## Dark Sectors with LHCb and Beyond

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#### Hidden Sectors



broken U(1) gauge symmetry in dark sector
allow mixing between dark and SM hypercharge fields

$$\mathcal{L} \supset -\frac{1}{4}F_{\mu
u}F^{\mu
u} - \frac{1}{4}F'_{\mu
u}F'^{\mu
u} + \frac{m_{A'}^2}{2}A'_{\mu}A'^{\mu} + g_eJ^{\mu}A_{\mu} + gg_eJ^{\mu}A'_{\mu}$$

#### Dark Photons



mass of the dark photon, m<sub>A'</sub>, and mixing, g, are free parameters
the dark photon couples like the photon, modified by g
if m<sub>A'</sub> < 2m<sub>DM</sub> then dark photon decays visibly

- what happens if **2** and **3** are relaxed?
- require  $m_{A'}$ , g, 12 fermion couplings, and an invisible width
- dark photon limits can be recast to any general vector model

## Parameter Space



## Lifetime



## Decay Products



## Search Strategies





- EM background free
- difficult to normalize



- sensitive to shorter lifetimes
- bump hunt on large EM background
- normalized from sidebands
- do both simultaneously for best of both worlds

## Production: Electron Bremsstrahlung



## Production: Proton Bremsstrahlung



### Production: Hadron Decays

![](_page_9_Figure_2.jpeg)

## Production: Electron-Positron Annihilation

![](_page_10_Figure_2.jpeg)

## Searching with LHCb *in theory* ...

Ilten, Soreq, Thaler, Williams, Xue Phys. Rev. Lett. **116**, no. 25, 251803 (2016)

Ilten, Thaler, Williams, Xue Phys. Rev. D **92**, no. 11, 115017 (2015)

## LHCb Detector

![](_page_12_Figure_2.jpeg)

## Good Backgrounds (prompt)

![](_page_13_Figure_2.jpeg)

## Signal (prompt and displaced)

![](_page_14_Figure_2.jpeg)

## Bad Backgrounds (prompt)

![](_page_15_Figure_2.jpeg)

 $N_{\rm signal}$  is not proportional to  $N_{\rm bad}$ LHCb mis-ID probability  $\approx 1$  out of 1000

## Production in Theory

![](_page_16_Figure_2.jpeg)

## Reach (prompt) in Theory

![](_page_17_Figure_2.jpeg)

## Bad Backgrounds (*displaced*)

![](_page_18_Figure_2.jpeg)

## Reach (displaced) in Theory

![](_page_19_Figure_2.jpeg)

#### Full Reach in Theory

![](_page_20_Figure_2.jpeg)

## Searching with LHCb in practice ...

LHCb Collaboration Phys. Rev. Lett. **120**, no. 6, 061801 (2018)

> LHCb Collaboration JINST **13**, no. 06, P06008 (2018)

LHCb Collaboration Phys. Rev. Lett. **124**, no. 4, 041801 (2020)

LHCb Collaboration JHEP **10**, 156 (2020)

## Real Data (prompt)

![](_page_22_Figure_2.jpeg)

Dark Sectors with LHCb and Beyond

## Limits (prompt)

![](_page_23_Figure_2.jpeg)

## Real Data (displaced)

![](_page_24_Figure_2.jpeg)

## **VELO** Sensors

![](_page_25_Picture_2.jpeg)

#### RF Foil

![](_page_26_Picture_2.jpeg)

## Bad Backgrounds (*displaced*)

![](_page_27_Figure_2.jpeg)

## Limits (displaced)

![](_page_28_Figure_2.jpeg)

# Dark Photons and beyond ...

Ilten, Soreq, Williams, Xue JHEP **1806**, 004 (2018)

Cid Vidal, Ilten, Plews, Shuve, Soreq Phys. Rev. D **100**, no. 5, 053003 (2019)

> Baruch, Ilten, Soreq, Williams JHEP 11, 124 (2022)

#### DARKCAST

• recast to any general model, e.g. 15 free parameters

![](_page_30_Picture_3.jpeg)

- available at gitlab.com/philten/darkcast
  - accompanying papers Serendipity in dark photon searches and Axial vectors in Darkcast

## The Recipe

- given  $(m, g_A)$  for model A, solve to find  $(m, g_B)$  for model B $\sigma_A(m, g_A)\mathcal{B}_A(m)\varepsilon(\tau_A(m, g_A)) = \sigma_B(m, g_B)\mathcal{B}_B(m)\varepsilon(\tau_B(m, g_B))$
- absolute cross-section can be tricky, ratios are easier

$$\frac{\sigma_A(m,g_A)}{\sigma_B(m,g_B)} \frac{\varepsilon(\tau_A(m,g_A))}{\varepsilon(\tau_B(m,g_B))} \frac{\mathcal{B}_A(m)}{\mathcal{B}_B(m)} = 1$$

branching fraction ratio: hidden local symmetries
cross-section ratio: hidden local symmetries

