

ARXIV: [2401.06843](https://arxiv.org/abs/2401.06843) , ARXIV: [2410.XXXXX](https://arxiv.org/abs/2410.XXXXX),

# Enhanced Dark Sector Production In Beam Dumps From Electromagnetic Cascades

RYAN PLESTID

NTN FELLOW, BURKE INSTITUTE, CALTECH

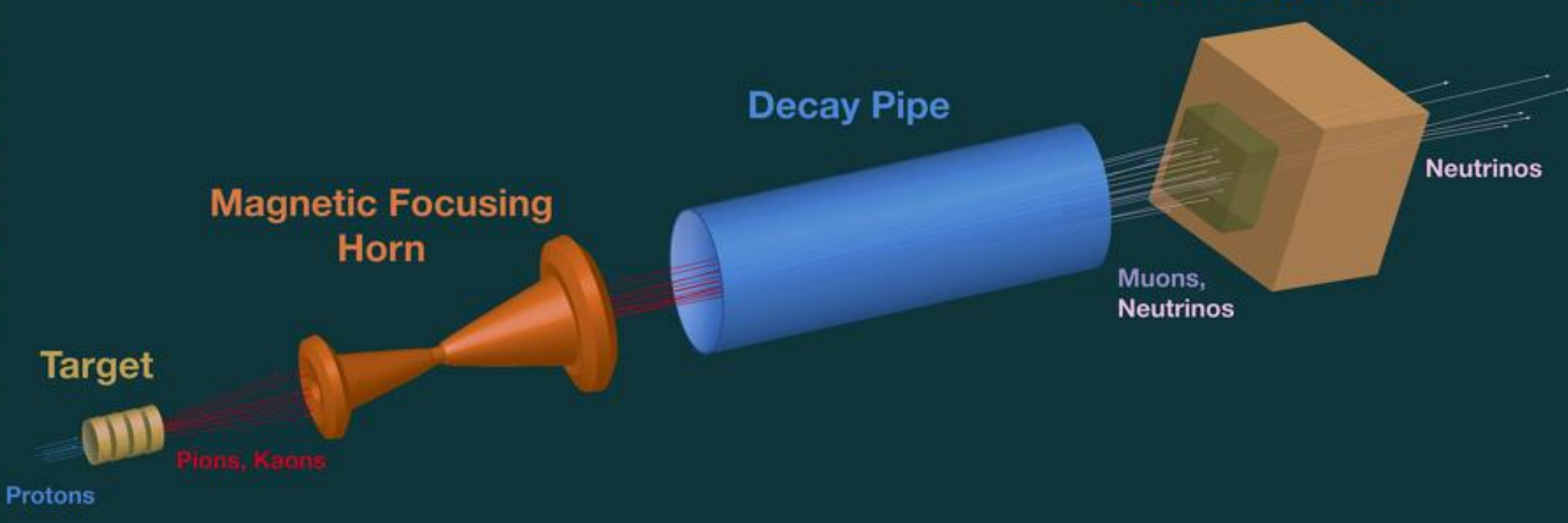
COLLABORATORS

K.J. KELLY, T. ZHOU (TAMU) | N. BLINOV (YORK) |  
| P.A.N. MACHADO, P.J. FOX (FERMILAB) |

DARK INTERACTIONS | VANCOUVER, BRITISH COLUMBIA | OCTOBER 2024

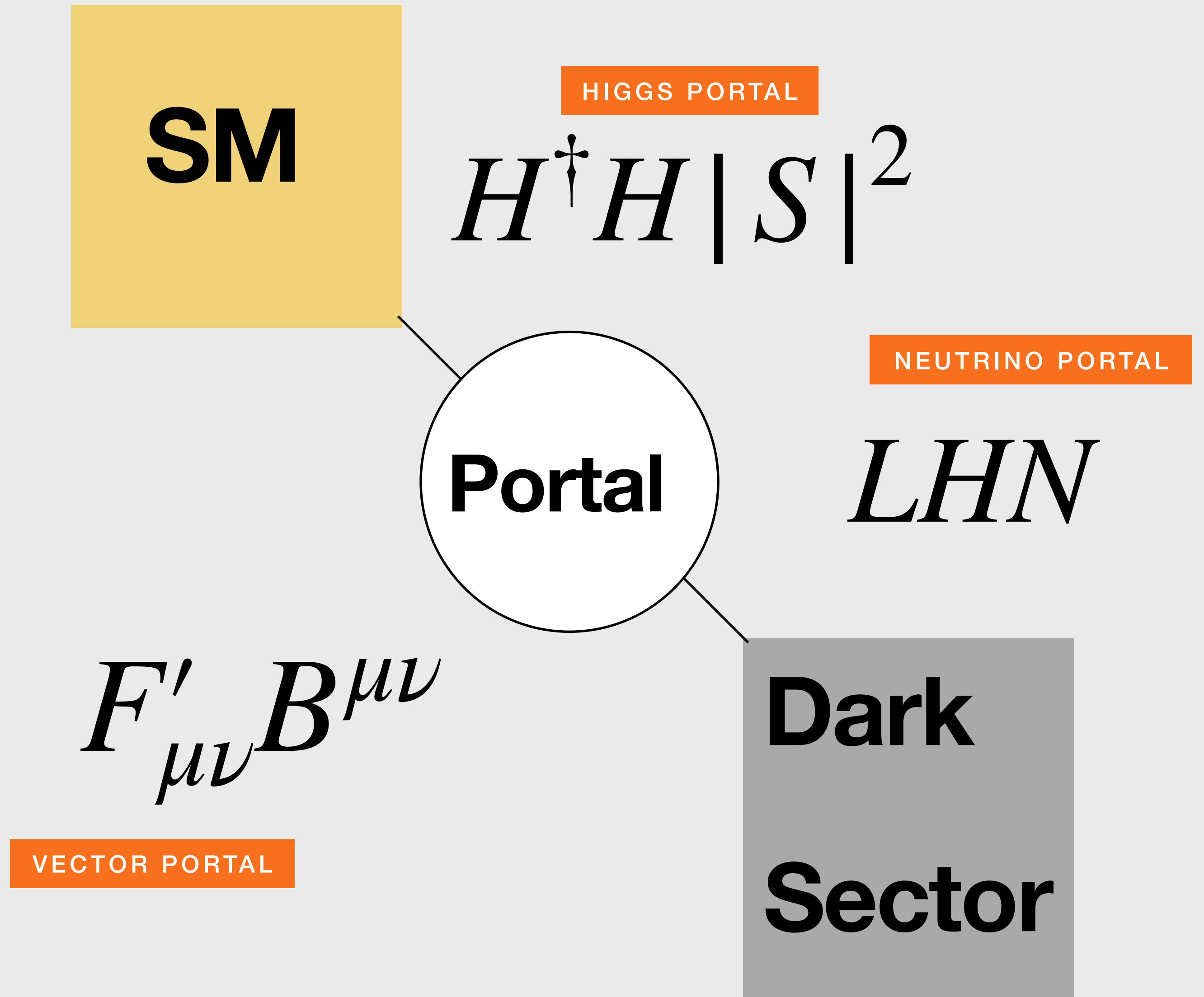
Caltech

Neutrino Theory Network



# Motivation & Context

- If light new physics exists it must be a gauge singlet.
  - Possibly complex dark sector (e.g. SM-like).
  - Few singlet operators available.
- Focus on "portals".

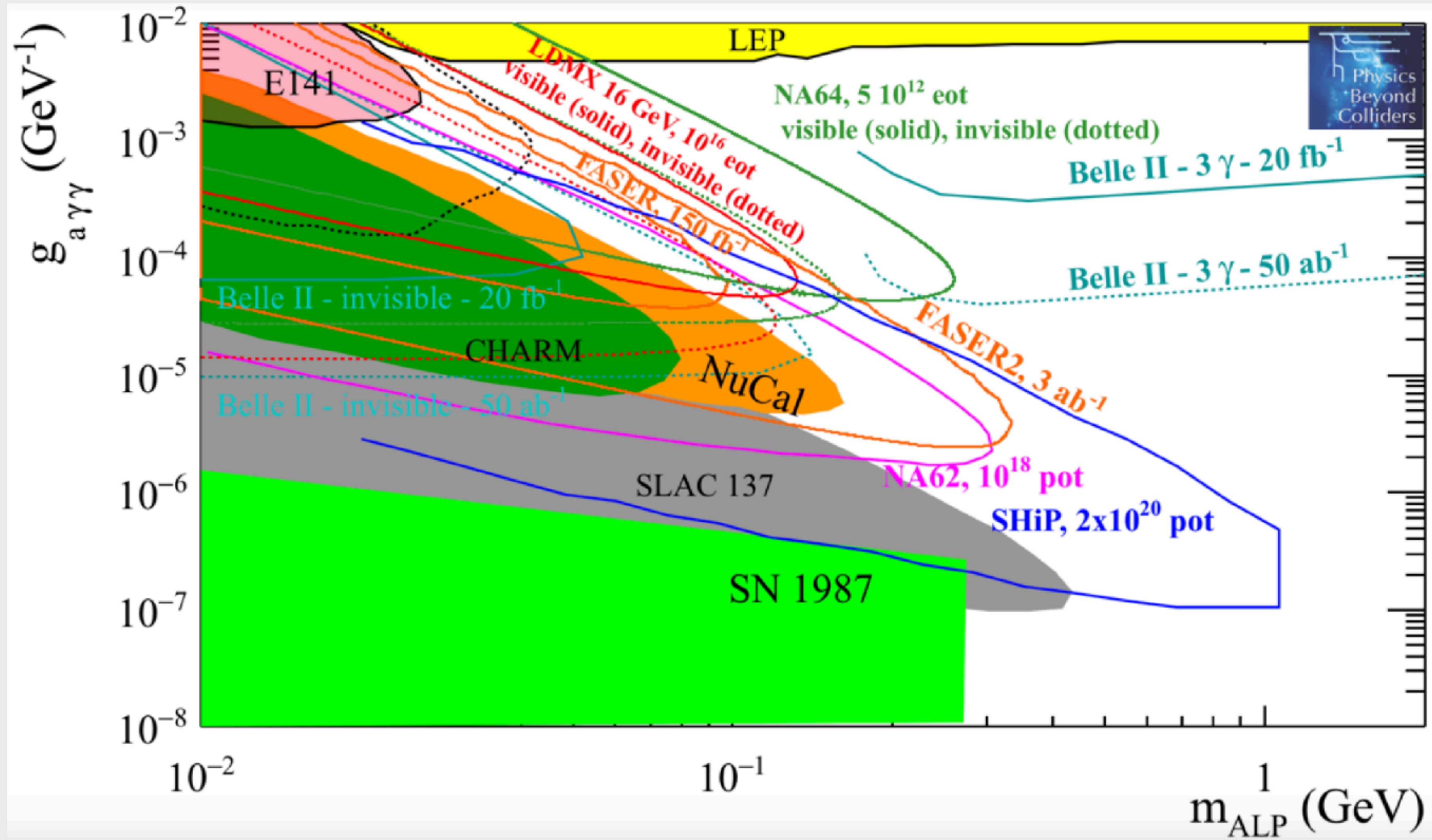




# Existing Searches

- Broad & competitive phenomenological landscape

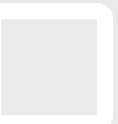
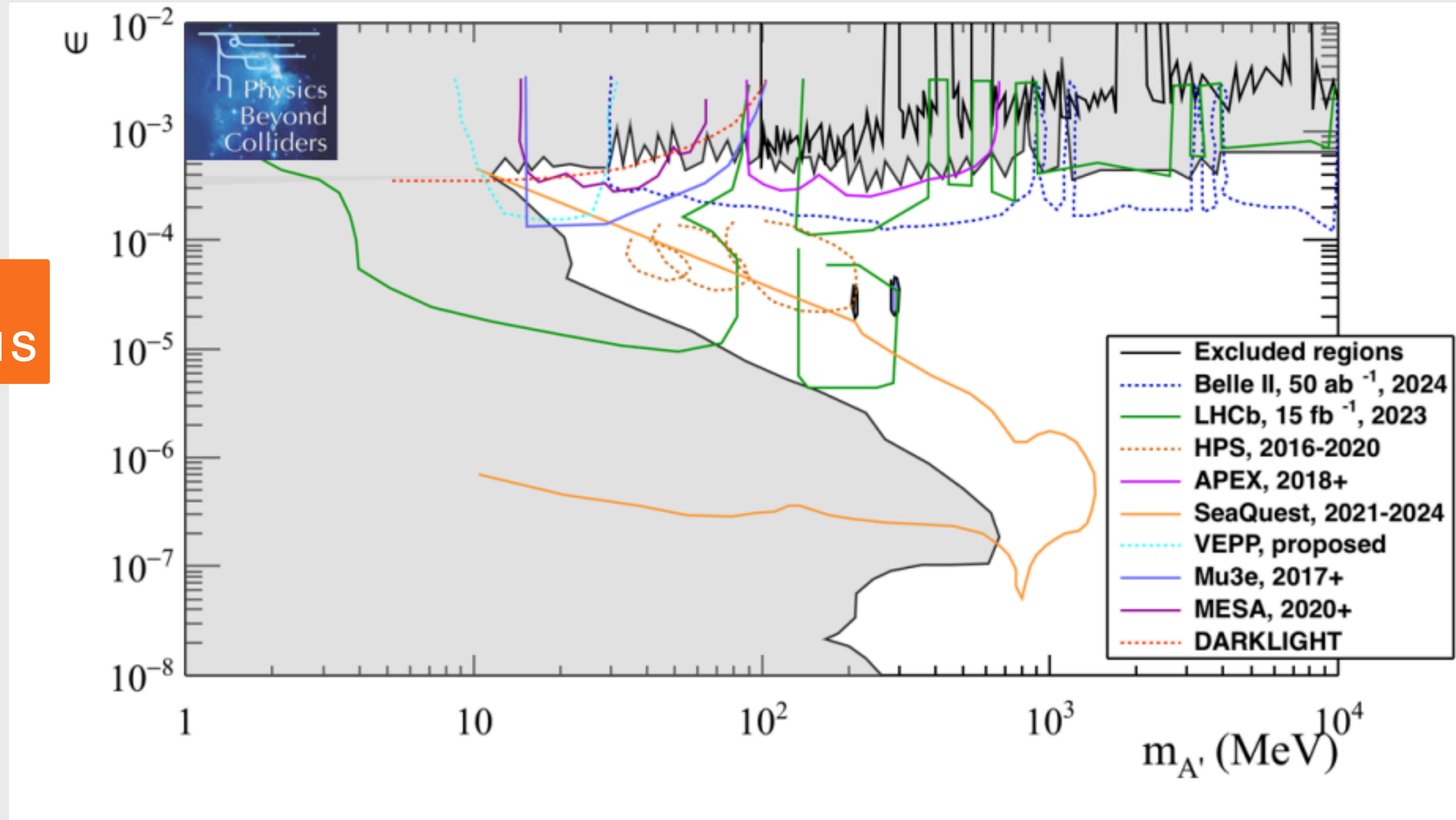
ALPS



# Existing Searches

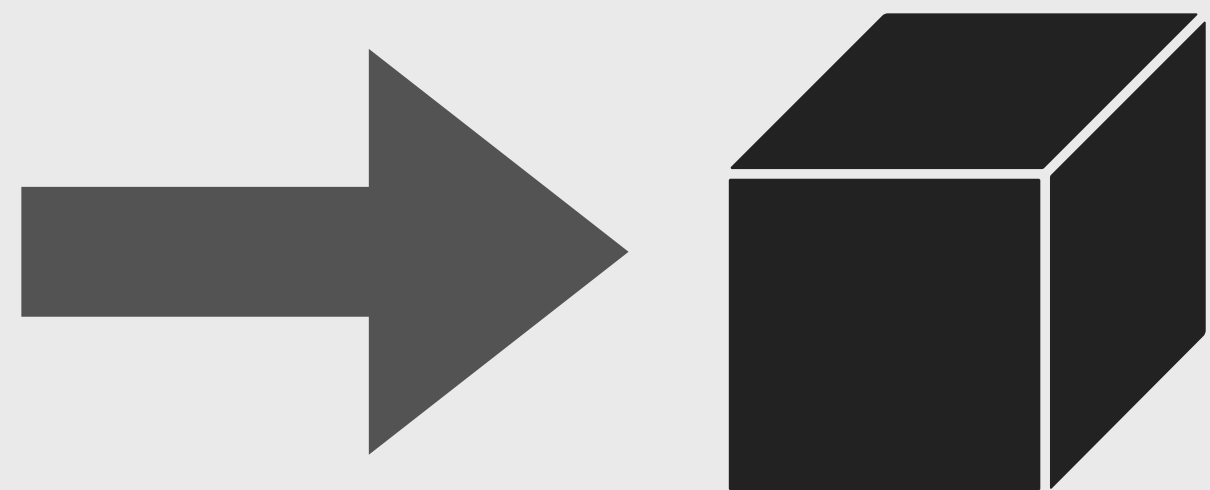
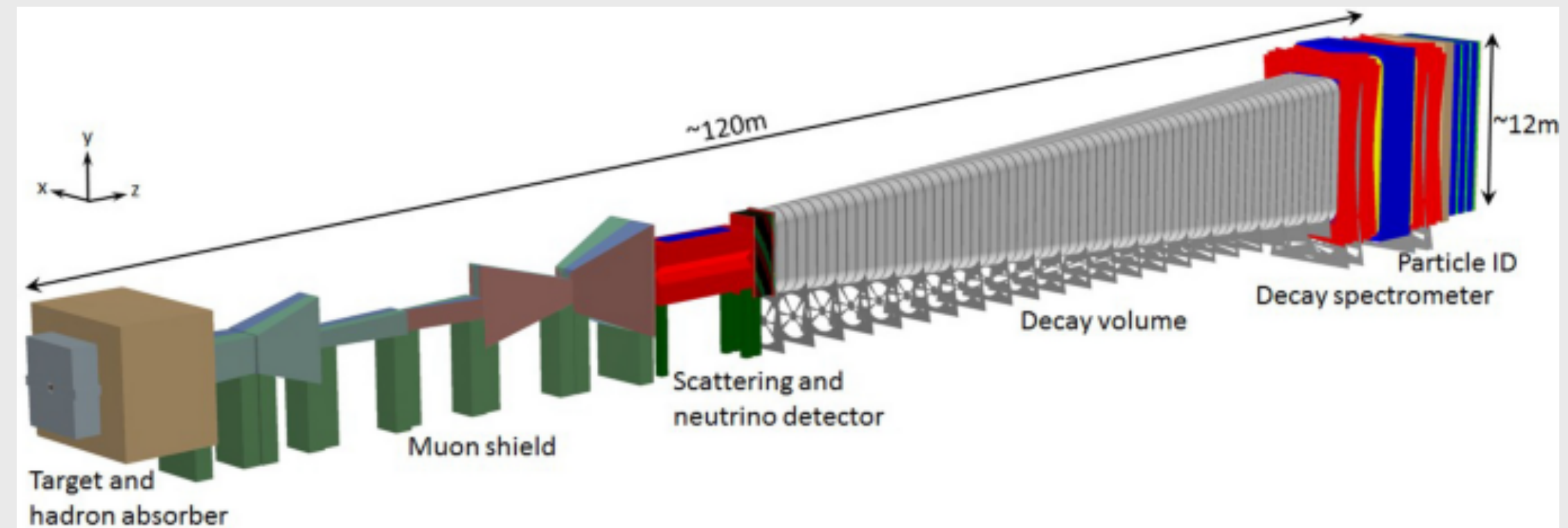
- Broad & competitive phenomenological landscape

