - First underground, low background environment testing of the new generation SuperCDMS detectors
- The TRIUMF/UBC group played a leading role in the 5-month long tower testing efforts



TRIUMF/UBC group standing on top of CUTE

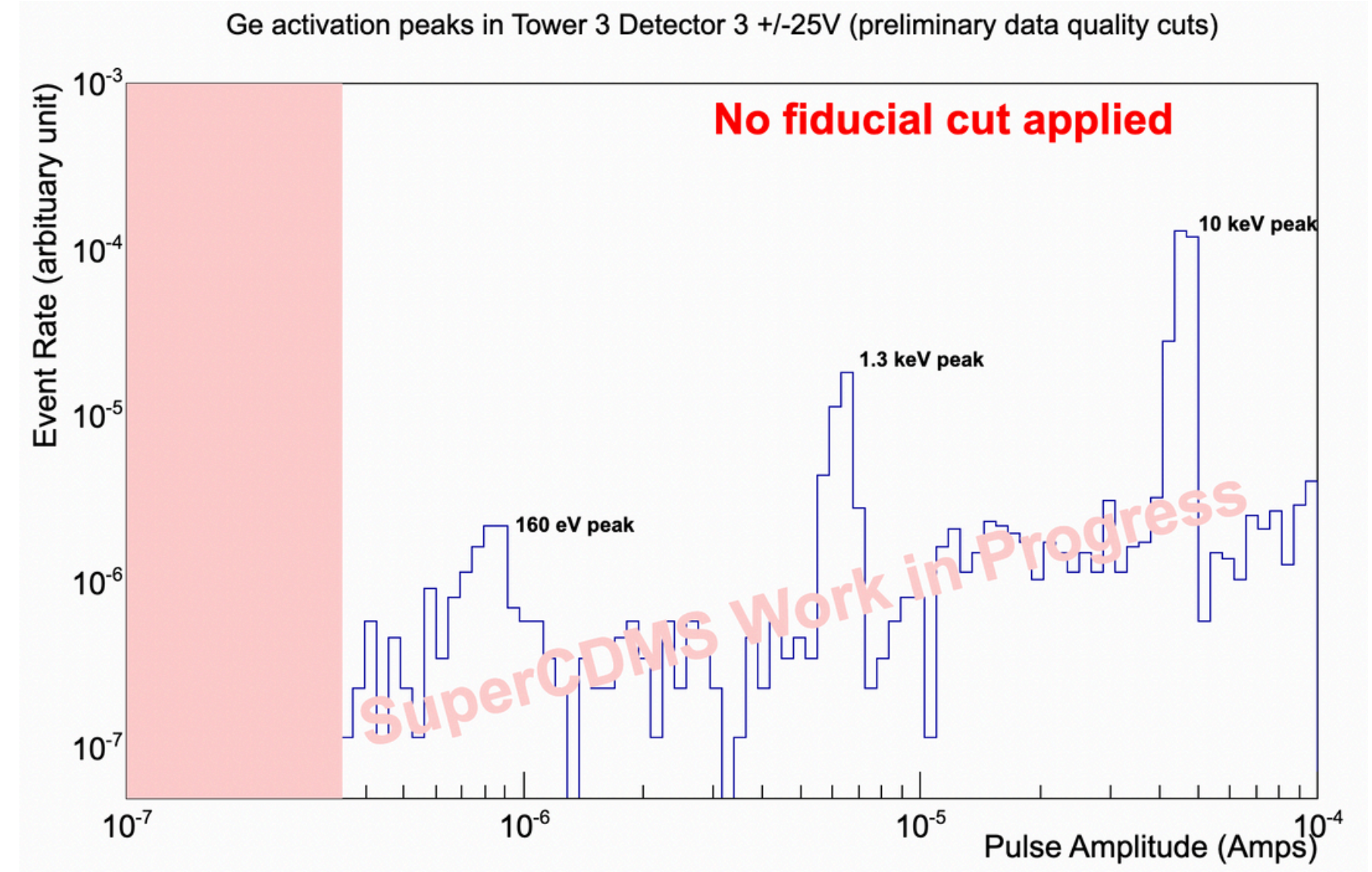
Tower Testing at CUTE: First Raw Pulses



At use is the SuperCDMS DAQ, developed at TRIUMF based on the MIDAS system

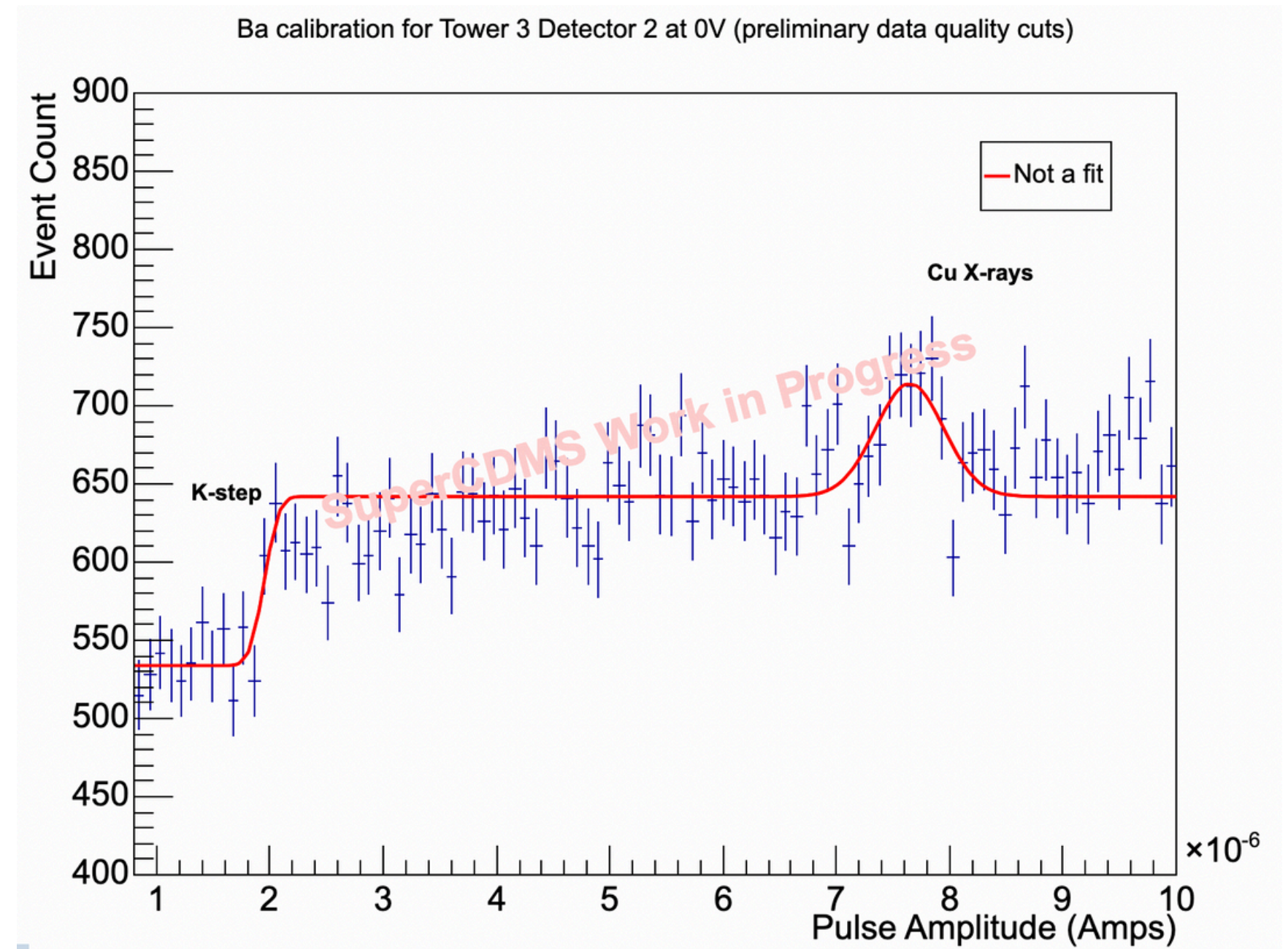
Tower Testing at CUTE: Ge Detector Calibration

