

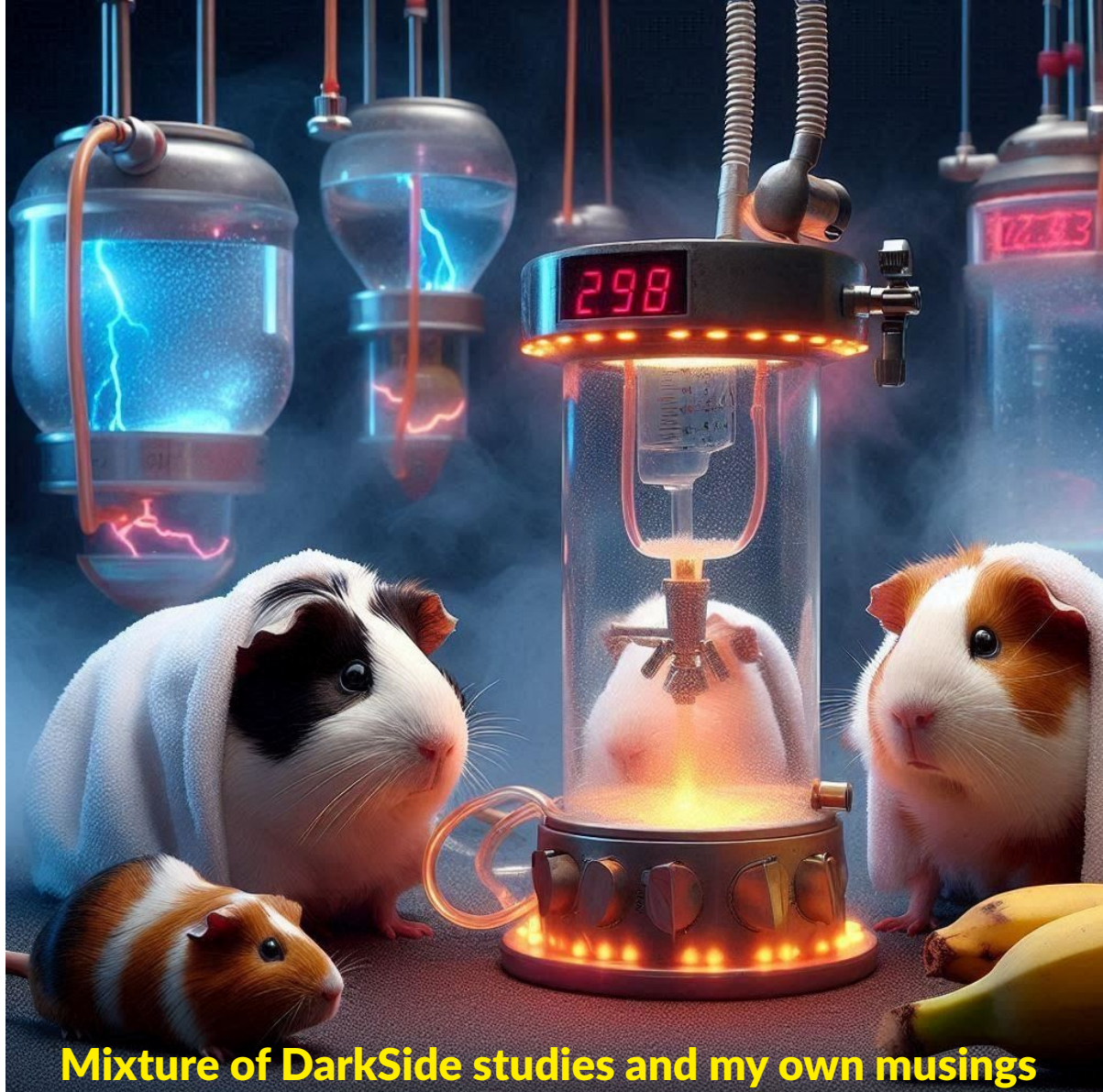


# Detecting low-mass dark matter with liquid argon

Shawn Westerdale  
with the GADMC  
GUINEAPIG Workshop  
20 August, 2024

UCR PHYSICS & ASTRONOMY

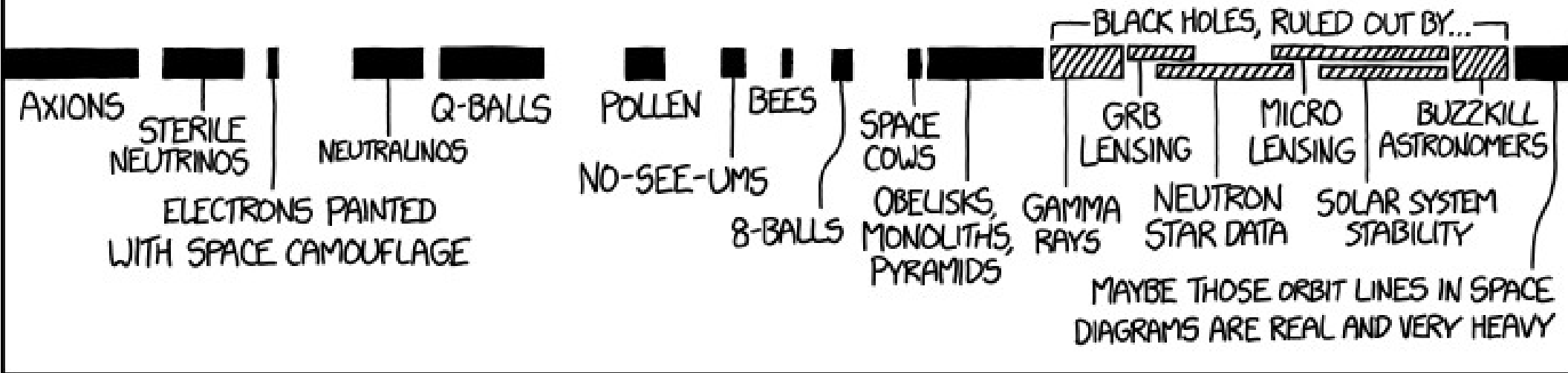
Bing AI, draw guineapigs making a LArTPC



Mixture of DarkSide studies and my own musings

# DARK MATTER CANDIDATES:

meV meV eV KeV MeV GeV TeV  $10^{-18}$ kg ng Mg mg g Kg TON  $10^6$ kg  $10^{12}$ kg  $10^{18}$ kg  $10^{24}$ kg  $10^{30}$ kg

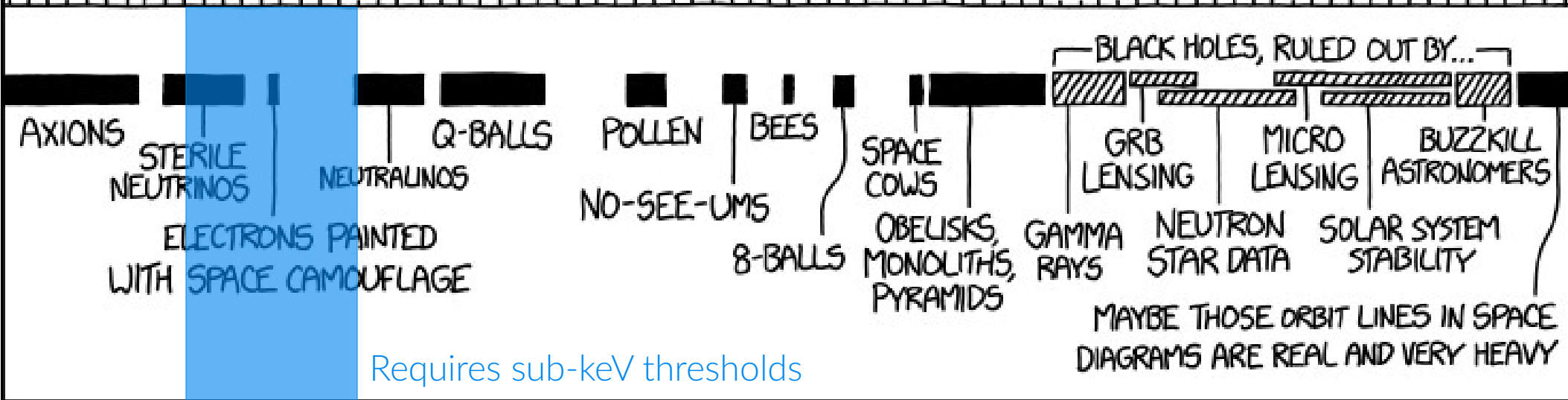


<u>Wave DM</u>	<u>Light DM</u>	<u>WIMP</u>	<u>Ultra-heavy DM</u>
Absorption EM resonance	Single scatter ER/NR Absorption	Single scatter NR	MIMPs Multi-scatter/tracks

Macroscopic DM  
Cosmic probes

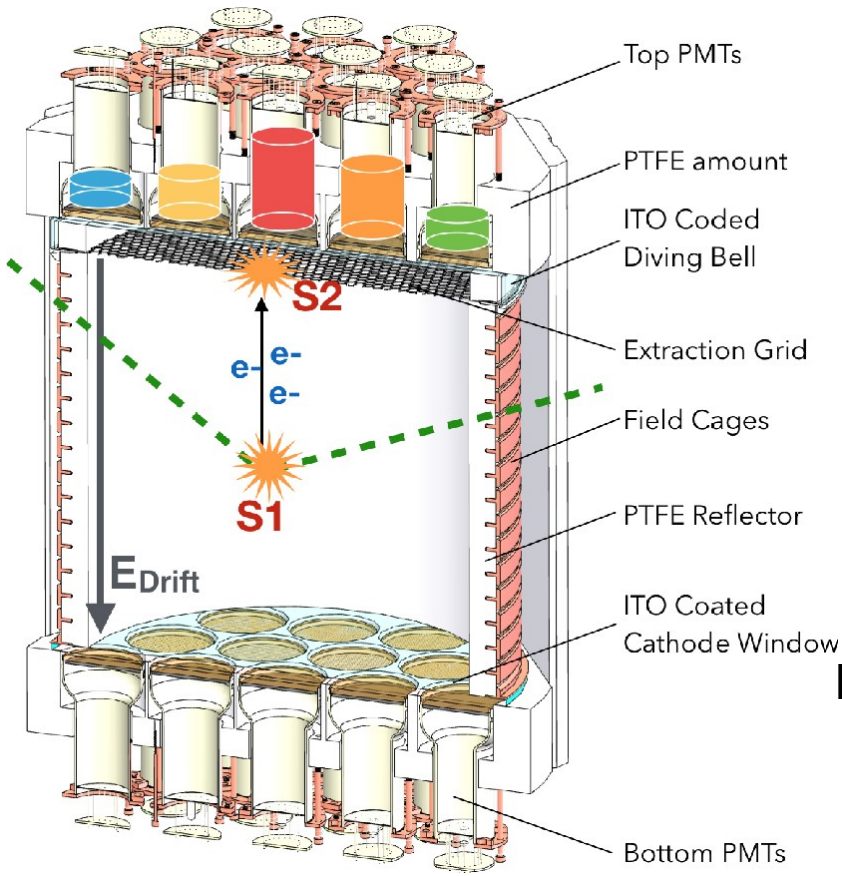
# DARK MATTER CANDIDATES:

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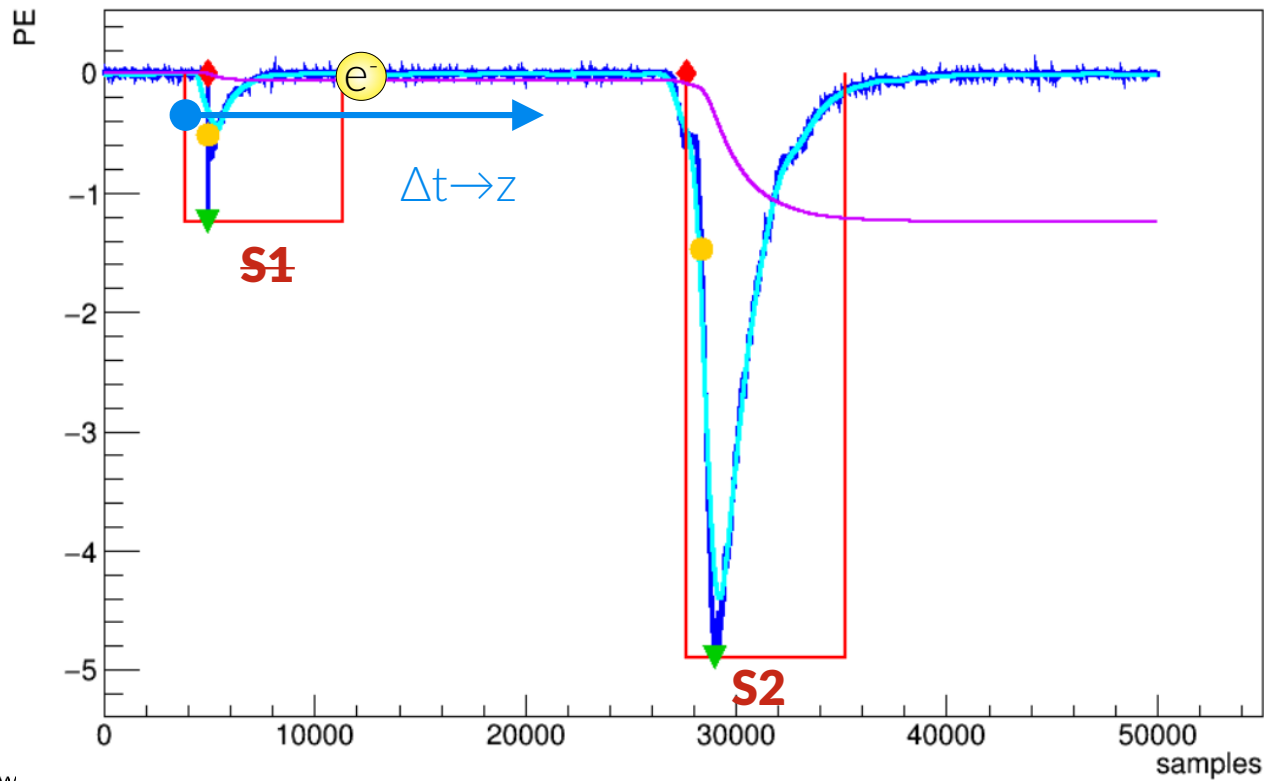


Wave DM	Light DM	WIMP	Ultra-heavy DM
Absorption EM resonance	Single scatter ER/NR Absorption	Single scatter NR	MIMPs Multi-scatter/tracks