DARK PHOTONS

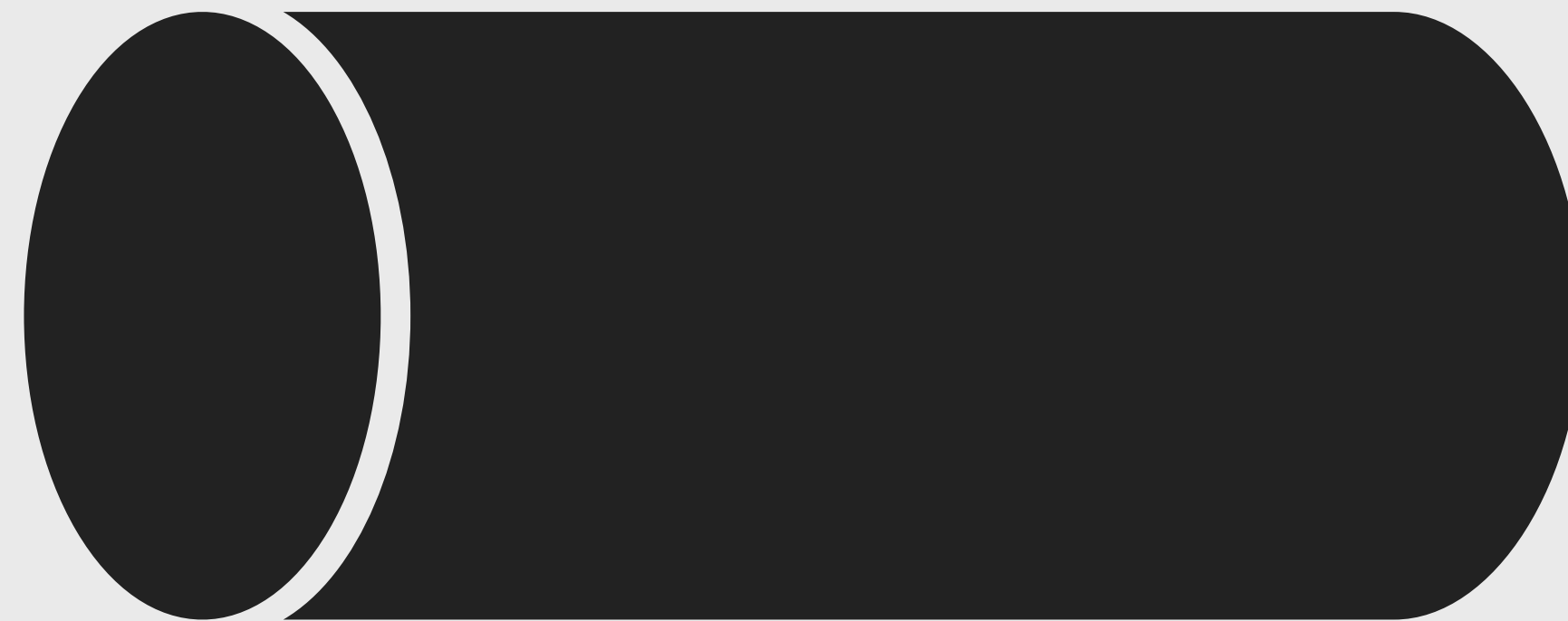


# How Do These Experiments Work

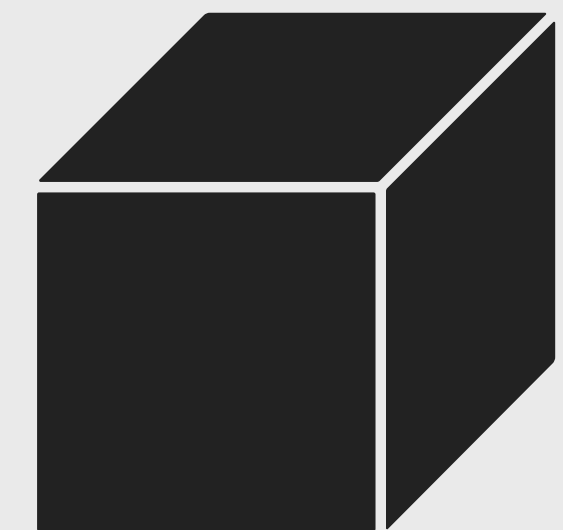
HOW TO MAKE  
A BSM BEAM



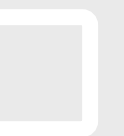
Beam target



pipe



detector



# Option 1: Meson Decays

IN NEUTRINO  
EXPERIMENTS

LONG LIVED

PROMPT

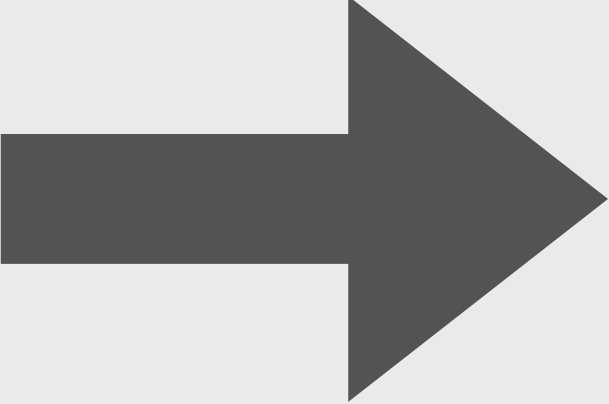
FROM BEAM  
STOP

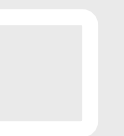
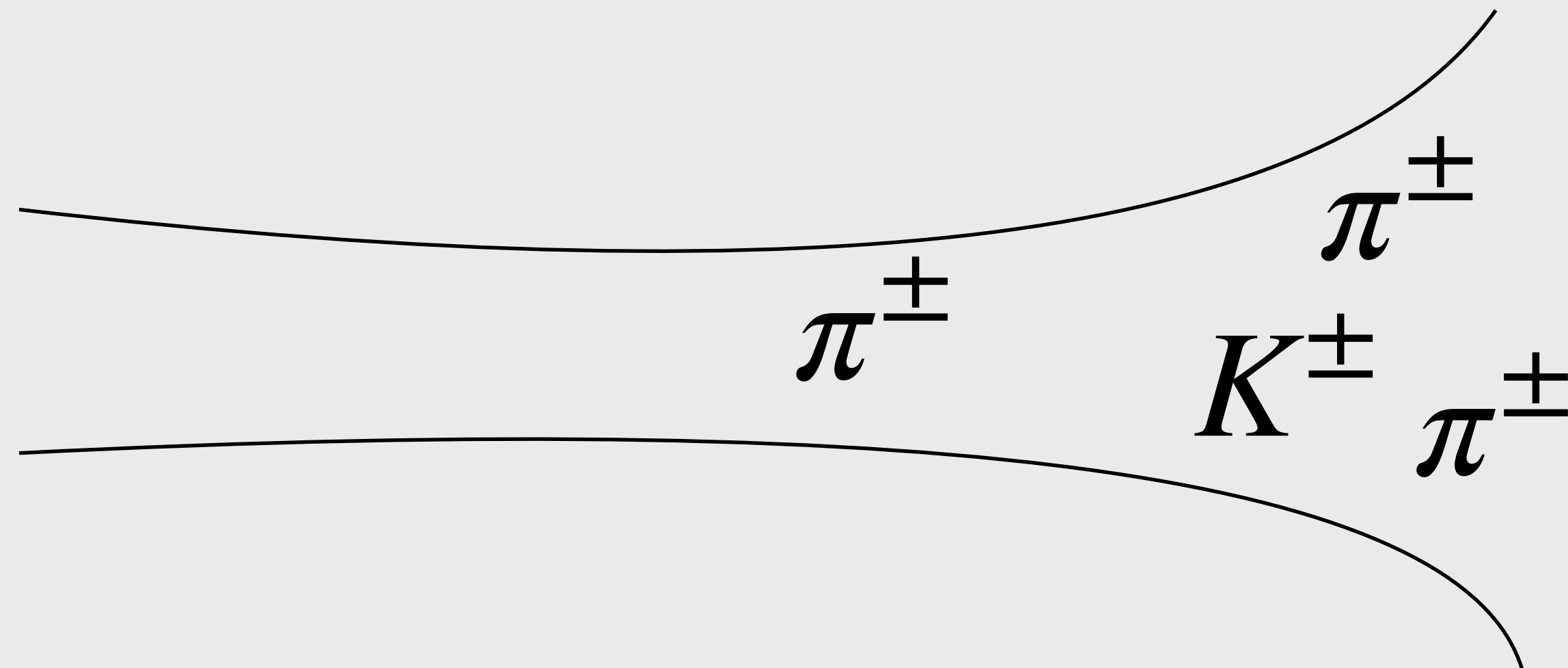
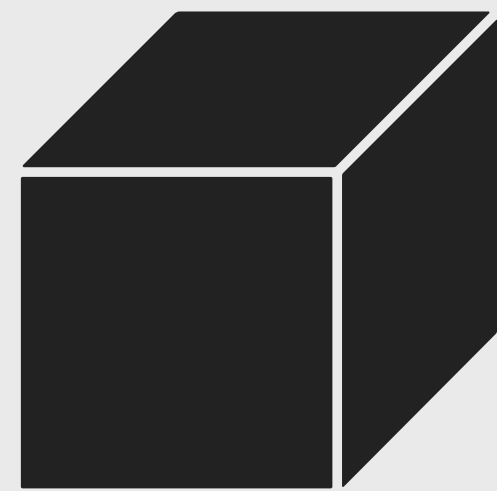
$$D^+ \rightarrow K^+ X$$

$$\pi^0 \rightarrow \gamma X$$

$$K^\pm \rightarrow \ell^\pm X$$

$\sim (10 - 100)$  GeV

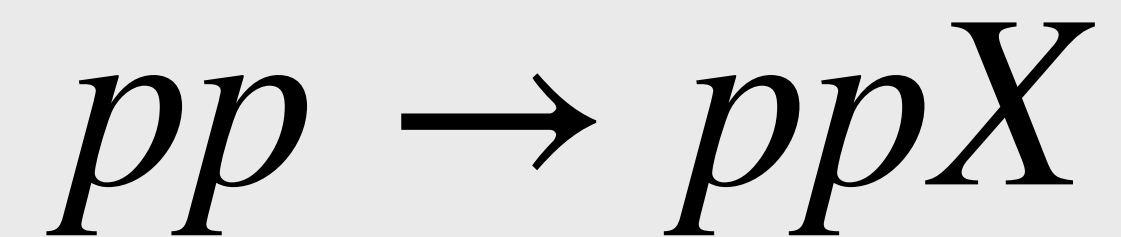
  
protons



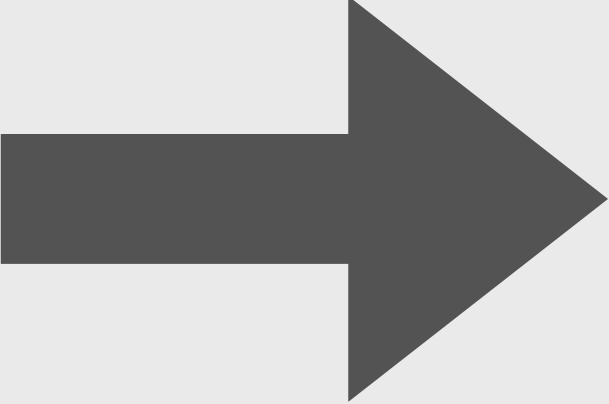


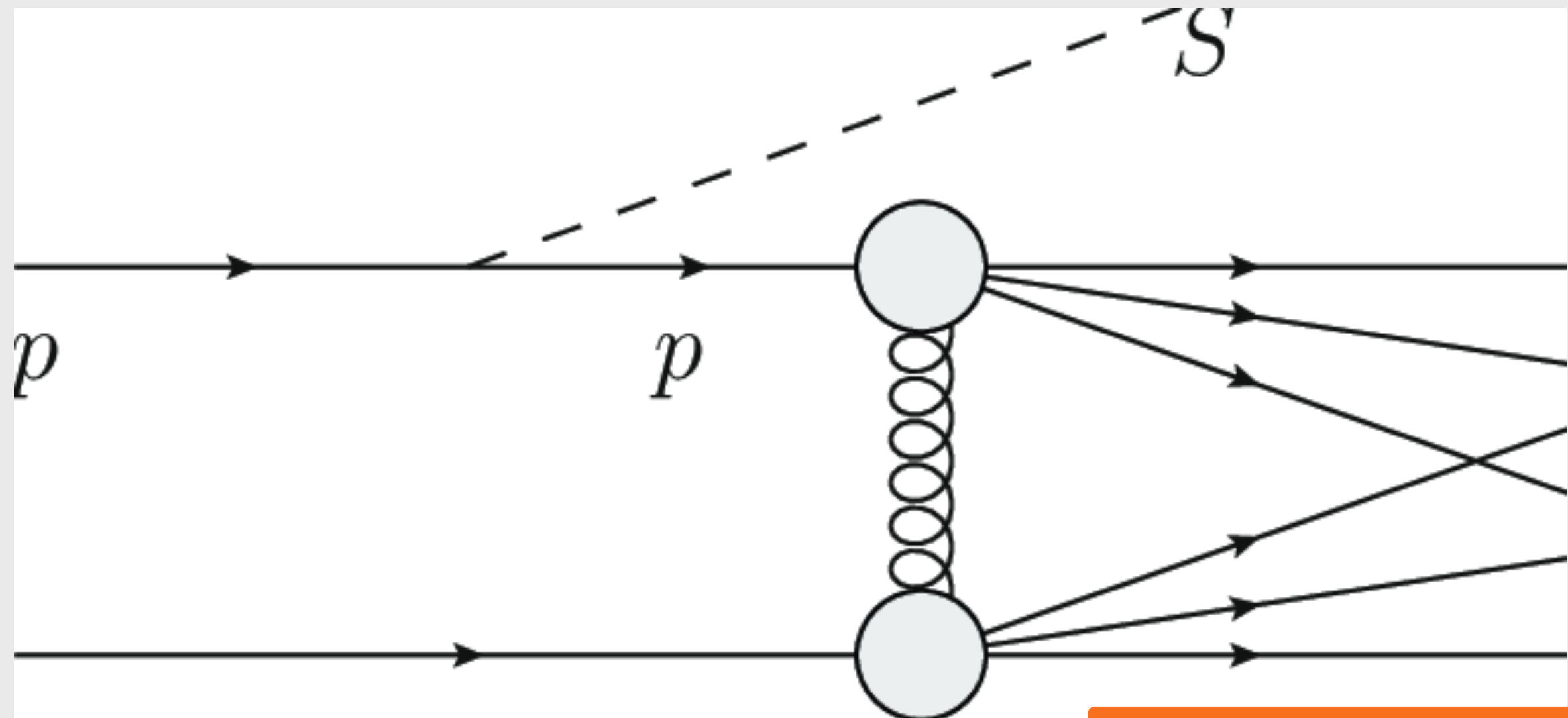
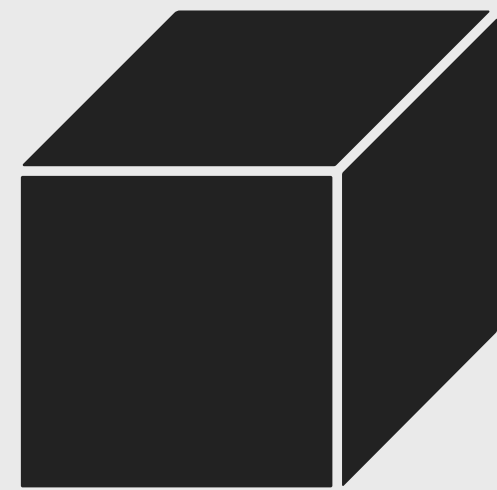
# Option 2: Primary Production

PROTON BREMMSTRAHLUNG

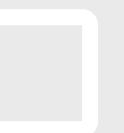


$\sim (10 - 100) \text{ GeV}$

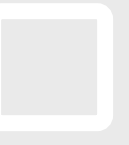
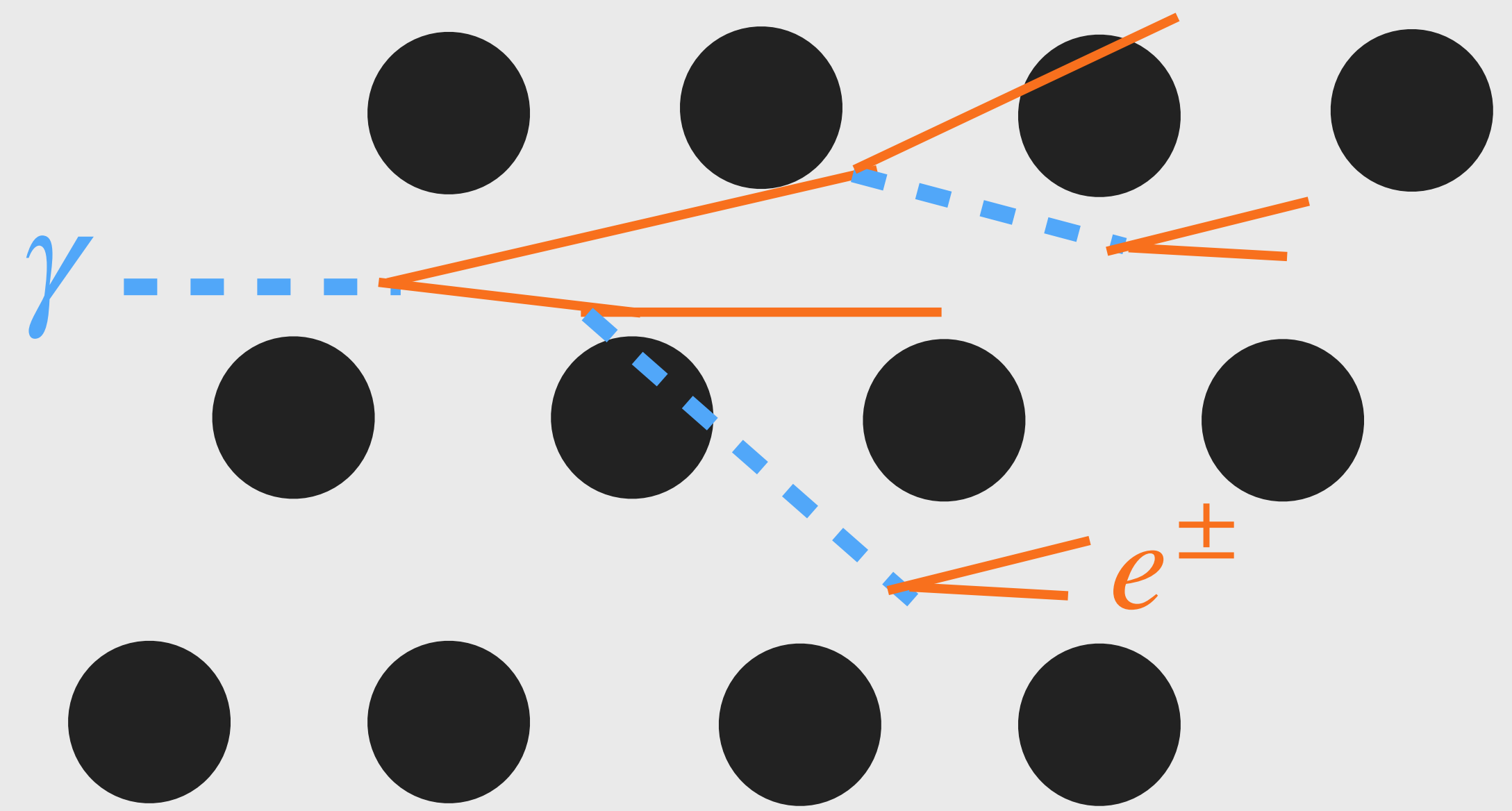
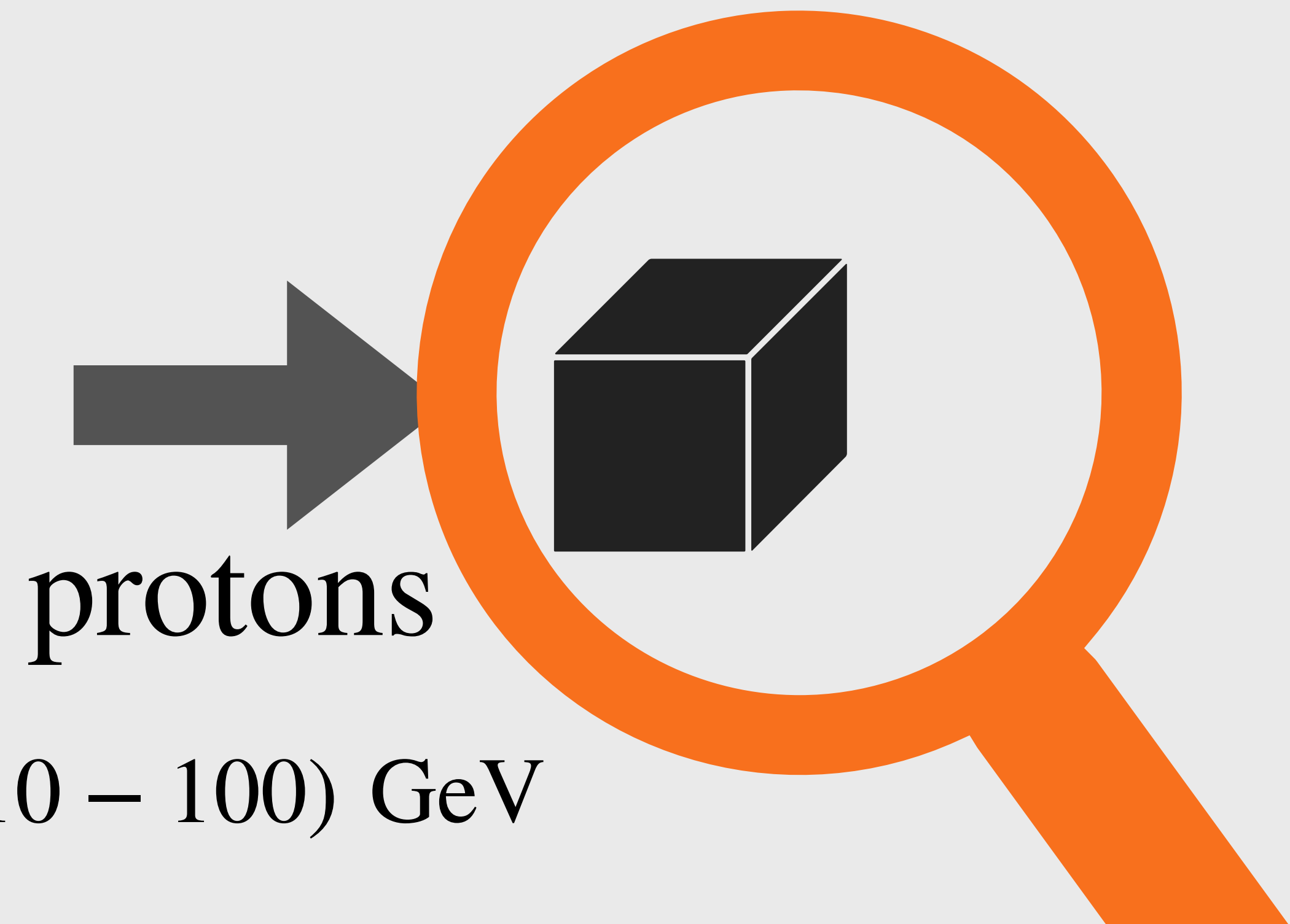
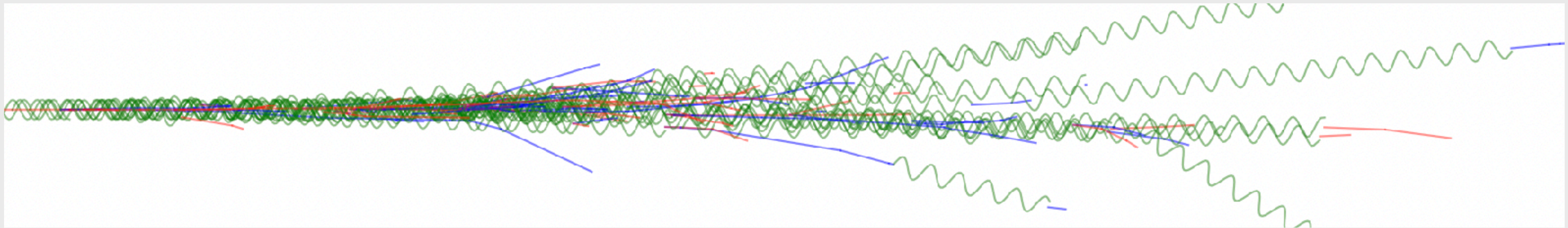
  