- The activation lines from ^{71}Ge electron captures are:
 - 10 keV, 1.3 keV, 160 eV
- From Dec. 16 to Dec. 21, we took ~90 hours of Ge detector calibration dataset



Tower Testing at CUTE: Si Detector Calibration

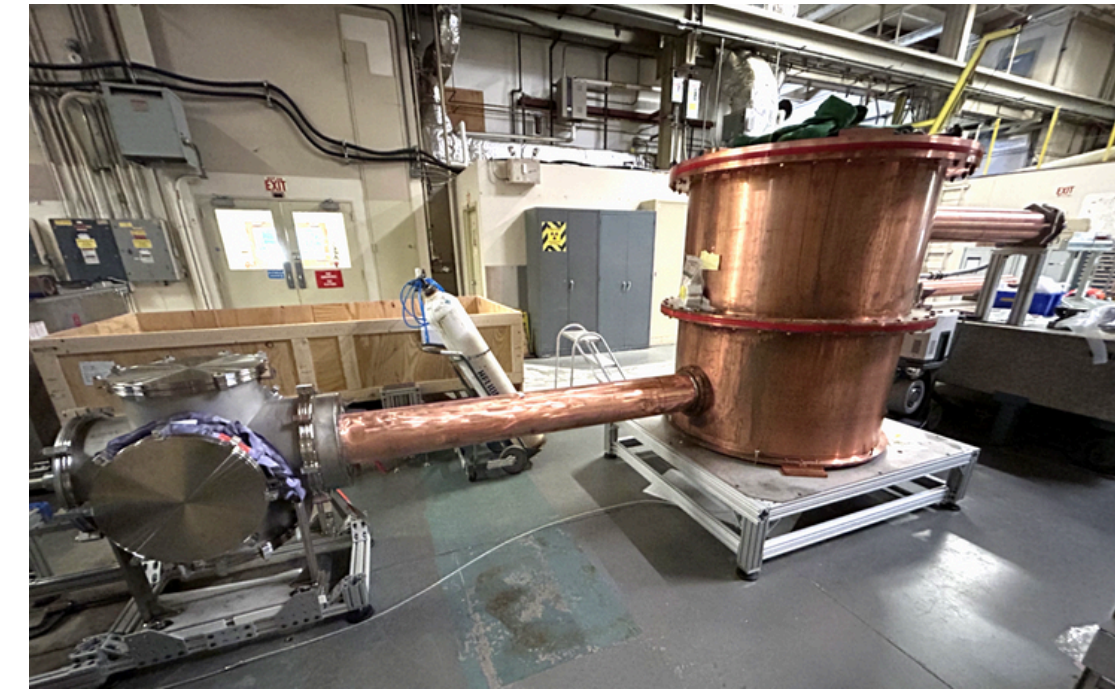
- On the other hand, Silicon detectors are much harder to calibrate due to lack of spectral lines
- Characteristic steps in energy spectrum caused by the binding energy of shell electrons can be used to calibrate the Si detector.
- **First proof-of-principle demonstration the calibration method for SuperCDMS Si detectors.**