Macroscopic DM Cosmic probes



Drawing of DarkSide-50



## Dual-phase LArTPCs

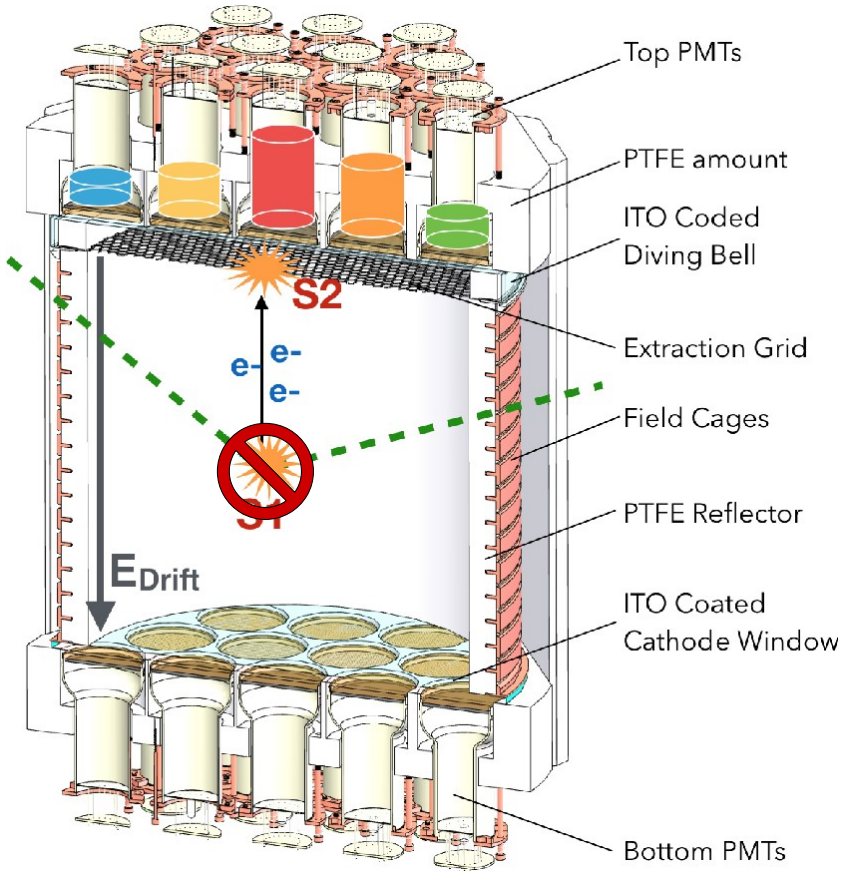
**S1:** Prompt scintillation, energy, PSD

Detect ~20% of photons

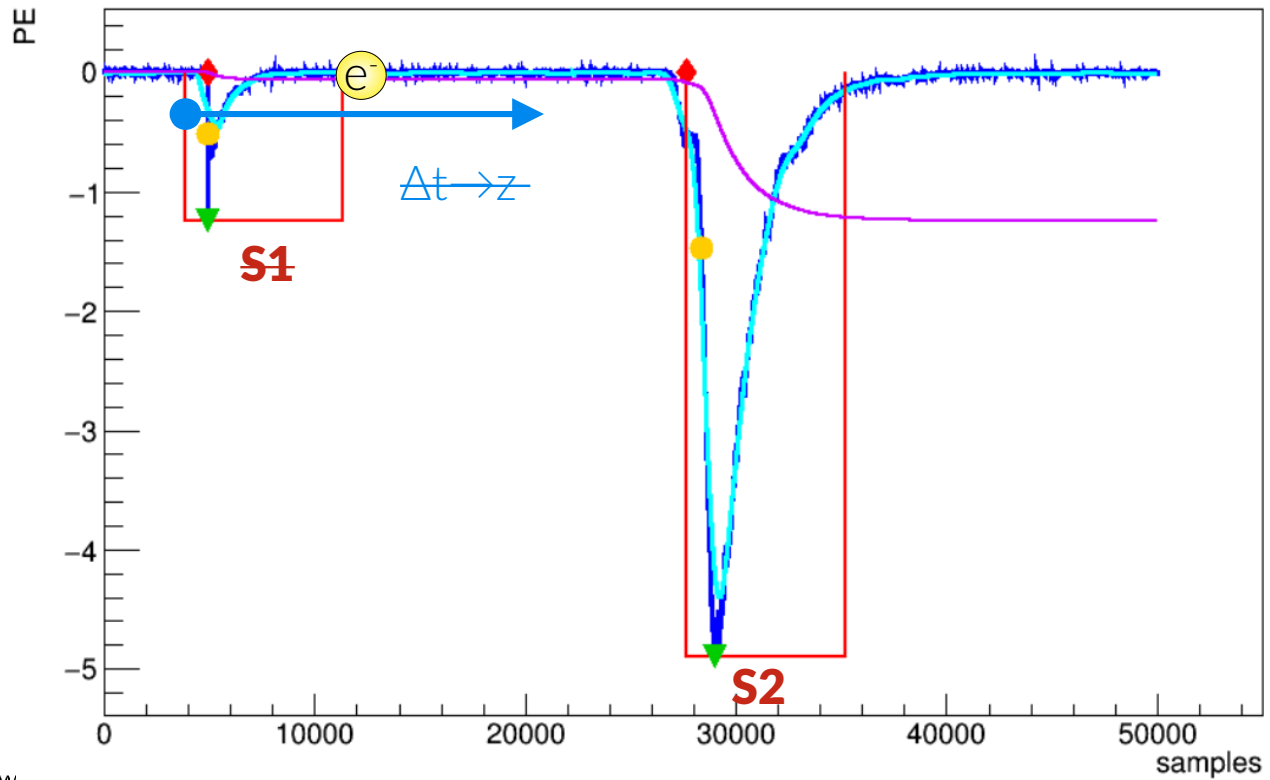
**S2:** Ionization electrons, position, n scatters

Detect nearly all electrons

Strong probe of ~100 GeV WIMPs, great bkgd discrimination



Drawing of DarkSide-50

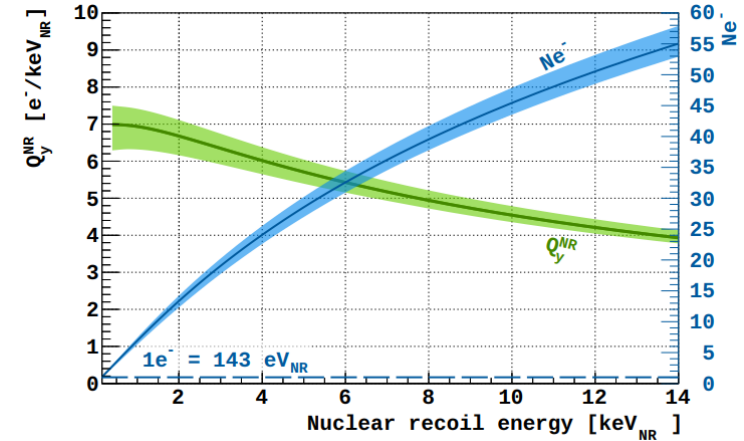
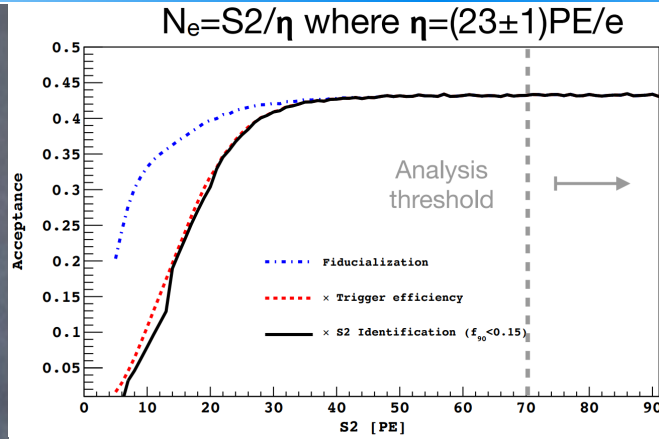


As a consequence, S2 can access lower energies than S1 → Low-threshold LArTPC via S2-only channel

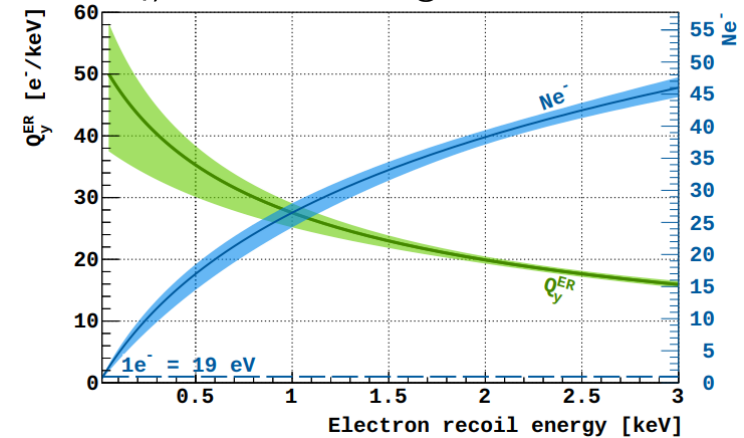
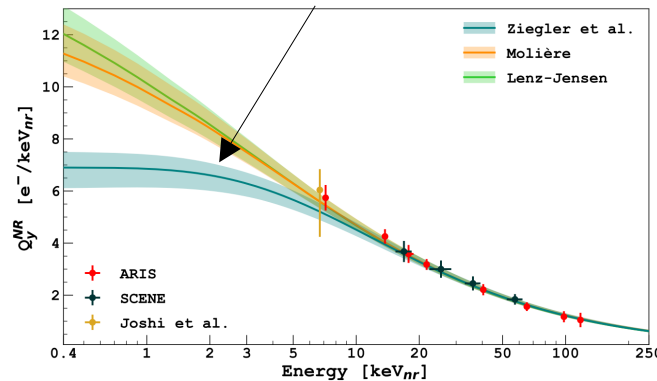
### Challenges:

- Cannot measure  $\Delta t$  → Can't fiducialize in  $z$
- Timing resolution limited by max  $e^-$  drift time
- No PSD → No ER discrimination
- Spurious electron backgrounds

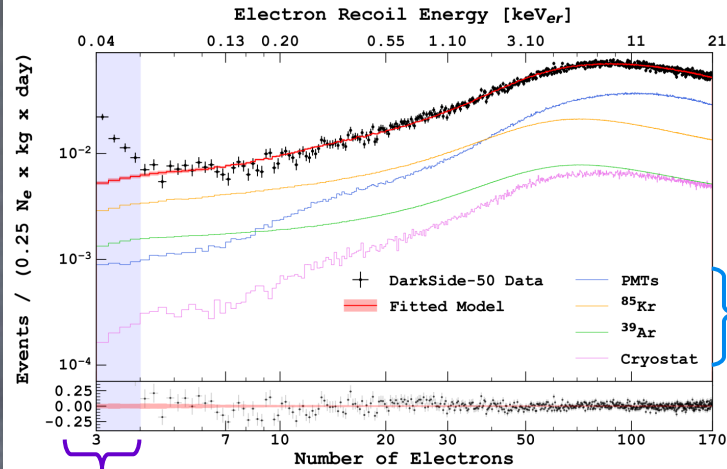
# Low-threshold S2-only analyses, but lose PSD and z position reconstruction



Need ex situ calibration to measure  $Q_y$  at low energies



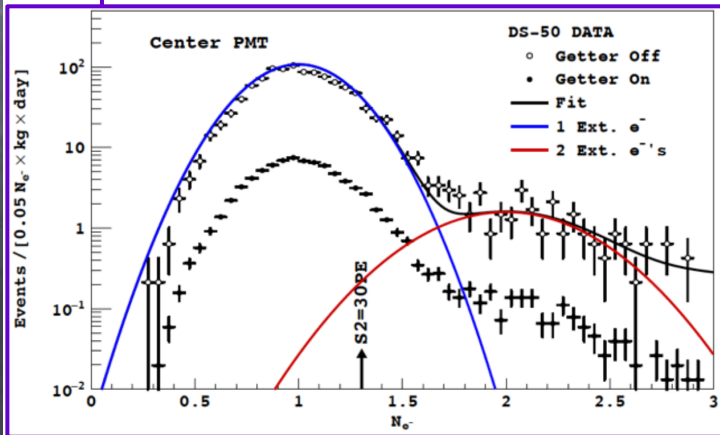
# Low-threshold S2-only analyses



## Loss of pulse shape discrimination (PSD)

Electromagnetic backgrounds from dominate  $\gamma$ -rays from internal radioactivity  
 $\beta$ -decays from  $^{39}\text{Ar}$  and  $^{85}\text{Kr}$

^Use underground/depleted Ar



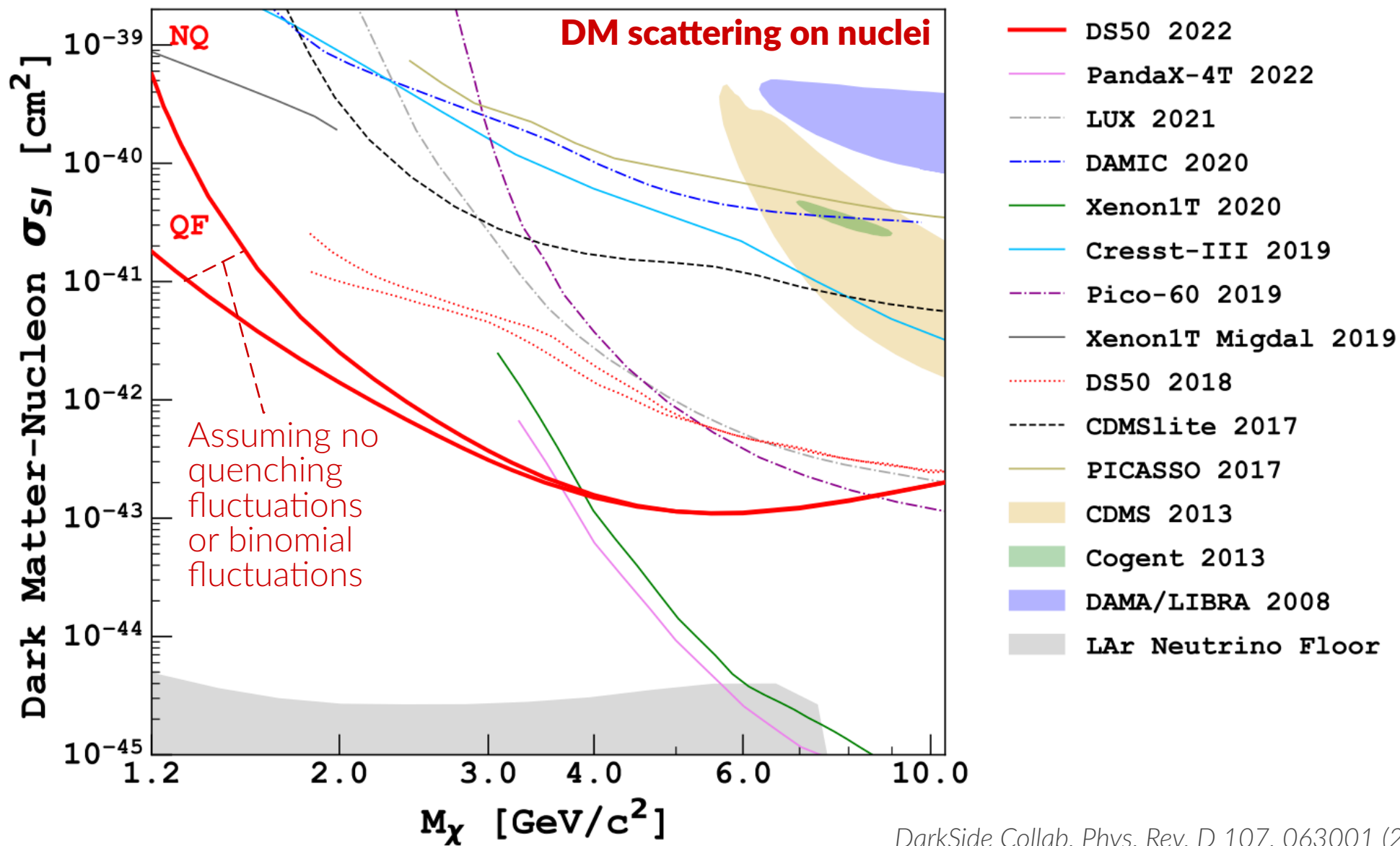
## Spurious e-'s dominate lowest energy bins

