



## Cryogenic Platforms at Fermilab: Recent Results and Outlook

Dylan J Temples, Fermilab

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GUINEAPIG Workshop @ UToronto

FERMILAB-SLIDES-24-0223-ETD



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**ENERGY**

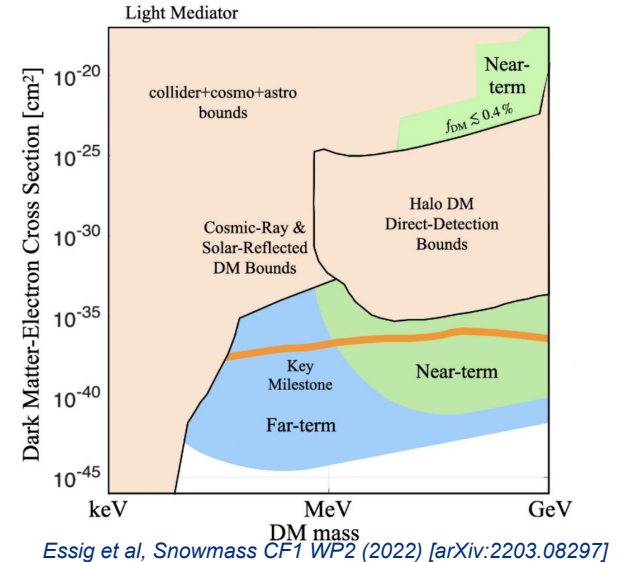
Office of Science

# GeV and Under Invisibles

Sub-GeV regime is an interesting parameter space to look for dark matter!

**Advantages:** high DM number density, larger cross-sections still unexplored

**Challenges:** low DM kinetic energy ( $\sim 1$  eV for  $1$  MeV/ $c^2$  DM), new backgrounds



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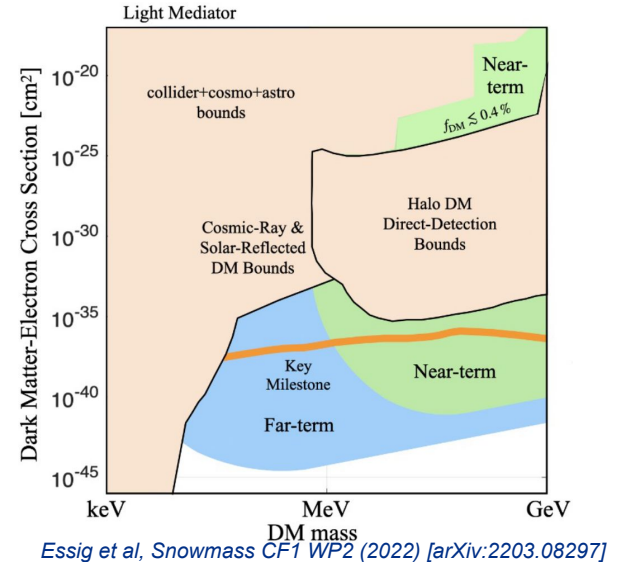
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What is needed to get there?

Energy threshold: sensitivity to sub-eV quanta

- Ballistic / athermal phonons ( $\sim 1$  meV)
- Cooper pairs / quasiparticles (10s-100s  $\mu$ eV)



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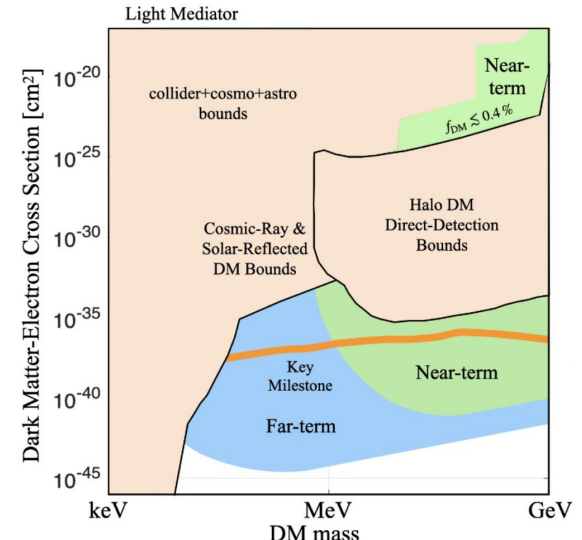
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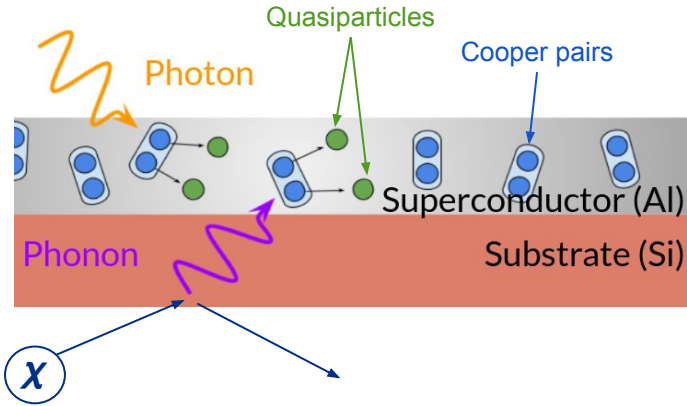
Backgrounds: new characterization and mitigation

- Radiogenic (“all-surface” detectors)
- Non-radiogenic (e.g., stress release)



Essig et al, Snowmass CF1 WP2 (2022) [arXiv:2203.08297]

# Sub-GeV DM Direct Detection with Superconducting Sensors



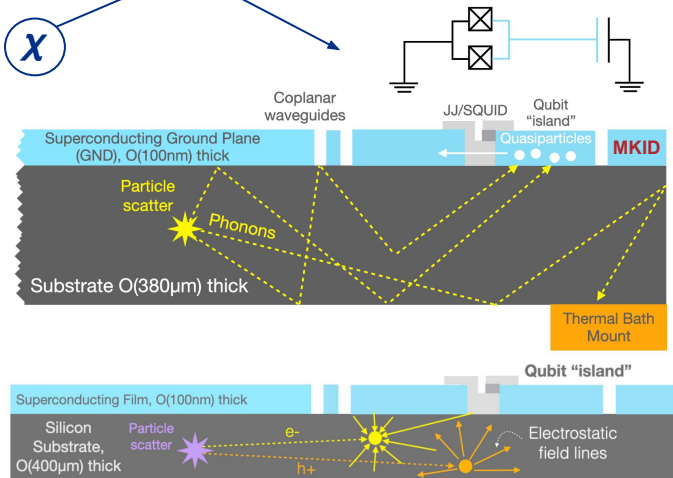
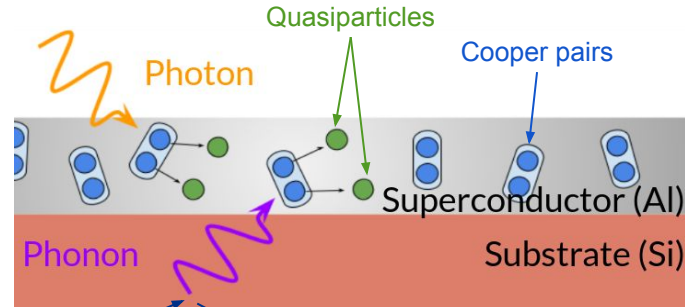
**Scattering target:** crystalline substrate (Si, Sapph, ++)

- DM-nuclear scattering
- DM-electron scattering / Bosonic DM absorption
- DM coupling into collective modes (phonons)

**Ballistic phonons** propagate to superconducting film  
→ break Cooper pairs into quasiparticles (QPs)

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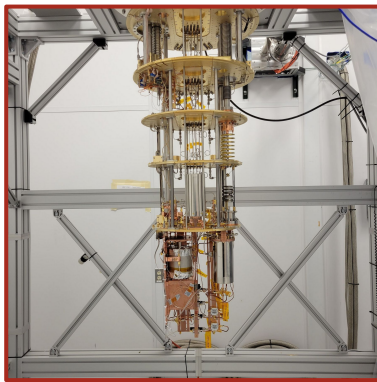
**Primary signal channel:** QP density ( $n_{qp}$ )

**Sensor architectures:**

- Kinetic Inductance Phonon-Mediated Detectors (KIPMDs)
  - $n_{qp}$  modulates resonant frequency of MKID
- Superconducting qubits
  - $n_{qp}$  sets rate of QP tunneling through junction
  - *Non-quasiparticle sensing schemes possible*
- TES, SNSPD, QCD (not in this talk)



- ❖ CryoConcept HexaDry dilution refrigerator (10 mK)
- ❖ 107 m rock overburden
- ❖ Moveable lead shield
- ❖ RF + DC payloads for qubits and DM detectors



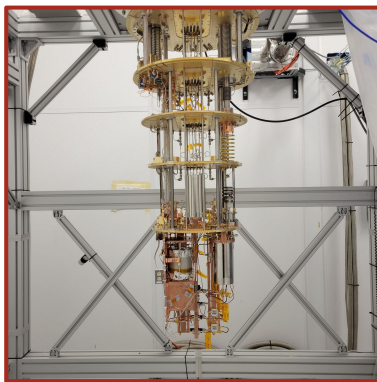
- ❖ Oxford Proteox dilution refrigerator (20 mK)
- ❖ Surface-level facility
- ❖ Qubit-based sensor characterization & calibration
- ❖ Scannable optical source



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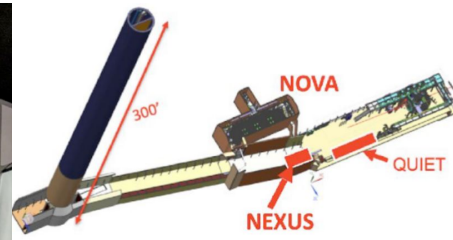
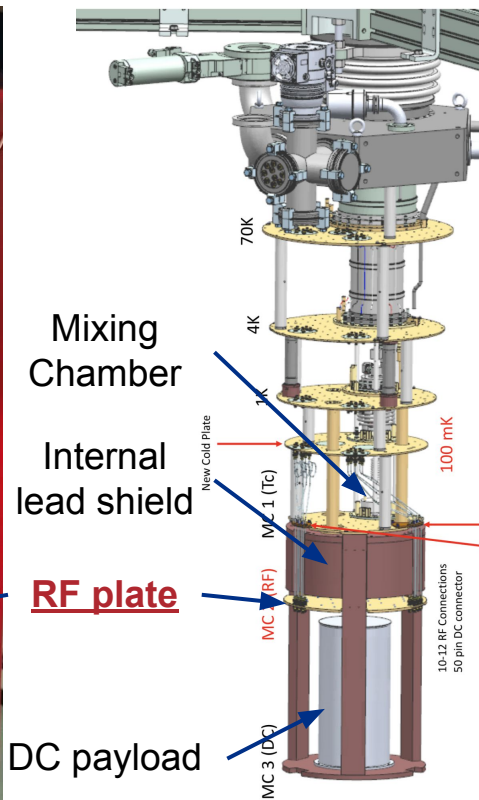
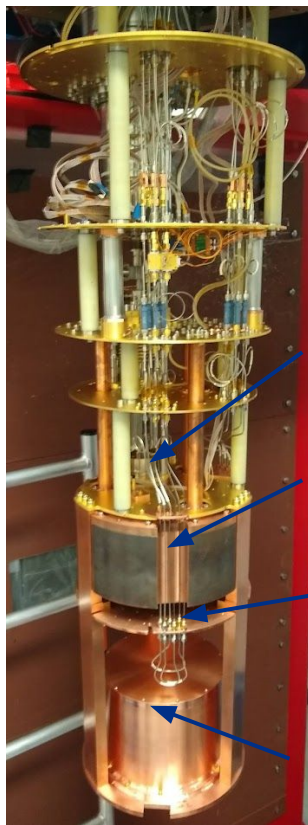


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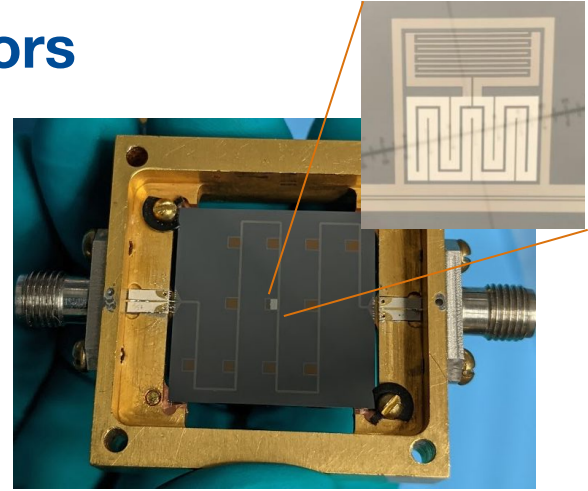
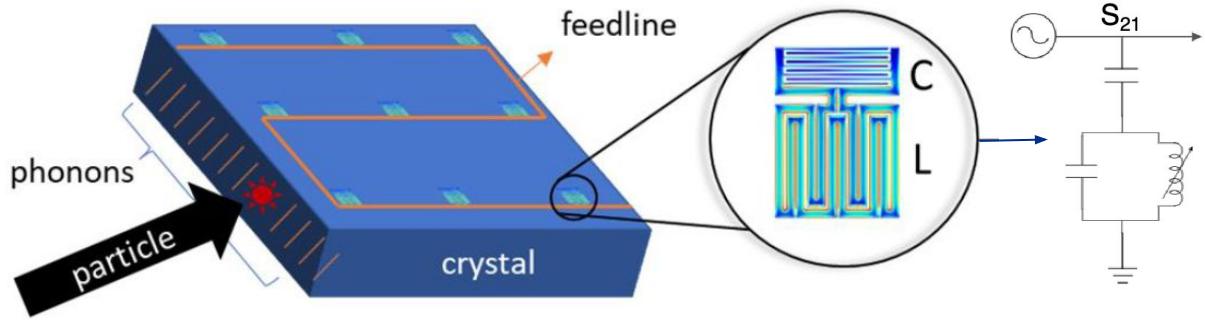




**107 m rock overburden:** reduction in cosmic muon flux by 2 orders of magnitude (7 muons/cm<sup>2</sup>/day)

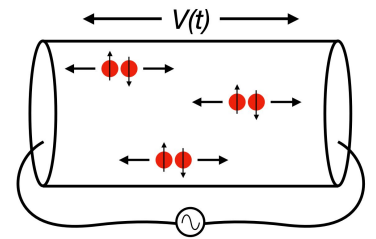


# Kinetic Inductance Phonon-Mediated Detectors



## Kinetic Inductance Detector (KID)-based microcalorimeters

- RLC resonator with nonlinear inductance (inertia of charge carriers)
- Probe complex transmission ( $S_{21}$ ) of RF stimulus at resonant frequency
- Frequency-domain multiplexability: simple scaling, pixelized readouts
- Path to sub-eV resolution on phonon channel!