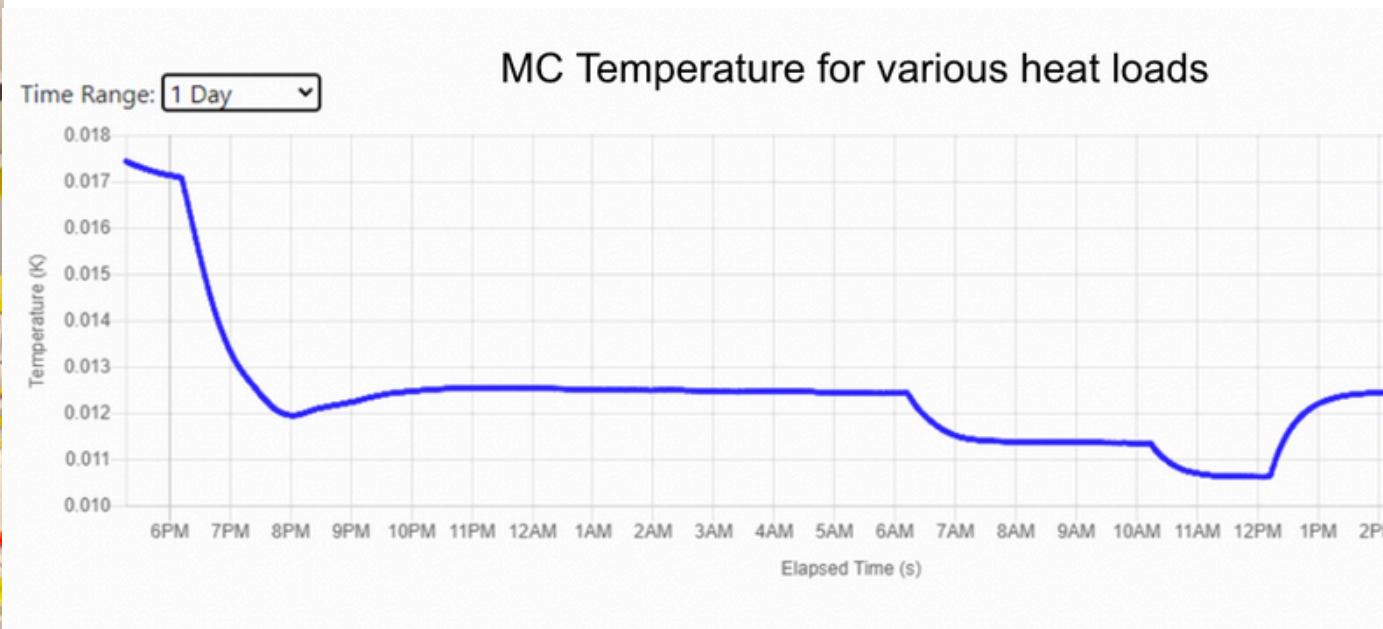
SuperCDMS Construction Progress during Last Year



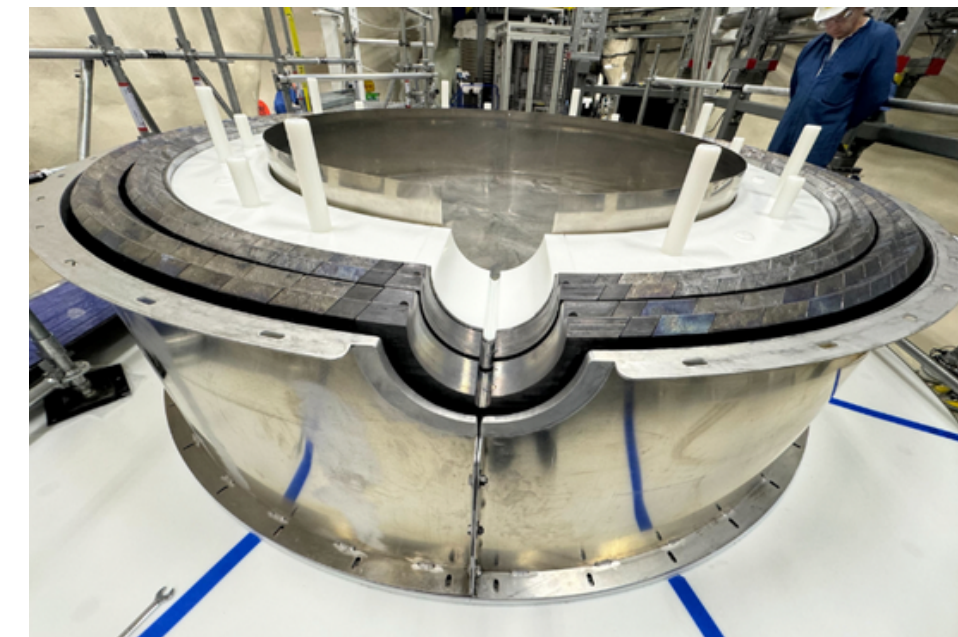
All four towers has been delivered underground at SNOLAB