protons

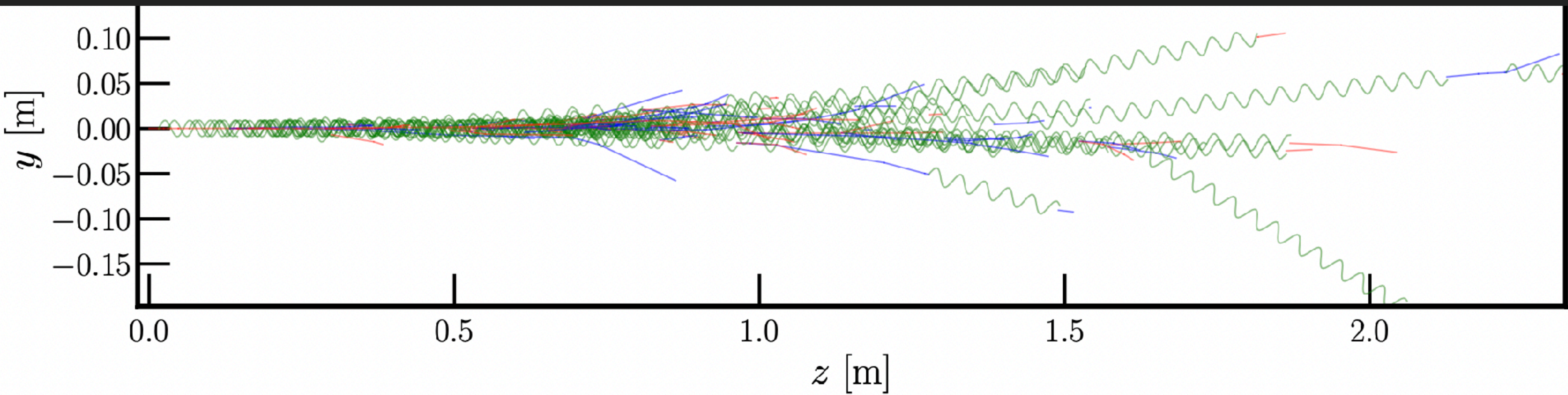


1904.10447



# Option 3: Secondary Production





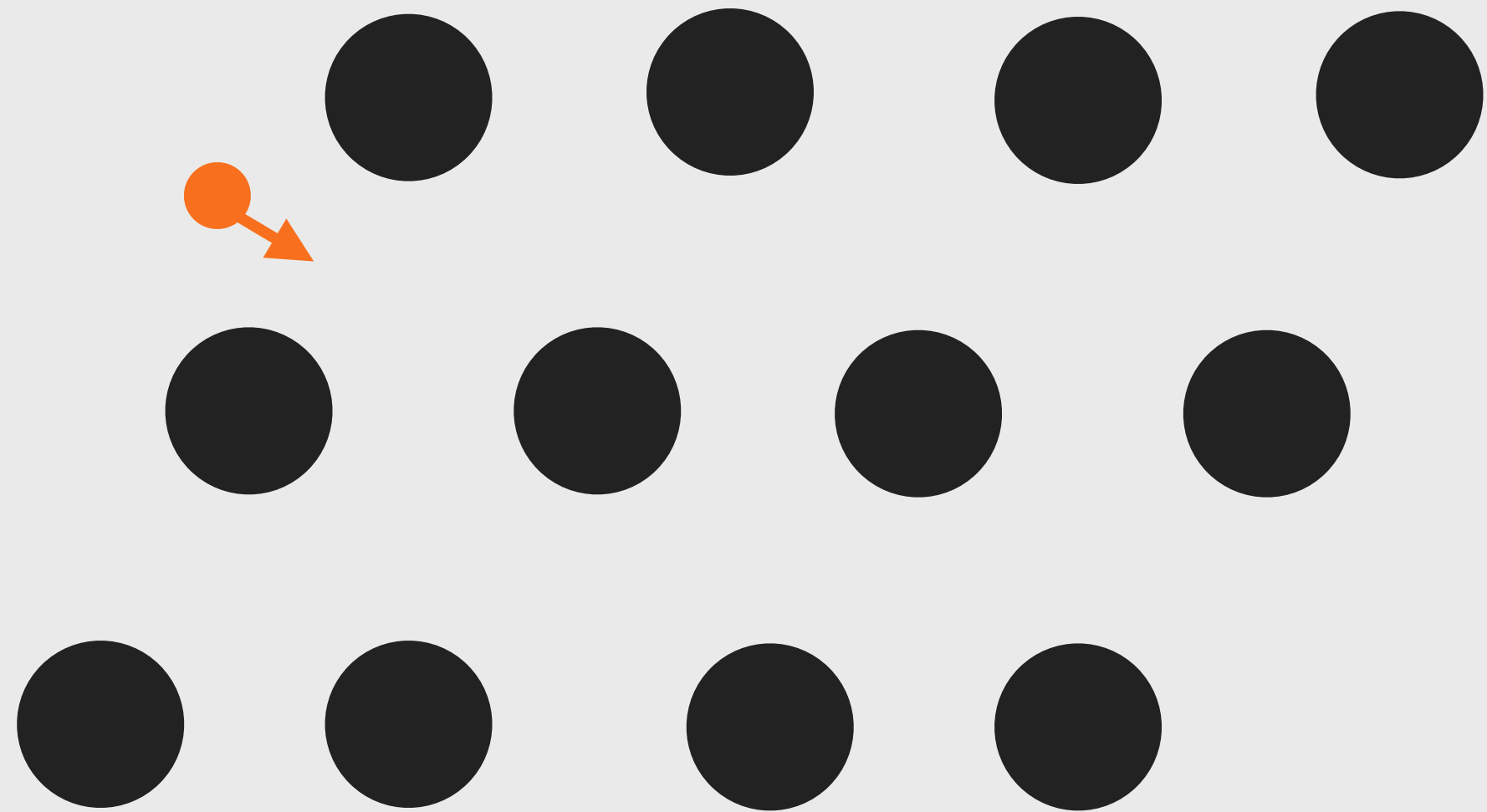
# Electromagnetic Secondaries

Electrophilic Light New Physics

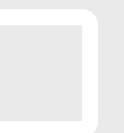


# Hadronic And Electromagnetic Cascades

- Consider a particle propagating through medium



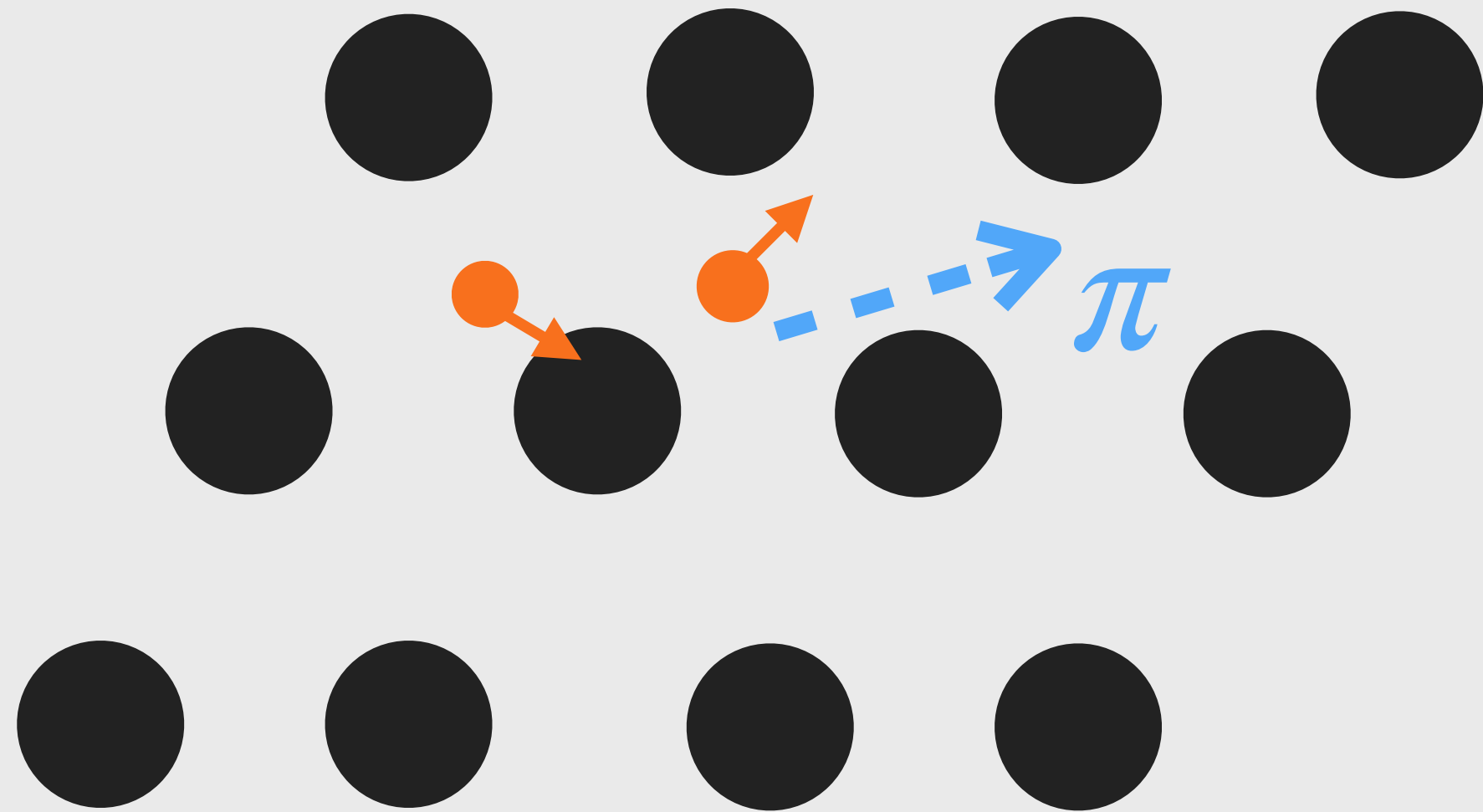
- Characteristic length between collisions  $\lambda_{\text{MFP}}$



# Hadronic And Electromagnetic Cascades

- Consider a particle propagating through medium

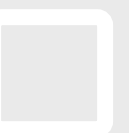
HADRONS



- Hadrons "down convert" energy into pions.
- Every generation is a new chance to make a BSM particle.
- Multiplicity of interactions grows with energy.

NEW RESOURCE BUT HARD TO STUDY SYSTEMATICALLY.

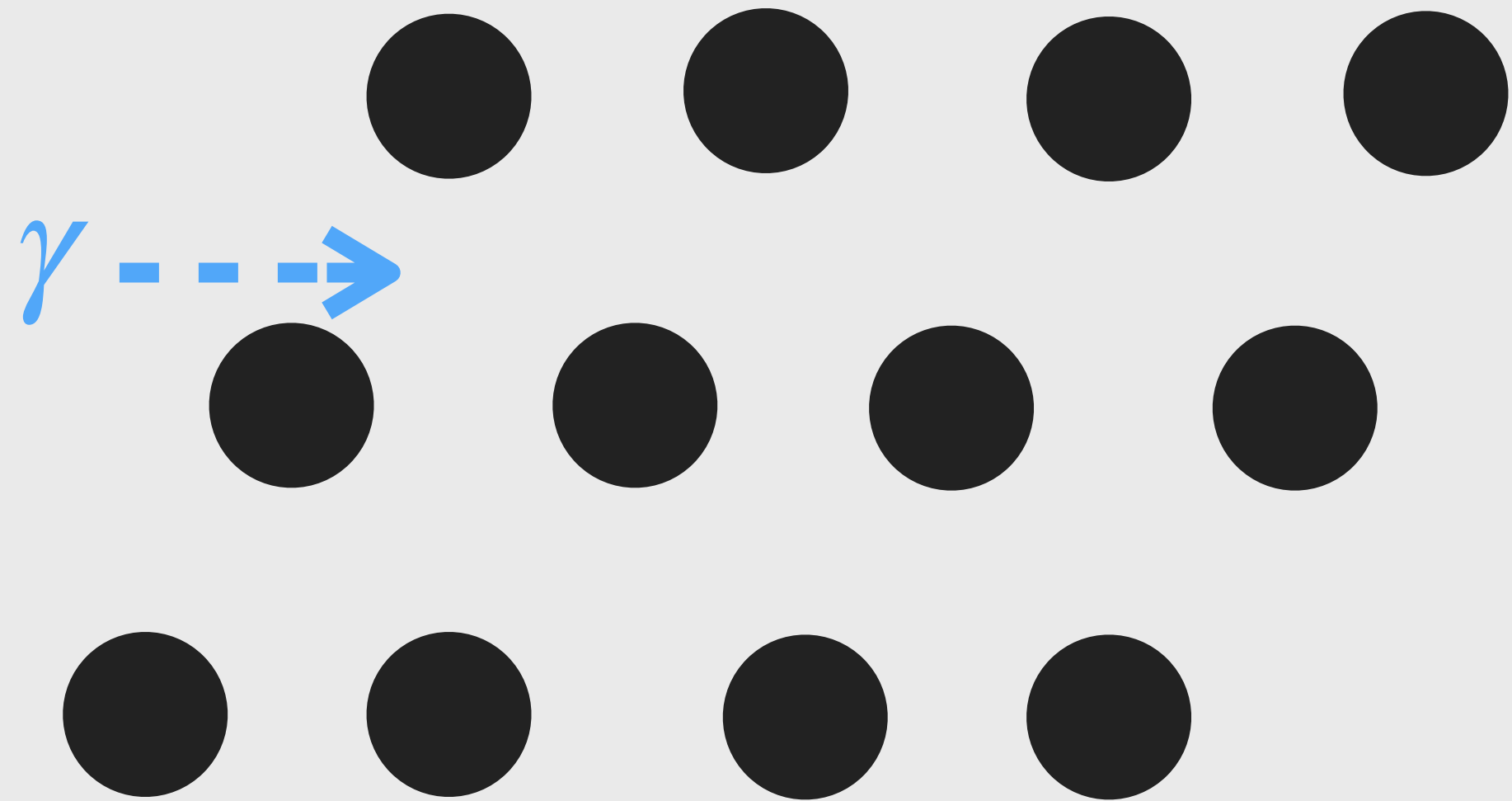
- Characteristic length between **hard** collisions  $X_H$



# Hadronic And Electromagnetic Cascades

- Consider a particle propagating through medium

## ELECTRONS & PHOTONS

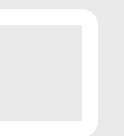


- Main reactions are



- Multiplicity of interactions grows with energy.

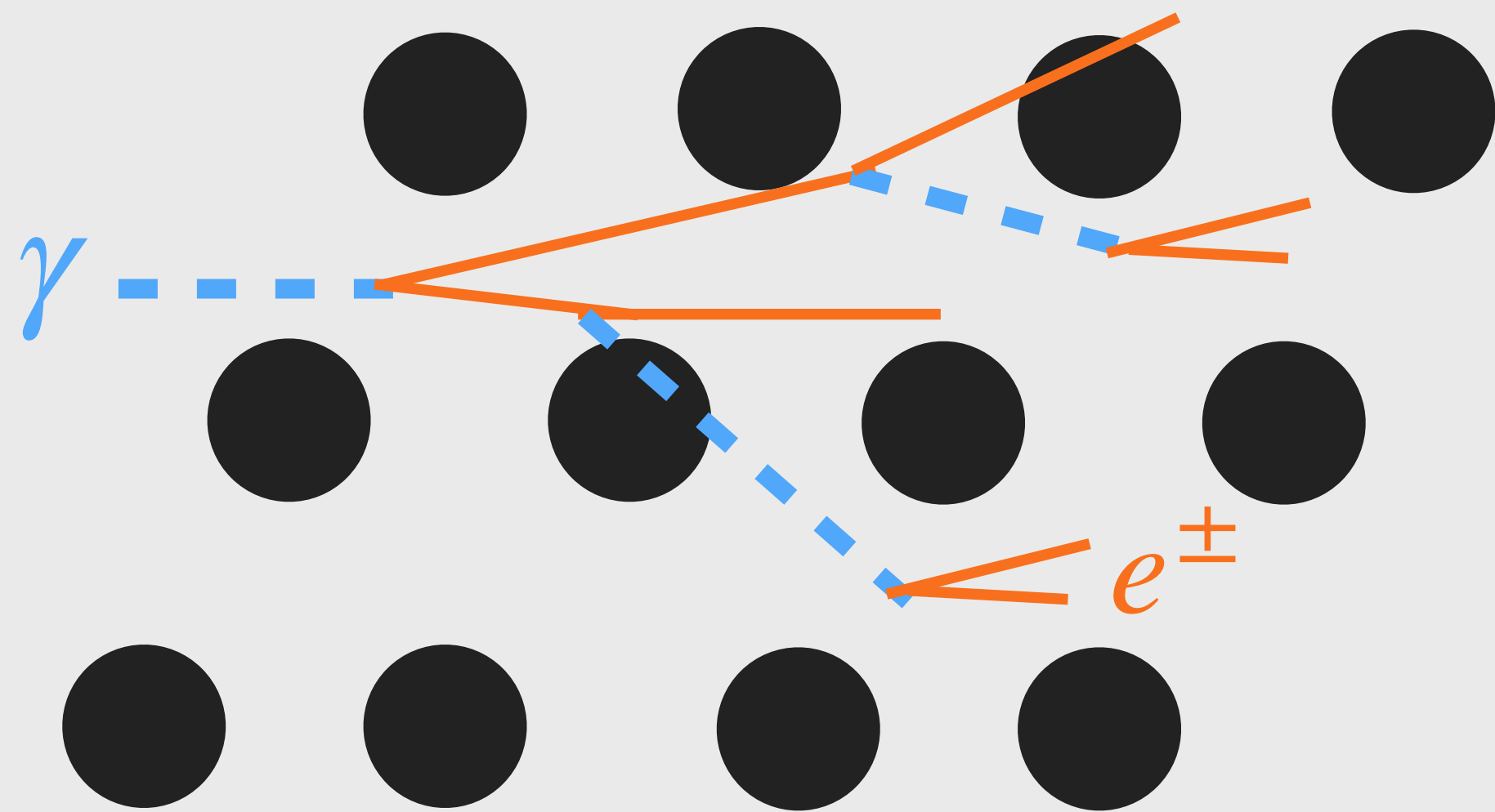
- Characteristic length between **hard** collisions  $X_0$  • Radiation length



# Hadronic And Electromagnetic Cascades

- Consider a particle propagating through medium

## ELECTRONS & PHOTONS



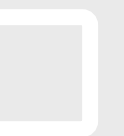
- Main reactions are

$$\gamma Z \rightarrow e^+ e^- Z \quad e^\pm Z \rightarrow e^\pm \gamma Z$$

- Multiplicity of interactions grows with energy.