Full origin being investigated

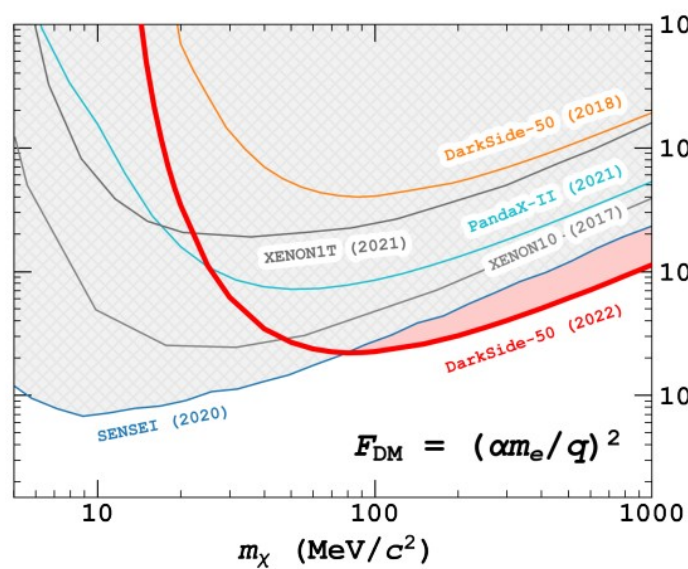
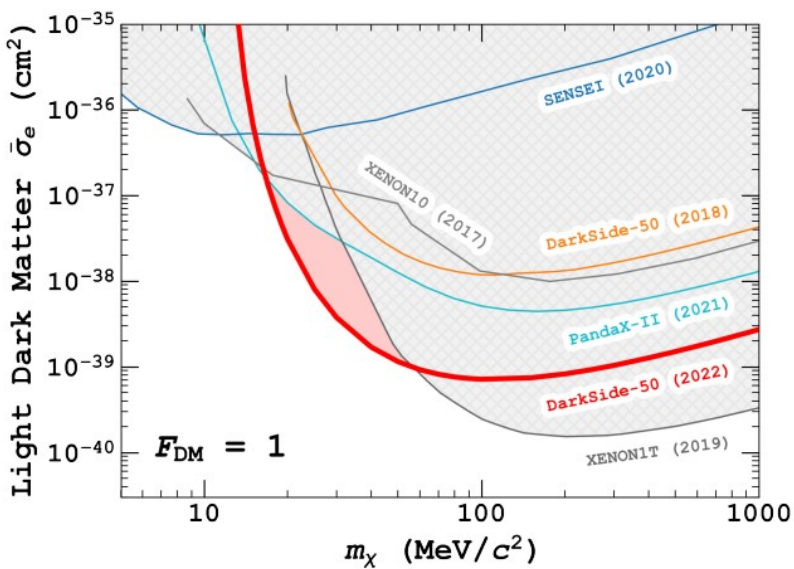
Correlations between SE and preceding S2 (time, xy position, S2 charge and  $t_{\text{drift}}$  vs.  $P_{\text{SE}}$ )

Possibly due to drifting electrons capturing on impurities and forming metastable states

Follow preceding S2 by O(5-50 ms)

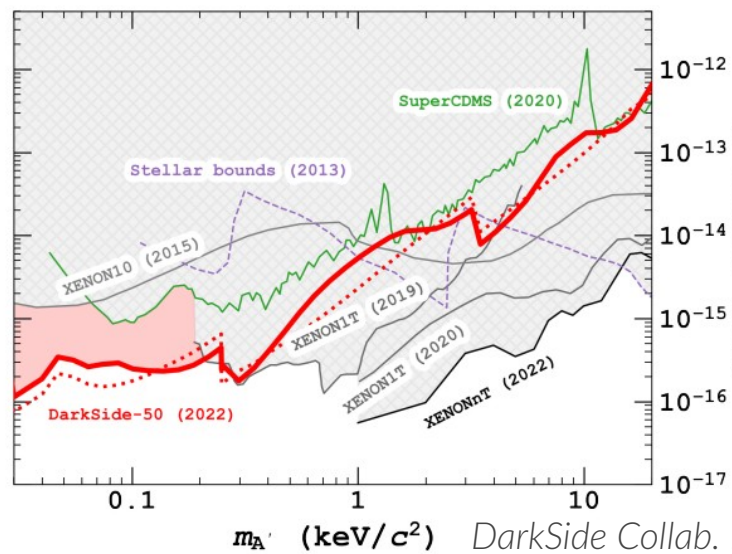
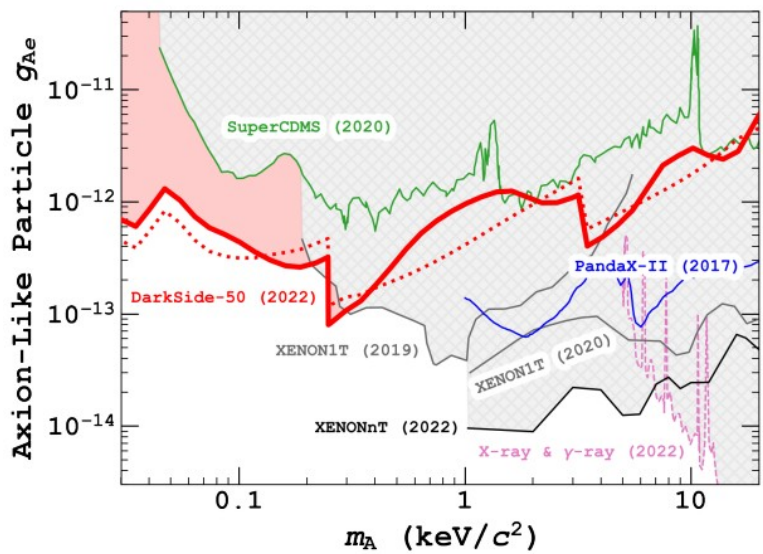






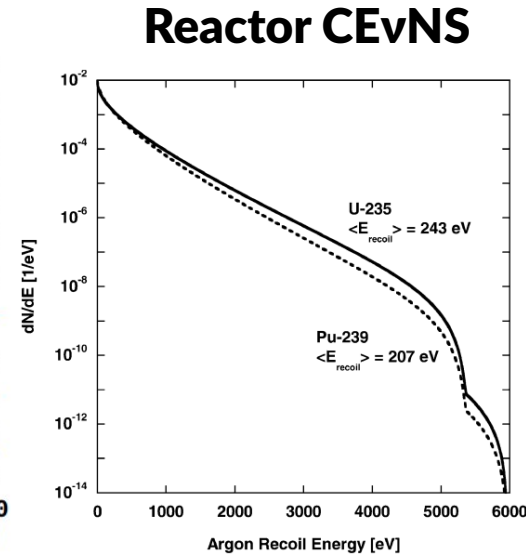
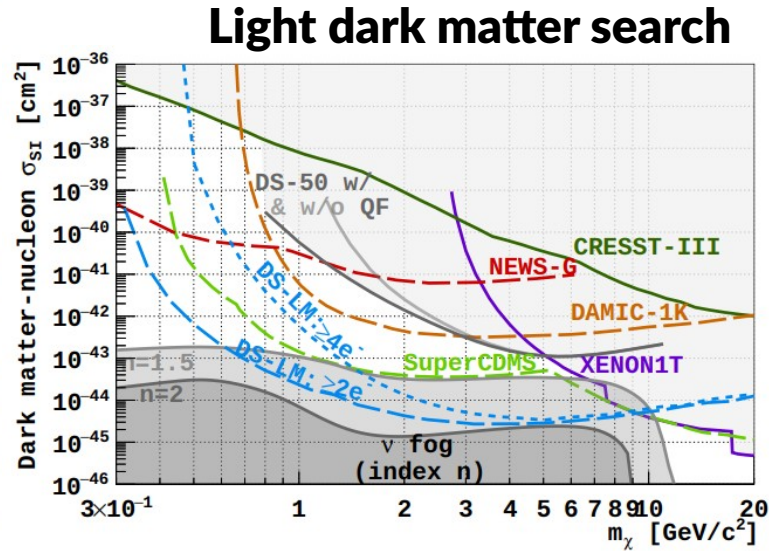
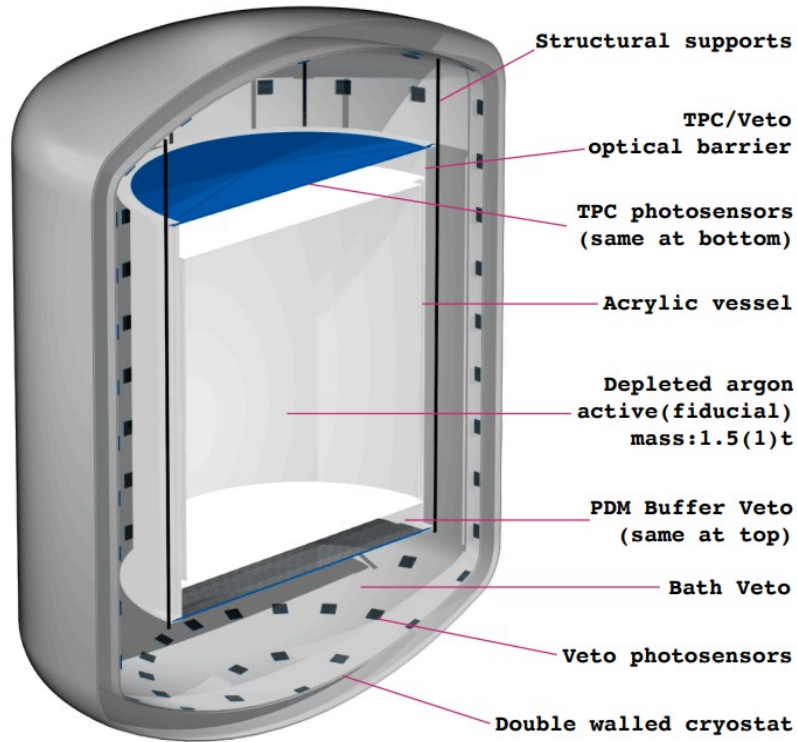
**DM w/ e-'s in final state**

**Scattering**

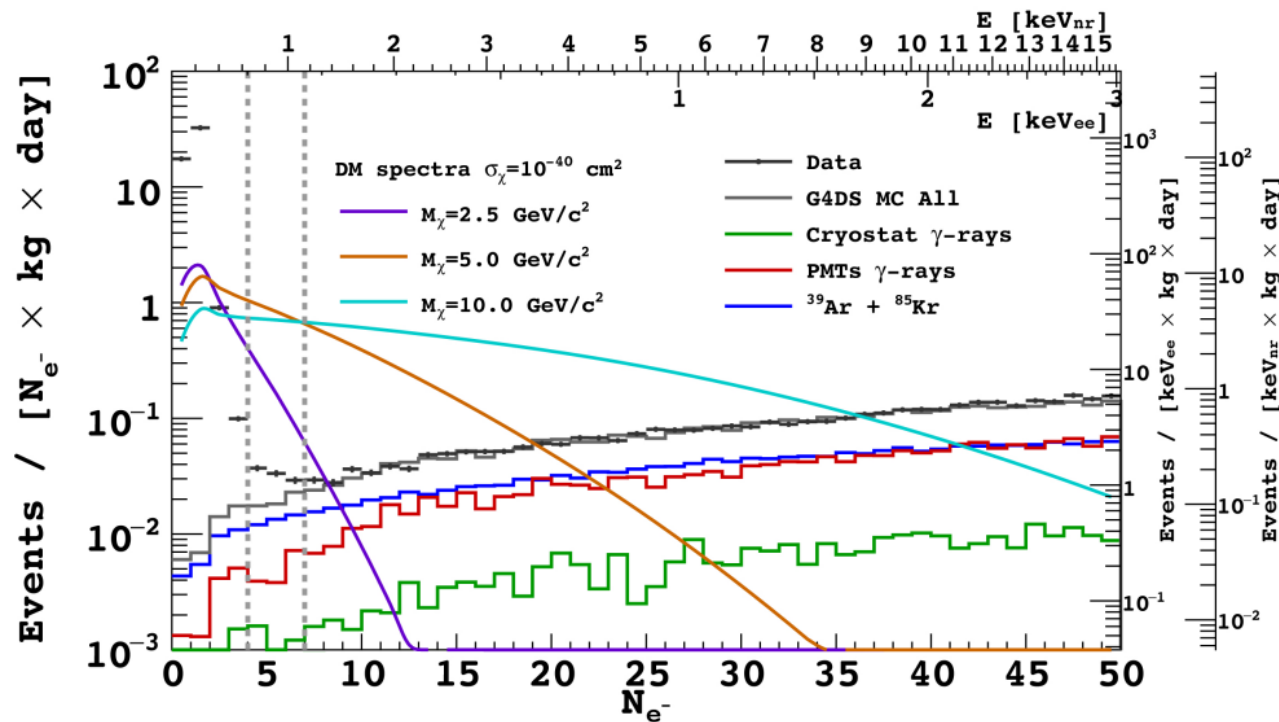


**Absorption**

# LArTPC in S2-only analyses: Powerful tool for low-threshold measurements



# How do we lower the threshold?

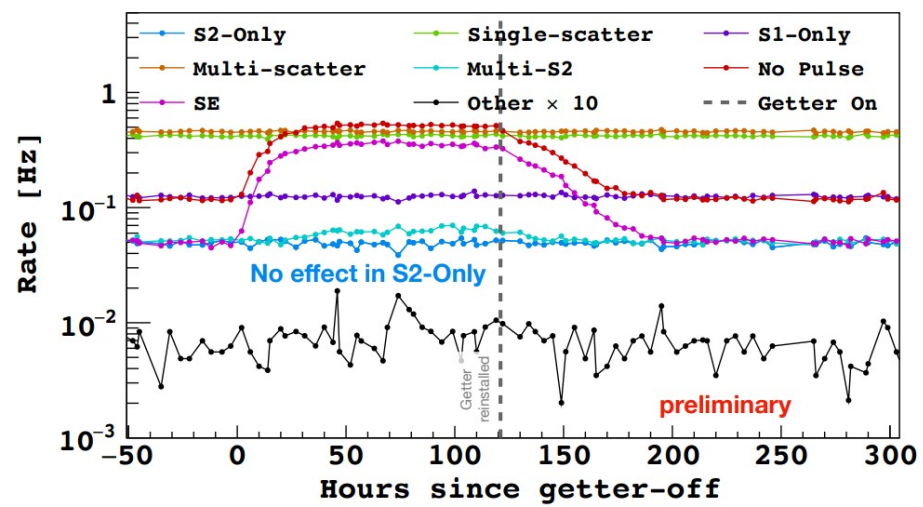
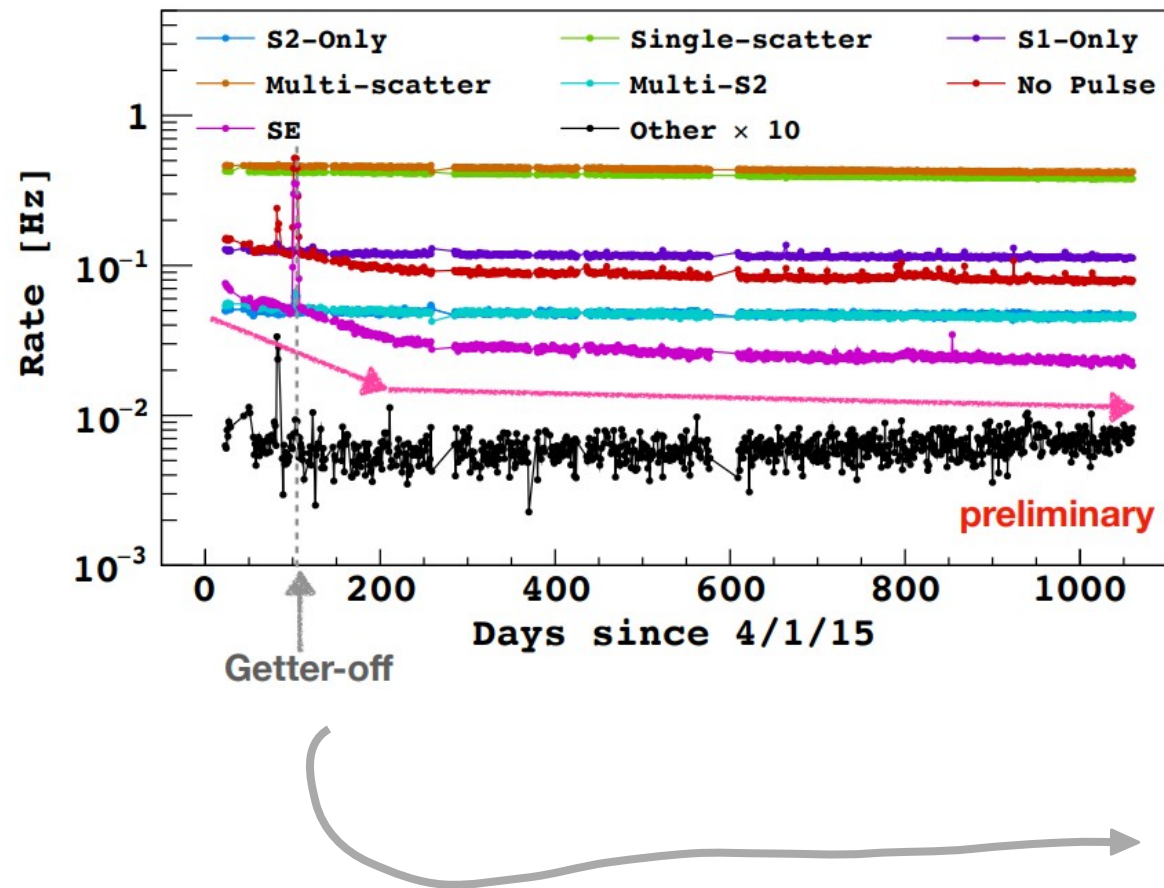