Cryostat pre-assembly accomplished at SLAC

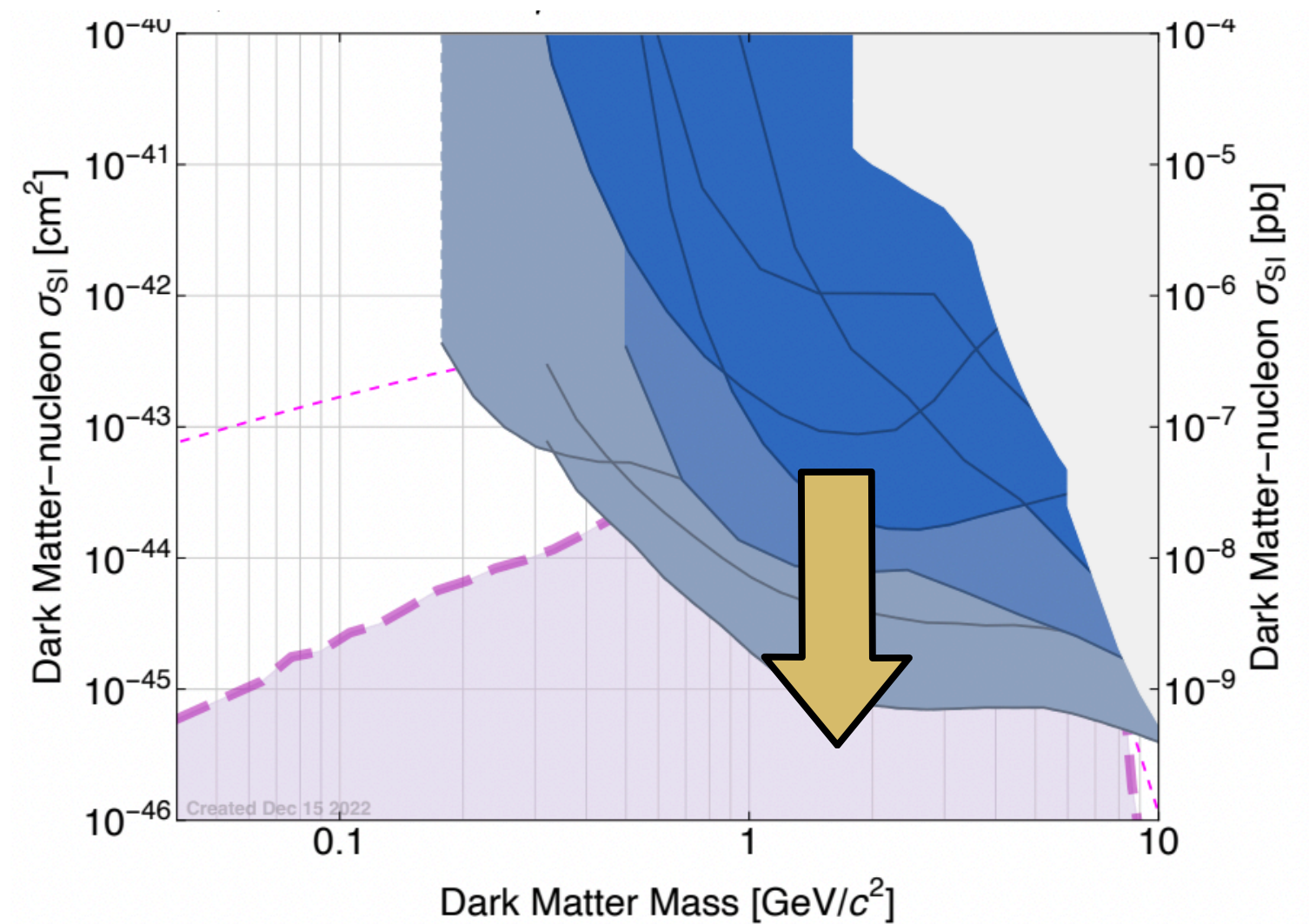
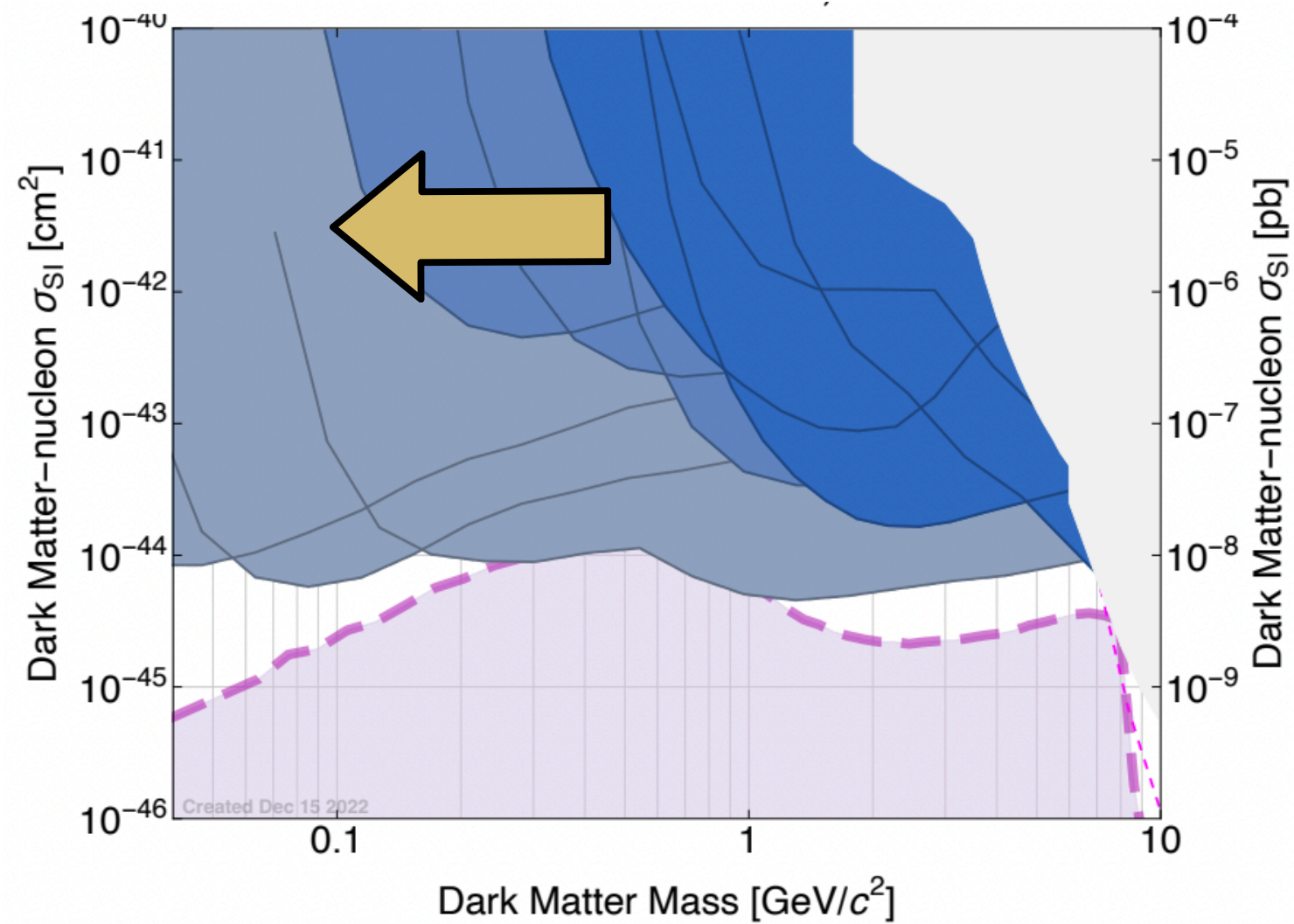