NEW RESOURCE FOR DARK SECTORS.  
CAN BE COMPUTED PERTURBATIVELY.

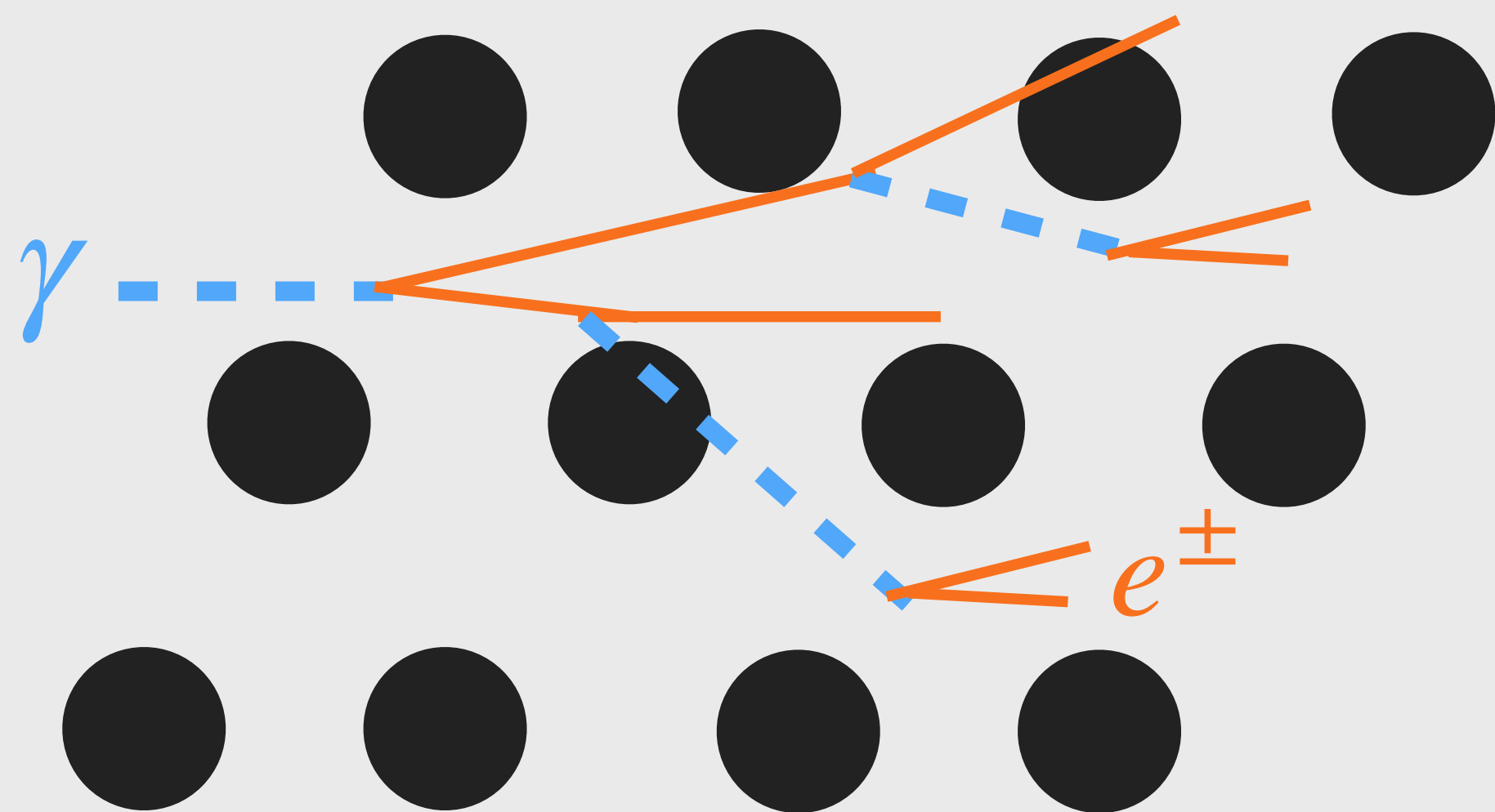
- Characteristic length between **hard** collisions  $X_0$  • Radiation length



# Hadronic And Electromagnetic Cascades

- Consider a particle propagating through medium

## ELECTRONS & PHOTONS



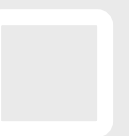
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# Previous Work On EM Secondaries

PHYSICAL REVIEW LETTERS **121**, 041802 (2018) **1807.05884**

## Novel Way to Search for Light Dark Matter in Lepton Beam-Dump Experiments

L. Marsicano,<sup>1,2</sup> M. Battaglieri,<sup>1</sup> M. Bondí,<sup>3</sup> C. D. R. Carvajal,<sup>4</sup> A. Celentano,<sup>1</sup> M. De Napoli,<sup>3</sup>  
R. De Vita,<sup>1</sup> E. Nardi,<sup>5</sup> M. Raggi,<sup>6</sup> and P. Valente<sup>7</sup>

PHYSICAL REVIEW D **102**, 075026 (2020)

**2006.09419**

## New production channels for light dark matter in hadronic showers

A. Celentano<sup>1</sup>, L. Darmé,<sup>2</sup> L. Marsicano,<sup>1</sup> and E. Nardi<sup>2</sup>

PHYSICAL REVIEW D **98**, 015031 (2018)

**1802.03794**

## Dark photon production through positron annihilation in beam-dump experiments

L. Marsicano,<sup>1,2</sup> M. Battaglieri,<sup>1</sup> M. Bondí,<sup>3</sup> C. D. R. Carvajal,<sup>4</sup> A. Celentano,<sup>1</sup>  
M. De Napoli,<sup>3</sup> R. De Vita,<sup>1</sup> E. Nardi,<sup>5</sup> M. Raggi,<sup>6</sup> and P. Valente<sup>7</sup>

## Event generation for beam dump experiments

Luca Buonocore,<sup>a,b</sup> Claudia Frugiuele,<sup>c</sup> Fabio Maltoni,<sup>d,e</sup> Olivier Mattelaer,<sup>d</sup> Francesco Tramontano<sup>b</sup>

**1812.06771**

**2108.03262**

PHYSICAL REVIEW D **104**, 115010 (2021)

## Extending the reach of leptophilic boson searches at DUNE and MiniBooNE with bremsstrahlung and resonant production

Francesco Capozzi<sup>1</sup>, Bhaskar Dutta,<sup>2</sup> Gajendra Gurung<sup>3</sup>, Wooyoung Jang,<sup>3</sup> Ian M. Shoemaker,<sup>1</sup>  
Adrian Thompson,<sup>2</sup> and Jaehoon Yu<sup>3</sup>

## Fully Geant4 compatible package for the simulation of Dark Matter in fixed target experiments ☆☆☆

**2101.12192**

M. Bondi<sup>a</sup>, A. Celentano<sup>a</sup>, R.R. Dusaev<sup>b</sup>, D.V. Kirpichnikov<sup>c</sup>, M.M. Kirsanov<sup>c</sup>,  
N.V. Krasnikov<sup>c,d</sup>, L. Marsicano<sup>a</sup>, D. Shchukin<sup>e</sup>

# Previous Work On Electromagnetic Cascades

PHYSICAL REVIEW LETTERS 121, 041802 (2018) 1807.05884

Event generation for beam dump experiments

- Despite multiple groups and a reasonable amount of activity, no systematic comparison between results has been made.
- Naive comparisons suggest large differences (orders of magnitude in some cases)
- Want a systematic analysis to resolve discrepancies.

# Previous Work On EM Secondaries

PHYSICAL REVIEW LETTERS 121, 041802 (2018) 1807.05884

Event generation for beam dump experiments

- Despite multiple groups and a reasonable amount of

## Dark fluxes from electromagnetic cascades

2401.06843

Nikita Blinov,<sup>1,2</sup> Patrick J. Fox,<sup>3</sup> Kevin J. Kelly,<sup>4,5</sup> Pedro A.N. Machado,<sup>3</sup> Ryan Plestid<sup>6</sup>

- want a systematic analysis to resolve discrepancies.

Dat

in

L. Marsicano,<sup>1,2</sup> M. Battaglieri,<sup>1</sup> M. Bondí,<sup>3</sup> C. D. R. Carvajal,<sup>4</sup> A. Celentano,<sup>1</sup>  
M. De Napoli,<sup>3</sup> R. De Vita,<sup>1</sup> E. Nardi,<sup>5</sup> M. Raggi,<sup>6</sup> and P. Valente<sup>7</sup>

M. Bondi<sup>a</sup>, A. Celentano<sup>a</sup>, R.R. Dusaev<sup>b</sup>, D.V. Kirpichnikov<sup>c</sup>, M.M. Kirsanov<sup>c</sup>,  
N.V. Krasnikov<sup>c,d</sup>, L. Marsicano<sup>a</sup>, D. Shchukin<sup>e</sup>

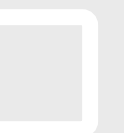
# Challenges With Far-Forward Detectors

SMALL ANGLES

$$\Phi(\theta, E) \text{ for } \theta < \theta_c$$



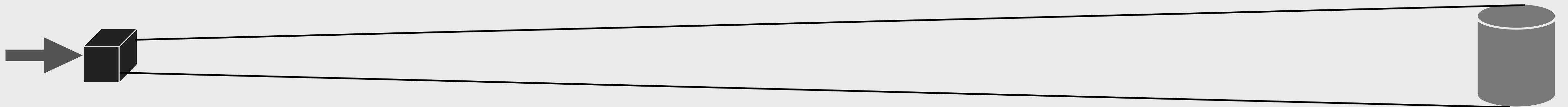
UNUSUALLY SENSITIVE TO ANGULAR SPREADING



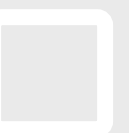
# Challenges With Far-Forward Detectors

LARGE HIERARCHY OF ENERGIES

$$E_{\text{beam}} \gg E_{\pi} \gg m_{\chi} \gg E_{\text{thr}}$$



NEED TO WORRY ABOUT MANY GENERATIONS IN A SHOWER

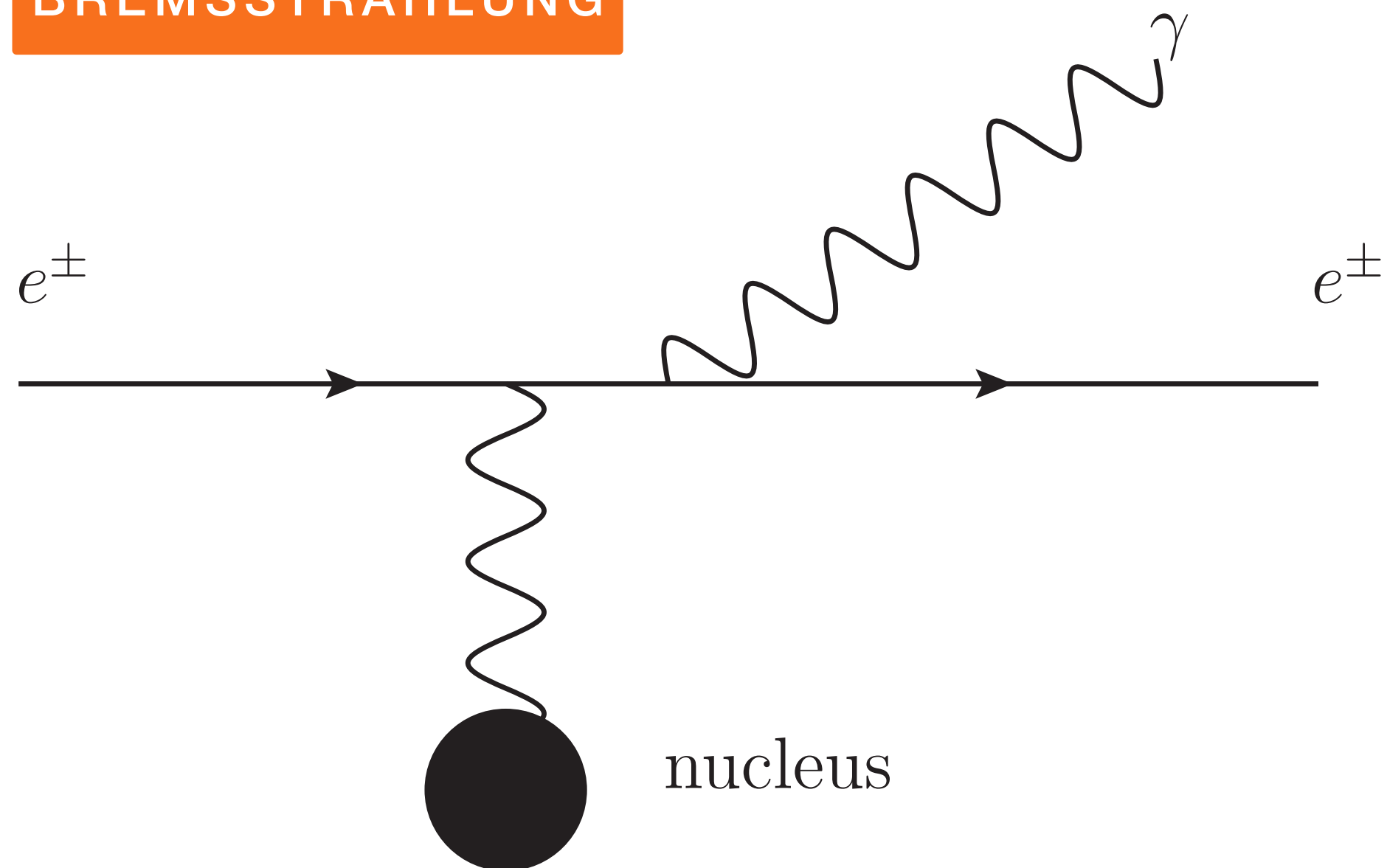


# Explicit Model: Dark Vector Boson

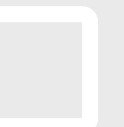
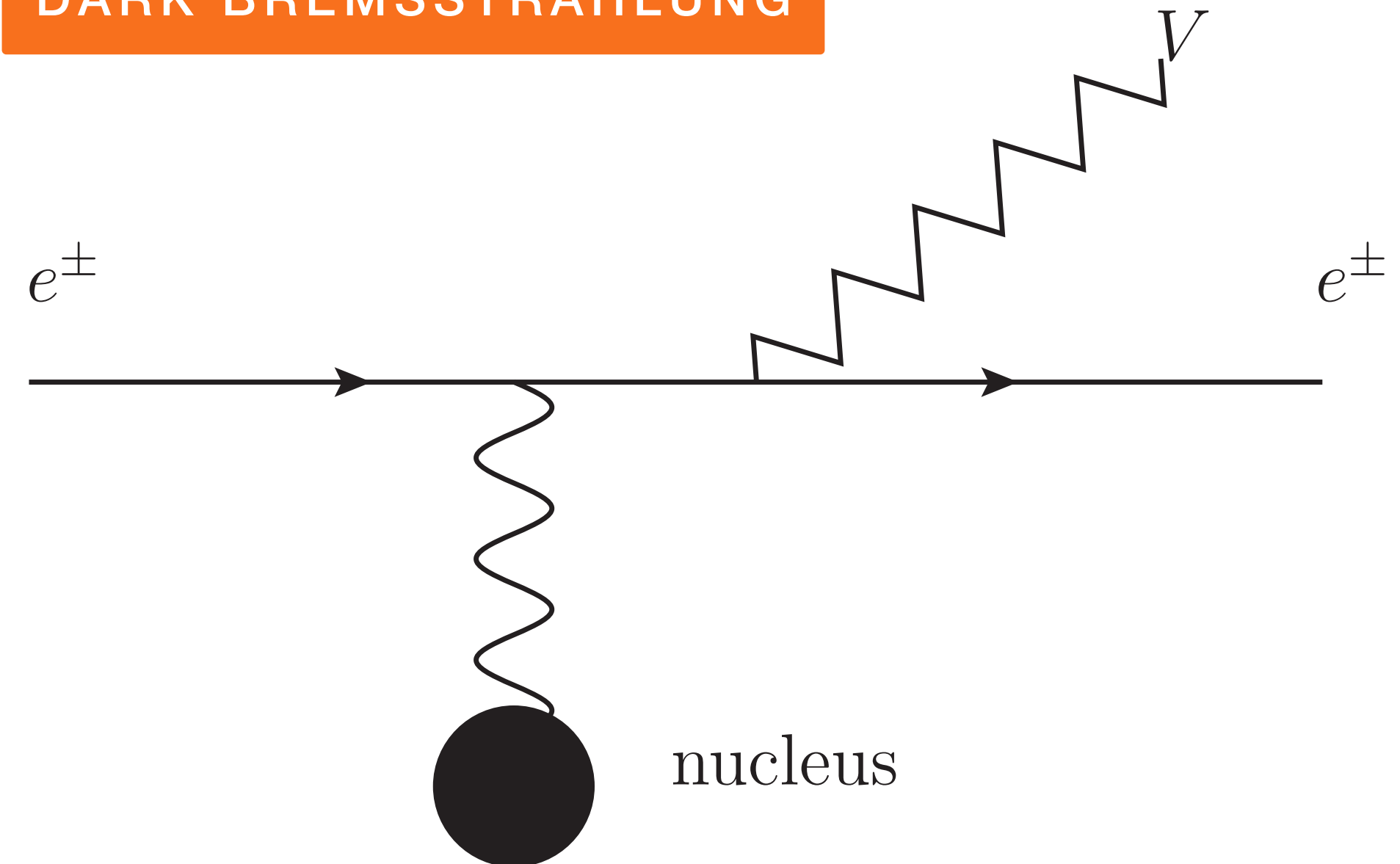
$$\mathcal{L} \supset g \bar{e} \gamma_{\mu} e V^{\mu}$$

- Vector boson of mass  $m_V$  couples to the electron vector current.