# Collaborators and Acknowledgements

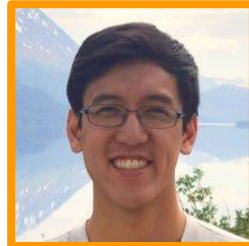
## Fermilab



**D. Temples**



**L. Hsu**



**O. Wen**



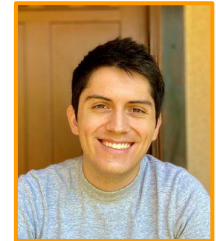
**K. Ramanathan**



**S. Golwala**



**T. Aralis**



**B. Sandoval**

## Caltech

## SLAC / Stanford



**N. Kurinsky**



**Z. Smith**



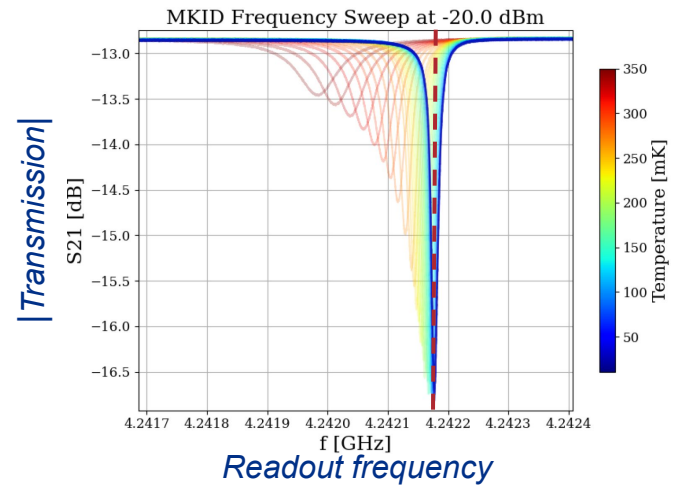
**H. Magoon**

**JPL:** R. Basu Thakur  
B. Bumble

**Undergrads:** S. Dang (Cornell)  
S. Kevane (Stanford)  
S. Ray (Northwestern)  
G. Spahn (UMN)  
C. Cap (Caltech)

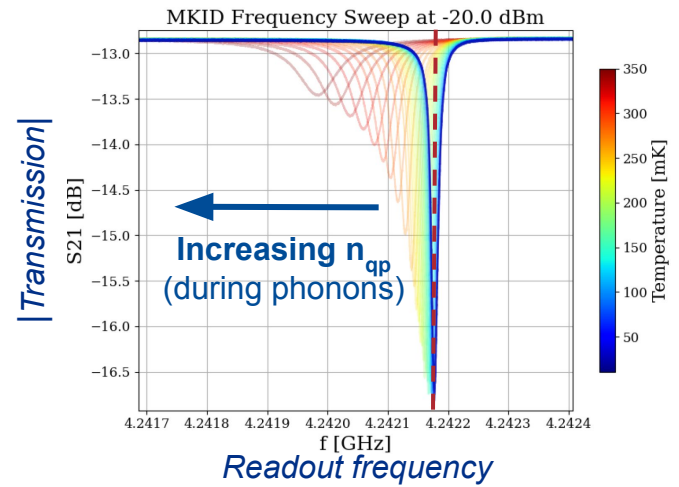
# KIPM Detector Readout

1. Park your radio at  $f_{r,0}(T_{op})$ 
  - Measure (in time) flat transmission + **noise**



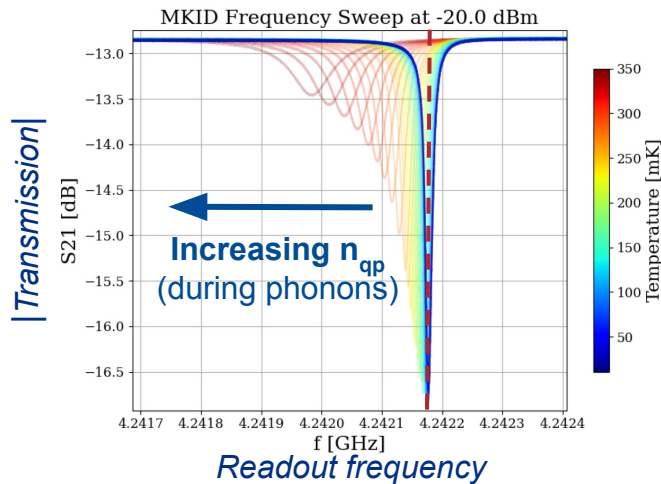
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2. Incident phonons  $\rightarrow$  prompt increase  $n_{qp}$ 
  - Shift  $f_r$ ,  $Q_r$  down



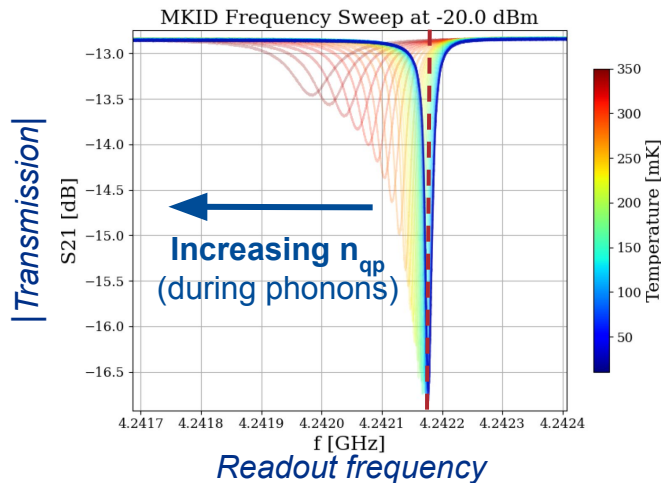
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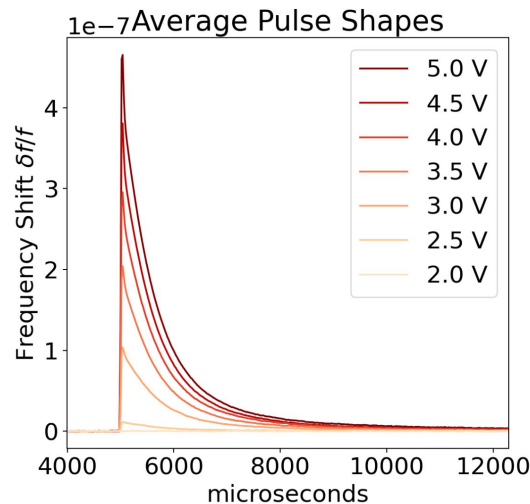
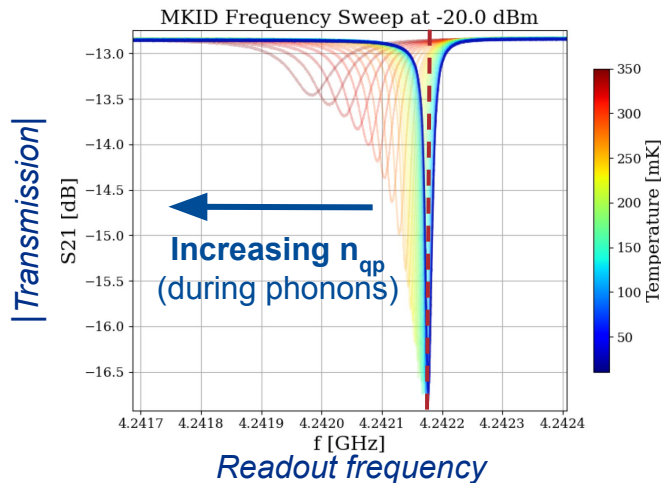
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5. Time domain response: pulses!
  - Frequency ( $S_{21}$  phase):  $\delta f/f$
  - Dissipation ( $S_{21}$  amplitude):  $\delta(1/Q)$
  - Pulse amplitude  $\propto n_{qp}$  ( $\propto$  Energy)





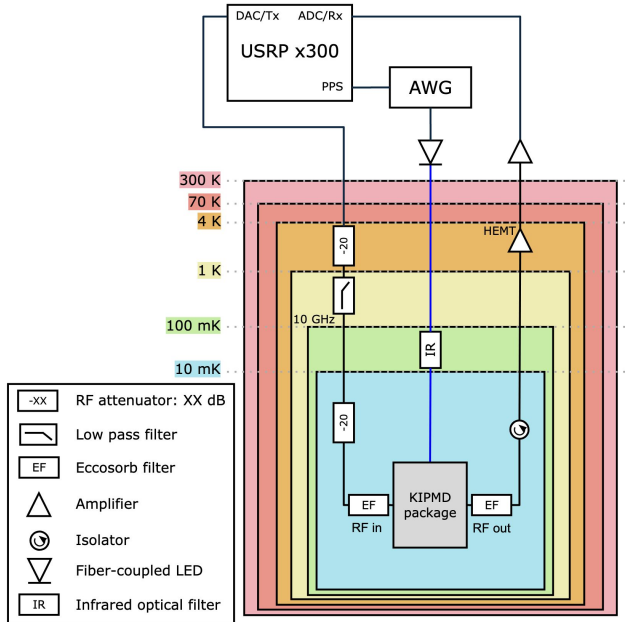
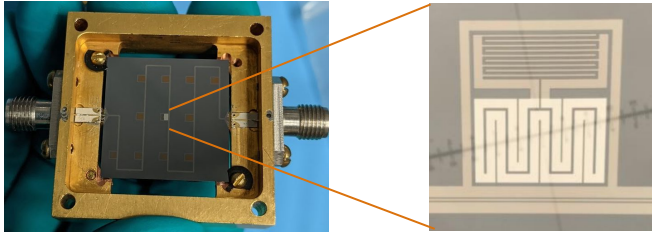
# KIPMD Calibration at NEXUS

Exploit Poisson statistics of photon bursts to infer total energy deposited

- Expose KIPMD substrate to pulsed optical source of varying optical power
- Phonon energy absorbed by sensor measured by change in  $n_{qp}$  (via  $\delta f/f$ )

Simultaneous measurement of absolute energy resolution and phonon collection efficiency

$$\sigma_E = \frac{\sigma_E^{abs}}{\eta_{ph}} = \frac{\sigma_{n_{qp}} V \Delta}{\eta_{ph}}$$



# Energy Resolution Model

## Two-component energy resolution model

- Intrinsic device noise ( $\sigma_0$ )
- Photon shot noise ( $\sigma_{LED}$ )  $\propto \sqrt{\bar{N}_\gamma}$

Cardani et al. *Supercond. Sci. Technol.* **31** 075002 (2018)

Cruciani et al. *Eur. Phys. J. C* **81**, 636 (2021)

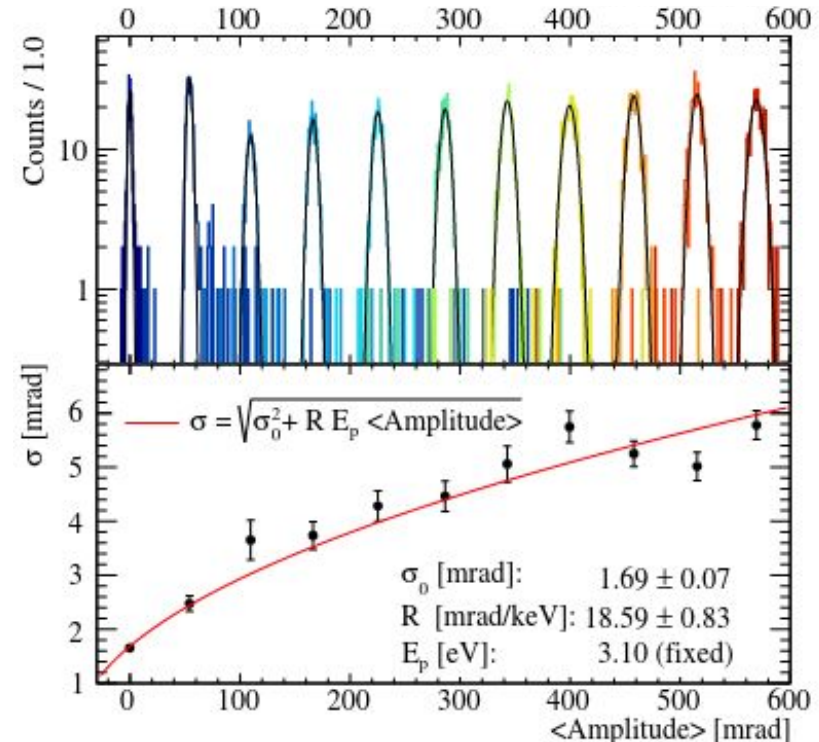
Cardani et al. *Appl. Phys. Lett.* **121**, 213504 (2022)

$$\sigma_E = \sqrt{\sigma_0^2 + \sigma_{LED}^2} = \sqrt{\sigma_0^2 + r \times \mu}$$