## Threshold is set by

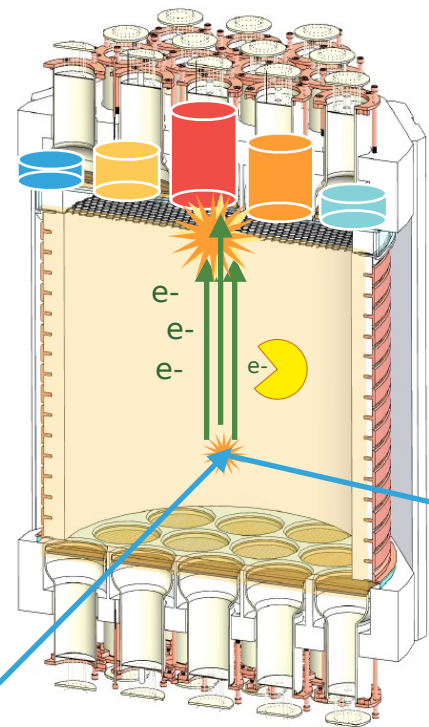
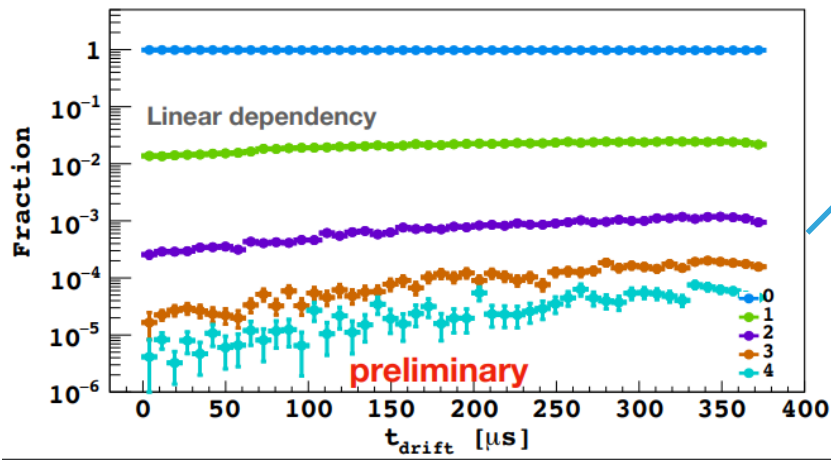
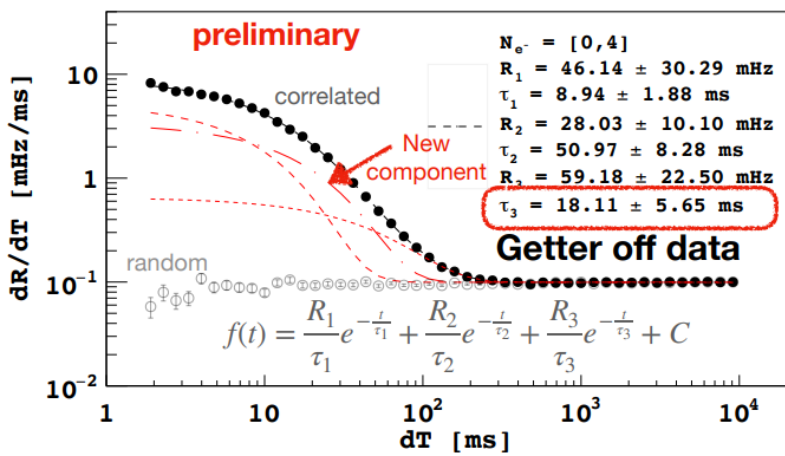
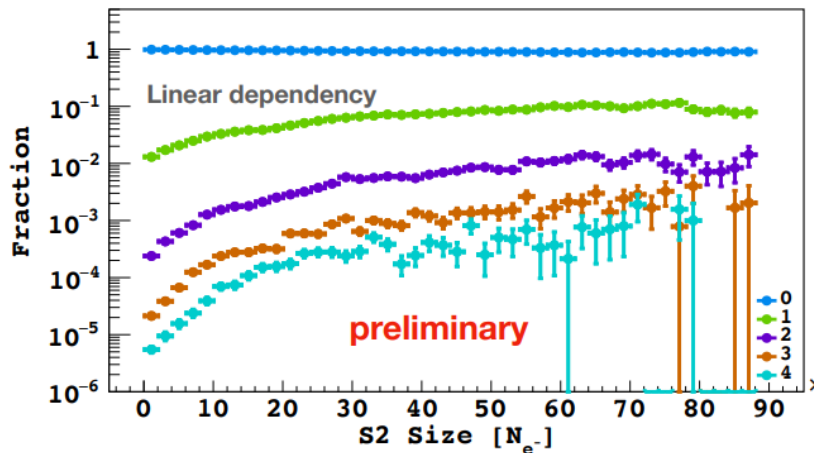
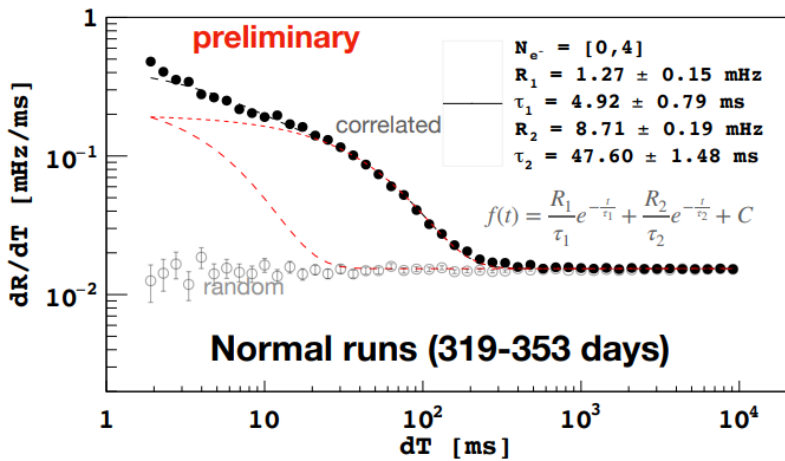
- Spurious electrons
- Medium's ionization energy
- Partitioning between ionization & scintillation photons
- heat

All of these are pathways for lowering the energy threshold

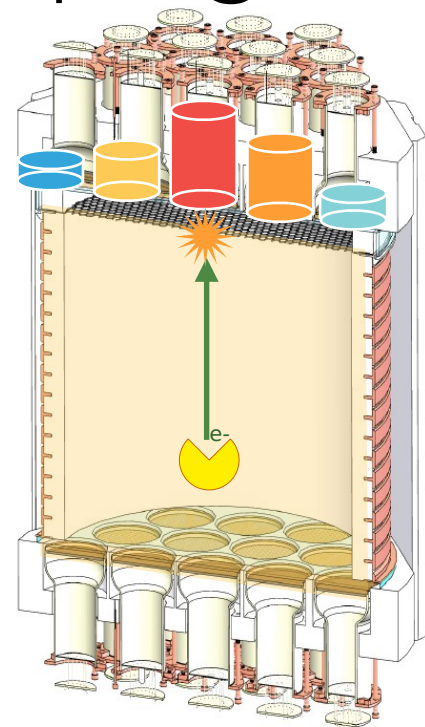
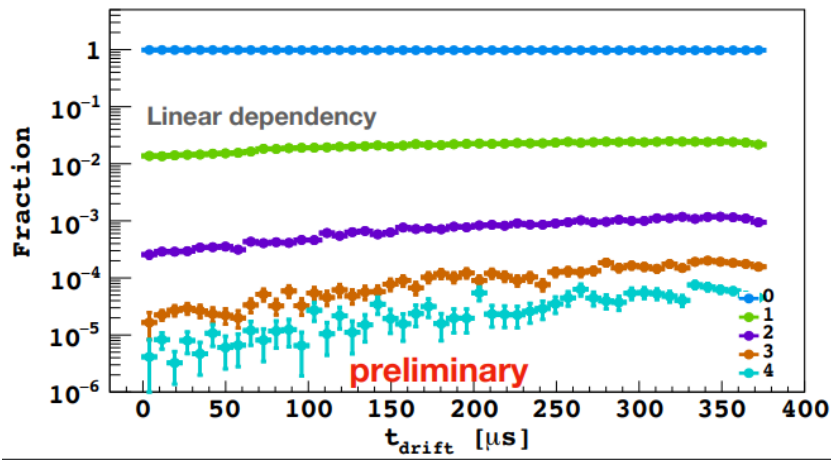
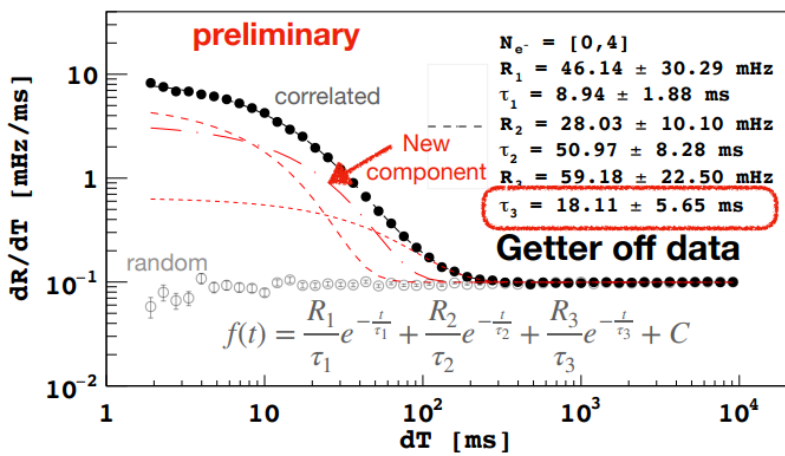
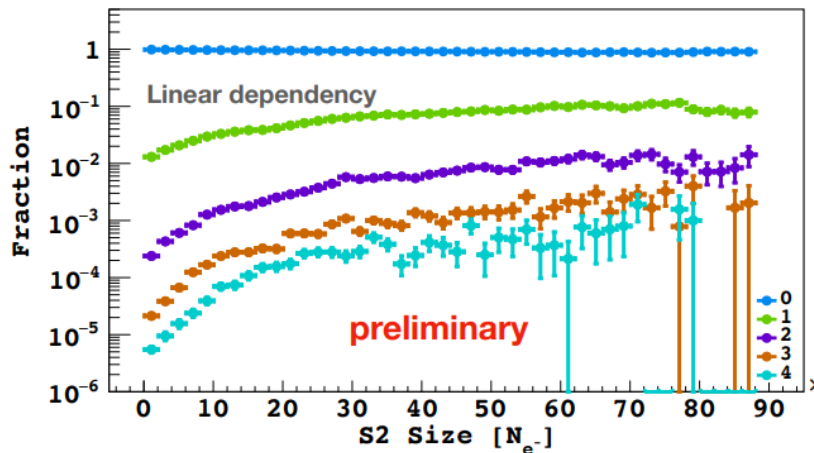
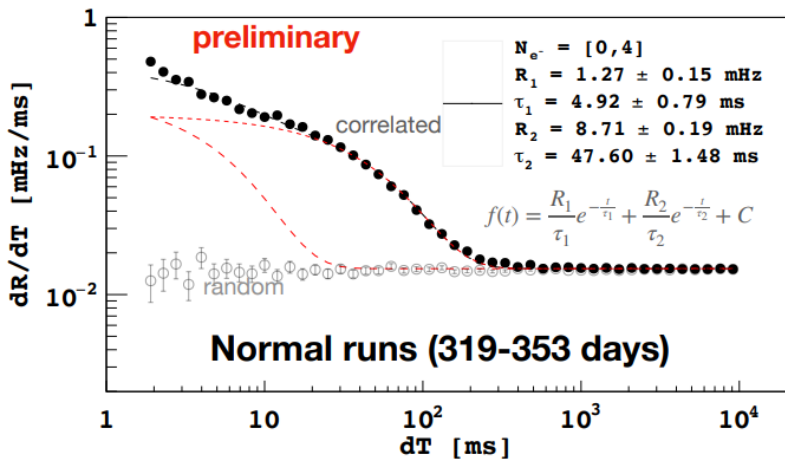
# SE rate changes over time and appears to be related to impurities