BREMSSTRAHLUNG



DARK BREMSSTRAHLUNG

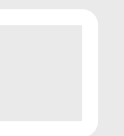


# Explicit Model: Dark Vector Boson

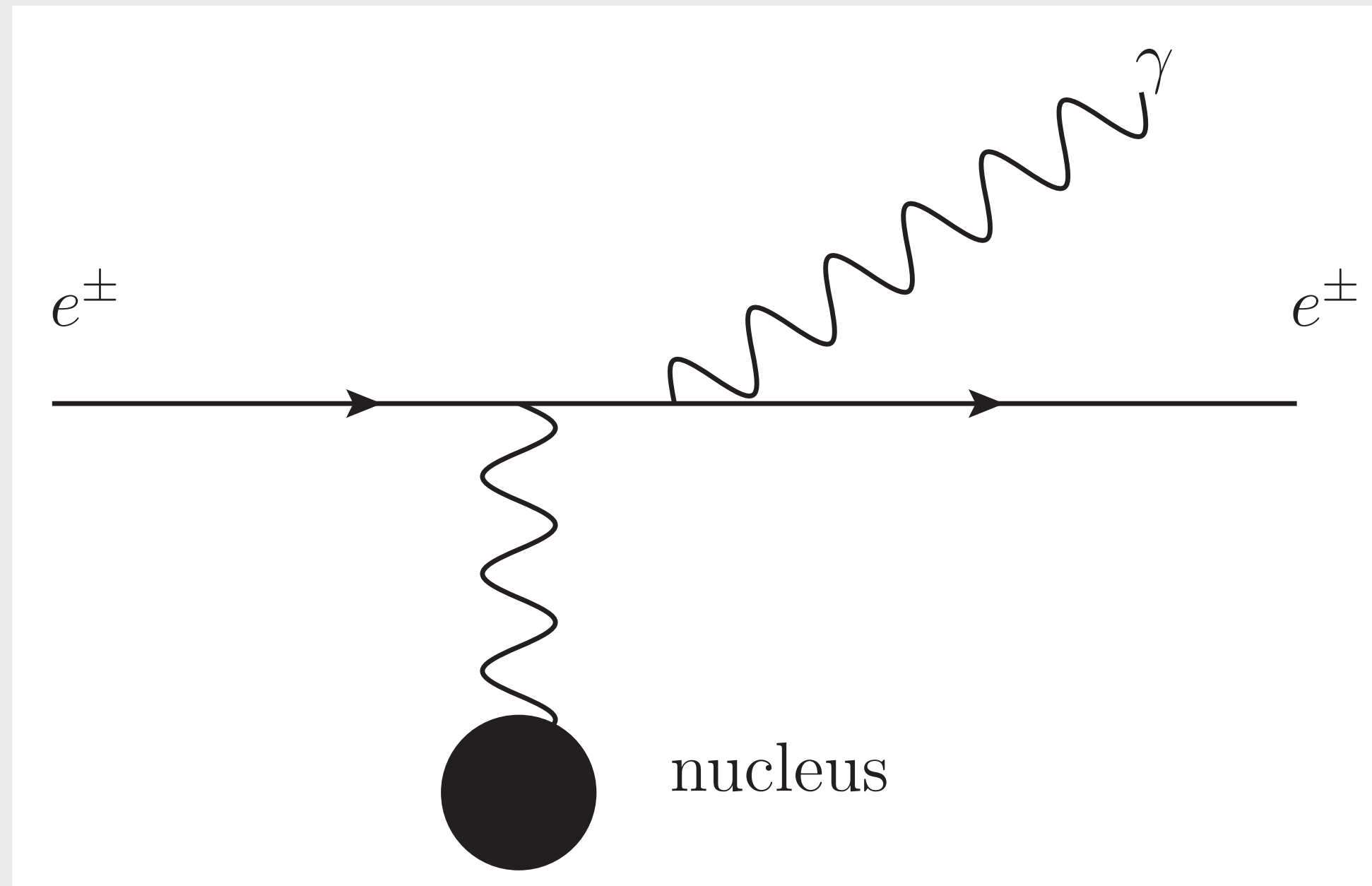
$$\mathcal{L} \supset g \bar{e} \gamma_{\mu} e V^{\mu}$$

- Our goal is to compute the flux from an EM cascade at the detector.

$\Phi_{\text{det}}(E_V)$

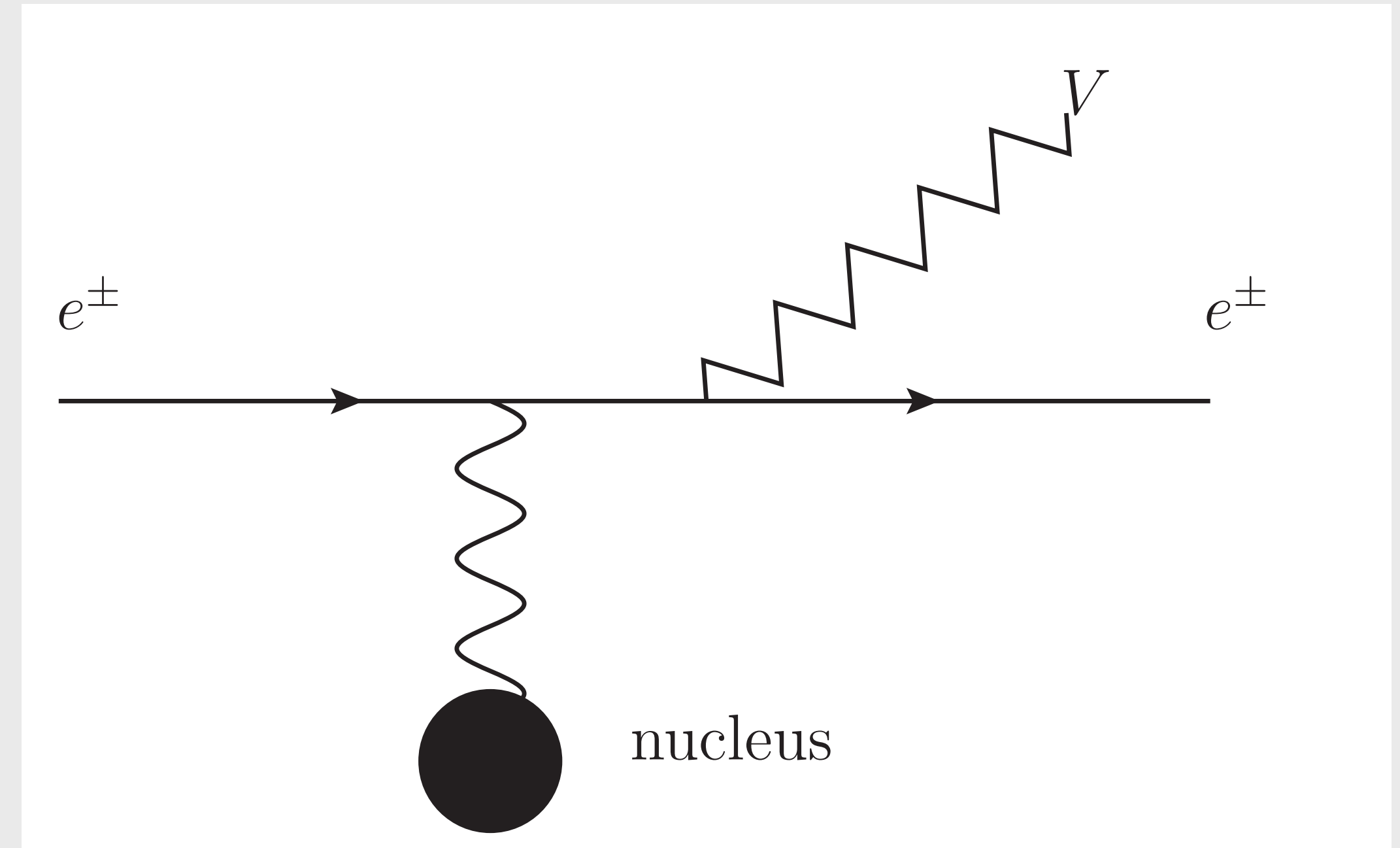


# SM Event $\rightarrow$ BSM Event



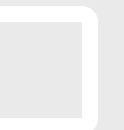
- Consider an event in a MC event record.

$$(\mathbf{p}, \mathbf{x})_e \rightarrow (\mathbf{p}', \mathbf{x})_e + (\mathbf{q}', \mathbf{x})_\gamma$$



- How do we use this to generate BSM event?

$$(\mathbf{p}, \mathbf{x})_e \rightarrow (\mathbf{p}', \mathbf{x})_e + (\mathbf{q}', \mathbf{x})_\nu$$





# SM Event $\rightarrow$ BSM Event

$$(\mathbf{p}, \mathbf{x})_e \rightarrow (\mathbf{p}', \mathbf{x})_e + (\mathbf{q}', \mathbf{x})_\gamma$$

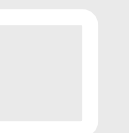
FOCUS ON PARENT

DRAW KINEMATICS  
FROM BSM DIST.

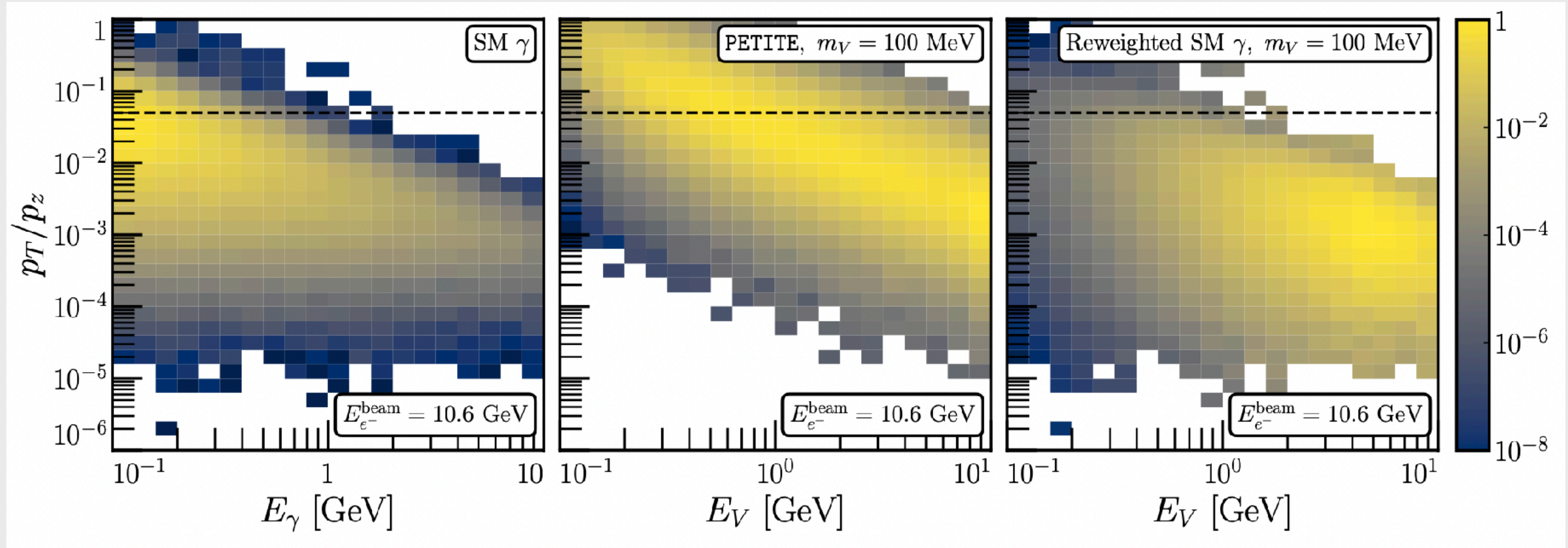
$$\frac{d\sigma}{d\Pi} = (2\pi)^4 \delta^{(4)}(\Sigma P) |\mathcal{M}_{e \rightarrow e\gamma}|^2$$

COMPUTE  
BRANCHING RATIO

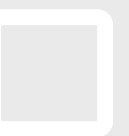
$$\text{BR} = \frac{\sigma_{\text{BSM}}}{\sigma_{\text{tot}}} \approx \frac{\sigma_{\text{BSM}}}{\sigma_{\text{SM}}}$$



# What Not To Do

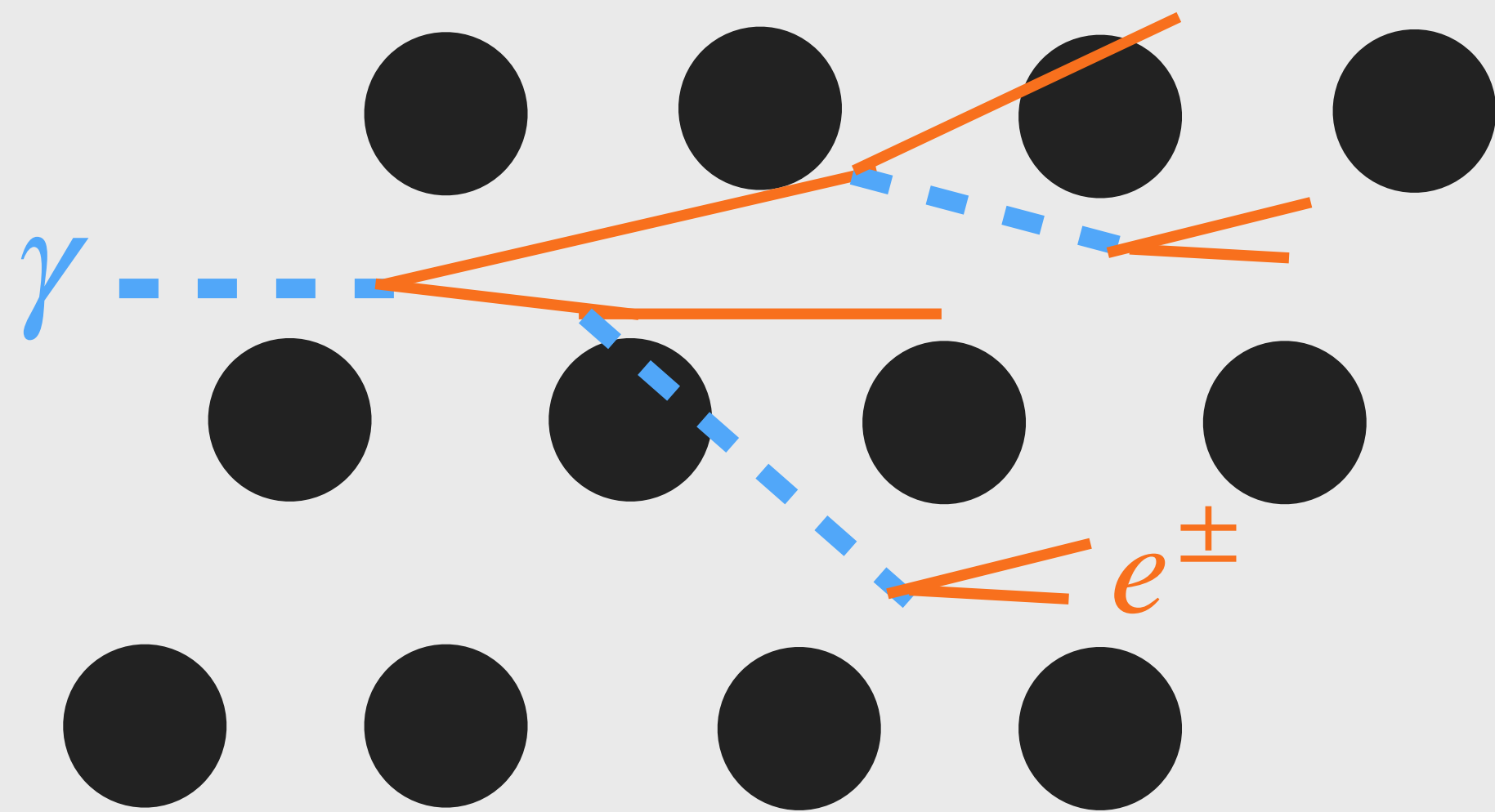


- Trying to turn daughter photons into daughter dark photons is dangerous because of different distributions.



# PETITE In A Nutshell

## PROCESSES INCLUDED

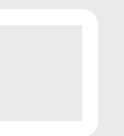
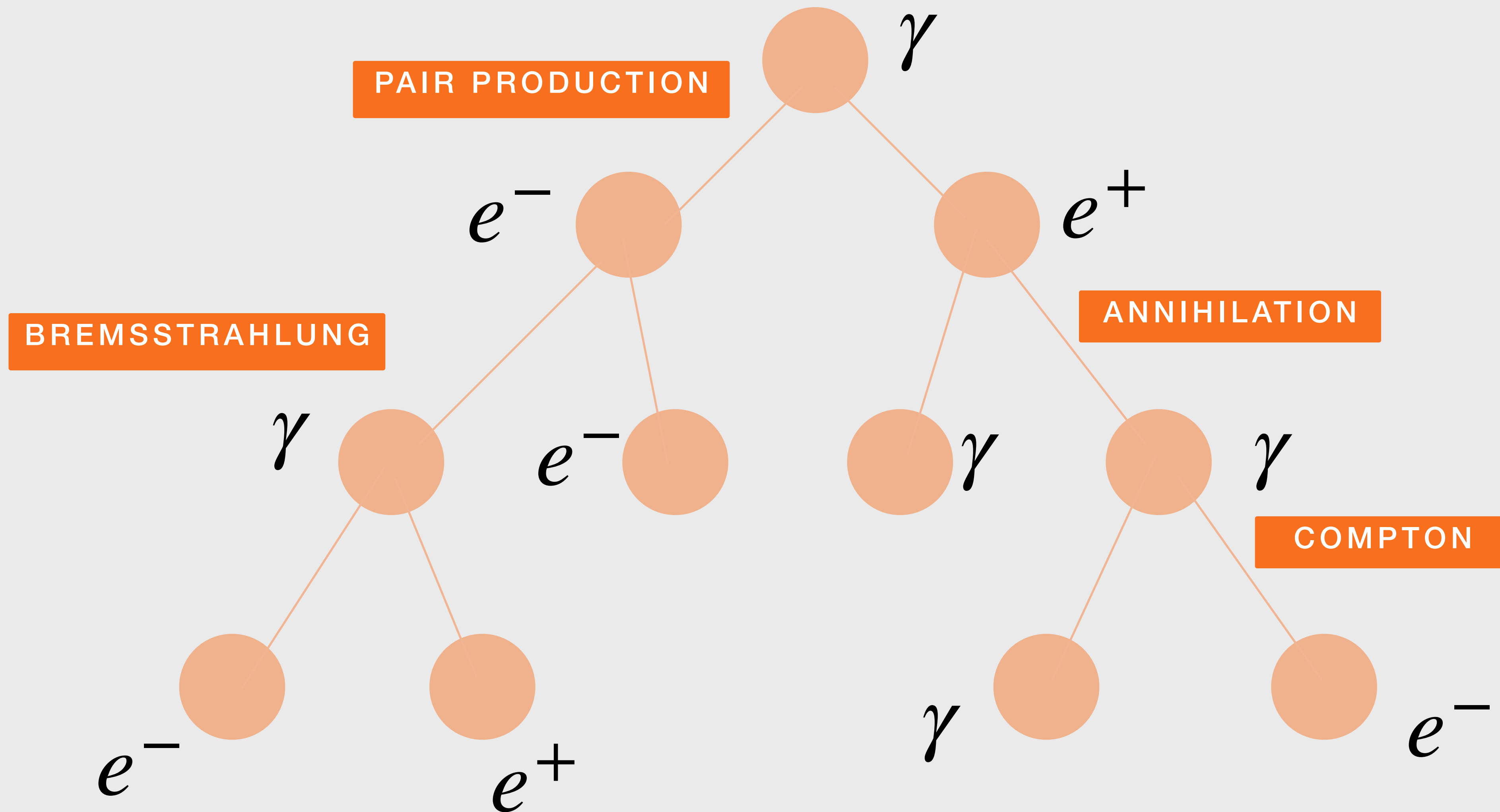


- Static nuclear centres source Coulomb fields.
- Electrons treated as a homogeneous gas of electrons at rest.
- Atomic screening included for bremsstrahlung and pair production.

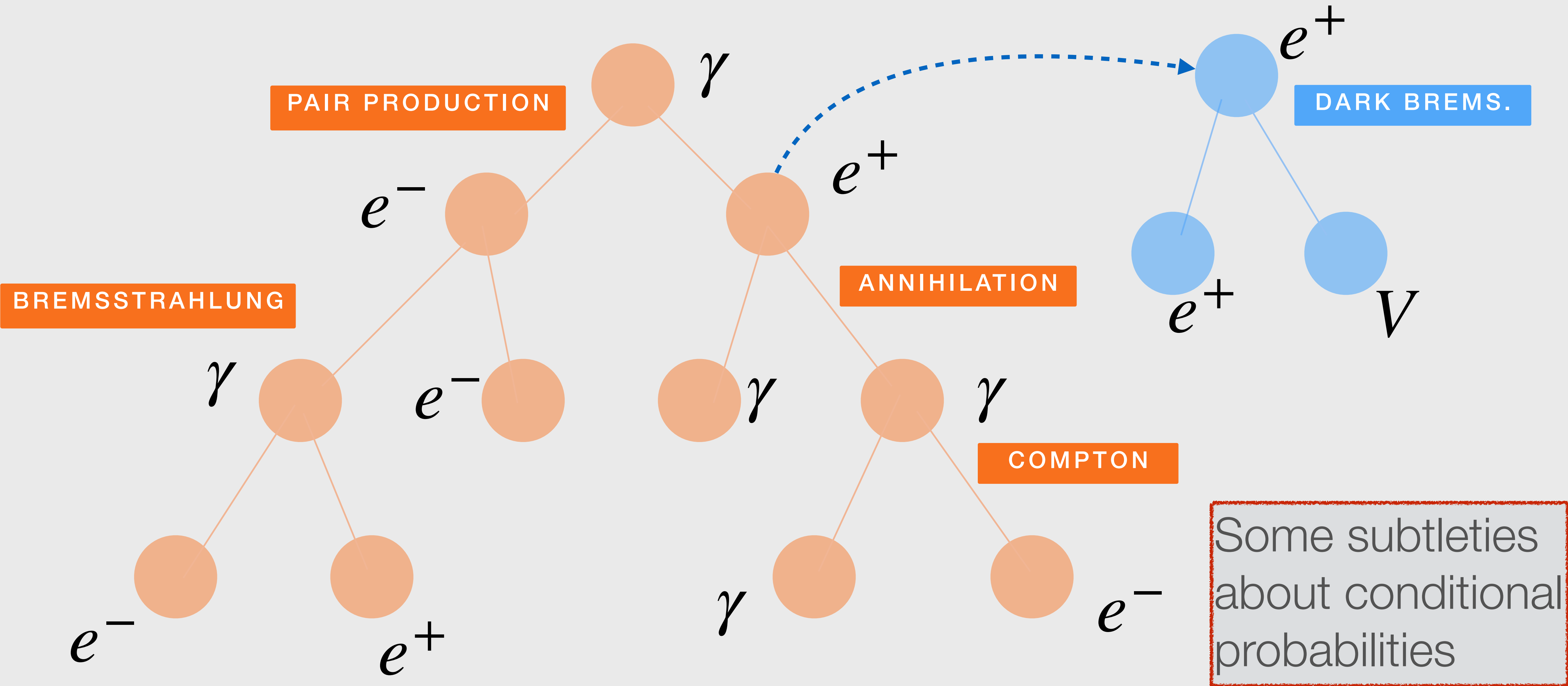
- $e^{\pm}Z \rightarrow e^{\pm}\gamma Z$
- $\gamma Z \rightarrow e^{+}e^{-}Z$
- $e^{\pm}e^{-} \rightarrow e^{\pm}e^{-}$
- $\gamma e \rightarrow \gamma e$
- $e^{+}e^{-} \rightarrow \gamma\gamma$

Continuous energy loss  
&  
Multiple Coulomb scattering.

