Beyond the Standard Model Interactions at Coherent CAPTAIN Mills

Dark Interactions 2024 Vancouver

Edward Dunton On behalf of the CCM Collaboration

October 16, 2024



Managed by Triad National Security, LLC for the U.S. Department of Energy's NNSA

• LOS Alamos NATIONAL LABORATORY

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The Coherent CAPTAIN Mills (CCM) Experiment





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Vector Portal Light Dark Matter





- Assumes mediator is at the weak scale
- Strong constraints from cosmology, astrophysics, and colliders

- Does not assume mediator is the Z
 boson
- Simple parameterization of DM + Mediator:
 - 4 free parameters
 - Mass of the dark photon $m_{_V}$
 - Mass of the dark matter m_{γ}
 - Mixing angle between SM and dark sector ε
 - Coupling between dark photon and dark matter α_p



Produce Dark Sector in Standard Model particle cascades in fixed target experiments

MiniBooNE beam dump mode first dedicated DM search in this way

Phenomenology: Vector Portal Light Dark Matter



Dark Sector Explanations of the MiniBooNE Anomaly



- 4.8σ excess at MiniBooNE target mode runs
- No excess in dump mode run
- If excess is due to new long-lived particles (LLPs) or light dark matter (LDM), it may be correlated to rare charged meson decays [*PRL* 129 (2022) 11, 111803]
- We can test this possibility at CCM



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Phenomenology: Dark Sector Coupling to Meson Decays (DSCMD)



	Scalar Mediator	Pseudoscalar Mediator	SM Pi0 Mediator	Vector Mediator	
Scalar IB1 (e)				 ✓ 	
Scalar IB1 (mu)		cion-4			
Pseudoscalar IB1 (e)	1	jimens			
Pseudoscalar IB1 (mu)				~	
Vector IB1 (e)	~	1	~	ns	
Vector IB1 (mu)	~	1	~	alo	
Vector IB2 (e+mu)	~	V (om	
Vector contact (e+mu)	(\checkmark)		-	Δn	

Primakoff / Photoconversion Scattering Model

- Wide variety of models combining decay model and scattering model.
 - So far limited focus, but scope is total.
- CCM (stopped pion) tests in a complementary way to MiniBooNE (focused pion) and similar experiments.

...and many more models!

- Axion-Like Particles (ALPs)
- Heavy Neutral Leptons (HNL)
- ATOMKI Anomaly
- CEvNS neutrino search
- v_eCCQE cross section measurements
- Mirror neutrons
- Etc...



Nuclear Instruments and Methods in Physics Research A 594 (2008) 373-381 Nuclear Instruments and Methods in Physics Research A 632 (2011) 101-108

 $\approx 26 \, \mathrm{nse}$

APTURI

Intense source muon neutrinos: target MCNP simulation flux 4.74x10⁵ v/cm²/s at 20 m

/DS



-Run detector in multiple locations.

The CCM Experiment:

LANSCE-Lujan Facility

-Room to deploy shielding, large overhead crane, power, etc

CCM120 beam and calibration run 2019 (L=20 m)



800 MeV

The CCM120 Detector: 2019 run

- Mounted in the 10 ton CAPTAIN cryostat.
- 7 tons LAr Fiducial volume, 3 tons LAr Veto (2-3 radiation lengths)
- 17 1" and 5 8" Veto PMTs instrument LAr Veto region.
- 120 R5912 PMT's, wavelength shifting TPB foils
 - LAr cold test of the entire SBND PDS system: 96 TPB coated + 24 uncoated PMT's, mounts, cables, feed thrus, HV, electronics, trigger, DAQ, calibration, simulations and data analysis. -> now installed in SBND



TPB coated PMTs Uncoated PMTs

TPB coated reflector foils. Maximize light output to detect coherent neutrino-nucleus scattering

6-week Engineering run in 2019 collected 1.79 × 10²¹ Protons on Target (POT).

The CCM200 Detector: 2021 to 2025 runs.

- Instrument all 120 cylinder 8" PMTs with new PMTs
- + Instrument 40 8" PMTs on each end cap => total **200 new 8" PMT's**.
- 40 1x1" vetos instrument outer region (double veto PMTs)
- Evaporative TPB foils (double efficiency of CCM120 foils) produced at U. Edinburgh.
- CCM200 detector built July 2021, initial test run done October-December 2021 collected 1.9 × 10²¹ POT.