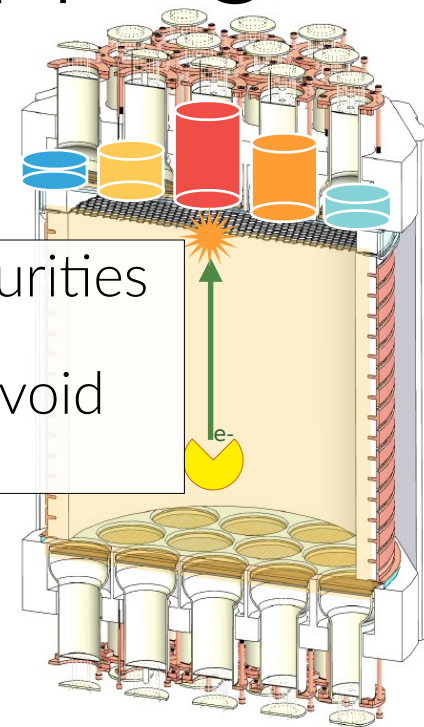
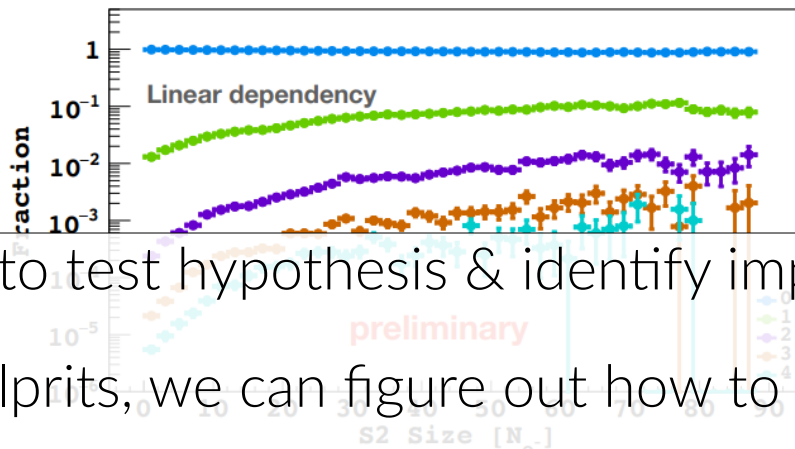
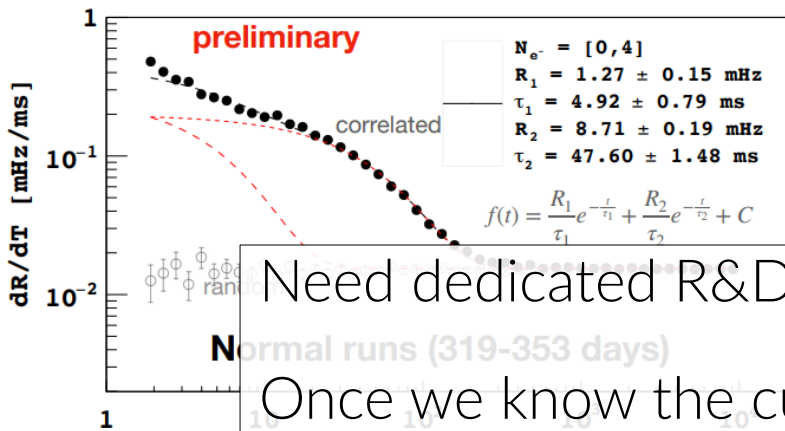
# SEs appear to be correlated in time, space, and charge with ordinary events



# Hypothesis: Dominant SE production mechanism from impurities trapping $e^-$

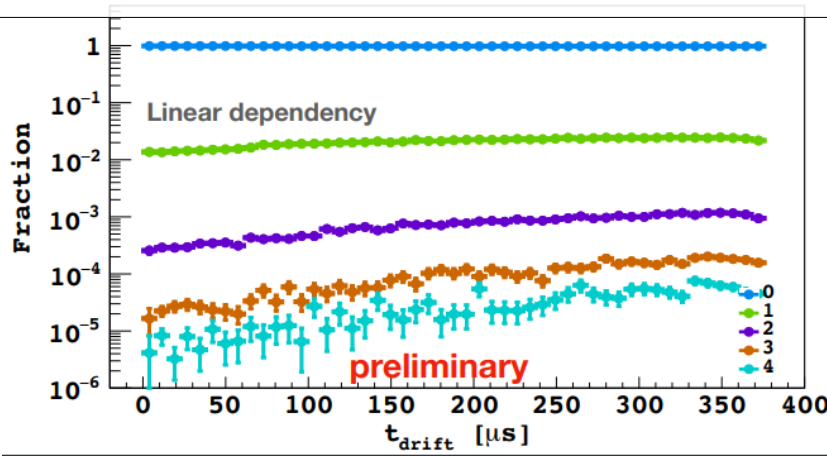
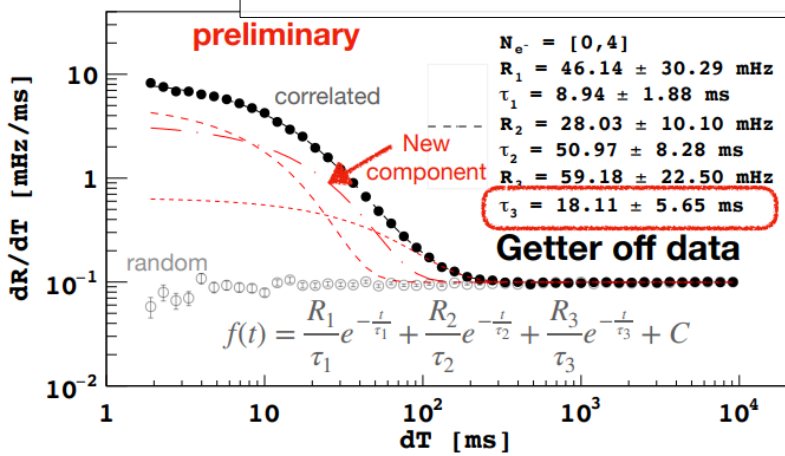


# Hypothesis: Dominant SE production mechanism from impurities trapping $e^-$



Need dedicated R&D to test hypothesis & identify impurities

Once we know the culprits, we can figure out how to avoid or remove them



# Other ideas: enhance detector response by doping LAr

Kubota et al. Phys. Lett. 49A, 5 (1974): 933

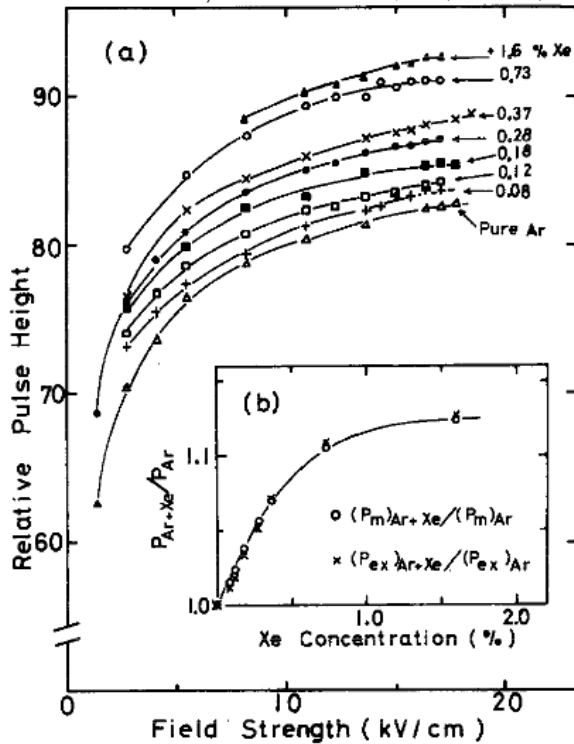


Fig. 1. a) Saturation curves of the pulse height produced by  $^{207}\text{Bi}$  conversion electron for the different concentration of Xe. b) Relative ionization yields as a function of doped-Xe concentration.

## Idea:

Xe ionization energy is 12.1 eV  
Ar ionization energy is 15.8 eV, W-value  $\sim 24$  eV  
Xe-doped LAr effective ionization potential  $\sim 10.7$  eV  
[B. Raz and J. Jortner, Proc. Roy. Soc. London A 317 (1970) 113.]

Even at low Xe concentration, each ionization takes less energy from the nucleus, and nuclei can continue to produce ions even down to lower energies

Nuclear recoils produce ionization through a cascade (one nucleus colliding with others, and so on)  $\rightarrow$  effects may be non-linear

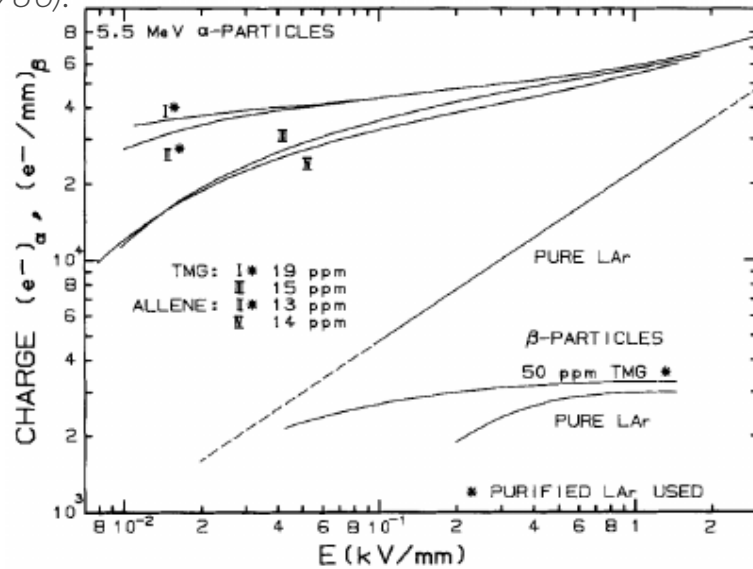
## R&D needs:

Need to measure  $Q_y$  of low-energy NRs as a function of Xe concentration to see the effect side

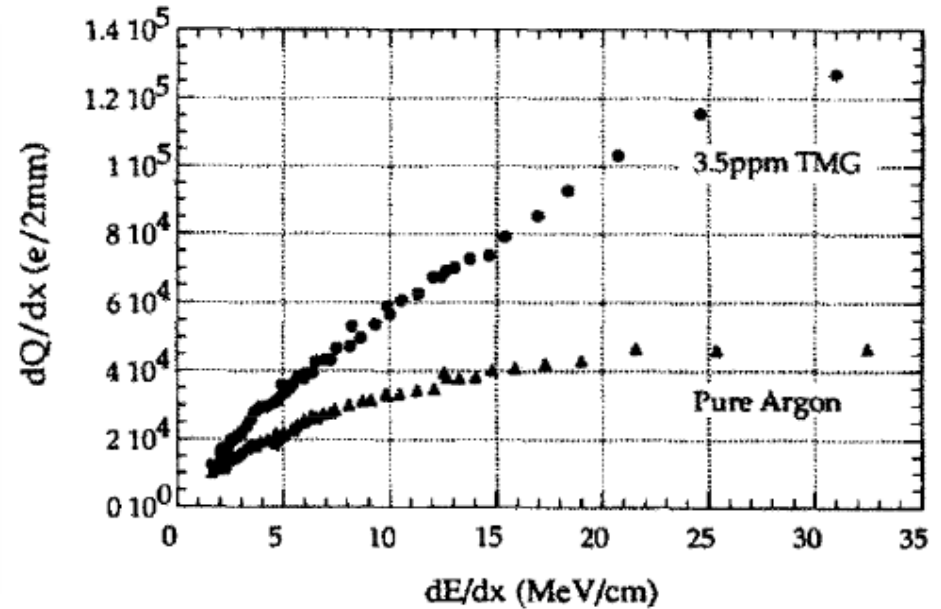


# Other ideas: Photo-sensitive dopants— potentially great improvements, more R&D

D. F. Anderson, *Nucl. Instrum. Methods Phys. Res. A* 245, 361 (1986).



P. Cennini et al., *Nucl. Instrum. Methods Phys. Res. A* 355, 660 (1995).

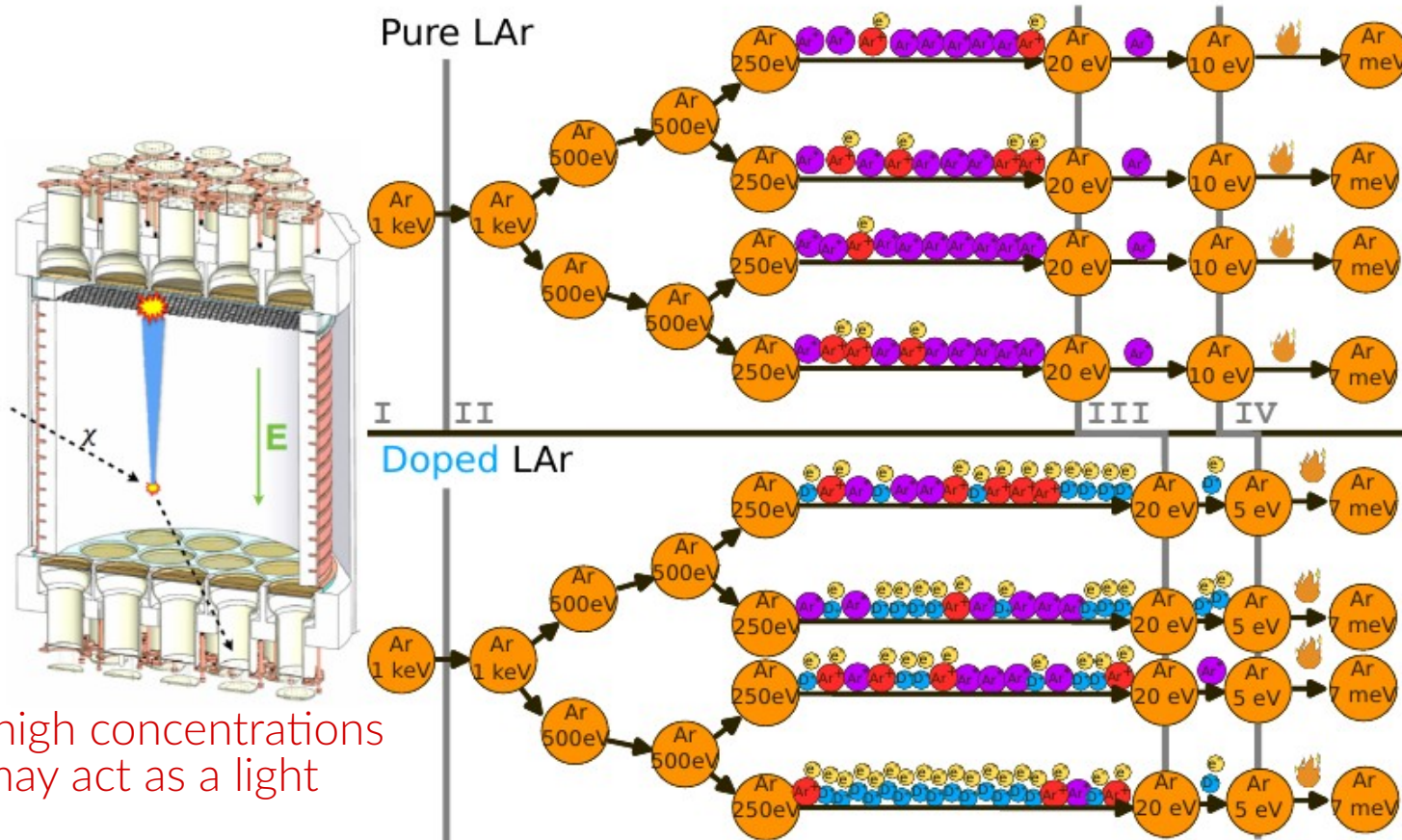


Dopants with 7.5–9.5 eV ionization energies have previously been studied in LAr, in the context of improving the resolution of high-energy ionization chambers

Larger effects at higher ionization energies → energy otherwise lost as heat goes into ionization

Turns inefficiently detected S1 into efficiently detected S2!