# SM Event Record $\rightarrow$ BSM Event Record



# SM Event Record $\rightarrow$ BSM Event Record



# Implemented In PETITE

📖 README

## PETITE

PETITE: Package for Electromagnetic Transitions In Thick-target Environments Monte Carlo generator for production of dark sector objects in thick-target experiments PETITE generates electromagnetic showers for incoming electron, positron or photon propagating through a dense medium, and includes the possibility of dark sector particle production.

## Installation

To install, from the top directory run

```
pip install .
```

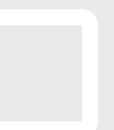
## Dependencies

PETITE, its tutorials and tools require the following packages: numpy 1.24, vegas ( $\geq 5.4.2$ ), cProfile, pickle, matplotlib, scipy, datetime, tqdm, copy, sys, random and functools. Using `pip install .` should install all requirements, but if needed, you can manually install these packages with

```
pip install <package_name>==<version_required>
```



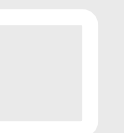
[HTTPS://GITHUB.COM/KJKELLYPHYS/PETITE](https://github.com/kjkellyphys/petite)

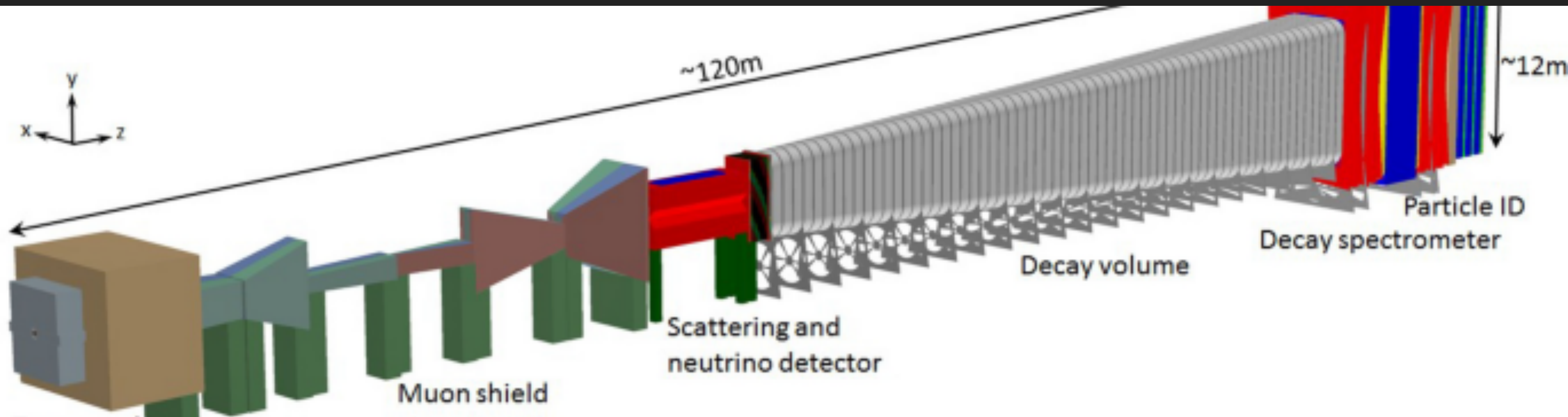


# Join A Growing Community!



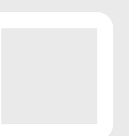
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# The Lifetime Frontier

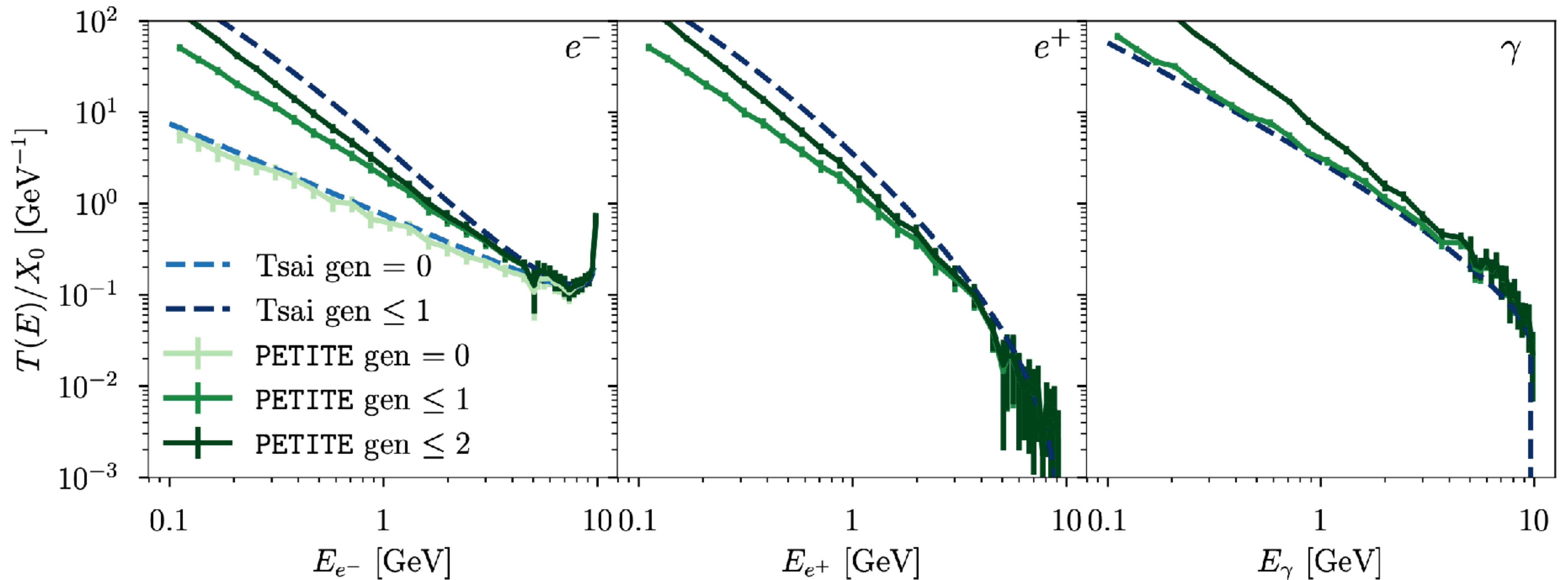
Lower Energies And Shorter Lifetimes





# Particle Multiplicity

PETITE Track Length vs Analytics (Bremsstrahlung and Pair Production Only)



# Lifetimes & Production Spectra

- Secondary, tertiary, etc. particles have lower energies.
- This means they have decay faster in the lab frame

$$\lambda = c\beta\gamma\tau \propto E_\nu$$

- In the long-lifetime limit the probability of decay goes like

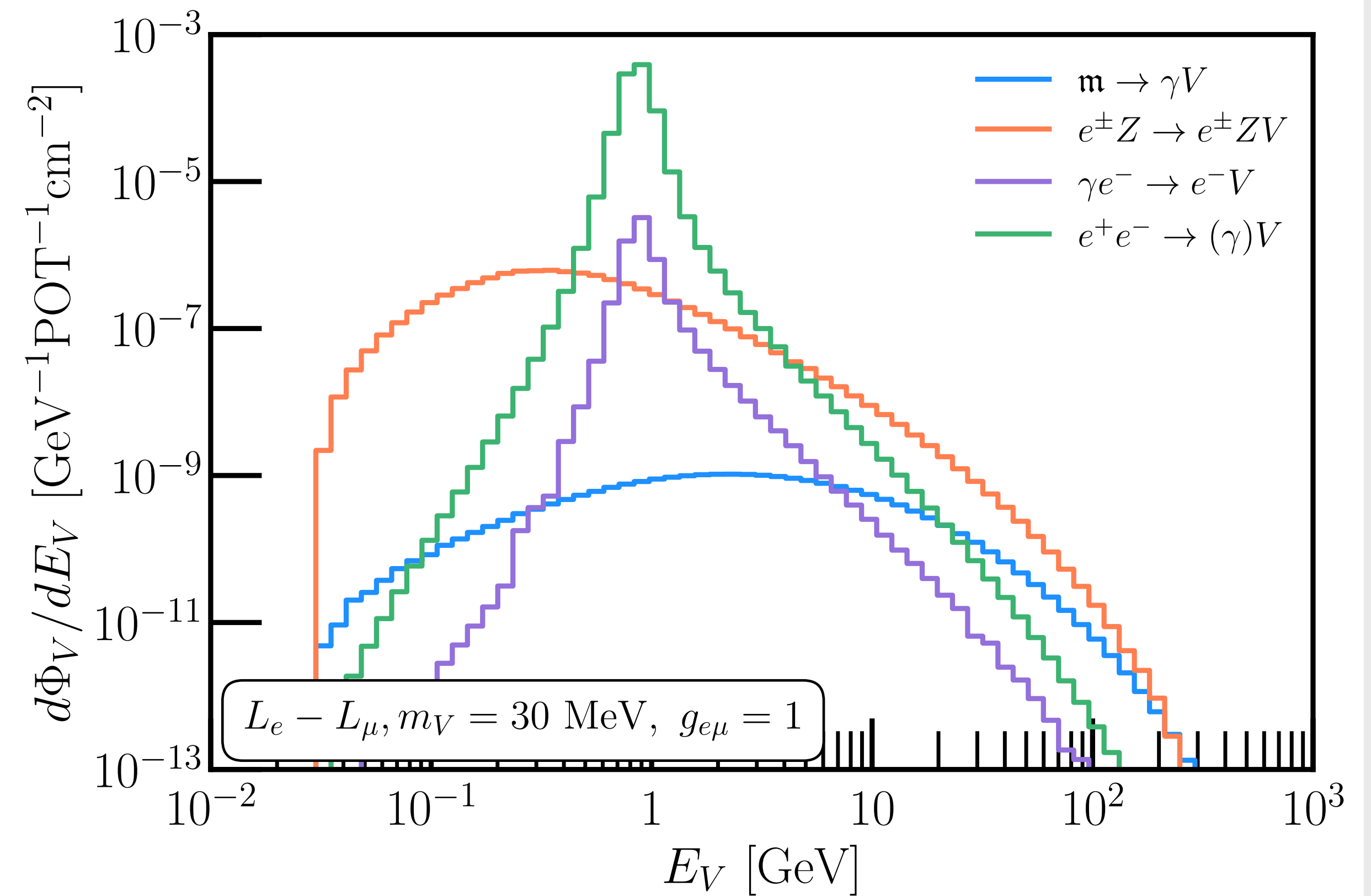
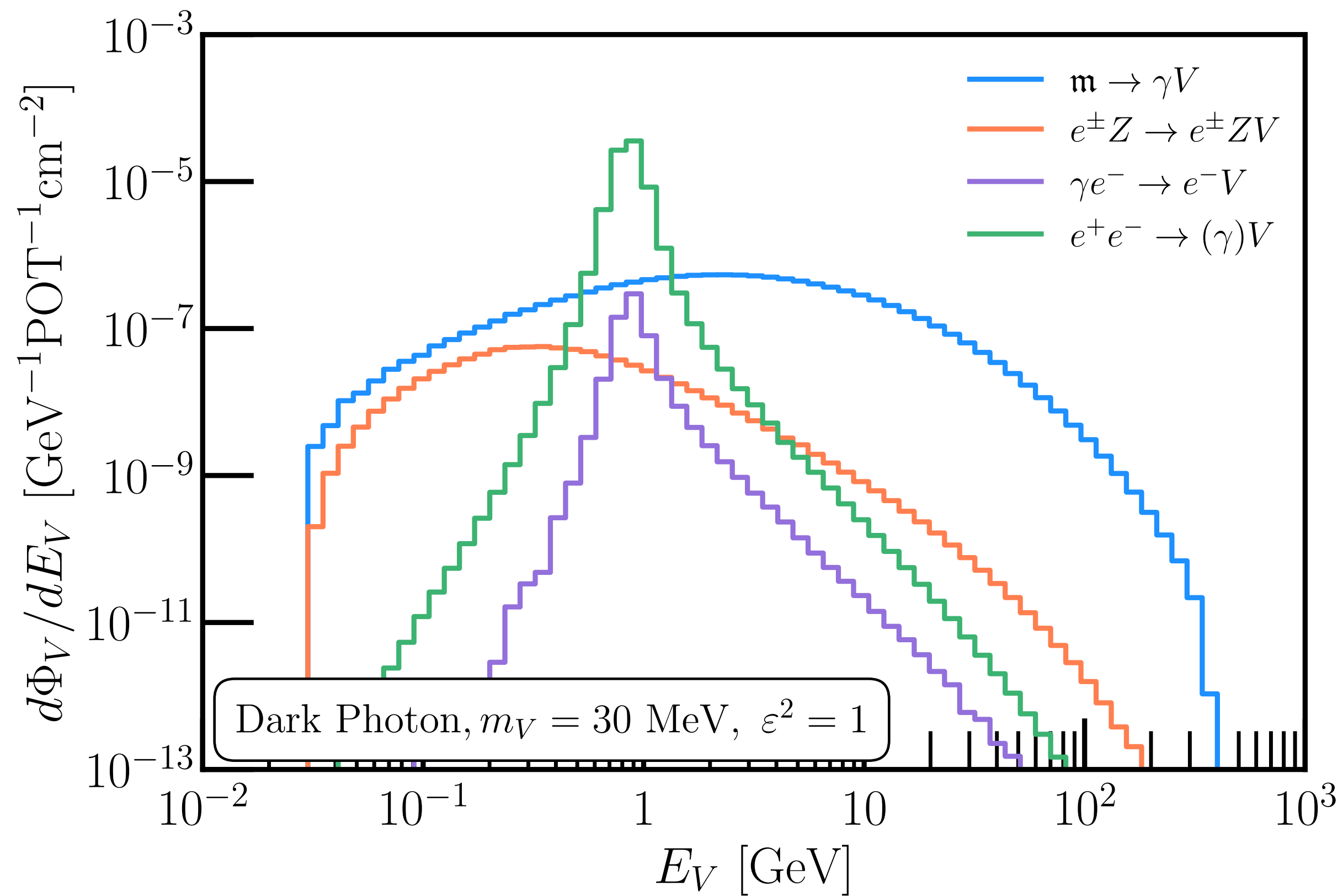
$$P_{\text{decay}} \simeq \frac{L_{\text{pipe}}}{\lambda}$$



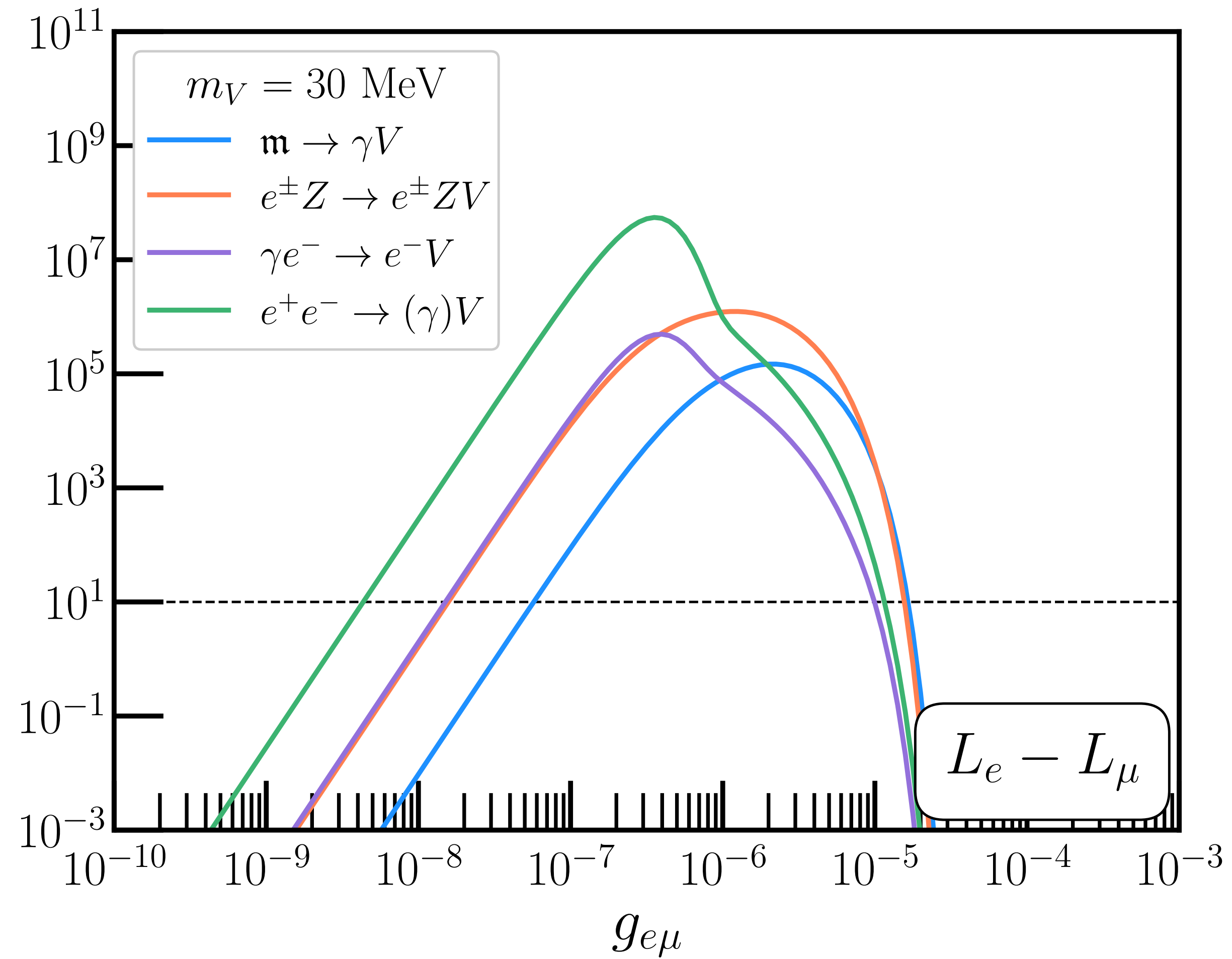
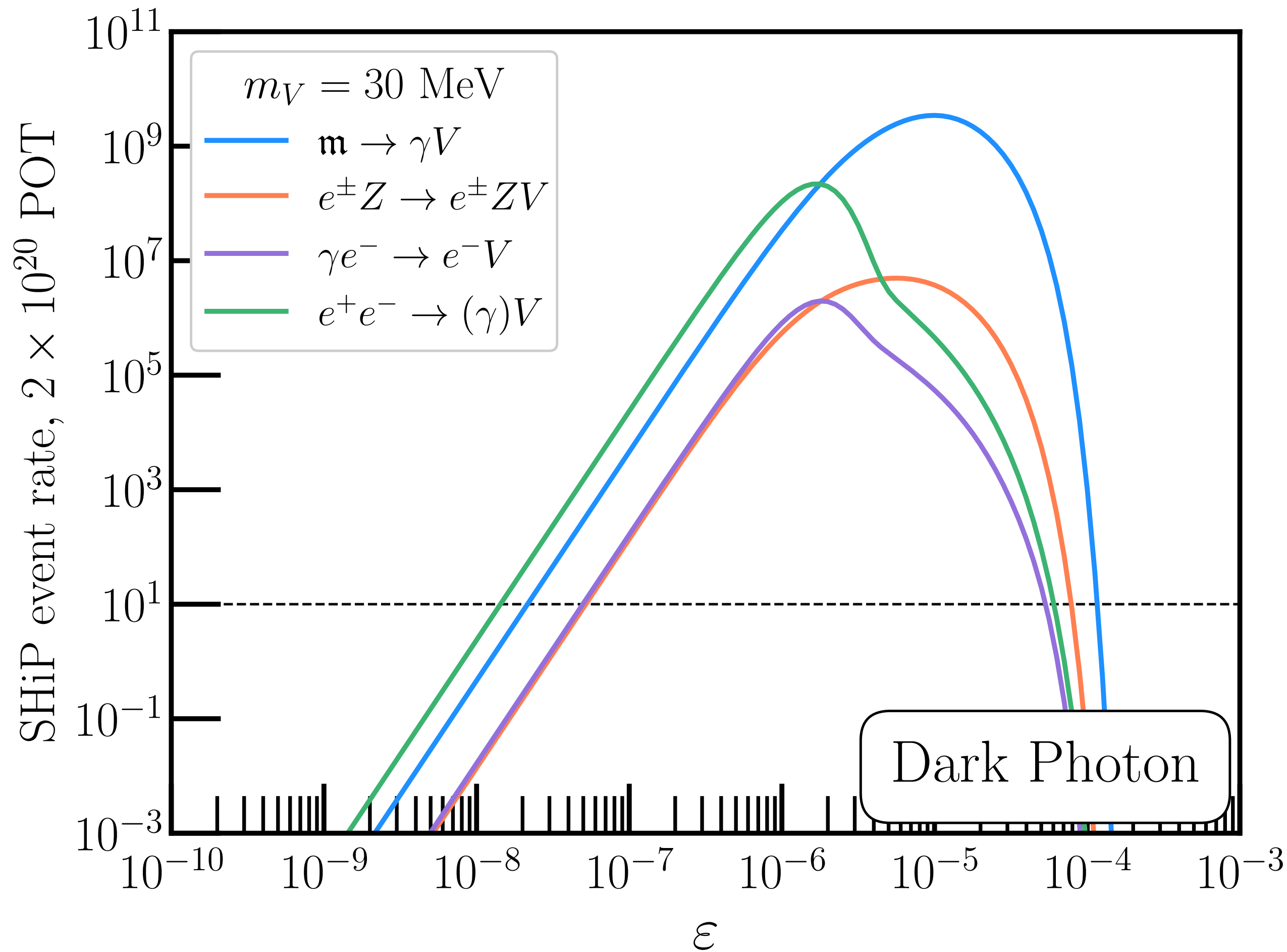
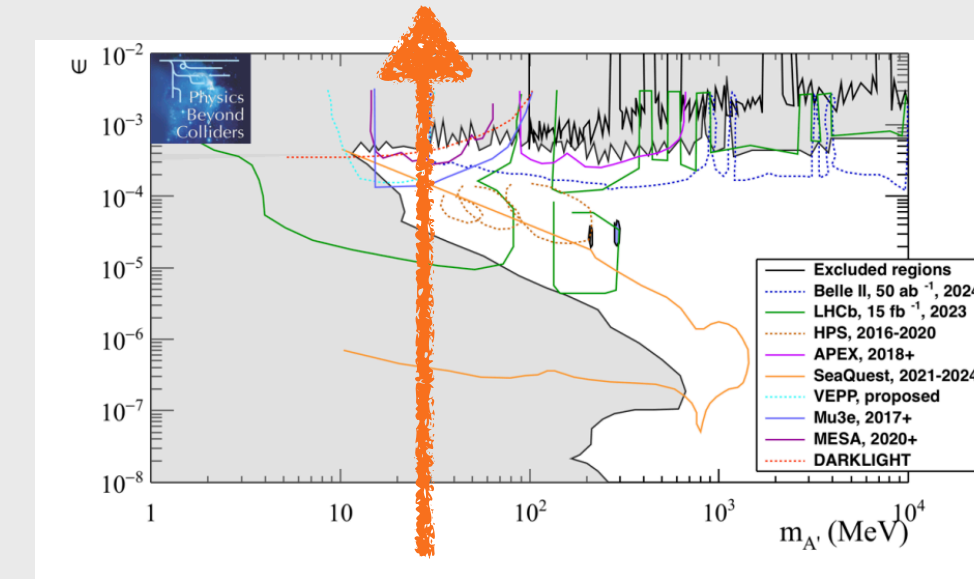
# Lifetimes & Production Spectra