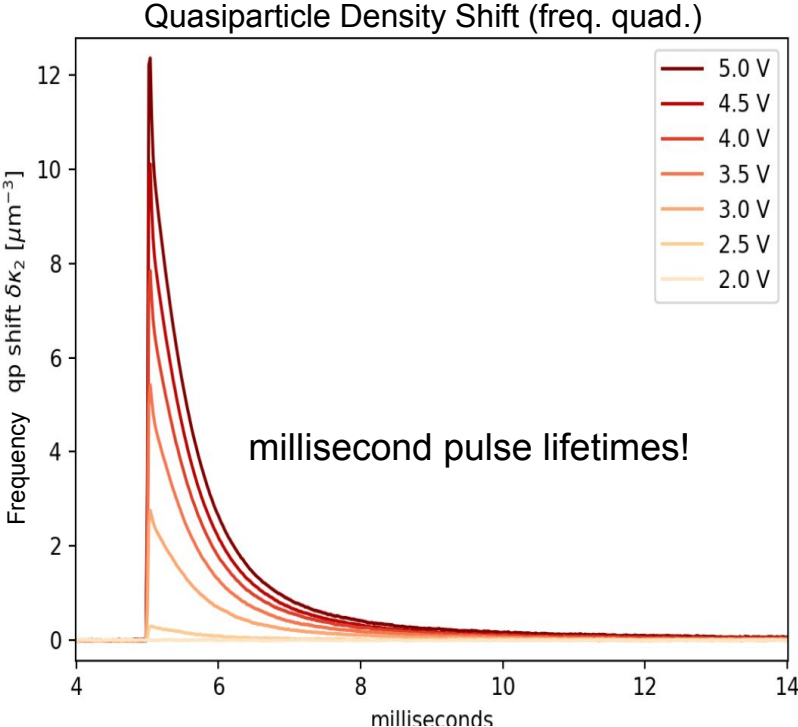
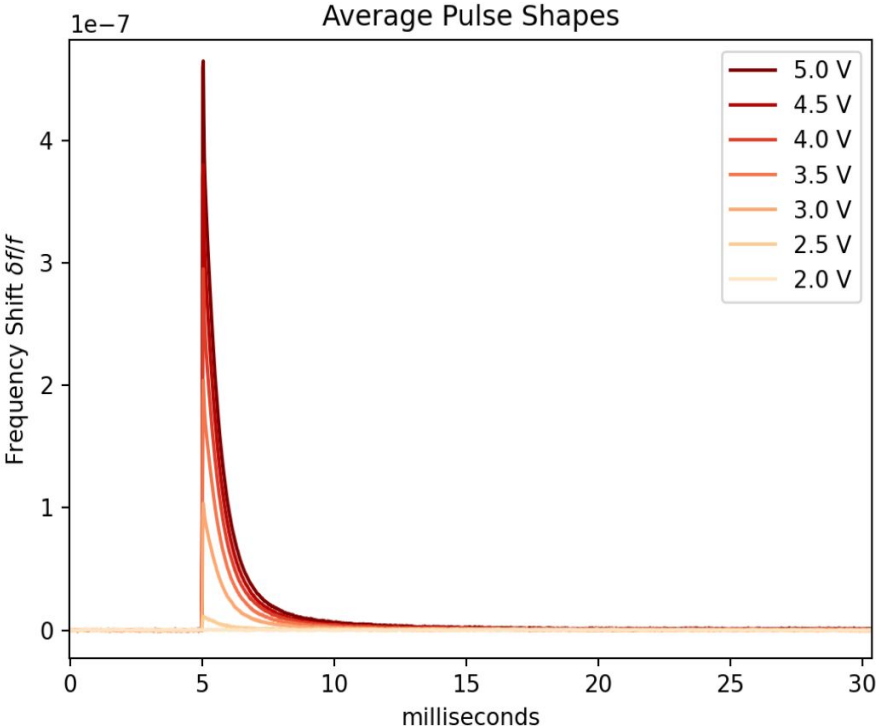
Responsivity per unit energy

Average pulse amplitude for a given  $\bar{N}_\gamma$

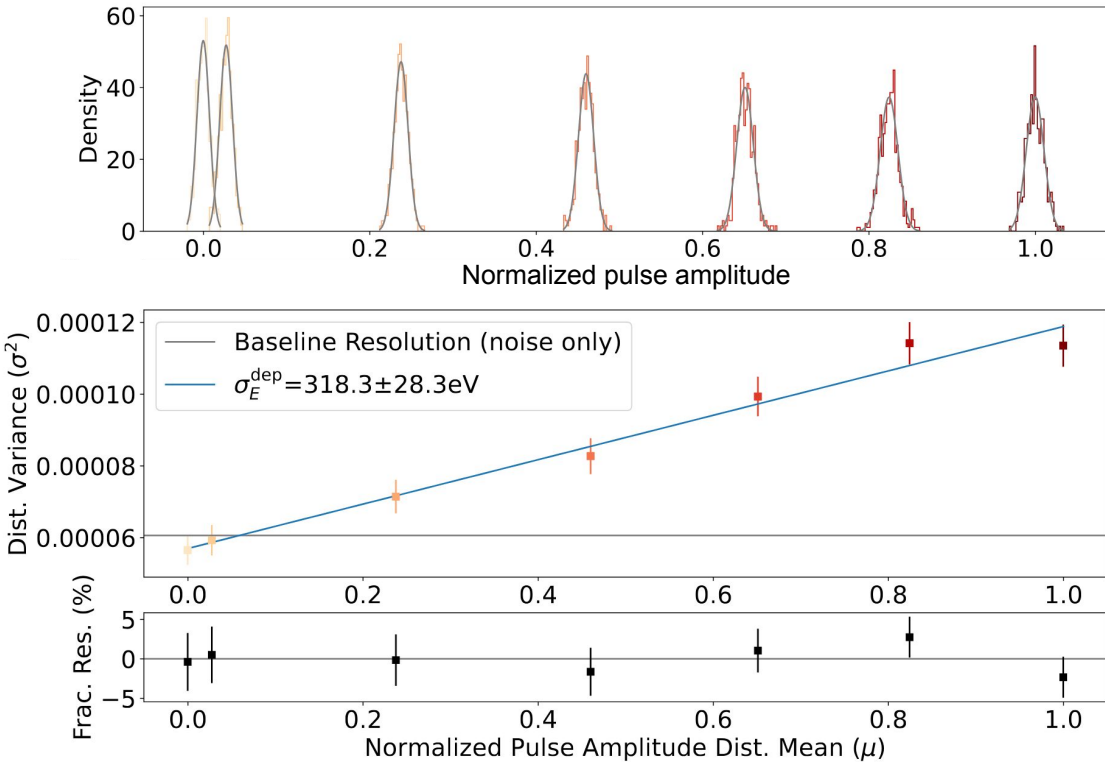
**Example:** Cardani et al. *Appl. Phys. Lett.* **121**, 213504 (2022)



# Average LED Pulses



# Device Performance: Energy Resolution



From pulse amplitude in  $\delta f/f$ :

$$“1.0” = s_{\delta f/f} = 4.65 \times 10^{-7}$$

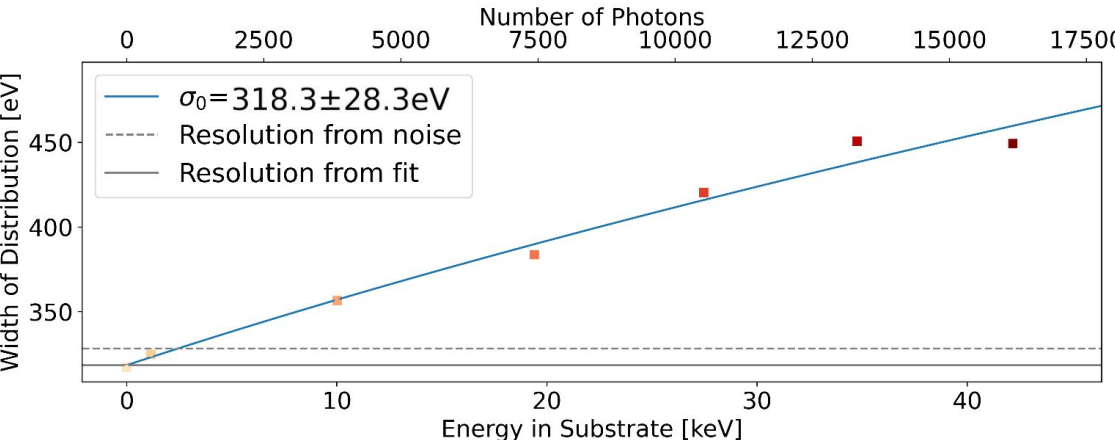
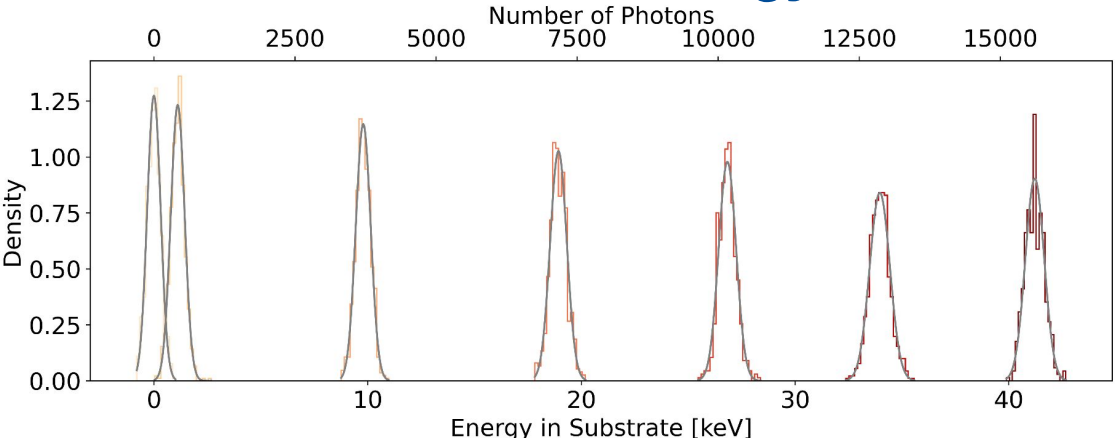
Resolution on qp density:

$$\sigma_{\text{qp}}^{\kappa_2} = \frac{2}{\alpha \kappa_2} (\sigma_0 s_{\delta f/f}) = 0.47 \mu\text{m}^{-3}$$

Resolution on energy absorbed by sensor:

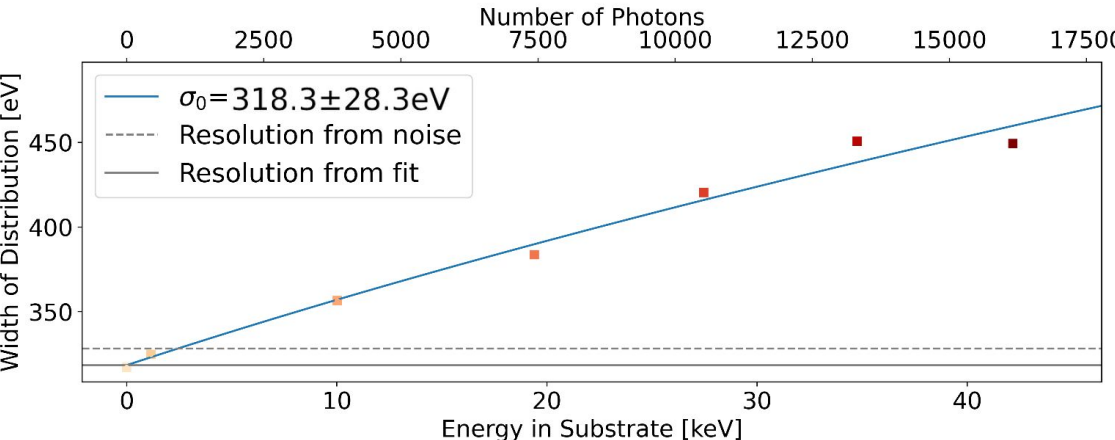
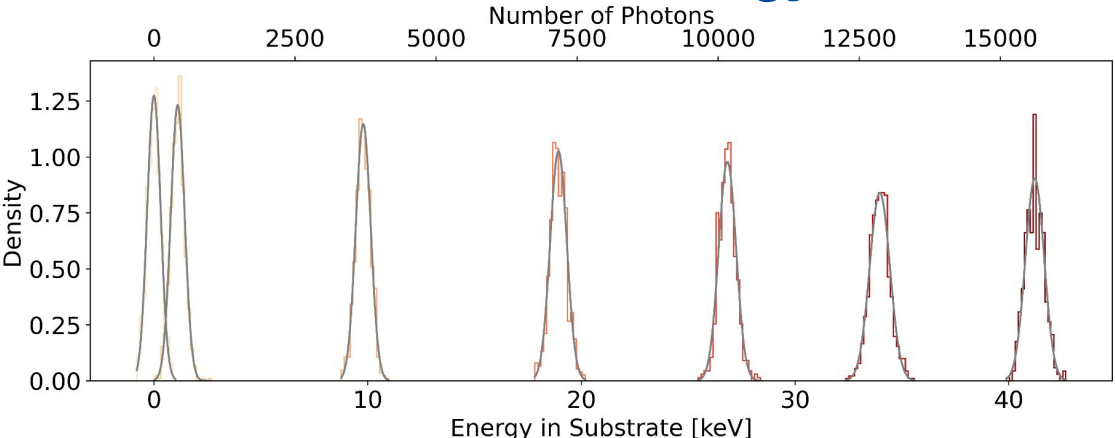
$$\sigma_E^{\text{abs}} = \sigma_{\text{qp}} V \Delta = 2.5 \pm 0.4 \text{ eV}$$