# Potential benefits from doping



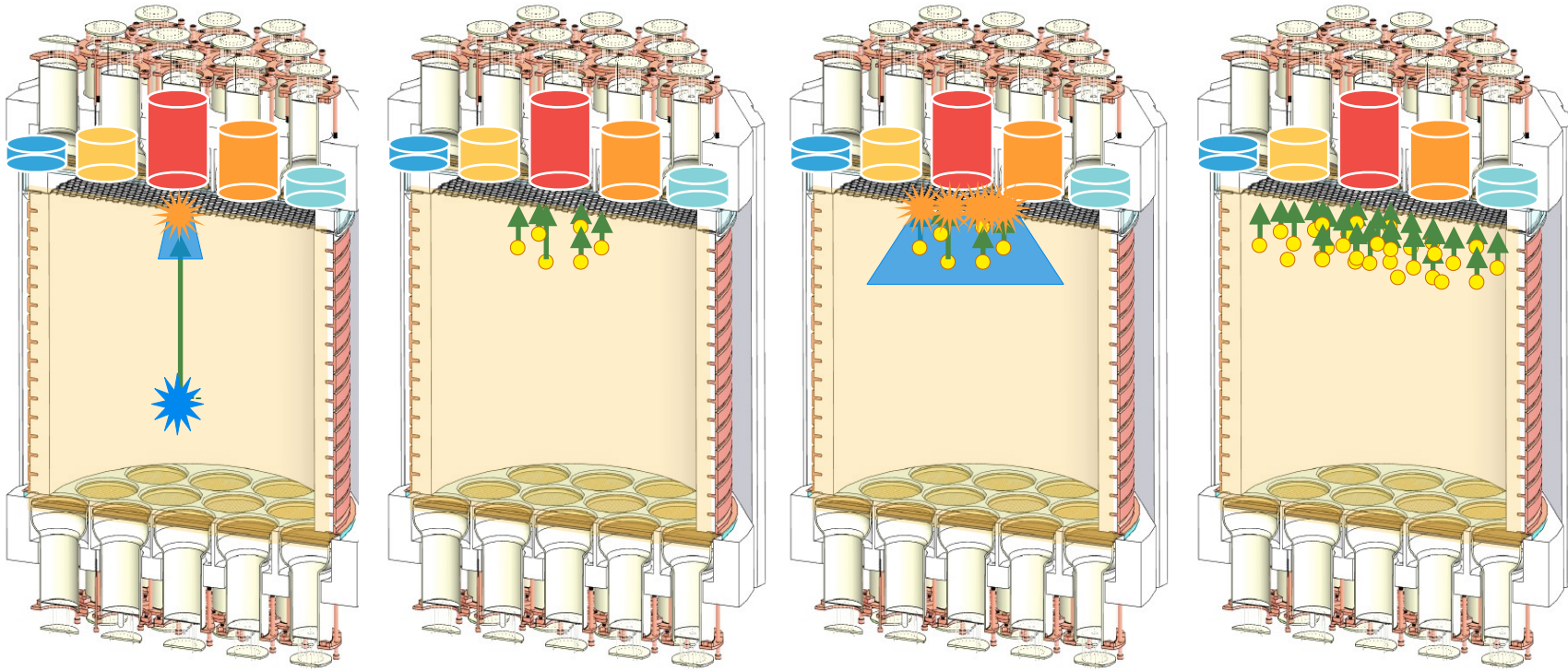
At sufficiently high concentrations H in dopants may act as a light DM target!

# Some potential dopants

Material <sup>d)</sup>	$I_g$ [eV] <sup>a)</sup>	Dipole moment [debyes] <sup>b)</sup>	Estimated pressure 90 K [Torr] <sup>b)</sup>	Charge collected* <sup>c)</sup> (LAr $\equiv$ 1)		Concentration [ppm]	
				0.1 kv mm <sup>-1</sup>	1.0 kv mm <sup>-1</sup>		
TEA	(C <sub>2</sub> H <sub>5</sub> ) <sub>3</sub> N	7.50	0.66	–	2.2	1.3	47
TMA	(CH <sub>3</sub> ) <sub>3</sub> N	7.82	0.612	$3 \times 10^{-8}$	3.4	1.6	110
TMT	(CH <sub>3</sub> ) <sub>4</sub> Sn	8.25/8.76	–	$4 \times 10^{-12}$	3.0	1.6	1.5
Cyclohexene	C <sub>6</sub> H <sub>10</sub>	8.95	–	–	2.1	1.3	3.6
1.3-butadiene	C <sub>4</sub> H <sub>6</sub>	9.06	0	$4 \times 10^{-7}$	4.6	1.9	17
Cis & Trans 2 butene	C <sub>4</sub> H <sub>8</sub>	9.13	0 (trans)	$5 \times 10^{-8}$	3.6	1.6	72
TMG	(CH <sub>3</sub> ) <sub>4</sub> Ge	9.2/9.29	–	$3 \times 10^{-10}$	7.4 (9.8)	2.6 (2.7)	15
Isobutylene	C <sub>4</sub> H <sub>8</sub>	9.23	0.5	$5 \times 10^{-7}$	4.9	1.8	16
Methyl mercaptan	CH <sub>3</sub> SH	9.44	1.52	$2 \times 10^{-8}$	2.0	2.0	15
Pentene (technical)	C <sub>5</sub> H <sub>10</sub>	9.5	–	$1 \times 10^{-9}$	3.1	1.5	7
Allene	C <sub>3</sub> H <sub>4</sub>	9.53	0	$1 \times 10^{-5}$	6.5 (8.7)	2.5 (2.7)	14
TMS	(CH <sub>3</sub> ) <sub>4</sub> Si	9.86	0.525	$8 \times 10^{-9}$	4.6	1.8	5.8
DME	(CH <sub>3</sub> ) <sub>2</sub> O	10.0	1.30	$5 \times 10^{-8}$	3.6	1.4	14

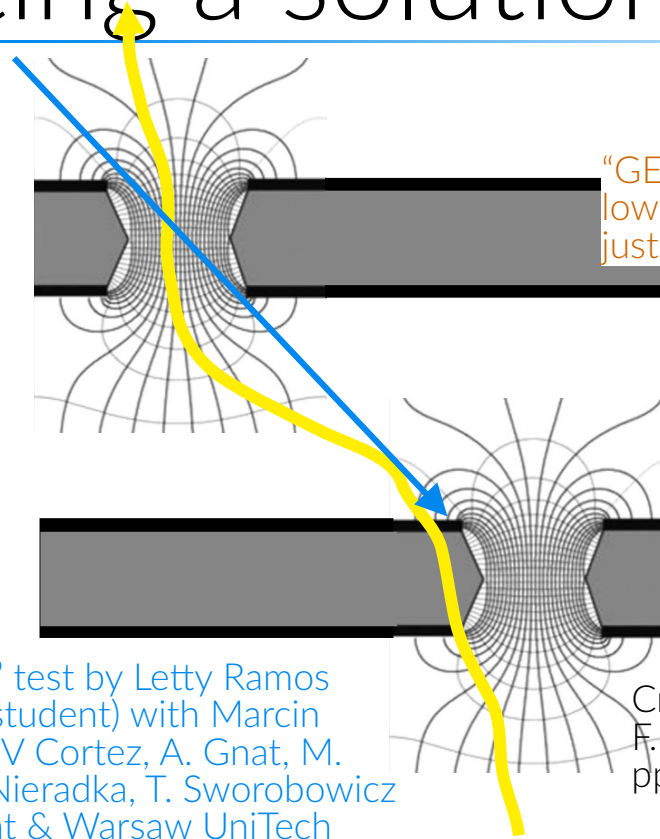
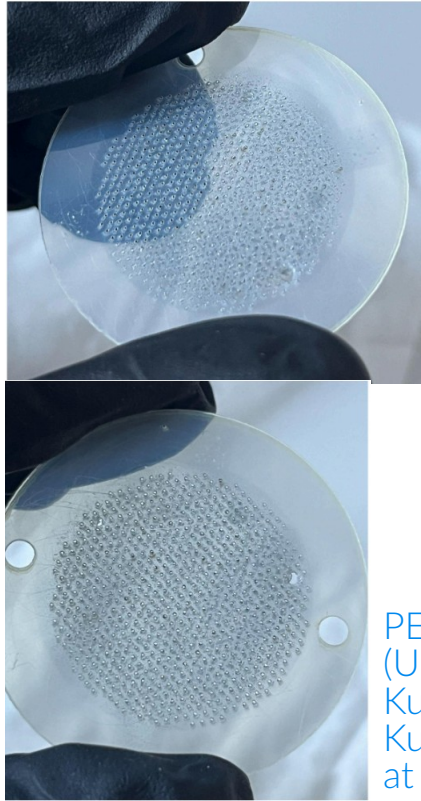
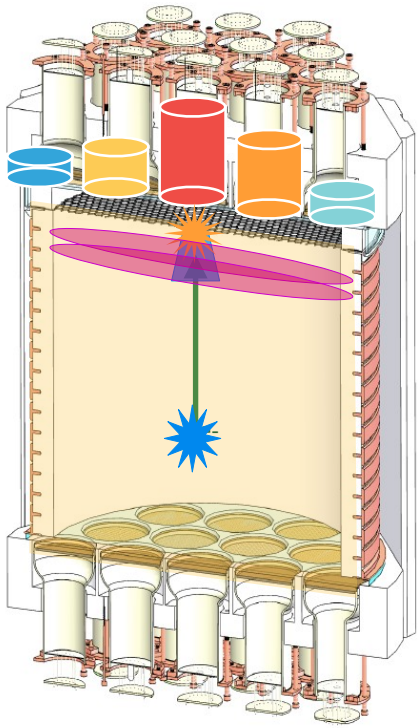
D. F. Anderson, NIM. A 245, 361 (1986).

# Challenge for photosensitive dopants: S2 feedback



Need to prevent S2 from re-entering liquid or shift S2 to below dopant's ionization energy

# Challenge for photosensitive dopants: Testing a solution



“GEM” in LAr at voltage too low for electroluminescence just directs electron traffic

PEN “GEM” test by Letty Ramos (UCR PhD student) with Marcin Kuzniak, AFV Cortez, A. Gnat, M. Kuzwa, G. Nieradka, T. Sworobowicz at AstroCent & Warsaw UniTech

Crudely adapted from F. Sauli NIM A 805 (2016) pp 2-24

Two GEM-like structures (without the Gas and without the Electron Multiplication) to form a “grid” that is transparent to electrons but opaque to VUV photons

# R&D needs

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## **Modeling and reducing SEs**

Dedicated R&D to identify primary mechanisms

Improved chemical purity, likely at the ppb level and better

Investigate other subdominant SE production mechanisms

## **Calibration**

Calibrate charge and scintillation yield of low-energy ERs and NRs

Need to understand the **fluctuations** about mean yields at low energy

## **Doping LAr**

Calibrate doped LArTPCs to low-energy ERs and NRs vs. dopant concentration

Develop high-purity doping techniques—need to add dopant w/o adding impurities

Assess and improve the stability of dopant in LAr

For photo-sensitive dopants, need to prevent S2 feedback

# A Saga Begins

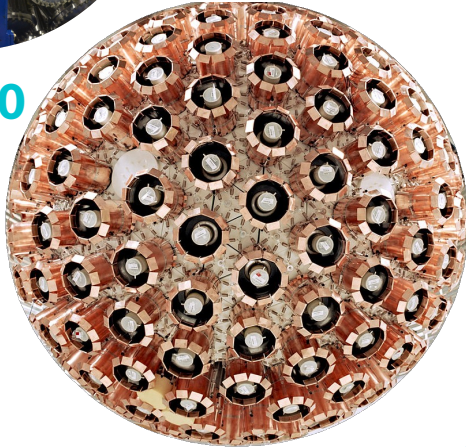
ArDM  
Canfranc



MiniCLEAN  
SNOLAB

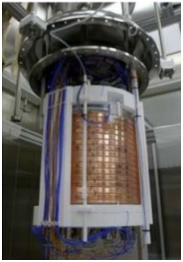


DEAP-3600  
SNOLAB

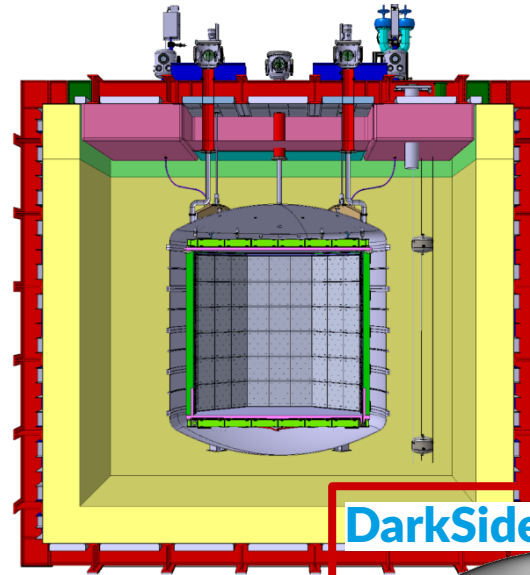


DarkSide-50  
LNGS

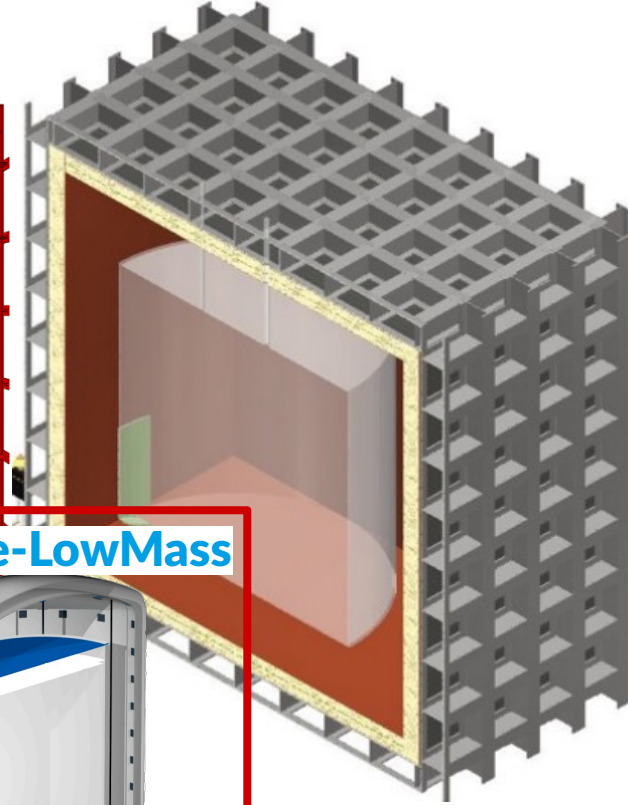
DS-10  
LNGS



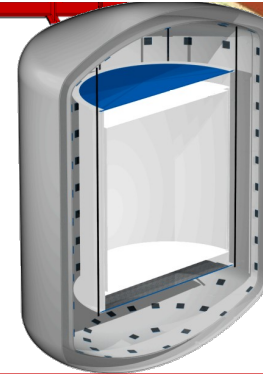
DarkSide-20k  
LNGS



Argo  
SNOLAB



DarkSide-LowMass



2011

2013

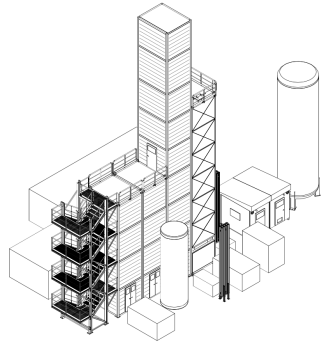
2016

2026

2030+

# Radiopure underground argon

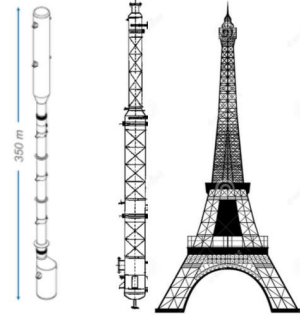
## Urania



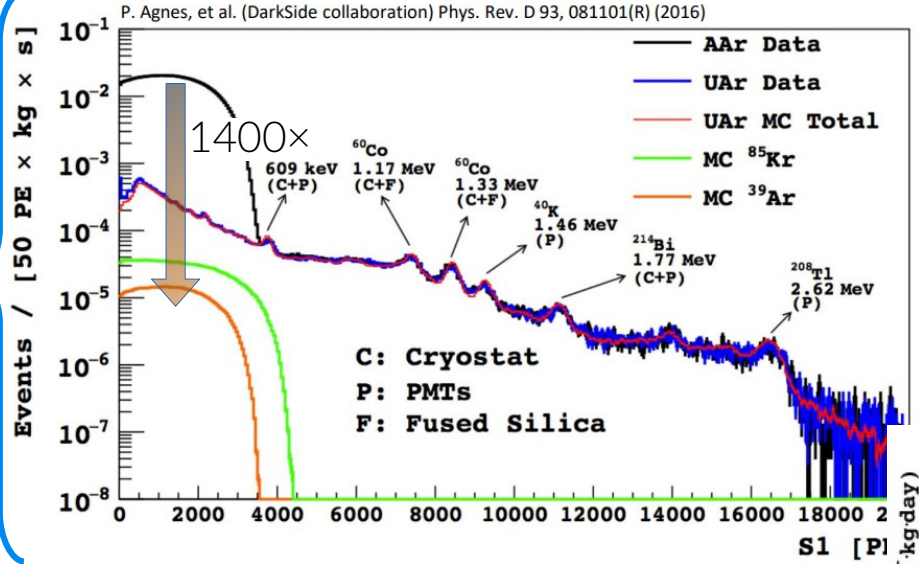
UAr extraction plant  
250 kg/d  
99.99% purity