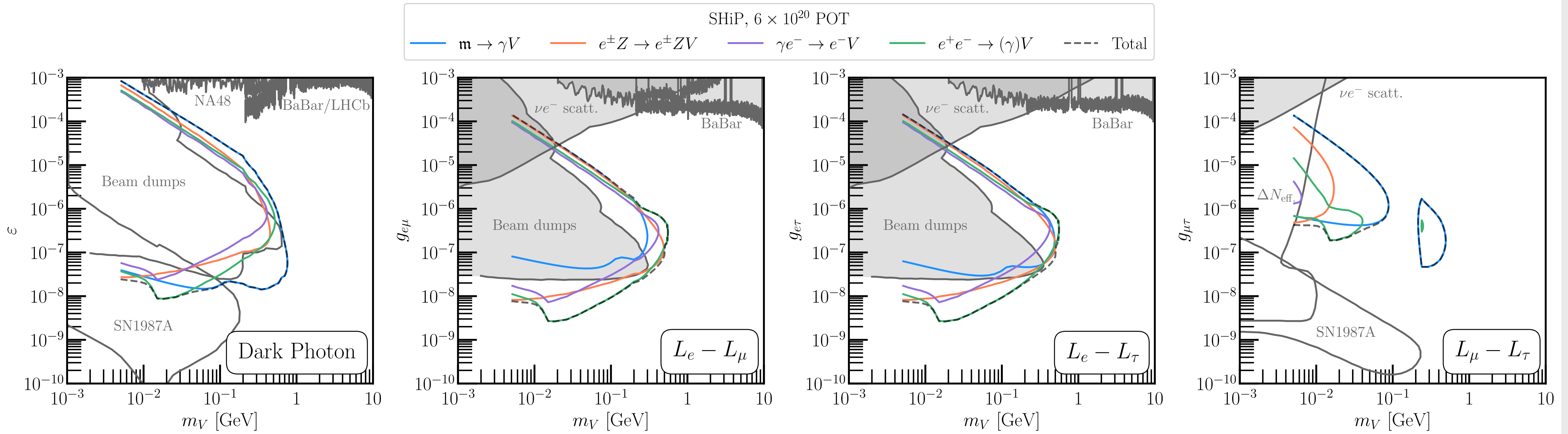
DARK PHOTON

ELECTROPHILLIC



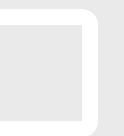
# Impact For LLP Searches





# Application To SHiP

Enhancements At The Lifetime Frontier



# Quick Intro To SHiP

PROTON BEAM

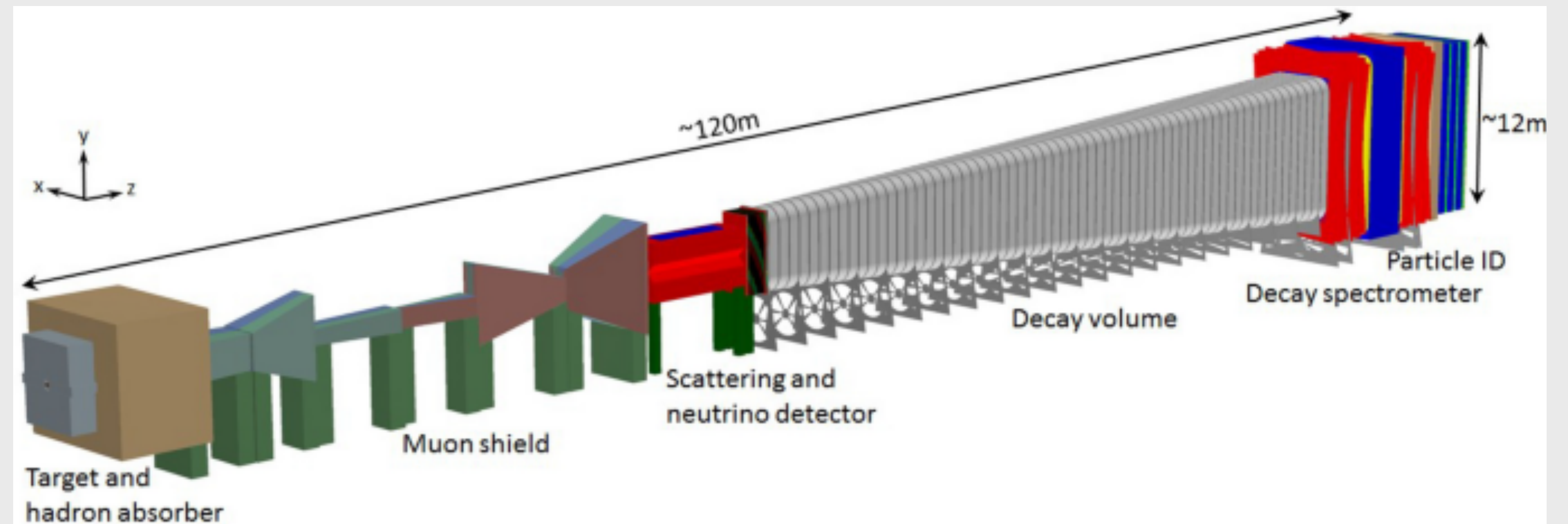
400 GeV

PROTONS ON TARGET

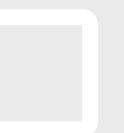
$6 \times 10^{20}$

DECAY PIPE

$\sim 30 - 80$  m



- Uses SPS beam at CERN.
- Excellent sensitivity to long-lived particles.
- Also has a scattering detector and tau-neutrino program.