# Device Performance: Energy Resolution



Date	Feb 15, 2023	Jul 13, 2023
$\sigma_E^{abs}$ [eV]	$2.1 \pm 0.2$	$2.8 \pm 0.3$
$\sigma_E$ [eV]	$318 \pm 28$	$315 \pm 28$
$\eta_{ph}$ [%]	$0.66 \pm 0.1$	$0.89 \pm 0.11$

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**Energy resolution primarily limited by poor phonon collection efficiency!**

	CALDER 2021	BULLKID 2022
$\sigma_E^{dep}$	34 eV	26 eV
$\eta_{ph}$	~12 %	24%
$\sigma_E^{abs}$	4.1 eV	6.25 eV
	Cruciani et al. <i>Eur. Phys. J. C</i> <b>81</b> , 6366 Cardani et al. <i>Supercond. Sci. Technol.</i> <b>31</b> 075002 (2018)	Cardani et al. <i>Appl. Phys. Lett.</i> <b>121</b> , 213504



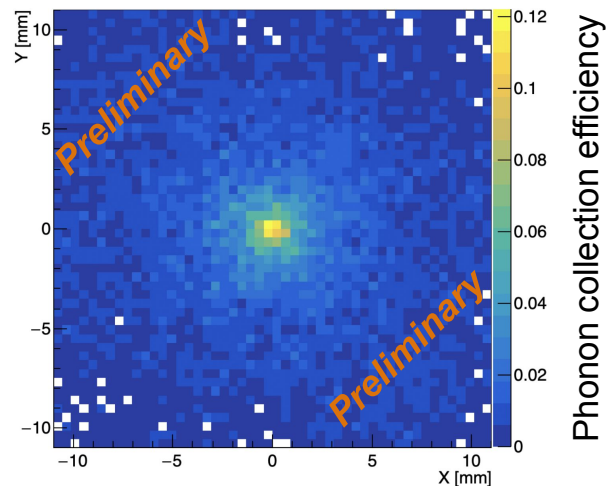
# Phonon Loss Mechanisms

- Phonon egress at device mounts
- Down-conversion in high-gap metals (e.g., Nb)
  - Recent simulation work indicates more down-conversion than initially expected
- Surface-mediated down-conversions (“fill fraction” 2% in prototype, 0.1% now)



**Can be investigated with device geometry/packaging and simulation!**

# KIPMDs - G4CMP Simulations

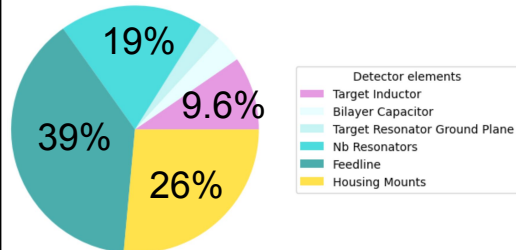
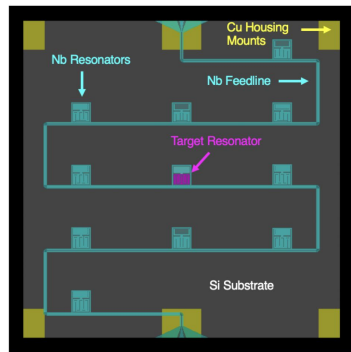


Goal: understand phonon loss contribution from various structures

- Requires phonon absorption probability at interface as simulation input

Simulation predicts larger  $\eta_{ph}$  than observed

- Ongoing investigation into signal generation and parameter inputs (using theory vs sweep)
- Expanding G4CMP processes (community effort led by Ryan Linehan @ FNAL)



Share of phonons lost to detector elements

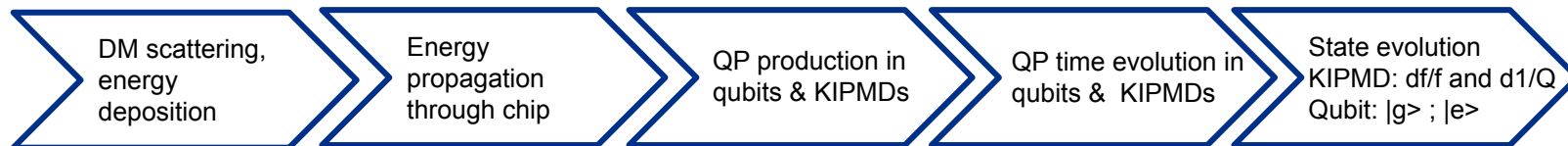
Working on joint calibration with SLAC DMQIS group using steerable cryogenic optical system

- Position dependent calibration allows better reconstruction of phonon absorption prob.



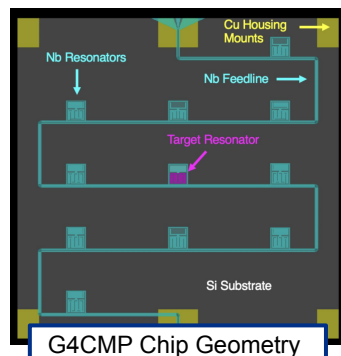
# Simulating Sensor Signals from Energy Deposits in Substrate

Need a “sensor response” simulation to complete sims chain!



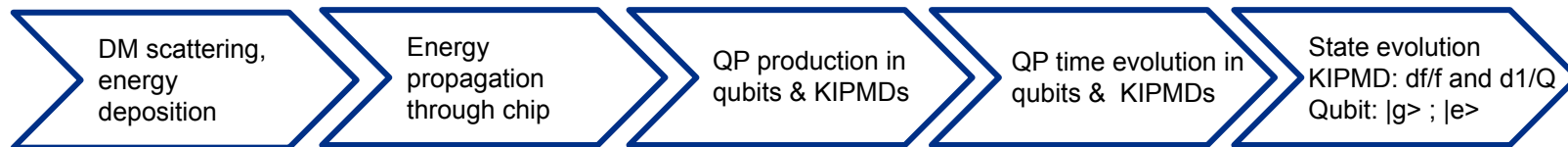
## G4CMP simulation

- Geant4-based
- Phonon and e/h pair tracking
- Simple QP modeling
- Extensions being developed by community



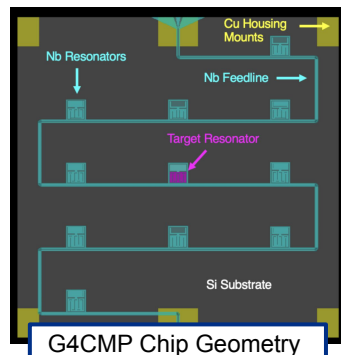
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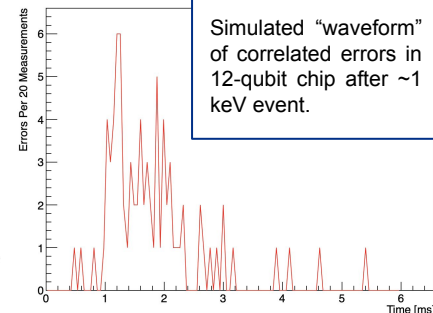
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G4CMP Chip Geometry

## Quantum Device Response (QDR)

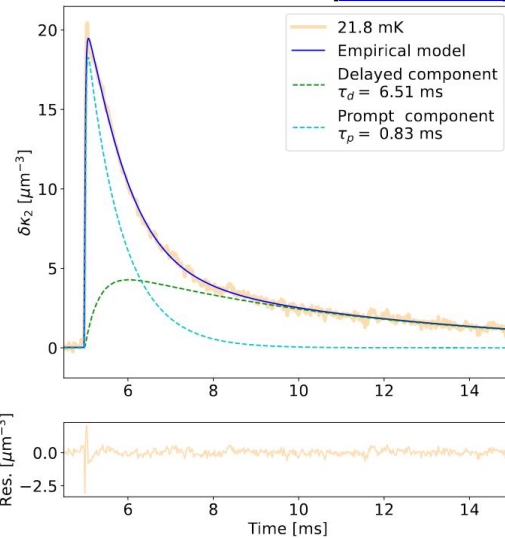
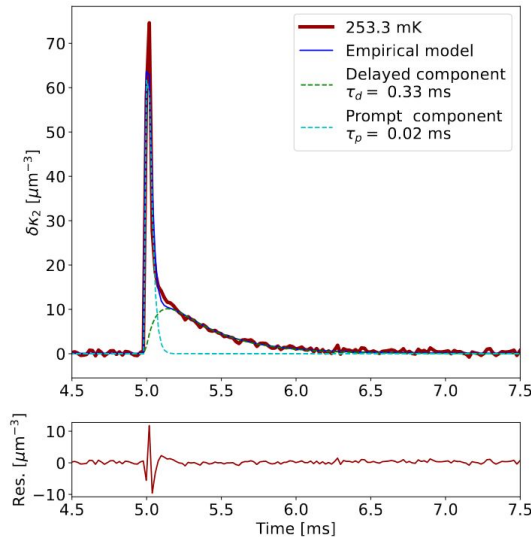
- Folds in detection scheme, critical readout parameters
- Flexible: models multiple sensor types (MKIDs, Transmons), even on same chip!



# QDR Application: Understanding Pulse Shapes

Separating phonon dynamics from quasiparticle dynamics in MKID sensors

DJT, O. Wen, et al. [[arxiv:2402.04473](https://arxiv.org/abs/2402.04473)]



In simulation:

- Known quasiparticle lifetime
- Control over phonon time profile

1. Generate phonon bursts in G4CMP
2. Generate readout waveforms in QDR
3. Pipe through analysis framework to extract the fall time constants as is done with data

Allows us to confirm (or refute) the origin of the two fall-time constants and their dependence on device temperature.

Two fall time constants: phonon lifetime & quasiparticle lifetime?

# QDR Application: Estimating energy thresholds of qubit sensors

We can reconstruct an in-chip energy using  $\eta_{ph}$  and reconstructed single-qubit energies:

$$E_{r,chip} = \frac{\sum_i E_{r,i}}{\eta_{ph,sp}}$$

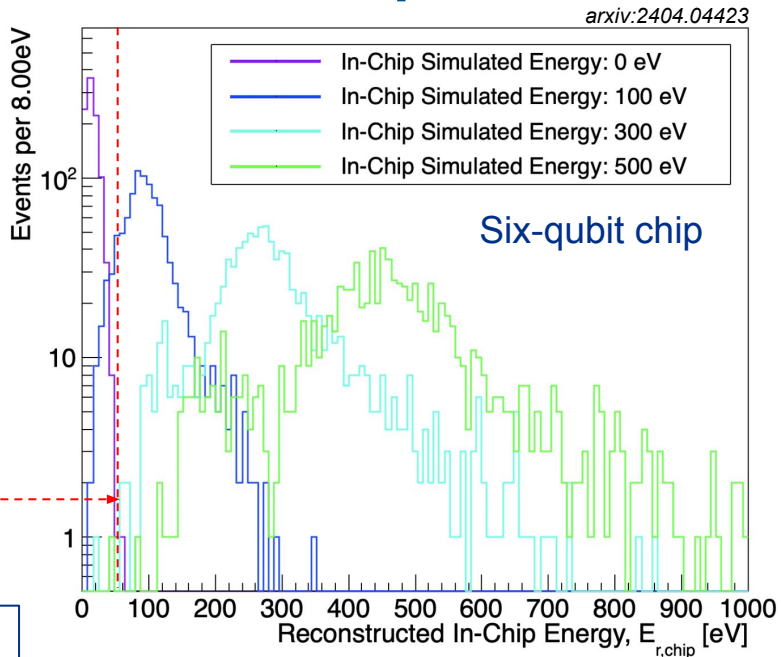
This helps us understand how to find our in-chip threshold:

Equivalent to  $5\sigma_{E,abs}$  for a sum of  $N_q$  qubits' energies

$$E_{thr,chip} \simeq \frac{\sigma_{E,abs}}{\eta_{ph,sp}} \left[ \frac{N_q}{\sqrt{2\pi}} + aN_q^b \right]$$

(with  $a \sim 4.27$ ,  $b \sim 0.44$ )

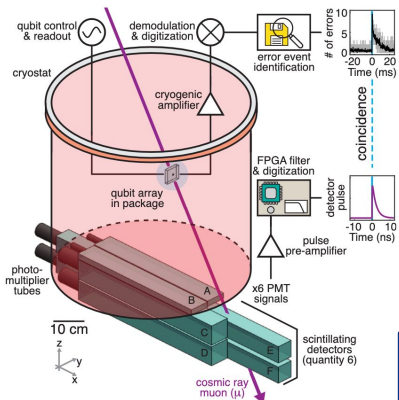
1. For “standard” transmon qubit designs, find **in-chip threshold in range of 40 eV–1200 eV**.
2. With collection fins and QP traps (and no added noise), could reduce to **O(100 meV)**.