Standalone test of the dilution fridge demonstrated base temperature



Shield base installation completed

Science Reach of SuperCDMS SNOLAB

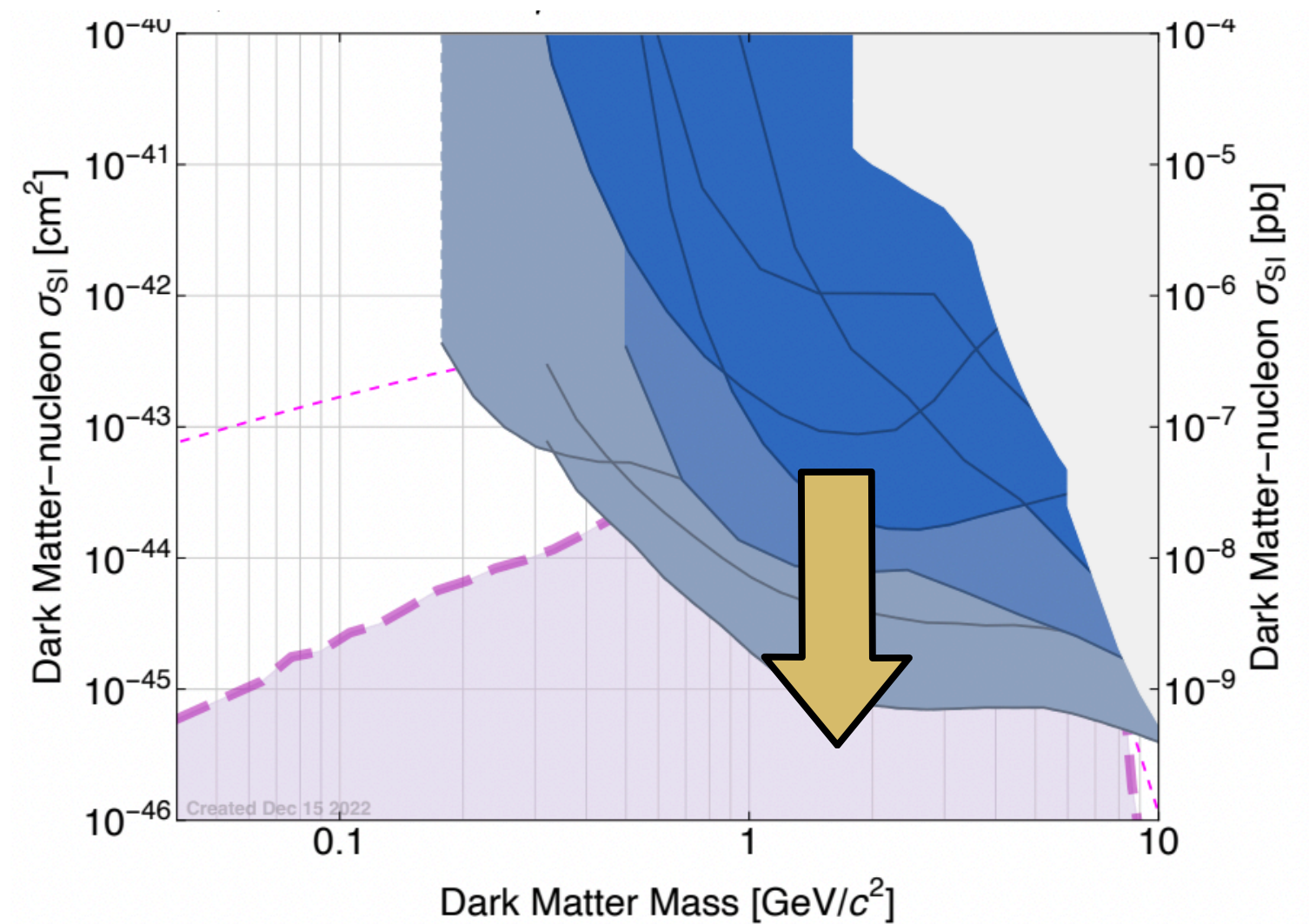
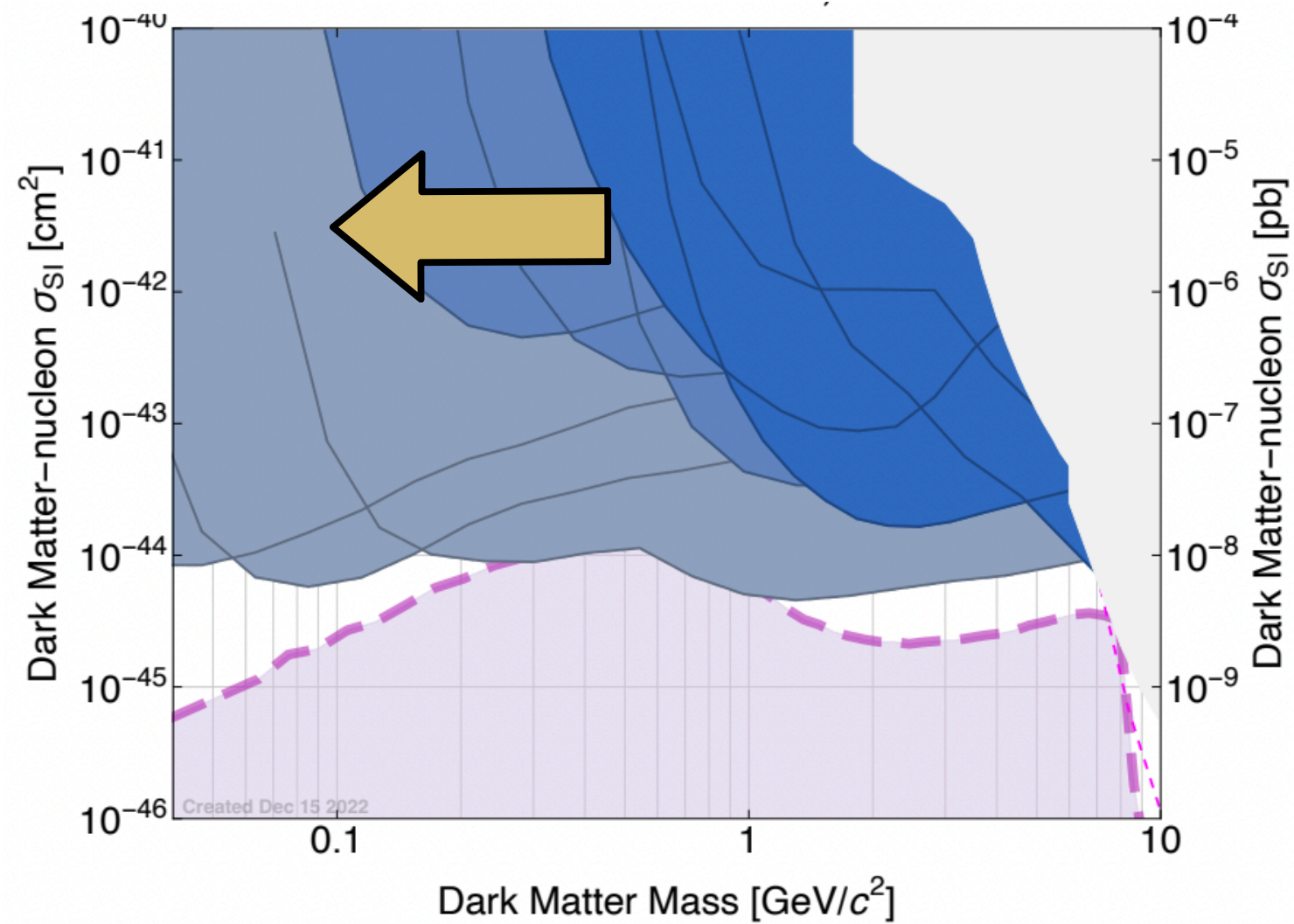


arXiv:2203.08463 SuperCDMS Collaboration

Dark blue: current experiment sensitivity

Lighter colors: improved science reach based on experience-guided technological advancement

Science Reach of SuperCDMS SNOLAB



arXiv:2203.08463 SuperCDMS Collaboration

Background level achieved at the SuperCDMS SNOLAB will not be limiting our dark matter sensitivity for the next decade or more.

CONCLUSIONS

- 1.** Direct detection of dark matter is one of the most promising methods we have to investigate BSM physics.
- 2.** SuperCDMS SNOLAB is sensitive to a variety of dark matter models ranging over 10 orders of magnitude in dark matter mass.
- 3.** We successfully ran one SuperCDMS High Voltage tower at CUTE last year, marking the first time these detectors are operated in an underground, low background environment.
- 4.** Preliminary results indicate great potential with these detectors, enabling us to explore exciting science in the coming months and beyond.
- 5.** Stay tuned!

 thank you

