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## **DIRECTLY DETECTING THE IRREDUCIBLE** MILLICHARGED BACKGROUND





## WHERE STHENEW PHYSICS?



# WHAT'S HAPPENING ON THE DARK SIDE?





#### WHAT'S THE DEAL WITH WHAT'S HAPPENING THE STANDARD MODEL? **ON THE DARK SIDE?**





Neutrino masses? Strong CP problem? Hierarchy problem?

Experimental anomalies?

NECESSITATE UV COMPLETION!

## THERE ARE PROBABLY NEW PARTICLES BEYOND THE STANDARD MODEL AND THESE NEW PARICLES MAY BE DARK MATTER





#### DETERMINE PARTICLE TYPE



#### DETERMINE PARTICLE TYPE



#### DETERMINE INTERACTION TYPE



#### DETERMINE PARTICLE TYPE



#### DETERMINE INTERACTION TYPE



ONLY A LIMITED NUMBER OF RENORMALIZABLE OPERATORS!

#### DETERMINE PARTICLE TYPE



#### DETERMINE INTERACTION TYPE



ONLY A LIMITED NUMBER OF RENORMALIZABLE OPERATORS!

- 1. Vector Portal
- 2. Axion Portal
- 3. Neutrino Portal
- 4. Scalar Portal

#### DETERMINE PARTICLE TYPE



#### DETERMINE INTERACTION TYPE



ONLY A LIMITED NUMBER OF RENORMALIZABLE OPERATORS!

1. Vector Portal

 $\frac{\kappa}{2} F_{\mu\nu} F^{\mu\nu} + q_{\chi} A'_{\mu} \bar{\chi} \delta^{\mu} \chi$ 

#### DETERMINE PARTICLE TYPE



#### DETERMINE INTERACTION TYPE



ONLY A LIMITED NUMBER OF RENORMALIZABLE OPERATORS!

1. Vector Portal

 $\frac{\kappa}{2} F_{\mu\nu} F^{\mu\nu} + q_{\chi} A'_{\mu} \bar{\chi} g^{\mu} \chi$ 

Dark U(1) that mixes with the SM U(1)

#### DETERMINE PARTICLE TYPE



#### **DETERMINE INTERACTION TYPE**



#### ONLY A LIMITED NUMBER OF **RENORMALIZABLE OPERATORS!**

1. Vector Portal

 $\frac{\kappa}{2} F_{\mu\nu} F^{\mu\nu} + q_{\chi} A'_{\mu} \bar{\chi} g^{\mu} \chi$ 

Dark U(1)that mixes with the SM U(1)

Additional particles charged under this U(1)



### VECTOR PORTAL



#### **BSM SECTOR**



### VECTOR PORTAL



#### **BSM SECTOR**

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#### **BSM SECTOR**

## MILICHARGED\* PARTICLE (MCP)

**IF THE DARK PHOTON IS** LIGHT, FIELD REDEFINITIONS **RESULT IN AN** EFFECTIVE MILLICHARGE



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# Effective Millicharge

\*nothing milli about millicharged





## 1. HOW DO MCPS IMPACT OBSERVABLES?

## 2. REGARDLESS OF WHETHER MCPS ARE DARK MATTER, CAN WE USE THE DARK MATTER EXPERIMENTAL PROGRAM TO LOOK FOR THEM?



## 1 HOW DO MCPS IMPACT OBSERVABLES?

## 2. REGARDLESS OF WHETHER MCPS ARE DARK MATTER, CAN WE USE THE DARK MATTER EXPERIMENTAL PROGRAM TO LOOK FOR THEM?



## MCPS IN ASTROPHYSICAL SYSTEMS LIGHT MCPS CAN BE PRODUCED IN STARS THROUGH <u>PLASMON</u> DECAYS









B B DISPERSES LIKE MASSIVE PARTICLE A  $TT_{T} = \omega_{p}^{2} = \omega_{z}^{2} - k^{2}$  $TT_{L} = \omega_{p}^{2} = \omega_{z}^{2}$ 