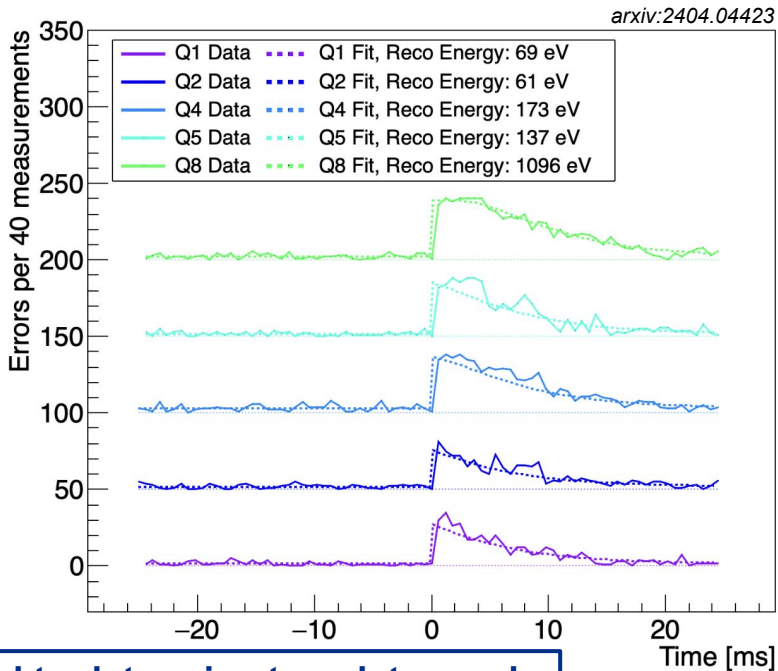
# QDR Application: Extracting Muon Energy from Qubit Errors

Applying sensor/chip energy reconstruction techniques to literature data gives reasonably self-consistent results!

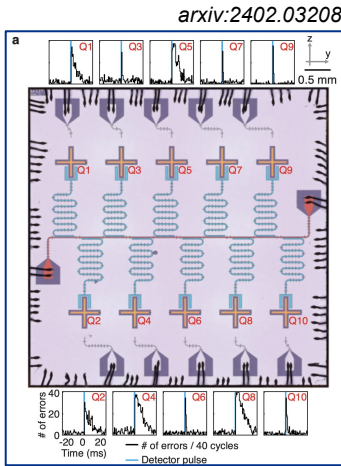
- Single-qubit ML fits
- Best-guess  $\eta_{\text{ph,sp}}$  given chip design
- In-chip energy  $\sim 440$  keV
- Builds confidence in energy reconstruction technique



Harrington et al. (2024) [arxiv:2402.03208](https://arxiv.org/abs/2402.03208)



[arxiv:2404.04423](https://arxiv.org/abs/2404.04423)



[arxiv:2402.03208](https://arxiv.org/abs/2402.03208)

MIT (Harrington) waveforms fit using this technique

**QDR used to determine templates and anchor to real energy scale**



# Qubits: Charge Sensing at NEXUS

Unique features of NEXUS allow us to directly control radiation environment

- Ionizing radiation shown to induce errors in qubit chips
- Investigate spatial and temporal correlations in a qubit chip under different ionizing radiation conditions

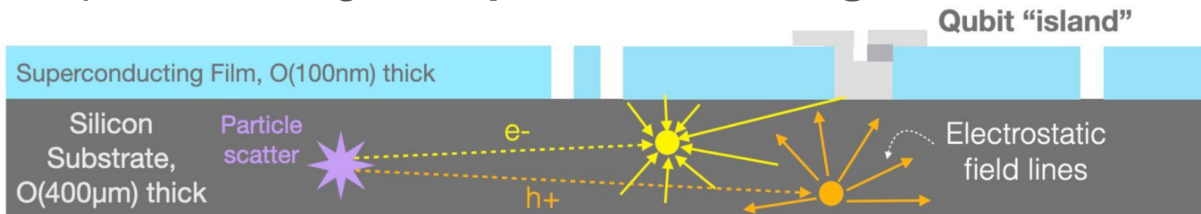
Goal: Understand and exploit radiation effects on qubits to optimize design for **quantum sensing**

# Qubits: Charge Sensing at NEXUS

Unique features of NEXUS allow us to directly control radiation environment

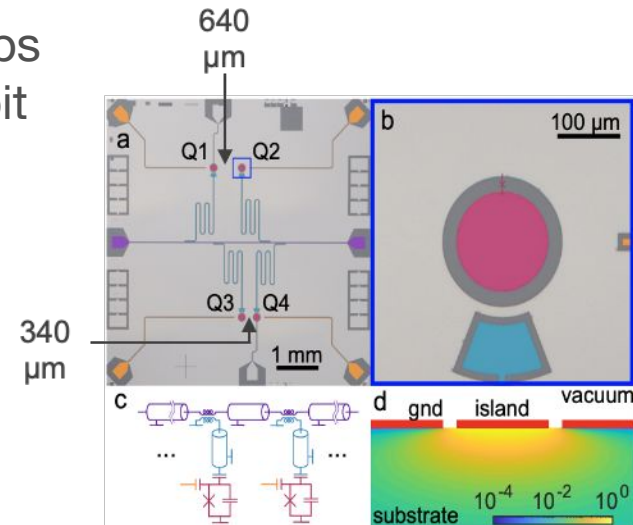
- Ionizing radiation shown to induce errors in qubit chips
- Investigate spatial and temporal correlations in a qubit chip under different ionizing radiation conditions

Goal: Understand and exploit radiation effects on qubits to optimize design for **quantum sensing**



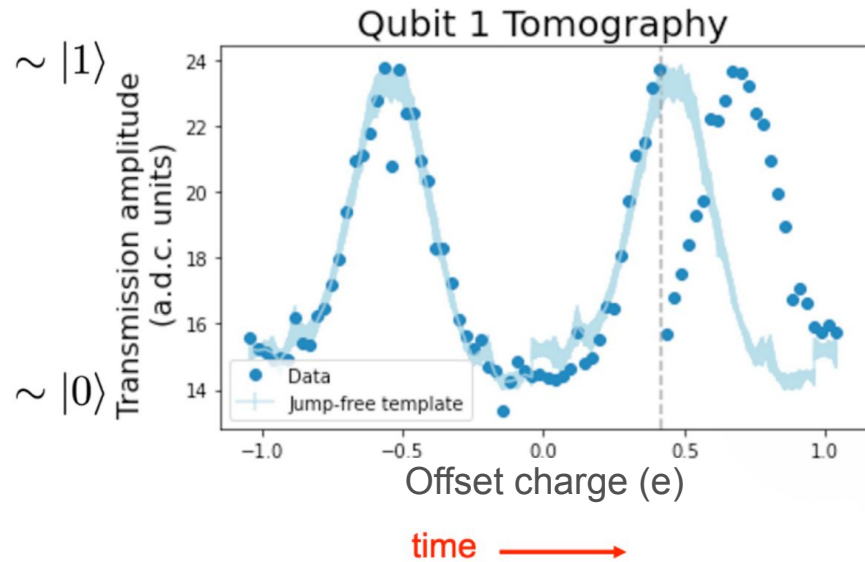
**Weakly charge-sensitive qubits:** sense changes in local electric field via “charge jumps”

Charge trapping: in absence of applied electric field, burst effects can be “locked-in” on timescales of hours -- days



Qubit device used in this work  
(McDermott group, UW Madison)  
Wilén et al, Nature 594, 369 (2021)

# Qubits: Charge Sensing at NEXUS

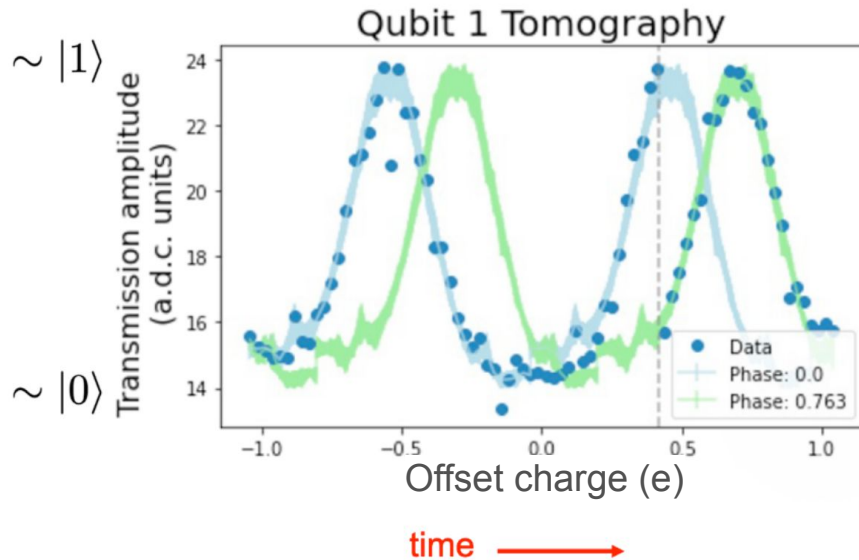


**Ramsey Tomography:** maps “offset charge” to probability of measuring qubit in excited state

- Qubit state periodic in offset charge induced by an applied voltage bias



# Qubits: Charge Sensing at NEXUS

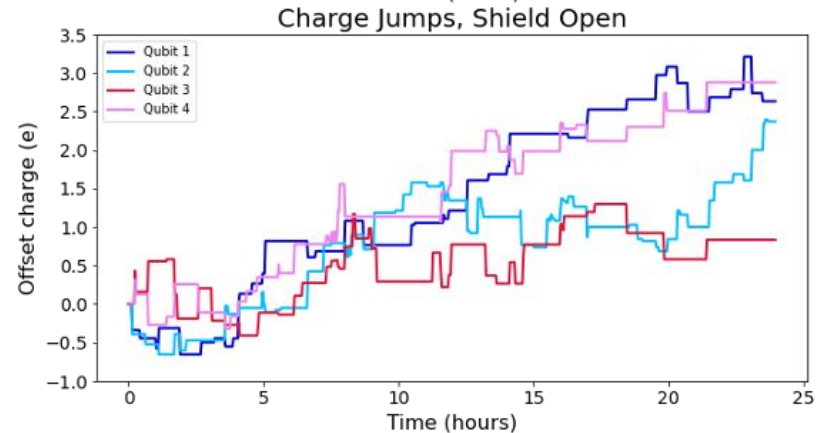
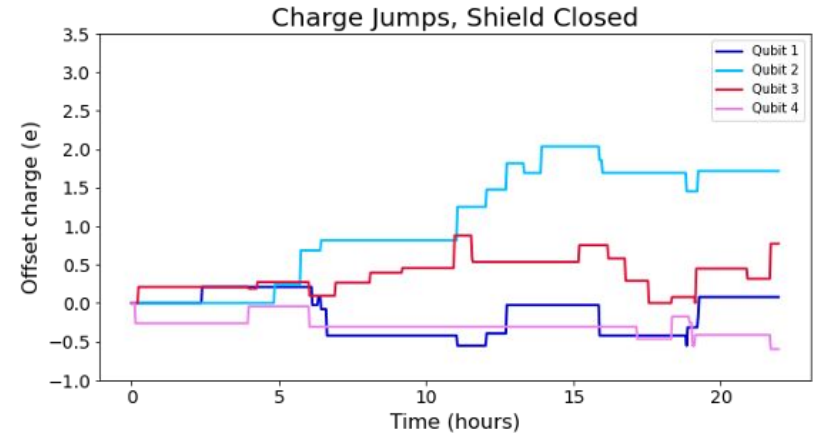