## Aria

Seruci-I Seruci-II

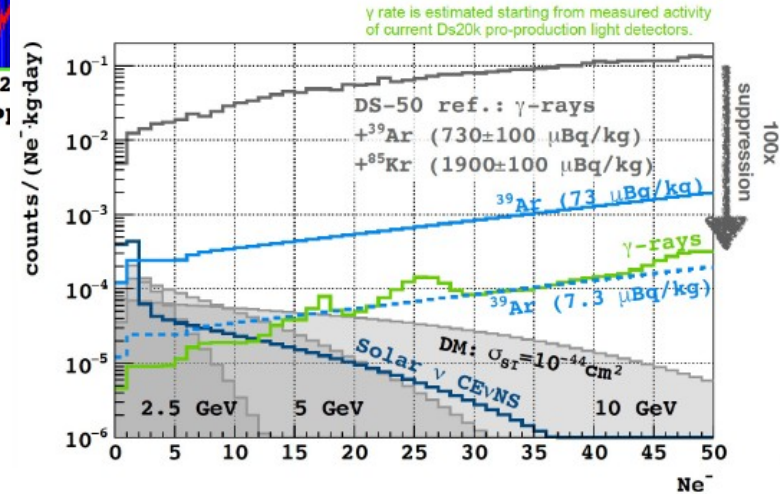


Seruci-0



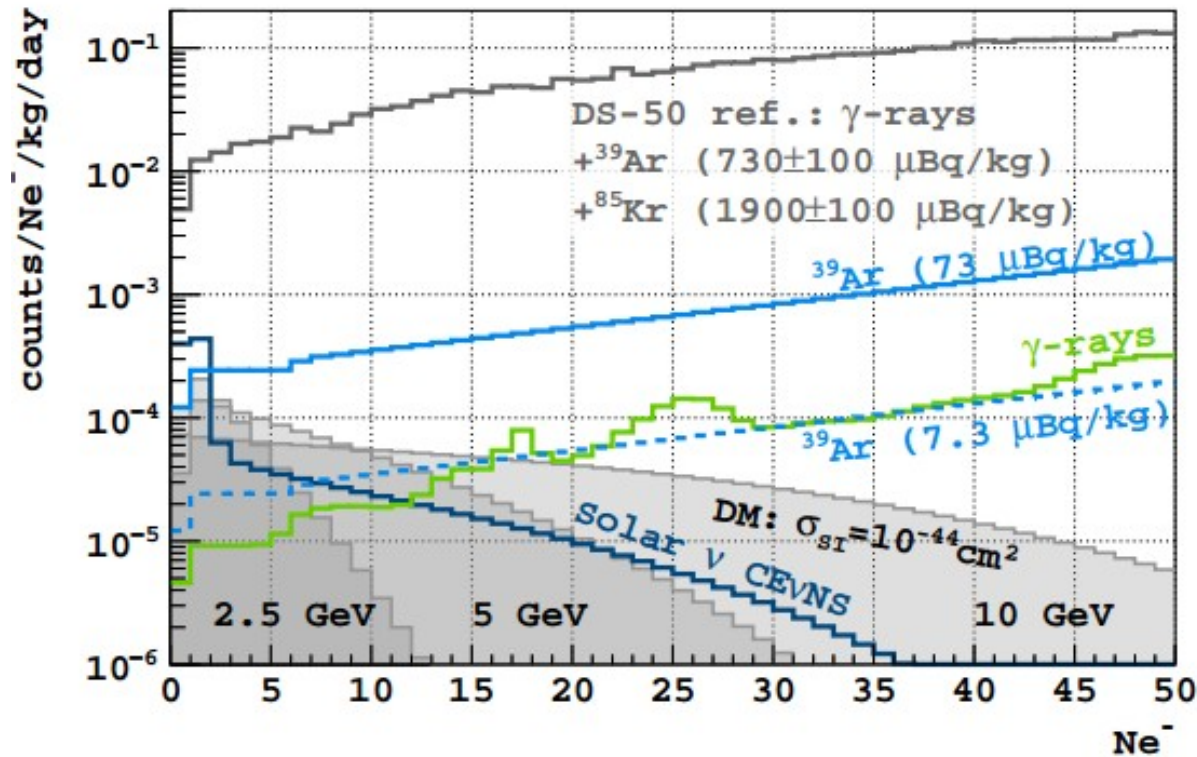
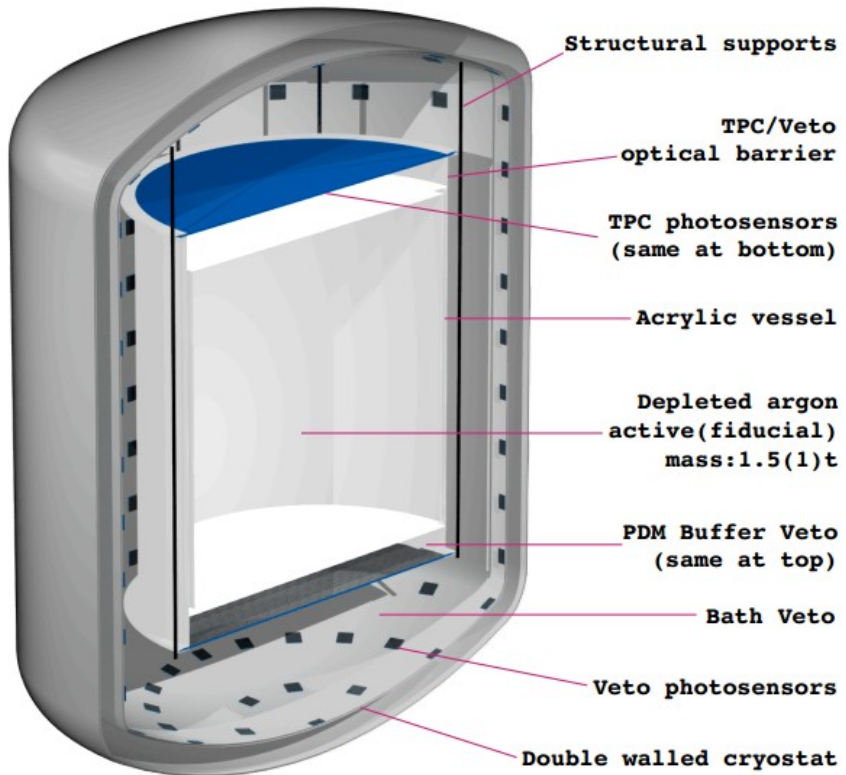
A hipster dark matter scientist who was into argon while it was still underground

**Chemical purification:**  $10^3\times$  reduction at O(1 tonne/day)  
 **$^{39}\text{Ar}$  depletion:**  $10\times$  at O(10 kg/day)





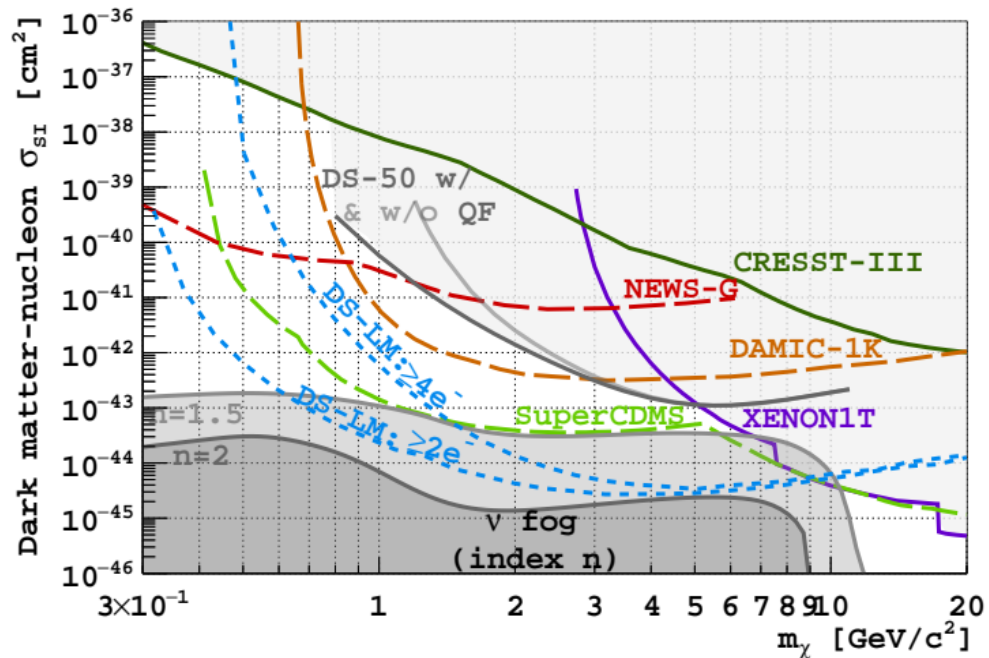
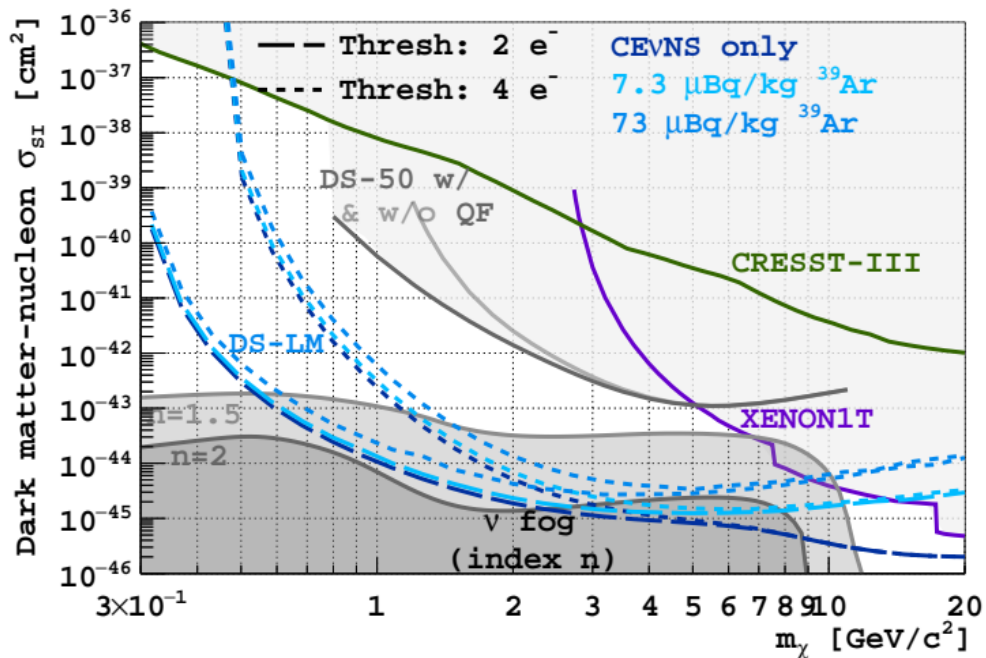
# DarkSide-LowMass: Optimized for S2-only analyses

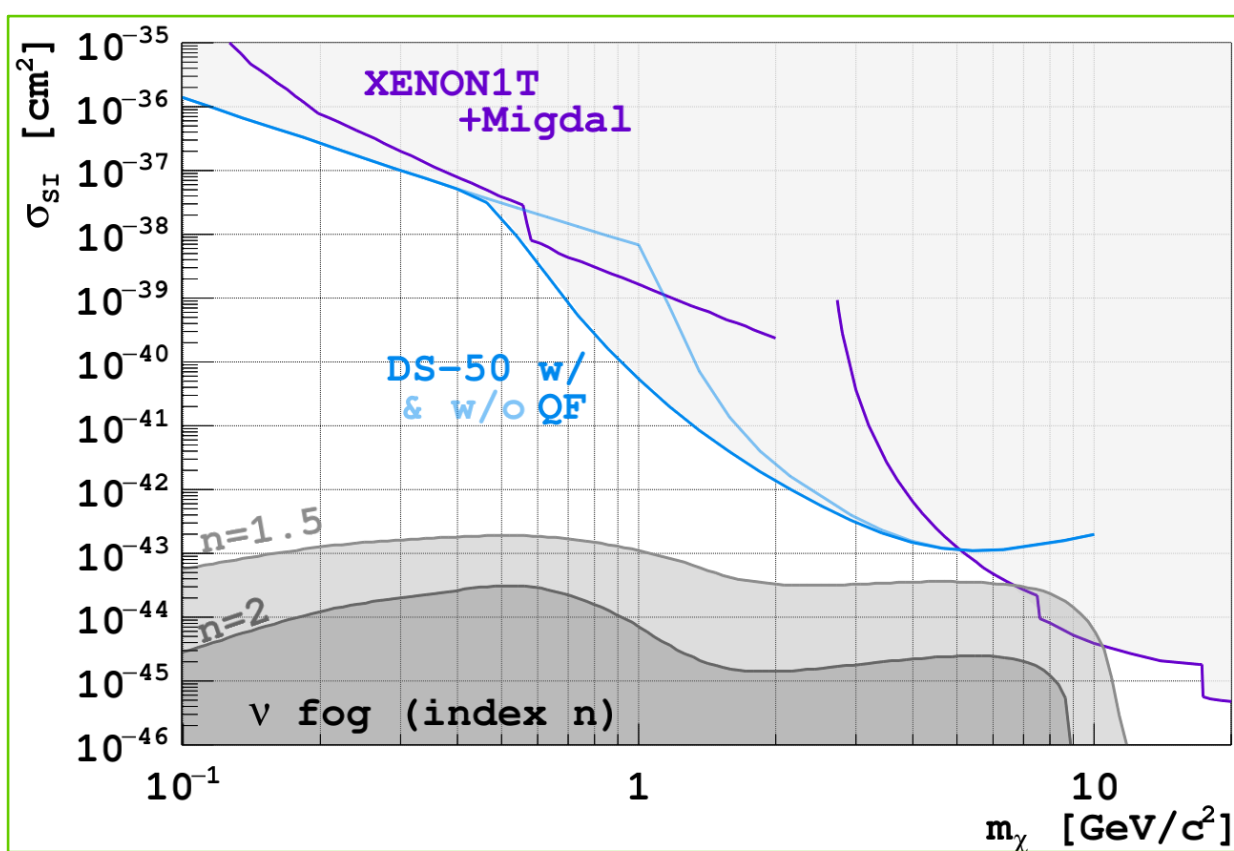


Cryostat immersed in water tank (not shown)