# Dark Vector Production In A Proton Beam Dump

## PRIMARY MESONS

$$\pi^0 \rightarrow \gamma V$$

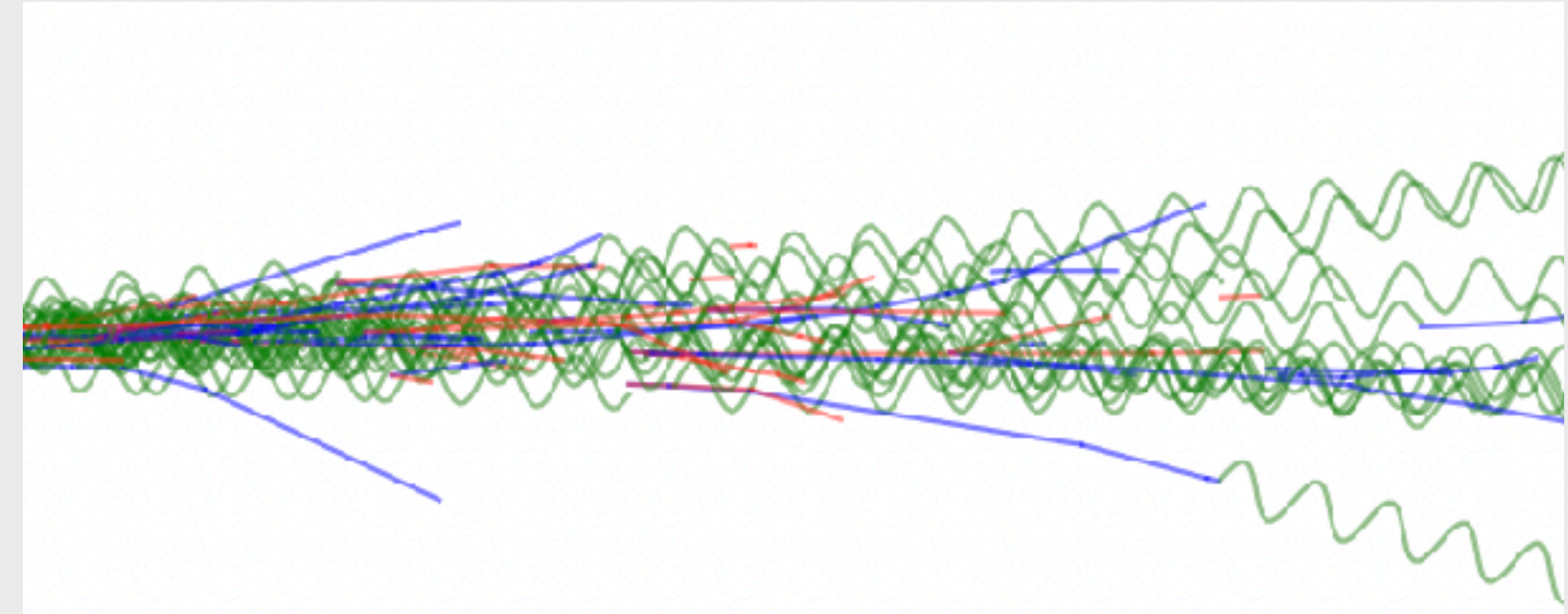
$$\eta \rightarrow \gamma V$$

## PROTON BREMS.

$$pA \rightarrow pAV$$

## PERTURBATIVE QCD

$$q\bar{q} \rightarrow V$$

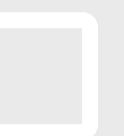


## CASCADE PRODUCTION

$$e^+e^- \rightarrow V(\gamma)$$

$$Ze^\pm \rightarrow Ze^\pm V$$

$$\gamma e^- \rightarrow e^- V$$

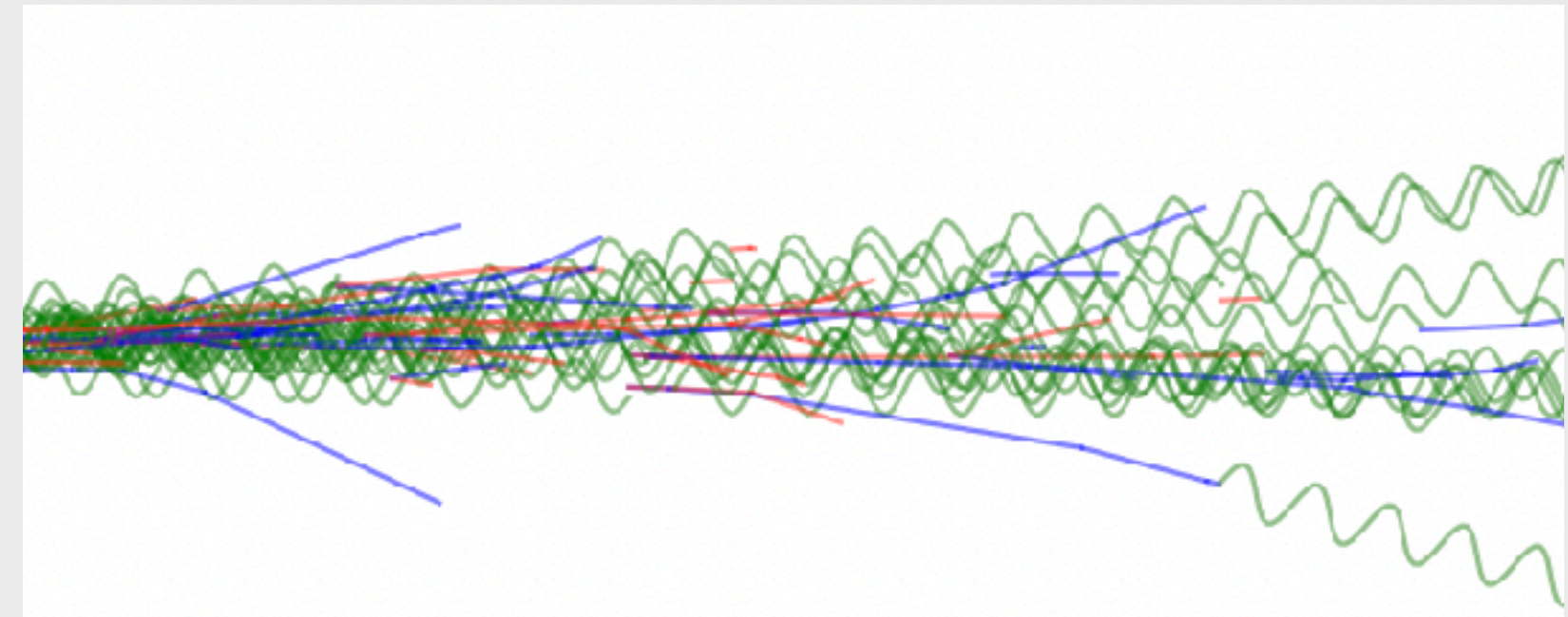


# Dark Vector Production In A Proton Beam Dump

## PRIMARY MESONS

$$\pi^0 \rightarrow \gamma V$$

$$\eta \rightarrow \gamma V$$



## CASCADE PRODUCTION

$$e^+e^- \rightarrow V(\gamma)$$

$$Ze^\pm \rightarrow Ze^\pm V$$

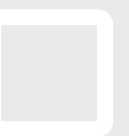
$$\gamma e^- \rightarrow e^- V$$

## PROTON BREMS.

$$pA \rightarrow pAV$$

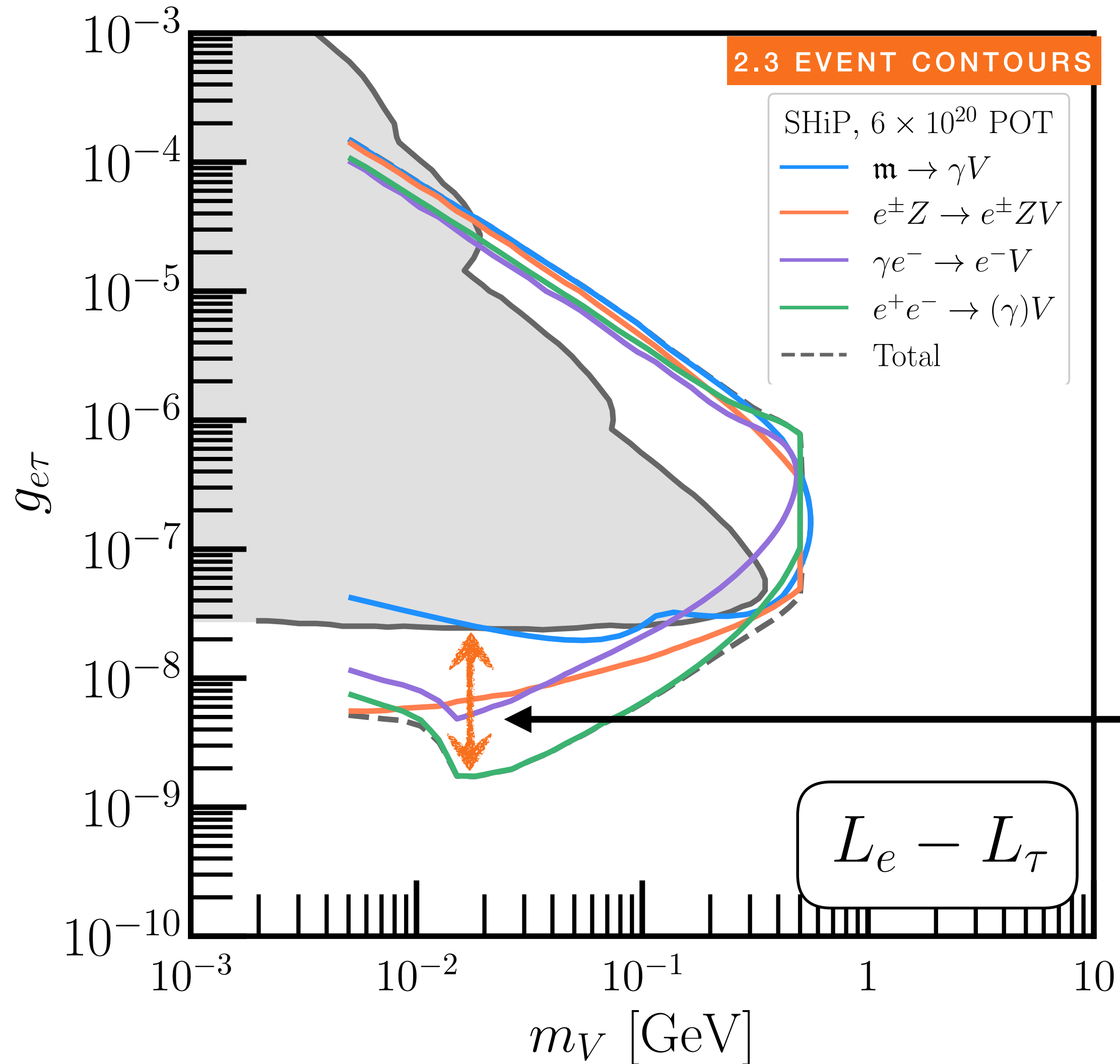
## PERTURBATIVE QCD

$$q\bar{q} \rightarrow V$$





# Gains In Sensitivity

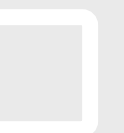


## Huge Shift In Flux

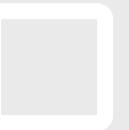
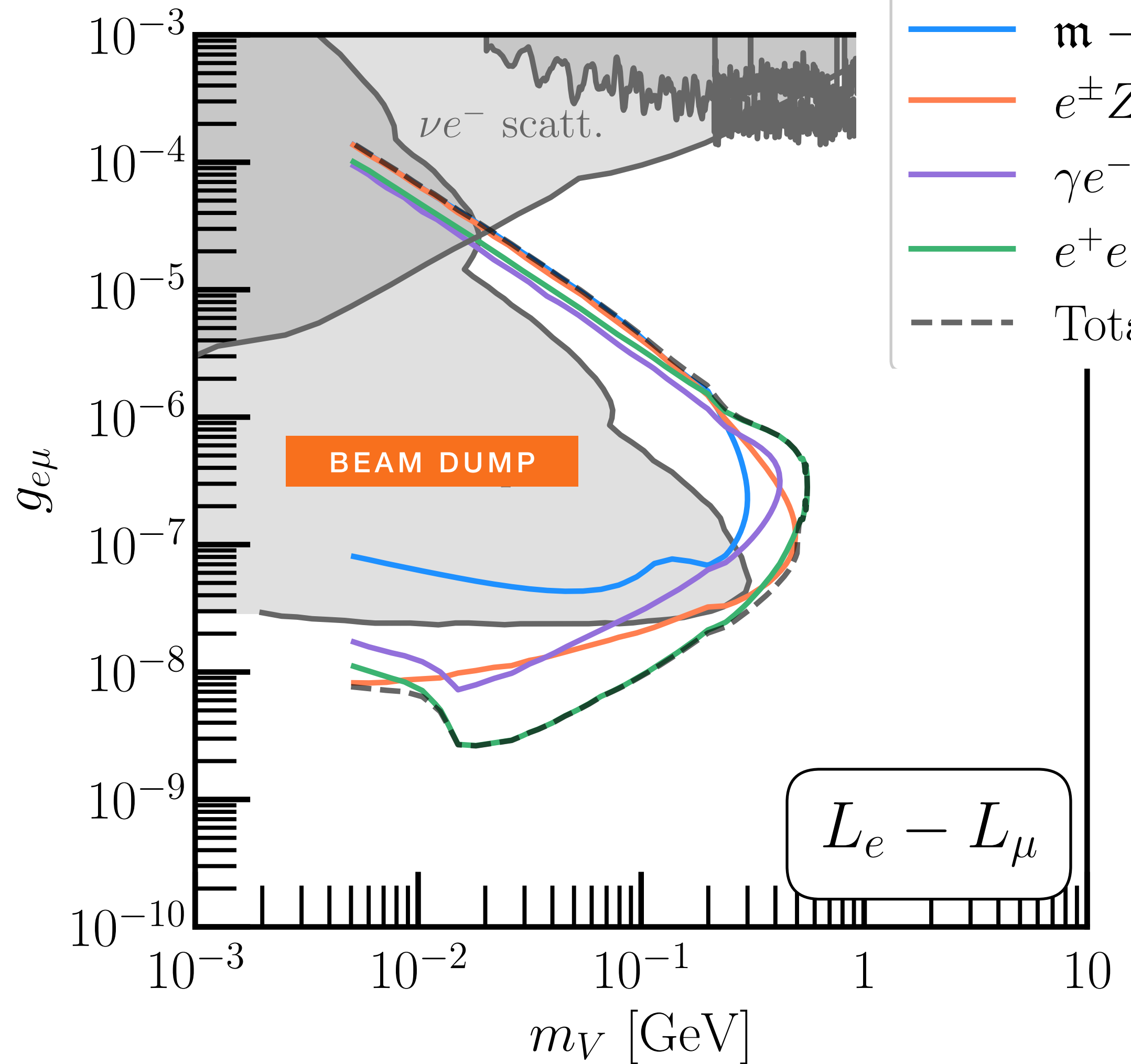
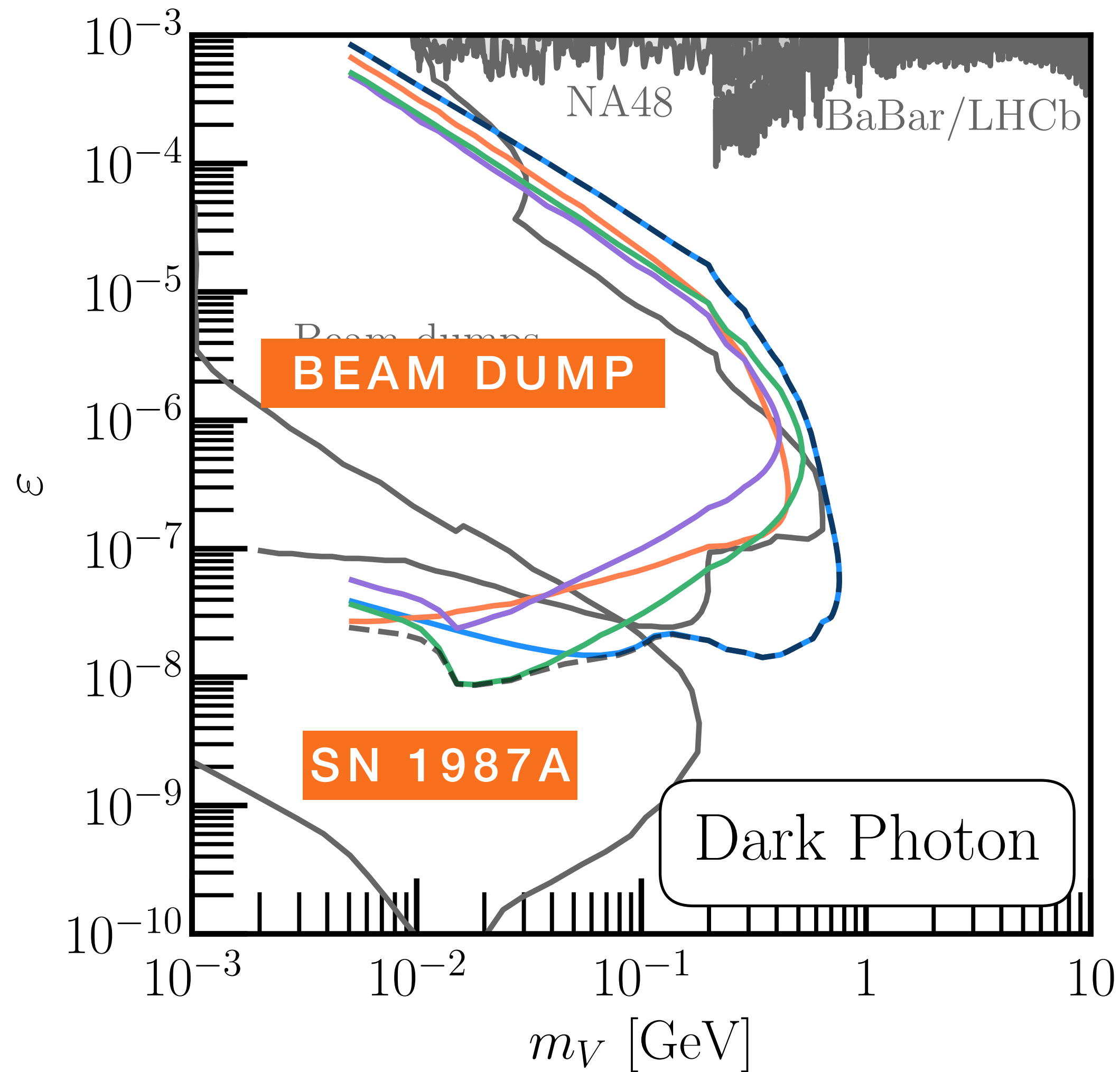
~ FACTOR OF 10 IN  $g_{e\tau}$

RATE  $\sim g_{e\tau}^4$

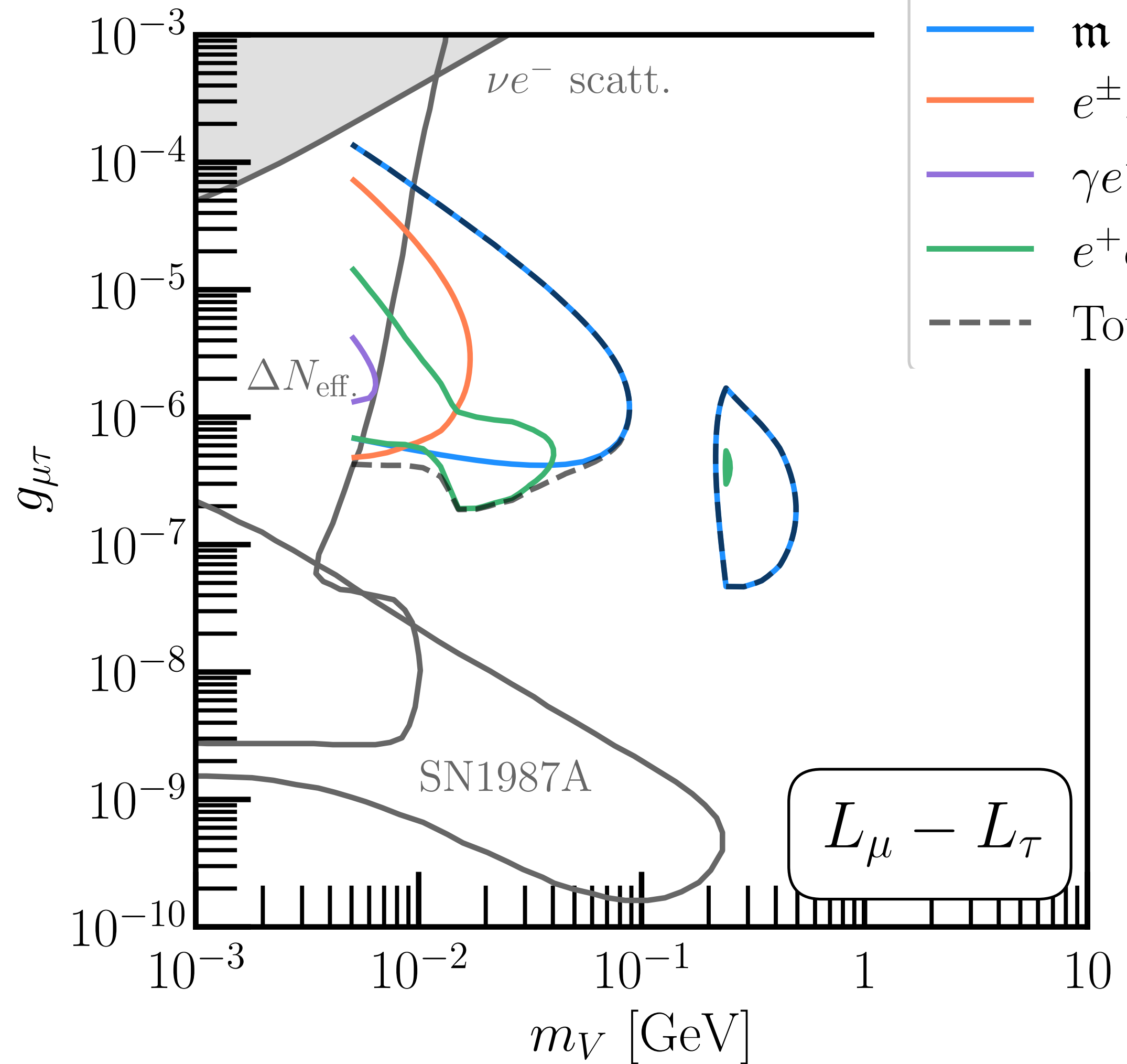
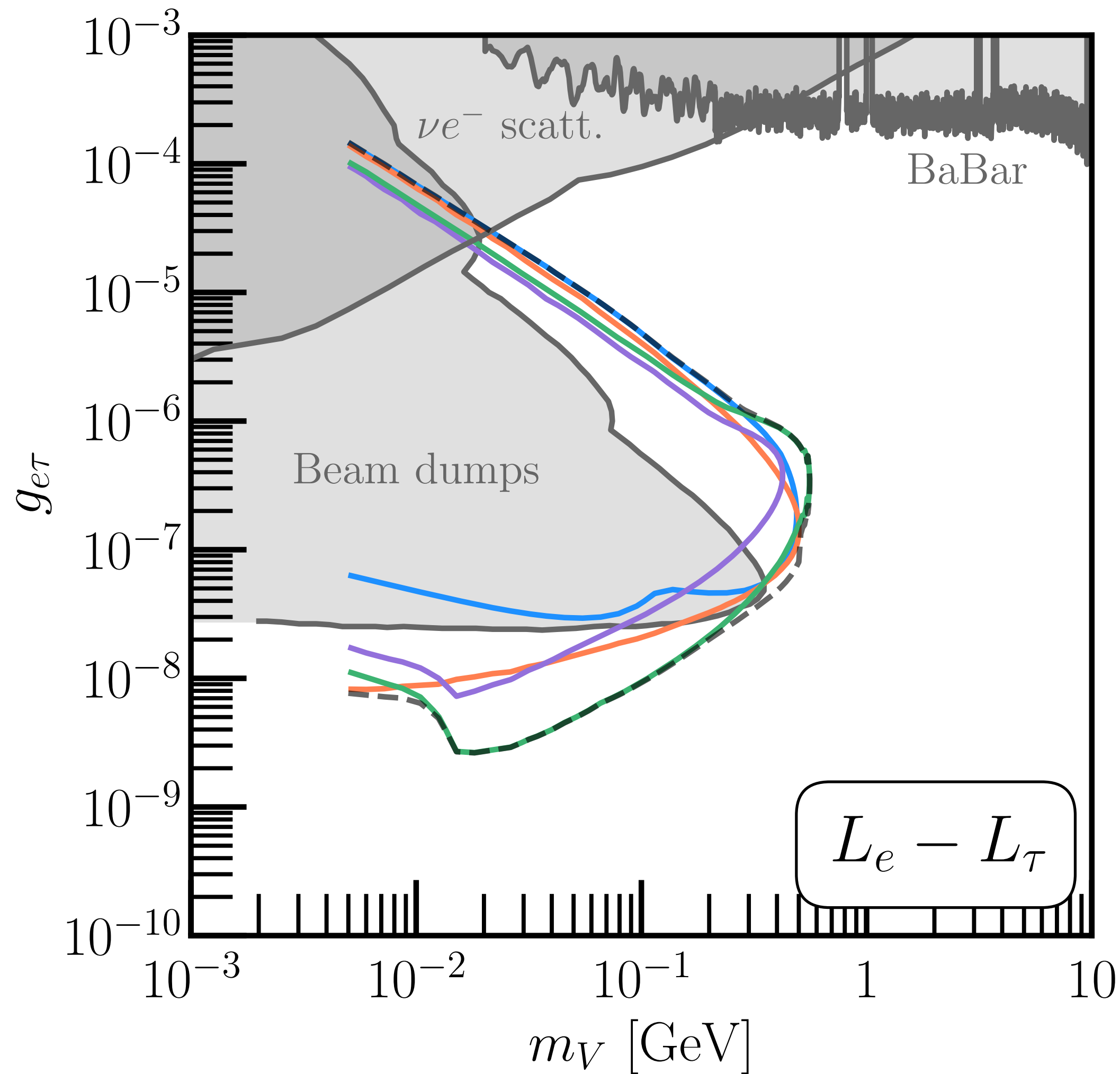
~ FACTOR OF  $10^4$  IN RATE!



# Enhanced Sensitivity At SHiP

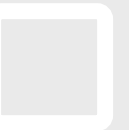


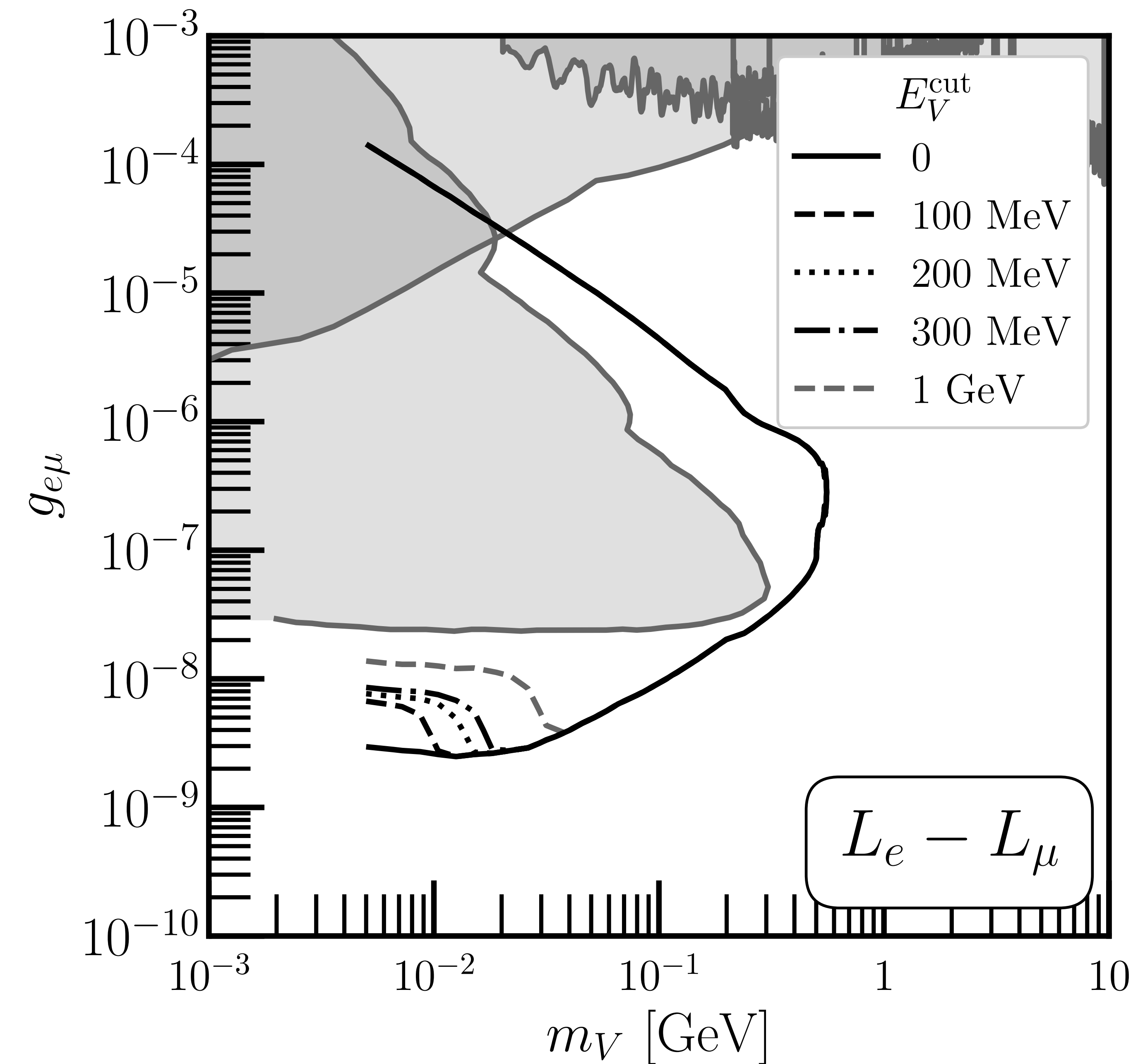
# Other Leptophilic Models



10 EVENT CONTOURS

- SHiP,  $6 \times 10^{20}$  POT
- $m \rightarrow \gamma V$
- $e^\pm Z \rightarrow e^\pm ZV$
- $\gamma e^- \rightarrow e^- V$
- $e^+ e^- \rightarrow (\gamma)V$
- - - Total



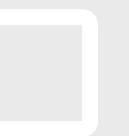


# Energy Threshold Dependence

$$e^+e^- \rightarrow V(\gamma)$$

$$E_{\text{res}} \simeq \frac{m_V^2}{2m_e}$$

- Motivates exploring lower thresholds at SHiP



A panoramic view of a city skyline, likely Vancouver, with a dense cluster of high-rise buildings in the foreground. In the background, a range of large, rugged mountains is partially covered in snow under a clear blue sky. The water of a bay or harbor is visible at the bottom of the frame.

# Conclusions & Outlook

# Conclusions

- SHiP can substantially extend its sensitivity to dark vectors by adding EM cascade production.
- There is an open source tool **PETITE** that can be used immediately to run simulations.
- Gains are per- $\pi^0 \implies$  robust against hadronic modelling.
- Biggest gains in sensitivity come at long-lifetimes.
- The smaller boost helps particles decay in the decay pipe.