**Ramsey Tomography:** maps “offset charge” to probability of measuring qubit in excited state

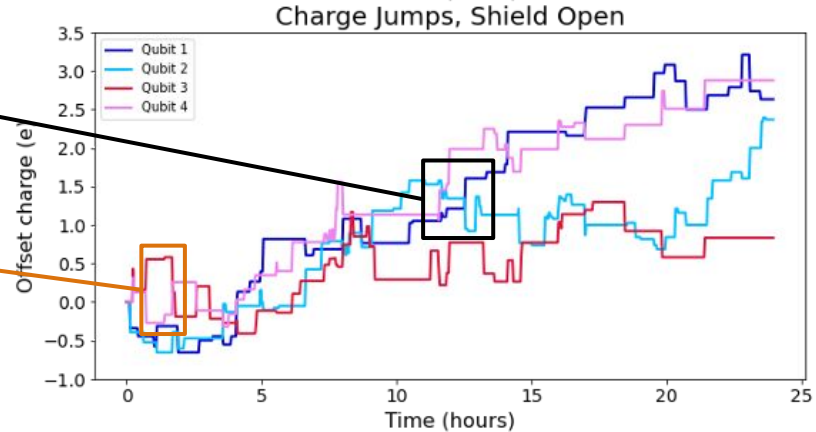
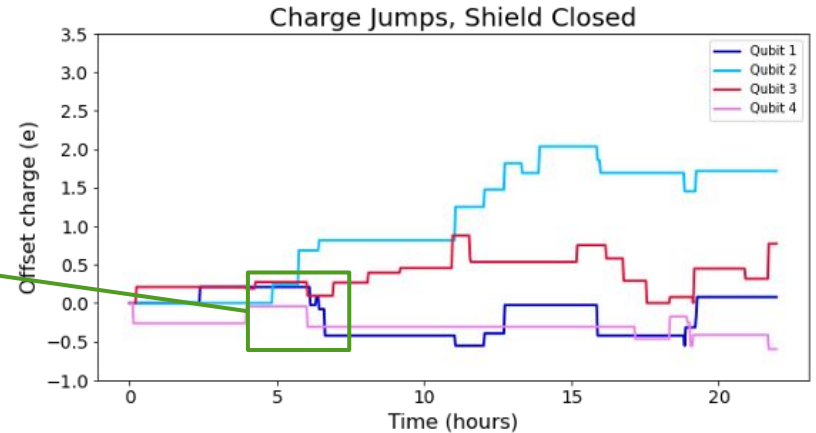
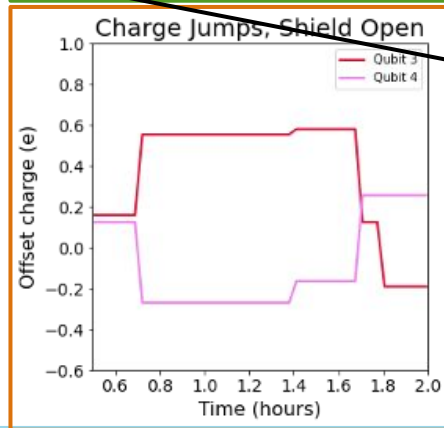
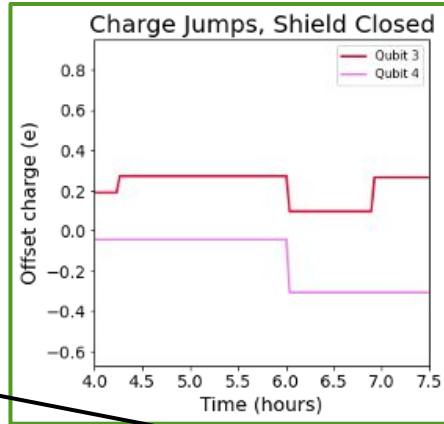
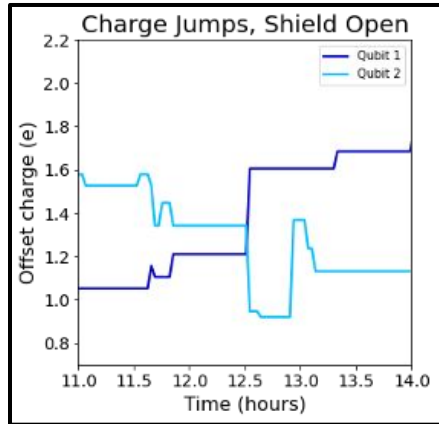
- Qubit state periodic in offset charge induced by an applied voltage bias
- “Charge jump”: Prompt change in local electric field environment leads to qubit state change
- Resolved as phase shift in a Ramsey tomography scan

# Qubits: Charge Sensing at NEXUS

- NEXUS → muon rate negligible, dominated by  $\gamma$ s (controlled by shield)
- Repeated tomography scans over ~20-hour periods of lead shield open vs shield closed
- Impact of shielding on charge jump rate is obvious!
- Can compare our charge jump rates to measurement of same qubit chip on surface [Wilen et al, Nature 594, 369 (2021)]
- Working on development of G4CMP + QDR implementation



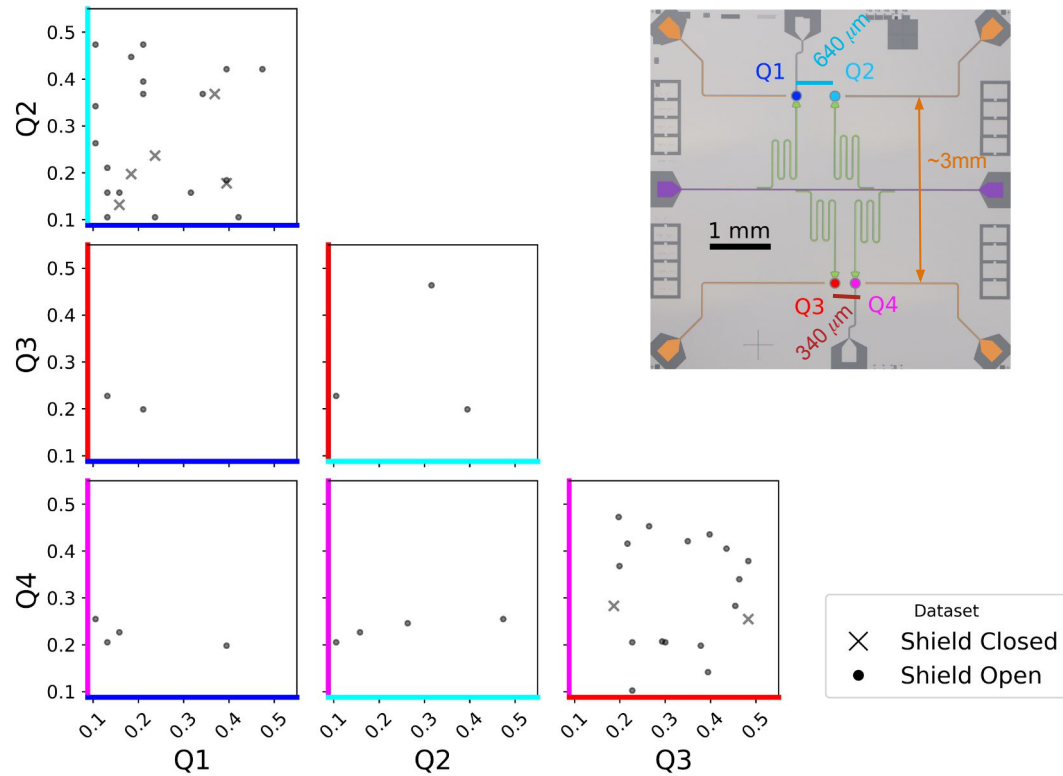
# Qubits: Charge Sensing at NEXUS



# Qubits: Charge Sensing at NEXUS

Underground & shielded operation:

**We observe no correlated errors between qubits spaced  $\sim 3\text{mm}$  apart in 22 hours**



# Qubits: Charge Sensing at NEXUS

Compare to surface rate: 1.35 mHz

- Wilen et al, Nature 594, 369 (2021)

Measured  $\gamma$  flux with witness LMO detector 10 cm away from qubit (in fridge)

- Shield closed: 13x reduction in rate

	Shield Open	Shield Closed	Units
Average Rate	$0.51^{+0.05}_{-0.04}$	$0.19^{+0.04}_{-0.03}$	mHz

	Livetime (hr)	Q1	Q2	Q3	Q4
SO	23.9	$0.42^{+0.09}_{-0.08}$	$0.60^{+0.11}_{-0.09}$	$0.52^{+0.10}_{-0.08}$	$0.51^{+0.11}_{-0.09}$
SC	22.1	$0.20^{+0.07}_{-0.05}$	$0.19^{+0.07}_{-0.05}$	$0.19^{+0.07}_{-0.05}$	$0.16^{+0.07}_{-0.05}$

**Charge jump rates (mHz) for individual qubits in each dataset. Statistical errors are provided as systematic errors are an order of magnitude smaller.**

	Q1-Q2 640 $\mu\text{m}$	Q3-Q4 340 $\mu\text{m}$	Q1-Q3 3195 $\mu\text{m}$	Q1-Q4 3330 $\mu\text{m}$	Q2-Q3 3180 $\mu\text{m}$	Q2-Q4 3240 $\mu\text{m}$
SO	$0.27^{+0.09}_{-0.07}$	$0.29^{+0.09}_{-0.07}$	$0.03^{+0.04}_{-0.02}$	$0.08^{+0.06}_{-0.04}$	$0.05^{+0.05}_{-0.03}$	$0.08^{+0.06}_{-0.04}$
SC	$0.10^{+0.07}_{-0.04}$	$0.04^{+0.05}_{-0.03}$	< 0.03	< 0.04	< 0.03	< 0.04

**Correlated charge jump rates (mHz) for individual qubits in each dataset. Separation distances listed below qubit numbers. Statistical errors are provided as systematic errors are an order of magnitude smaller.**

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- Shield closed: 13x reduction in rate
- Charge jump rate reduction: 4.2x

**Excess rate of charge jumps!**

Not explained by  $\gamma$  flux

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Average Rate	$0.51^{+0.05}_{-0.04}$	$0.19^{+0.04}_{-0.03}$	mHz
Corrected $\gamma$ Rate	$0.34^{+0.07}_{-0.06}$	$0.02^{+0.06}_{-0.05}$	mHz
Calculated Excess Rate	$0.17^{+0.04}_{-0.03}$		mHz

	Livetime (hr)	Q1	Q2	Q3	Q4
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**Charge jump rates (mHz) for individual qubits in each dataset. Statistical errors are provided as systematic errors are an order of magnitude smaller.**

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**Correlated charge jump rates (mHz) for individual qubits in each dataset. Separation distances listed below qubit numbers. Statistical errors are provided as systematic errors are an order of magnitude smaller.**

Potential explanations:

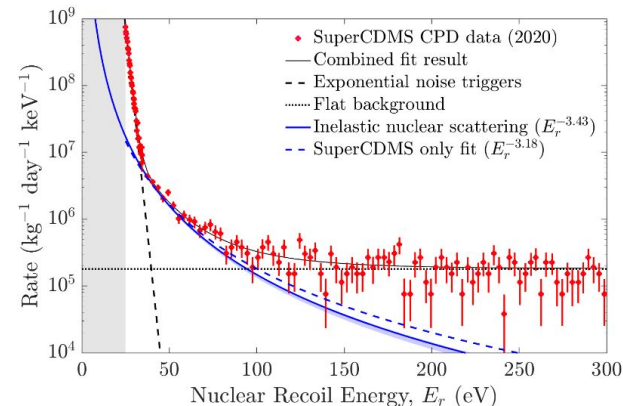
- Local source of ionizing radiation
- Changes in electric field, non-radiogenic origin
- Stress release knocking free trapped charges

# Segue: Low Energy [Phonon] Excess

Steeply rising low-energy background

Observed in cryogenic phonon detectors  
& superconducting qubits

1. **Non-ionizing:** no NTL amplification.
2. **Power law:** energy spectrum follows a power law.



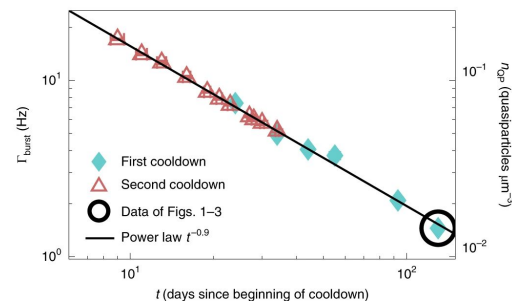
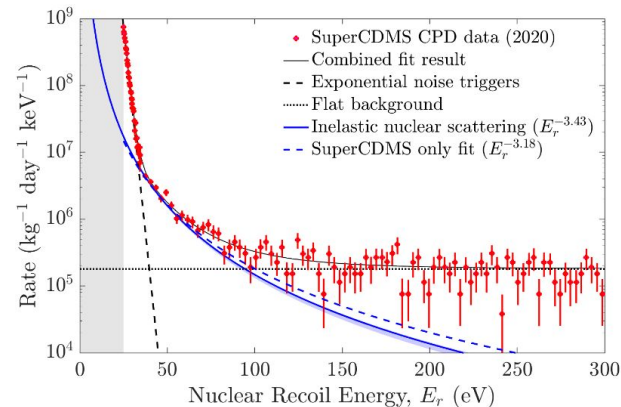


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1. **Non-ionizing:** no NTL amplification.
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3. **Time since cooldown:** background decays (long time constant) with time since reaching mK temperature.
  - a. Can increase rate by warming up and cooling back down



Mannila et al, Nat. Phys. 18, 145–148 (2022)