#### DISPERSES LIKE A MASSIVE PARTICLE







#### DISPERSES LIKE A MASSIVE PARTICLE

 $m_{\gamma,\mathrm{eff}}^2 \sim$ 













#### PLASMONS CAN DECAY



HYDROGEN BURNING ENVELOP DEGENERATE HELIUM CORE

#### **RED GIANT BRANCH**



#### **RED GIANT BRANCH**



#### **RED GIANT BRANCH**

#### **STAR BECOMES BRIGHTER AND REDDER**



#### **RED GIANT BRANCH**

#### **STAR BECOMES HELIUM FLASH BRIGHTER AND REDDER** (TIP OF THE RED GIANT BRANCH)

## MCPS IN ASTROPHYSICAL SYSTEMS ENERGY LOSS THROUGH MCPS COOLS THE STAR AND DELAYS HELIUM FLASH



#### **RED GIANT BRANCH**

#### **HELIUM FLASH** (TIP OF THE RED GIANT BRANCH **BRIGHTER AND REDDER IS BRIGHTER!**)



**PLASMON DECAYS ARE AN EFFICIENT MECHANISM OF ENERGY LOSS IN STARS!** 



Fung, SH, Schutz et al (2023)

PLASMON DECAYS **ARE AN EFFICIENT MECHANISM OF ENERGY LOSS IN STARS!** 

#### **INCREASE THE BRIGHTNESS OF** THE TRGB!



Fung, SH, Schutz et al (2023)

PLASMON DECAYS ARE AN EFFICIENT MECHANISM OF ENERGY LOSS IN STARS!

#### INCREASE THE BRIGHTNESS OF THE TRGB!

$$10^{-8}$$

$$10^{-9}$$

$$10^{-10}$$

$$10^{-11}$$

$$10^{-12}$$

$$10^{-13}$$

$$10^{-14}$$

$$10^{-15}$$



#### Fung, SH, Schutz et al (2023)

## MUCH LIKE STARS TODAY, THE EARLY UNIVERSE WAS ALSO A HOT MESS....



## MCPS WILL BE PRODUCED IN THE EARLY UNIVERSE THROUGH FREEZE-IN





### FREEZE-IN ABUNDANCE THROUGH PLASMON DECAYS



#### ABUNDANCETHROUGH PLASMON





#### ABUNDANCETHROUGH PLASMON ▋**╛**∶**】╡***╪***╱┫<del>╡</del><mark>╝</mark>║╲║**





#### MCP Mass = 40 keV25 $10^{0}$ $m_{\gamma,\rm eff}^2 \propto \alpha T^2$ $\sim$ Abundance $10^{-1}$ kinematically blocked Plasmon decays MCP $10^{-2}$ $Y_{\chi}^{\gamma^*} = 2 \times 10^{-10} \left(\frac{Q}{10^{-12}}\right)^2$ 1 **MeV** if $m_{\chi} \gtrsim 10 \text{ keV}$ $m_{\chi}$

Decays stop when plasma frequency is smaller than the MCP mass



## FREEZE-IN ABUNDANCE THROUGH PLASMON DECAYS



 $Y_{\chi}^{\gamma^*} = 3 \times 10^{-8} \left(\frac{Q}{10^{-12}}\right)^2 \quad \text{if } m_{\chi} \lesssim 10 \text{ keV}$ 

## FREZEN ABUNDANCE THROUGH PLASMON DECAYS



 $Y_{\chi}^{\gamma^*} = 3 \times 10^{-8} \left(\frac{Q}{10^{-12}}\right)^2 \quad \text{if } m_{\chi} \lesssim 10 \text{ keV}$ 

 $m_{\gamma,\text{eff}}^2 \sim \frac{\alpha n_e}{m_e}$ 





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#### $Y_{\chi}^{\gamma^*} = 3 \times 10^{-8} \left(\frac{Q}{10^{-12}}\right)$ if $m_{\chi} \lesssim 10 \text{ keV}$

If MCPs are very light, electron freezeout quenches the production!