 $V \in (\rho, \omega, \phi, K^*, \bar{K}^*)$  generated from  $U(3)_V$ 

**3** efficiency ratio: define proper time fiducial region with  $t_0$  and  $t_1$ 

$$\varepsilon(\tau) = e^{-t_0/\tau} - e^{-t_1/\tau}$$

## Widths

• width can be calculated perturbatively for fermions

$$\Gamma_{ff}(\boldsymbol{m}, \boldsymbol{g}) = \frac{\boldsymbol{g}^2 c_f \left(\boldsymbol{x}_V^f\right)^2}{12\pi} \boldsymbol{m} \left(1 + \frac{m_f^2}{\boldsymbol{m}}\right) \sqrt{1 - 4\frac{m_f^2}{\boldsymbol{m}}}$$

- but ... below 2 GeV this prediction is no longer reliable
- use data instead!

$$\Gamma_{\text{hadrons}}(\boldsymbol{m}, \boldsymbol{g}) = \Gamma_{\mu\mu}(\boldsymbol{m}, \boldsymbol{g}) \mathcal{R}(\boldsymbol{m}) + \frac{g^2 \boldsymbol{m}}{4\pi} \bigg[ (x_A^u - x_A^d)^2 a_1(\boldsymbol{m}^2) + (x_A^s)^2 \Theta(\boldsymbol{m}^2 - 4m_K^2) \times \left( \frac{1}{4} a_1(\boldsymbol{m}^2) + a_1^s(\boldsymbol{m}^2) - \cos(\phi) \sqrt{a_1(\boldsymbol{m}^2)a_1^s(\boldsymbol{m}^2)} \right) \bigg]$$

- $\mathcal{R}(\underline{m})$  is  $\sigma(ee \to \text{hadrons}) / \sigma(ee \to \mu\mu)$
- +  $a_1^{(s)}(m)$  are  $\tau$  axial spectral functions from ALEPH

#### Beyond Dark Photons

## The Data!

![](_page_33_Figure_2.jpeg)

#### Beyond Dark Photons

## The Data!

![](_page_34_Figure_2.jpeg)

#### BBoson

![](_page_35_Figure_2.jpeg)
### A Special Case

- *true* muonium is a  $\mu^+\mu^-$  state, not yet observed!
- different spin configurations, most abundant are  ${}^1S_0$  and  ${}^3S_1$
- ${}^1S_0 \to \gamma\gamma$  and  ${}^3S_1 \to e^+e^-$



$$E_B \approx m_\mu \alpha^2 / 4 = 1.41 \,\mathrm{keV}$$

$$m_{\mathcal{TM}} \approx 2m_{\mu} - E_B \approx 211 \,\mathrm{MeV}$$

$$g_{TM} \approx \alpha^2/2 \approx 2.66 \times 10^{-5}$$

$$\tau_{\mathcal{TM}} \approx \frac{6}{\alpha^5 m_{\mu}} \approx 1800 \text{ fs}$$

### Zooming In



# LHCb and beyond ...

CODEX-b Collaboration Eur. Phys. J. C **80**, no. 12, 1177 (2020)

CODEX-b Collaboration arXiv:2406.12880 [physics.ins-det]

### Mind the Coverage Gap



• there is a large gap in search parameter space

### Mind the Coverage Gap



• the LHC accesses unique high COM

### A Modest Proposal



- need low background large volume detector
- use space around LHCb  $\rightarrow$  CODEX-b proposal

### Dark Photon Again



• example coverage from non-minimal dark photon production via the SM Higgs

### Building It!



- CODEX- $\beta$  approved as time limited LHCb R&D project
- independent of LHCb, including members from CMS and ATLAS

### **CODEX-b** Collaboration



- CODEX-b collaboration is always looking for collaborators!
- MATHUSLA and ANUBIS are also excellent off-axis opportunities

### Conclusions



- LHCb search with  $\approx 4\%$  of Run 2 + Run 3 data is promising
- DARKCAST at gitlab.com/darkcast/releases
- true muonium is within reach at LHCb
- off-axis (transverse) detectors necessary for LLPs at high  $\sqrt{s}$

## Appendix

### New Physics in TM



### Data Taking



### Hidden Symmetries

- but what about flavor dependent couplings?
- use hidden local symmetries framework for VMD
- vector mesons  $V\in(\rho,\omega,\phi,K^*,\bar{K}^*)$  are gauge bosons of hidden  $U(3)_V$  symmetry
- vertices take the form  $PV_iV_j$  with P from the pseudoscalar nonet  $P\in(\pi,\eta,\eta',K,\bar{K})$

$$\operatorname{Tr}(T_{V_i}, T_{V_j}, T_P)$$

- T are the meson generators, e.g.  $T_{\omega} = \frac{1}{2}(1,1,0)$
- external gauge fields mix through V