# Segue: Low Energy [Phonon] Excess

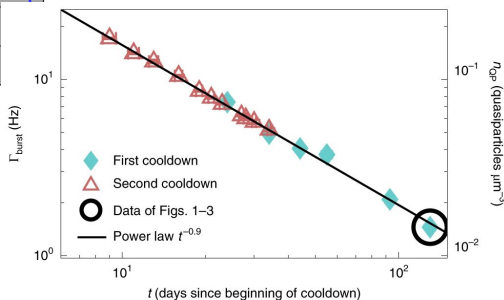
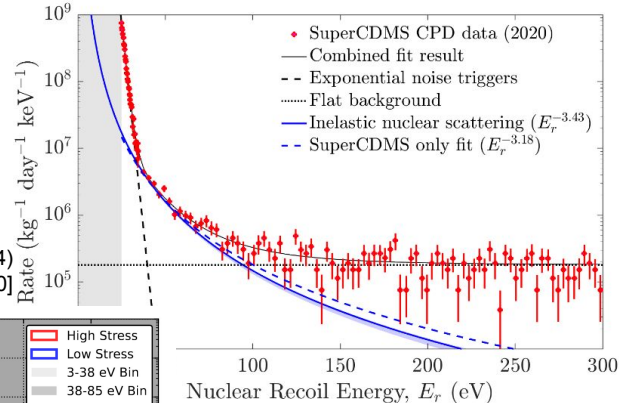
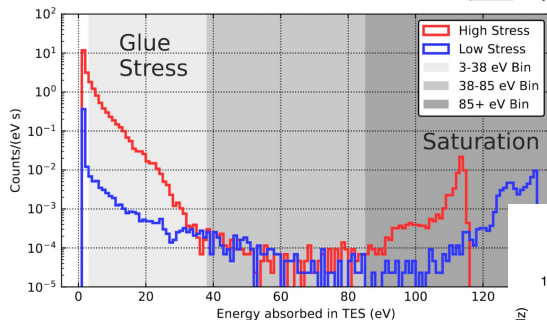
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1. **Non-ionizing:** no NTL amplification.
2. **Power law:** energy spectrum follows a power law.
3. **Time since cooldown:** background decays (long time constant) with time since reaching mK temperature.
  - a. Can increase rate by warming up and cooling back down
4. **Stress-dependent:** reducing mounting stress reduces background rate.
  - a. Stress buildup in thin film depositions can release phonons into substrate

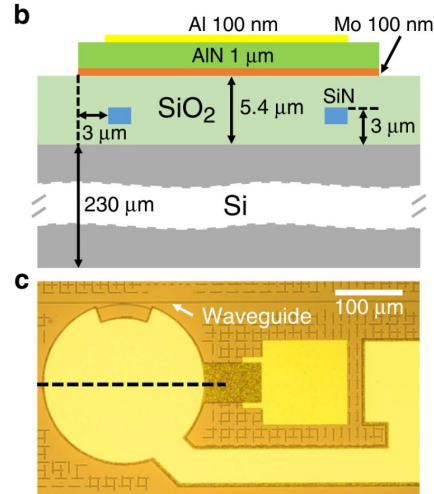
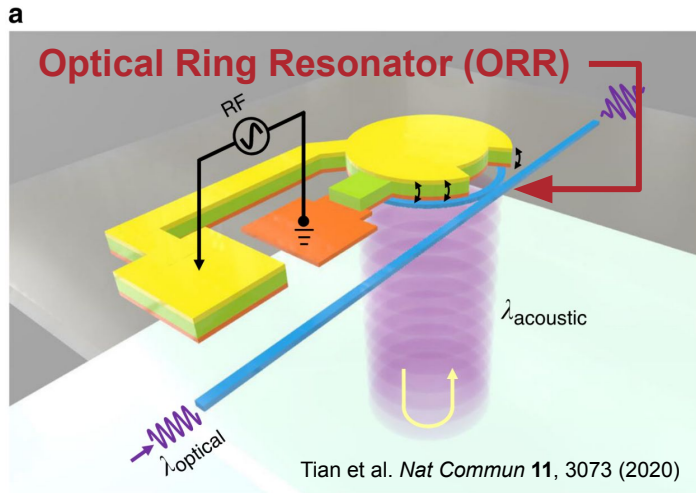
R. Romani (2024) [arXiv:2406.1542]

Anthony-Petersen et al, Nat. Commun. 15, 6444 (2024) [arXiv:2208.02790]



Mannila et al, Nat. Phys. 18, 145–148 (2022)

# Embedded SiN Optical Strain Sensors



**OxideMEMS Lab, Purdue**

*S. Bhawe (PI)*

*A. Attanasio*



**Stress-optical effect:** stress modulates the refractive index of the resonator.

→ Modulates transmission through waveguide for on-resonance  $\lambda$

Provides readout channel to directly probe crystal stress and substrate deformation.

Embedded sensors: surface free for deposition of primary sensors (qubit, MKID, TES, CCD).

# Embedded SiN Optical Strain Sensors

Photonics: microwave-optical transduction via piezo actuation

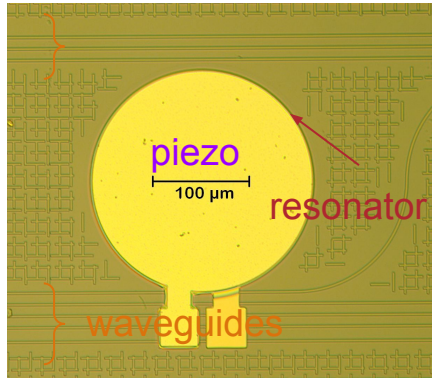
Tian et al. *Nat Commun* 11, 3073 (2020)

Sensing: micro-mechanical accelerometers with integrated test mass (Windchime)

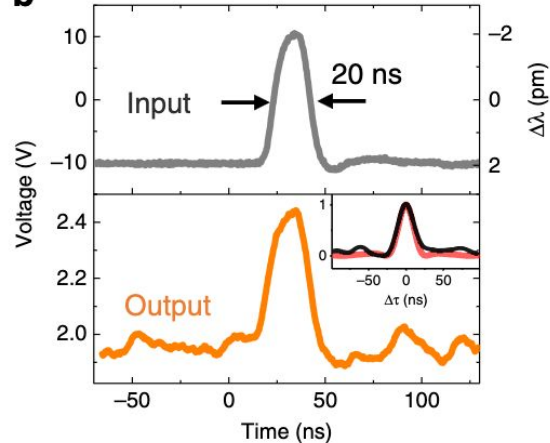
- Sub-ns optical response
- Optical Q  $\sim 10^6$ -- $10^8$
- Sensitivity:  $10^{-7}$  g/ $\sqrt{\text{Hz}}$  (accelerometer)

Piezo-actuated device

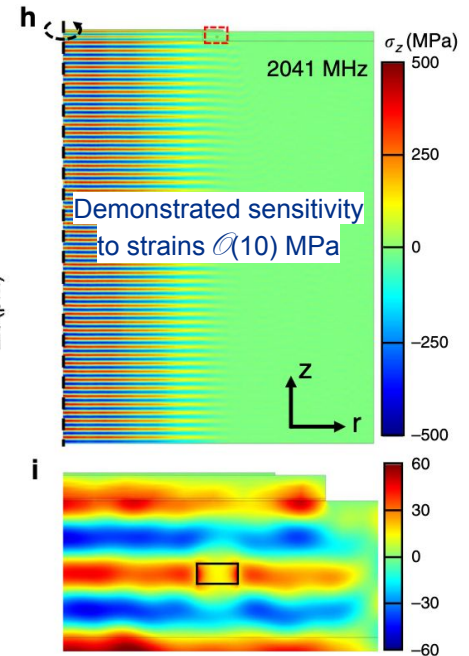
Accelerometer device

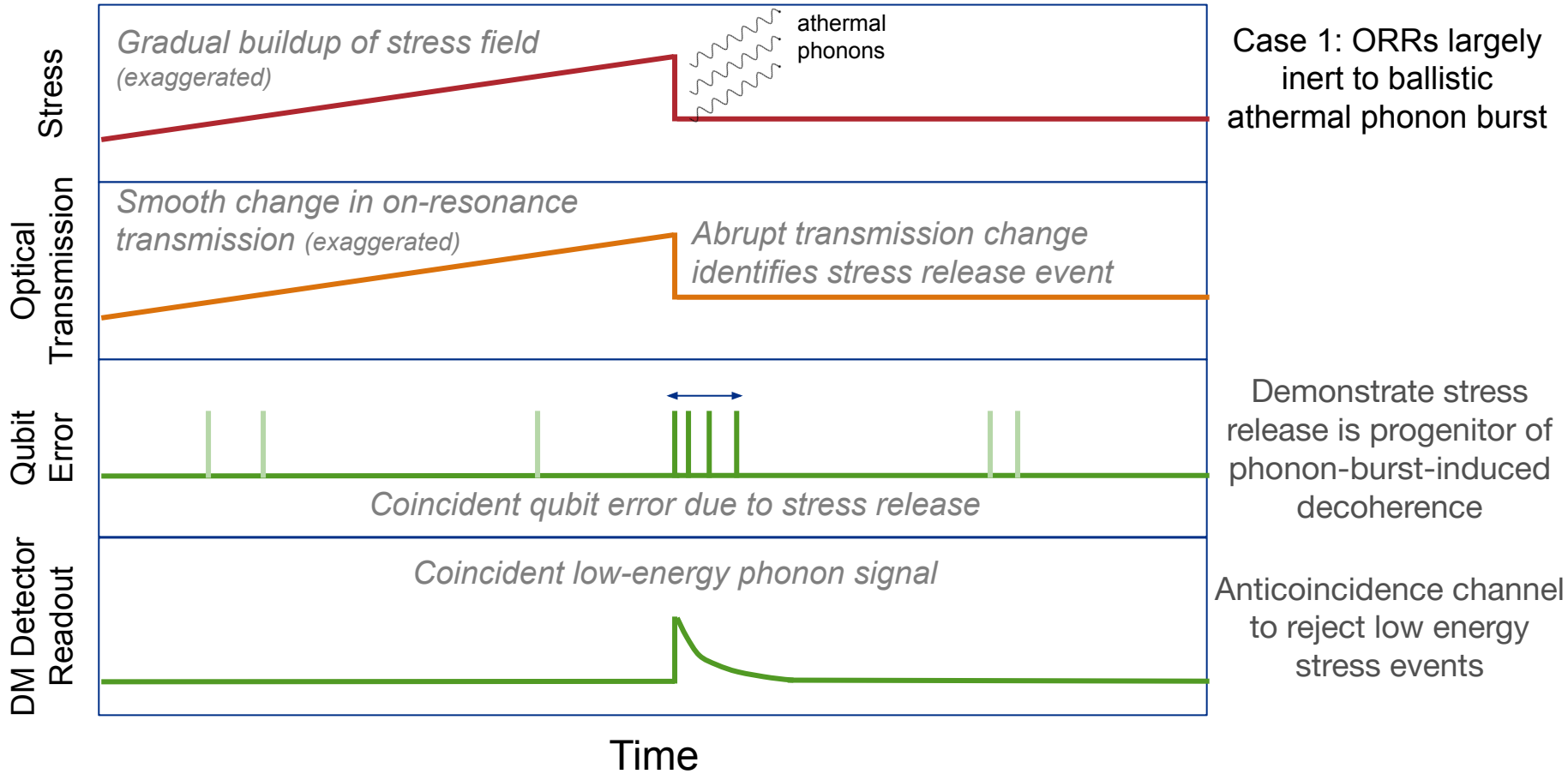


**b**



Device response to piezo impulse

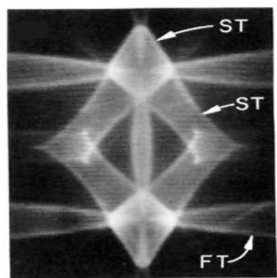
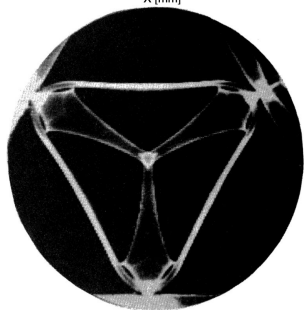
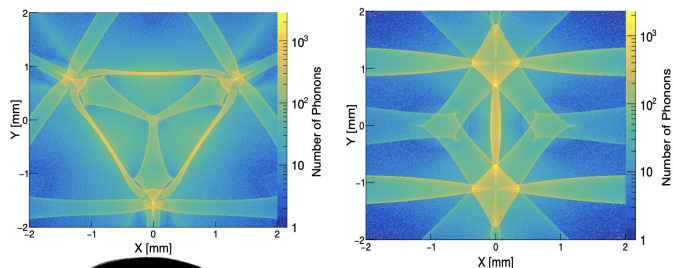




# ... and everything else!

## Modeling Athermal Phonons in Novel Materials using the G4CMP Simulation Toolkit

I. Hernandez<sup>a,b,\*</sup>, R. Linehan<sup>b,\*</sup>, R. Khatiwada<sup>a,b</sup>, K. Anyang<sup>a,b</sup>, D. Baxter<sup>b,d</sup>, G. Bratrud<sup>b,d</sup>, E. Figueroa-Feliciano<sup>d,b</sup>, L. Hsu<sup>b</sup>, M. Kelsey<sup>c</sup>, D. Temples<sup>b</sup>