Not shown: spurious electron backgrounds

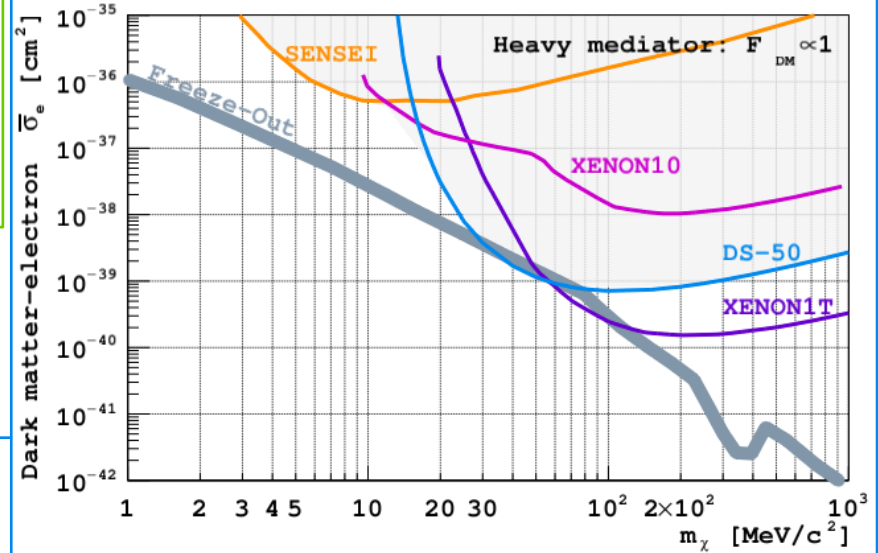
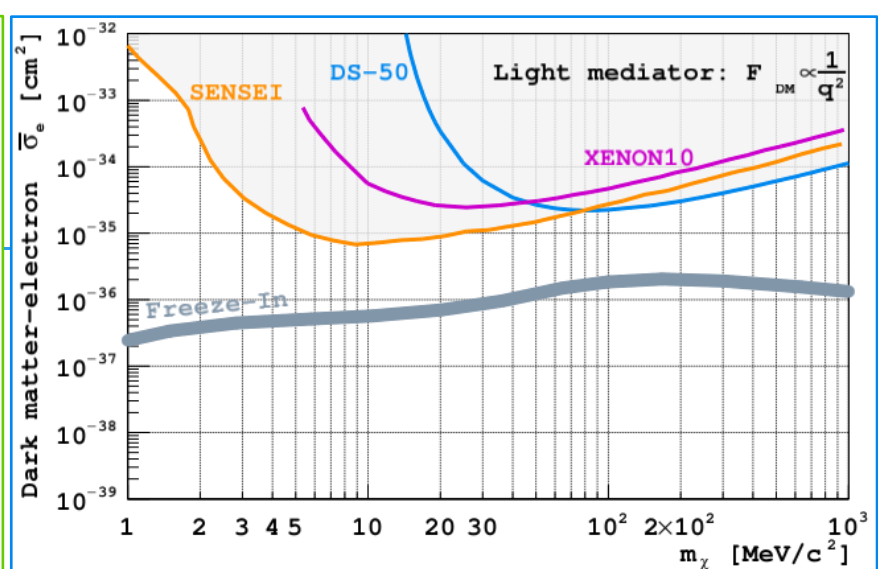
# DarkSide-LowMass: Sensitivity to the $\nu$ fog in 1 tonne year exposure

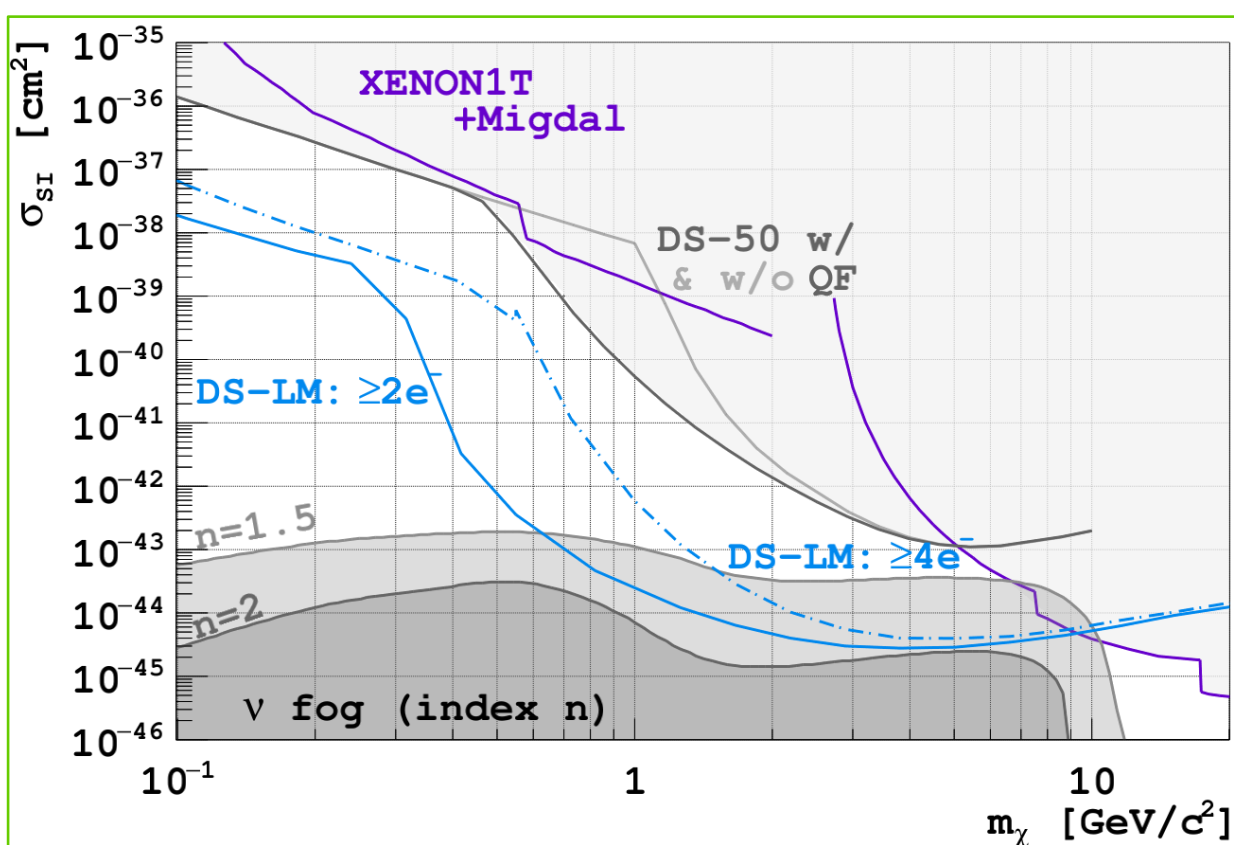




**DarkSide Collaboration.** "Search for dark matter-nucleon interactions via Migdal effect with DarkSide-50". arXiv:2207.11967 (2022)

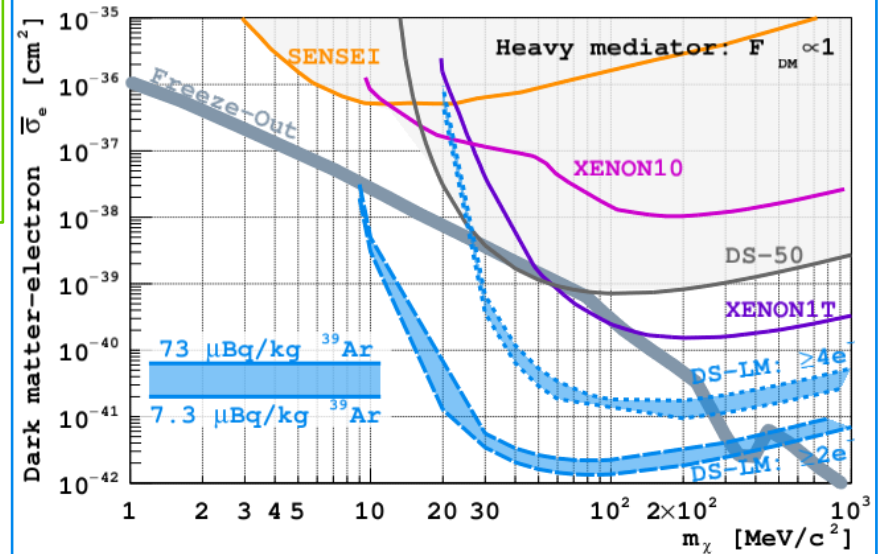
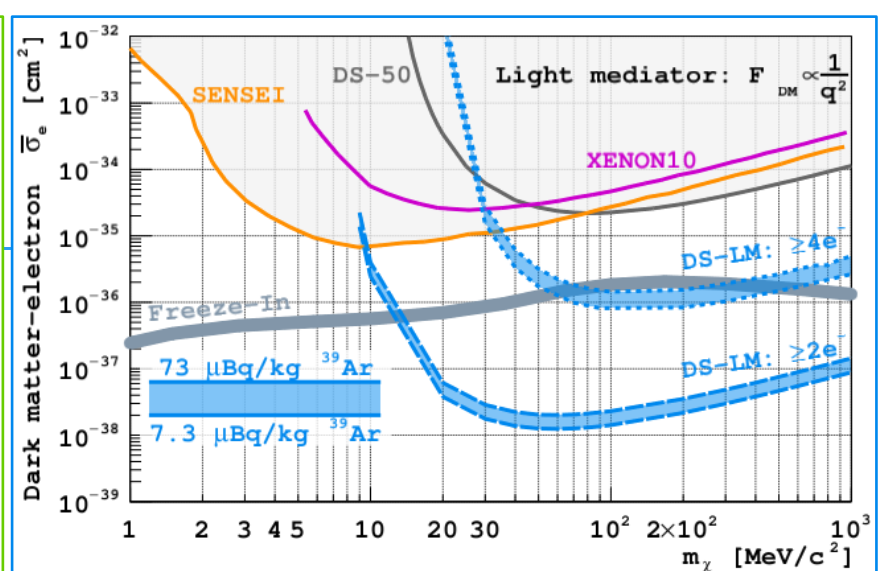
**DarkSide Collaboration.** "Search for dark matter particle interactions with electron final states with DarkSide-50". arXiv:2207.11968 (2022)





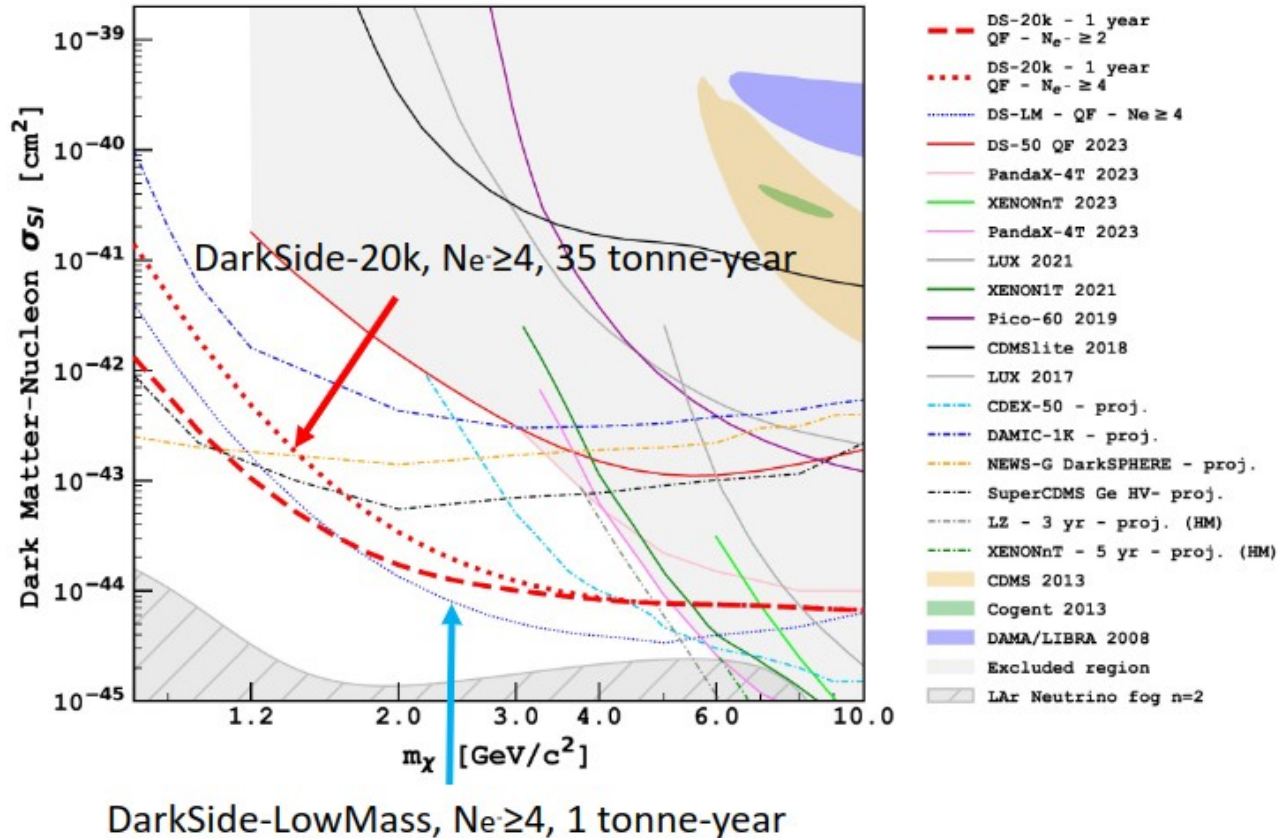
**Nuclear couplings:** Sensitivity to GeV-scale candidates down into the neutrino fog

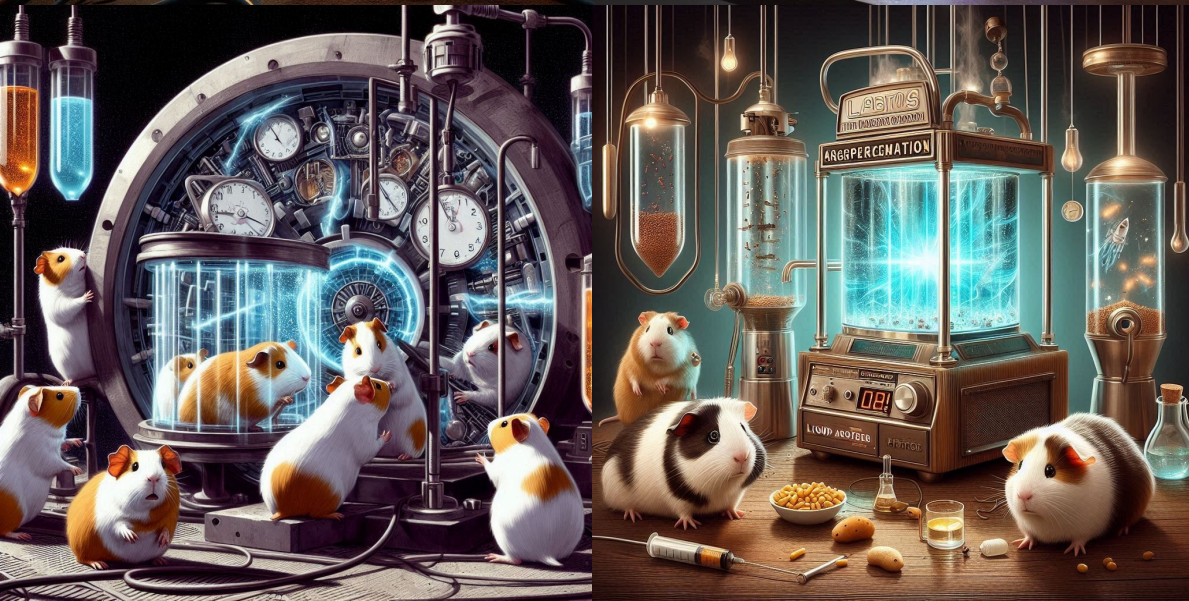
**Electronic couplings:** Sensitivity to sub-GeV candidates covering cross sections that explain the relic abundance in freeze-in and freeze-out production mechanisms



# DarkSide-20k

arXiv:2407.05813





END