 $\operatorname{Tr}(T_V, Q)$ 

• Q is the fermion coupling vector  $(Q_u, Q_d, Q_s)$ 

### Vector Decomposition



### B-L Boson



### B Boson



### Protophobic Boson



### Production Ratios

• electron-positron annihilation and electron bremsstrahlung

$$\frac{\sigma_A(m, g_A)}{\sigma_B(m, g_B)} = \frac{g_A{}^2 Q_A^{e\,2}}{g_B{}^2 Q_B^{e\,2}}$$

• proton bremsstrahlung

$$\frac{\sigma_A(m,g_A)}{\sigma_B(m,g_B)} = \frac{g_A{}^2(2Q_A^u + Q_A^d)^2}{g_B{}^2(2Q_A^u + Q_A^d)^2}$$

• hadron decays of the form  $X \to YA$ 

$$\frac{\sigma_A(\boldsymbol{m}, \boldsymbol{g}_A)}{\sigma_B(\boldsymbol{m}, \boldsymbol{g}_B)} = \frac{g_A^2 \sum_V \operatorname{Tr}(T_X, T_Y, T_V) \operatorname{Tr}(T_V, Q_A) \operatorname{BW}_V(\boldsymbol{m})}{g_B^2 \sum_V \operatorname{Tr}(T_X, T_Y, T_V) \operatorname{Tr}(T_V, Q_B) \operatorname{BW}_V(\boldsymbol{m})}$$

### LHCb Production Fractions

- templates taken from Monte Carlo and fit against LHCb result



### Efficiencies

- define proper time fiducial region with  $t_0$  and  $t_1$ 

$$\varepsilon(\tau) = e^{-t_0/\tau} - e^{-t_1/\tau}$$

- for prompt limits,  $t_0 = 0$  and  $t_1$  depends on the boost

$$t_1 = \frac{L_{\max}}{\gamma}$$

- for displaced beam-dump limits, relate  $t_0$  and  $t_1$ 

$$t_1 = t_0 + \frac{L_{\text{detector}}}{L_{\text{shield}}}$$

 $\rightarrow$  upper and lower limits are solutions, equate and solve for  $t_0$ :

$$\sigma(m, g_{\max})\mathcal{B}(m)\varepsilon\left(\tau(m, g_{\max})\right) = \sigma(m, g_{\min})\mathcal{B}(m)\varepsilon\left(\tau(m, g_{\min})\right)$$

### B-L Boson



### B Boson



### Protophobic Boson



### Mind the Gap



### What's in an Event?

• need final states that produce di-electrons

$\operatorname{meson}$	meson/event	$e^+e^-$ /event
$\pi^+$	$1.27  imes 10^1$	_
$\pi^0$	$7.08  imes 10^0$	$8.50  imes 10^{-2}$
$ ho^+$	$1.96  imes 10^0$	$2.36 imes10^{-2}$
$K^+$	$1.44 \times 10^0$	_
$ ho^0$	$1.02 \times 10^0$	$2.36 \times 10^{-5}$
$\omega$	$9.87 \times 10^{-1}$	$1.24\times10^{-2}$
n	$9.71 \times 10^{-1}$	—
p	$9.51  imes 10^{-1}$	—
$\eta$	$8.31  imes 10^{-1}$	$1.80  imes 10^{-2}$
$K_S^0$	$7.08  imes 10^{-1}$	$5.25  imes 10^{-3}$
$K_L^0$	$7.07 \times 10^{-1}$	_

### $e^+e^-\gamma$ Production



### Dissociation



### Detector Effects: Case (i)



### Detector Effects: Case (ii)



### **Discovery** Potential



### A Large Cube



- +  $10 \times 10 \times 10$  m baseline configuration
- use resistive plate chambers (1 mm spatial and 100 ps timing)
- sensitive to LLP masses via opening angle of LLP decay

### Raising the Shields



• active shielding is necessary for background reduction

### Raising the Shields



• target less than one event over entire run of CODEX-b

### Models Overview



• CODEX-b has sensitivity to a wide variety of general and specific models

### Dark Photon



• production via the SM Higgs makes this currently only possible at the LHC
## Higgs Portal

 $\lambda = 1.6 \times 10^{-3}$ 



• searches via B decays, limits from LHCb are lifetime constrained

#### Conclusions

# Heavy Neutral Leptons



• large portions of HNL space remain unexplored

### Conclusions

### Fermion Coupled Axion-Like Particles



• ALPs can both provide a dark matter candidate and "clean up" the strong CP problem