### FREEZE-IN ABUNDANCE THROUGH ANNIHILATIONS



 $1 \, \text{MeV}$  $\mathcal{Q}$  $Y_{\chi}^{e^+e^-} = 1 \times 10^{-9}$  $\left( \frac{10^{-12}}{10^{-12}} \right) \left( \max(m_{\chi}, m_e) \right)$ 



#### IF MILLICHARGED PARTICLES EXIST, THEY WILL ALWAYS HAVE BLEABUNDANCE BECAUSE OF FREEZEN







#### IF MILLICHARGED PARTICLES EXIST, THEY WILL ALWAYS HAVE CE BECAUSE OF FREFZEIN







#### IF MILLICHARGED PARTICLES **EXIST**, THEY WILL ALWAYS HAVE CE BECAUSE OF FREFZEIN D





#### IF MILLICHARGED PARTICLES **EXIST**, THEY WILL <u>ALWAYS</u> HAVE **BLE ABUNDANCE BECAUSE OF FREEZEIN**





#### IF MILLICHARGED PARTICLES EXIST, THEY WILL ALWAYS HAVE AN <u>IRREDUCIBLE ABUNDANCE</u> BECAUSE OF FREEZE-IN





#### IF MILLICHARGED PARTICLES **EXIST**, THEY WILL ALWAYS HAVE LE ABUNDANCE BECAUSE OF FREEZEIN

Part of the parameter space already excluded by astrophysics!





### 1. HOW DO MCPS IMPACT OBSERVABLES?

## 2. REGARDLESS OF WHETHER MCPS ARE DARK MATTER, CAN WE USE THE DARK MATTER EXPERIMENTAL PROGRAM TO LOOK FOR THEM?



### DIRECT DETECTION EXPERIMENTS ARE SENSITIVE TO THE ANBIENT DARK MATTER DENSITY



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# $\frac{R}{100} \propto 10^{-10}$

### DIRECT DETECTION EXPERIMENTS ARE SENSITIVE TO THE AMBIENT DARK MATTER DENSITY



#### CTION Event Rate

Ro

PDM MDM DM Number density

### DIRECT DETECTION EXPERIMENTS ARE SENSITIVE TO THE AND ENT DARK MATTER DENSITY



#### Event Rate

DM Number density

 $ho_{\rm DM}$ 

 $m_{\rm DM}$ 

Rœ

Interaction crosssection



### DRECT DETECTION EXPERIMENTS ARE SENSITIVE TO THE AMBIENT DARK MATTER DENSITY



For the MCP background, limits on  $\overline{\sigma}_{\rho}$  can be interpreted as limits on  $\bar{\sigma}_e \times \rho_{\rm MCP} / \rho_{\rm DM}!$ 

#### Event Rate

#### DM Number density

 $\rho_{\rm DM}$ 

 $m_{\rm DM}$ 

 $R \propto$ 

#### Interaction crosssection



#### DIRECT DETECTION EXPERIMENTS WILL BE SENSITIVE TO THE MCP PARAMETER SPACE





### DIRECT DETECTION EXPERIMENTS WILL BE SENSITIVE TO THE MCP PARAMETER SPACE



Nuclear recoil bounds recast for the case of light mediators (Hamby et al. PRD 2018)





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#### MCP BACKGROUND

Berlin et al, PRL (2020)

## DRECT DEFECTOR

- 0

60

SP. US

### MCP BACKGROUND

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No.

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**N** 

X

Berlin et al, PRL (2020)



#### DEFLECTOR

## 

0

60

 $\boldsymbol{\alpha}$ 

### MCP BACKGROUND

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Berlin et al, PRL (2020)



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### MCP BACKGROUND

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STA CO

Berlin et al, PRL (2020)





#### THE ALTERNATING ELECTRIC FIELD IN THE DEFLECTOR SETS UP AN Alternating MCP current in the detector!

## 16.4 SP CO 00 00 ••• SP. NP CC Berlin et al, PRL (2020)



#### DIRECT DEFLECTION EXPERIMENTS WILL BE SENSITIVE TO THE MCP PARAMETER SPACE

Idealized deflector setup, (Berlin et al, PRL 2020)





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## 1. MCPS ARE MINIMAL EXTENSIONS OF THE STANDARD MODEL 2. IF MCPS EXIST, THEY WILL BE IRREDUCIBLY **PRODUCED IN THE EARLY UNIVERSE THROUGH** FREEZE-N





## **3. DIRECT DETECTION EXPERIMENTS ARE SENSITIVE TO** THE IRREDUCIBLE MILLICHARGED BACKGROUND





#### ILES, SH & SCHUTZ: 2407.21096