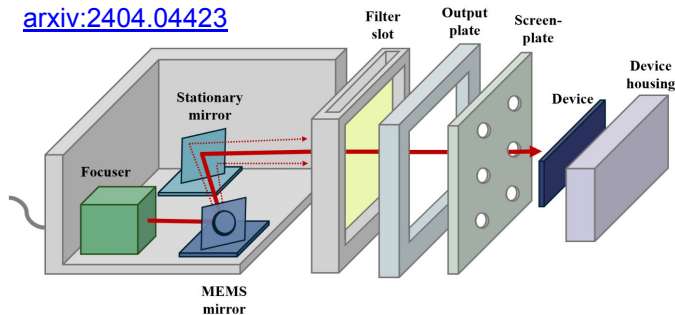


[arxiv:2408.04732](https://arxiv.org/abs/2408.04732)

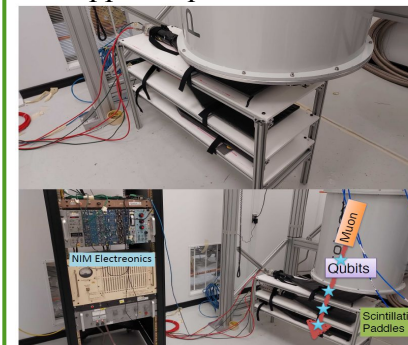
## Cryogenic optical beam steering for superconducting device calibration

K. Stifter<sup>1,2,3</sup>, H. Magoon<sup>1,2,3,4,5</sup>, A.J. Anderson<sup>1,6,7</sup>, D.J. Temples<sup>1</sup>, N.A. Kurinsky<sup>1,2,3</sup>, C. Stoughton<sup>1</sup>, I. Hernandez<sup>8,1</sup>, A. Nuñez<sup>2,3,4</sup>, K. Anyang<sup>8,1</sup>, R. Linehan<sup>1</sup>, M.R. Young<sup>1,6,7</sup>, P. Barry<sup>9</sup>, D. Baxter<sup>1,10</sup>, D. Bowring<sup>1</sup>, G. Cancelo<sup>1</sup>, A. Chou<sup>1</sup>, K.R. Diber<sup>6,7,1</sup>, E. Figueroa-Feliciano<sup>10,1</sup>, L. Hsu<sup>1</sup>, R. Khatiwada<sup>8,1</sup>, S.D. Mork<sup>1</sup>, L. Stefanazzi<sup>1</sup>, N. Tabassum<sup>2,3,4</sup>, S. Uemura<sup>1</sup>, B.A. Young<sup>11</sup>

[arxiv:2404.04423](https://arxiv.org/abs/2404.04423)

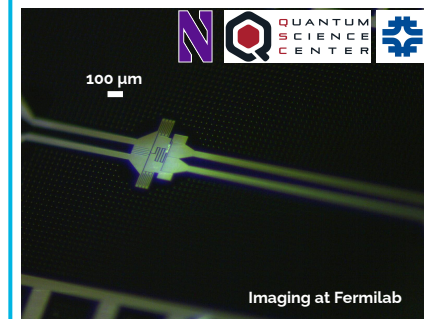


## Cosmic ray tagging with sapphire qubit at LOUD



K. Anyang, C. Veihmeyer, R. Linehan, R. Khatiwada

## Qubit design and fabrication



S. Sussman, R. Gaultieri

## QICK [arxiv:2311.17171](https://arxiv.org/abs/2311.17171)

Quantum Instrumentation Control Kit



- Fully integrated readout & control system for QIS, quantum networks, and superconducting detectors
- No extra room temperature hardware needed!

Experimental advances with the QICK (Quantum Instrumentation Control Kit) for superconducting quantum hardware

Chunyang Ding,<sup>1</sup> Martin Di Federico,<sup>2</sup> Michael Hatridge,<sup>3</sup> Andrew Houck,<sup>4</sup> Sebastien Leger,<sup>1</sup> Jeronimo Martinez,<sup>4</sup> Connie Miao,<sup>1</sup> David I Schuster,<sup>1</sup> Leandro Stefanazzi,<sup>2</sup> Chris Stoughton,<sup>2</sup> Sara Sussman,<sup>2</sup> Ken Treptow,<sup>2</sup> Sho Uemura,<sup>2</sup> Neal Wilcer,<sup>2</sup> Helin Zhang,<sup>2</sup> Chao Zhou,<sup>3</sup> and Gustavo Cancelo<sup>2</sup>

See demo in QUIET Lab Tour

QUANTUM SCIENCE CENTER

31

# Outlook: Next Year at QUIET

Magnetic shielded payload installed!

- 1000x reduction in field

## First device runs:

- Qubit chips galore!
- Superconducting Quasiparticle Amplifying Transmon (SQUAT)
- Kinetic Inductance Phonon-Mediated Detector
- Optical ring resonator strain sensor

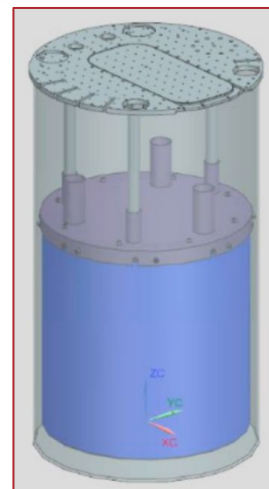
## Facility upgrades:

- Radiation shield (lead castle)
- Muon tagging system (for cosmic and NuMI beam)
- Full (remote) QICK control
- Expanding RF lines for ~100 qubit device

**Looking to run your device in a low-background facility?**

Talk to me or Dan Baxter

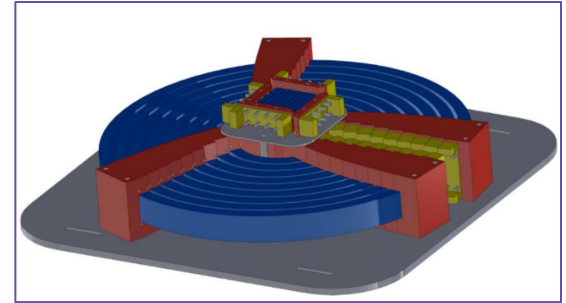
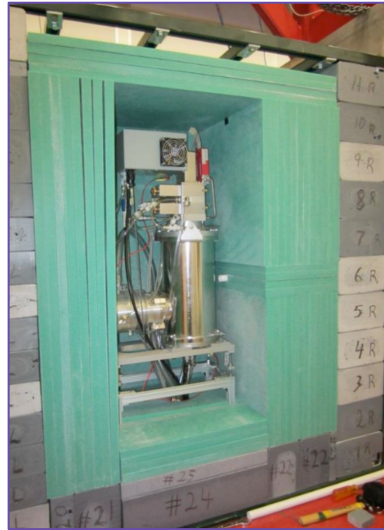
[dtemples@fnal.gov](mailto:dtemples@fnal.gov) | [dbaxter9@fnal.gov](mailto:dbaxter9@fnal.gov)



# Outlook: Next Year at NEXUS

## Continuing studies with charge-sensitive qubit

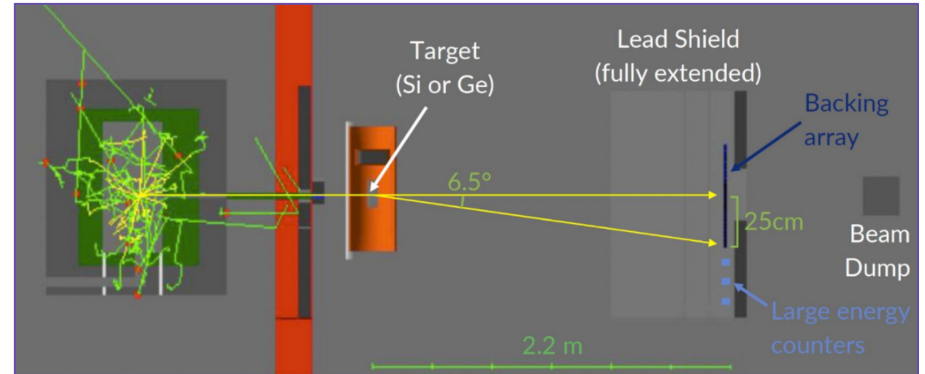
- QICK-216 w/ companion board -- lower noise
- Charge jump trigger for  $T_1$  measurements
  - Separately probe ionization + phonon channels
- Gamma source studies for charge jumps
- Investigate charge trapping



## DD neutron generator comes online!

- Adelphi DD-108:  $10^8$  [2.45 MeV] n/sec into  $4\pi$
- Borated poly shield w/ beam tube egress
- Backing array allows for fixed-angle scattering experiments

## DM searches with performant devices!







# CosmiQuantum @ FNAL

## FNAL Group

Daniel Baxter  
 Daniel Bowring  
 Gustavo Cancelo  
 Aaron Chou  
 Lauren Hsu  
 Sami Lewis\*  
 Ryan Linehan  
 Kelly Stifter\*  
 Sara Sussman  
 Dylan Temples  
 Sho Uemura  
 Matthew Hollister  
 Chris James  
 Hannah Magoon\*  
 Grace Wagner  
 Stella Dang

## NU Group

Enectalí Figueroa-Feliciano  
 Riccardo Gualtieri  
 Grace Bratrud  
 Arianna Colón Cesaní  
 Deeksha Sabhari  
 Shilin Ray

## IIT Group

Rakshya Khatiwada  
 Kester Anyang  
 Israel Hernandez  
 Jialin Yu

## External collaborators:

### Caltech

Sunil Golwala  
 Karthik Ramanathan  
 Osmond Wen  
 Taylor Aralis  
 Brandon Sandoval

### SLAC/Stanford

Noah Kurinsky  
 Kelly Stifter  
 Zoe Smith  
 Hannah Magoon  
 Elizabeth Panner

### QSC@Purdue

Alex Ma  
 Botao Du

### UW Madison

Robert McDermott  
 Sohair Abdullah

# Conclusions

- ❖ Superconducting phonon sensors leveraging quantum mechanics are a promising technology to probe the sub-GeV DM parameter space
  - KIPMDs: 2 eV sensor resolution provides path to eV-scale thresholds
  - Qubits: established a framework for energy reconstruction in relaxation readout mode
- ❖ Novel sensor architectures will require new calibration and background mitigation techniques
  - Qubits: excess rate of charge jumps inconsistent with  $\gamma$  flux
  - Stress release events may be identified with embedded optical strain sensors
- ❖ Device design optimization, low-background characterization, and simulation go hand-in-hand as we coalesce to a qubit-based detector architecture
- ❖ Lots of exciting work going on at Fermilab! Stay tuned.



Looking to run your device in a low-background facility?

Talk to me or Dan Baxter

[dtemples@fnal.gov](mailto:dtemples@fnal.gov) | [dbaxter9@fnal.gov](mailto:dbaxter9@fnal.gov)





**Thank You!**



Northwestern  
**NEXUS**  
Experimental  
Underground Site  
@Fermilab



U.S. DEPARTMENT OF  
**ENERGY**

Office of Science

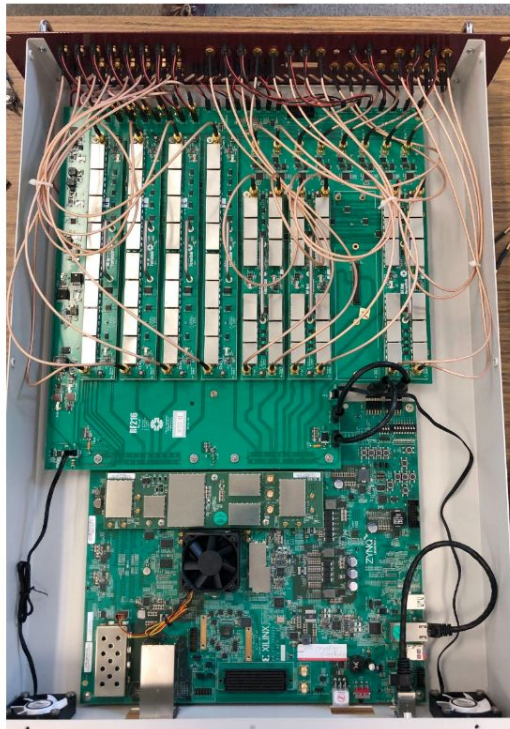
# Abstract

## **Cryogenic Platforms at Fermilab: Recent Results and Outlook**

The Cosmic Quantum group at Fermilab operates three cryogenic facilities dedicated to the development and calibration of superconducting low-threshold detectors and qubits. One of which (LOUD) is located at the surface, while the other two (NEXUS and QUIET) are located 100 m underground enabling low-background device characterization and rare event searches. Recently, we have demonstrated world-leading resolution in the quasiparticle channel for kinetic inductance phonon-mediated detectors as well as the lowest rate of spatially- and temporally- correlated errors (“charge jumps”) in superconducting qubit chips. We have additionally made significant progress in expanding the suite of tools for simulating signal production and readout of these devices. In this talk, I will review these recent results, discuss some nascent projects focused on enhancing sensitivity of these detectors to sub-GeV dark matter, and highlight activities in these facilities over the next year.

# QICK

## Quantum Instrumentation Control Kit



- **Fully integrated readout & control system for QIS, quantum networks, and superconducting detectors**
  - No extra room temperature hardware needed!
  - Has already been adopted by QIS groups around the world (including many of you)
- A factor of  $\sim 20$  cheaper compared to off-the-shelf equipment
- QICK team ongoing work includes paper on quantum measurement & readout fidelity, and major firmware upgrade

*See demo in QUIET Lab Tour*

## Dark Matter Detectors (HEP)

## Superconducting Qubits (QIS)