The CCM Detector: Using fast timing and shielding to remove beam-related neutrons



- Extensive 4m steel shielding around target and detector slows down neutrons
- EJ-301 n/ γ detector time of flight measurements indicate fastest neutrons ~20 MeV.
- Analysis of CCM120 data shows 210 nsec signal region free of neutrons.



- Laser/Diffuser for 213/532 nm calibrations to test TPB response for foils and PMTs and the LAr properties.
- LED calibrations for PMT gain/timing
- Co-57 source provide energy scale calibration 122 keV gamma-ray.
- Na-22 source provide energy scale calibration 2.2 MeV gamma-rays.
- Radioactive sources provides position reconstruction calibration.

CCM=Optical Detector

Light Properties need to be very well known.

CCM120 Searches: Energy Regimes

- 2 Types of Searches: Electromagnetic and Nuclear Recoil
- 3 Energy Regimes for Searches:
 - < 1 MeV Light Dark Matter (nr like)
 - Threshold 3 PE ≈ 200 keV up to Maximum 18 PE ≈ 2 MeV.



A. A. Aguilar-Arevalo et al. "First dark matter search results from Coherent CAPTAIN-Mills". In: *Physical Review D* 106.1 (July 2022).

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 - From minimum 10 PE ≈ 500 keV up to Maximum 200 PE ≈ 10 MeV.



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 - > 10 MeV Dark Sector (em like)
 - From 200 PE ≈ 10 MeV.
- CCM200 uses similar energy regimes, with different PE scales and a lower threshold.



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CCM200 2021 Searches: Energy Scale

- The 2021 Data set had a very different energy scale from 2019.
- Using the Na22 calibrations at the center of the detector, we found a 3:1 improvement in energy scale for CCM200.
- We also mapped out various positions inside the detector along the central axis, giving us some information on the positiondependent smearing.
- Using a 40ns prompt integral, we found ~30 PE/MeV.



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CCM120 Analysis: Beam Related Background Free Region of Interest (ROI)

- Timing of massive importance to CCM efficiency.
- We define a ROI that includes only steady state (measurable in prebeam!) backgrounds.
- Directly measured T0 with the "FP3 detector."
- **T0 = -530 ns** DAQ time
 - Defines start of ROI for all searches.
- Determined speed of light particles from target to arrive 150 ns before neutron events seen in CCM



CCM200 2021 Analysis: Region of Interest (ROI)

- The start of the 2021 ROI was the same as in 2019: -530 ns.
- The end, however, was delayed an additional 40-60 ns, depending on the T_0.
 - Determined by a significant change in cut efficiency and event rate (seen below)
- 2021 had a final ROI of 210 ns (representing 84% of prompt BSM signal), versus 150 ns in 2019 (<50% of prompt BSM).



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All Searches: Data Quality / Pre-selection Cuts

- Data Quality cuts meant to ensure the events were useable.
 - Number of Hits
 - Previous Event
 - Veto
 - Fiducial Volume
 - Event Time
 - Beam Current Monitor (BCM)



(a) DAQ Time

Data Quality Efficiency Table

Cut	numHits	previousEvent	Veto	fiducialVolume	eventTime
Efficiency	0.510	0.238	0.236	0.149	0.0037

CCM NR Search: Event Characteristics

- In CCM, LDM events will involve coherent-like scattering of Argon nuclei
- Nuclear-recoil type events,
 - No triplet light
 - Additional nuclear recoil quenching factor of ~0.25.

fron

- => Short in length and low in visible energy.
- Cuts:
 - Length < 44ns: a cut on the maximum length of the event (PID).
 - Niceness/Charge Uniformity: cut on charge distribution evenness between PMTs.



CCM120 E/M Search: Signal Cuts

 In addition to the Data quality cuts from the LDM search (Previous Event, numHits, Veto, Niceness) we added 4 cuts for the E/M search based on signal characteristics. **Cut Efficiencies after Preselection Cuts**



- Length > 38ns: Electron like ALP events have triplet light and long length.
- Radius < 80cm: A stricter position cut due to the isotropic nature of the signal.
- **Coated > 60%:** More than 60% of the light seen is seen by TPB coated PMTs in scintillation events.
- maxPMT < 20%: No more than 20% of the total light is seen by a single PMT in simulated e/m like events or calibration source events..

CCM200 E/M Search: Signal Cuts

- The 2021 Analysis used the same data quality cuts as in 2019.
- However, the event selection cuts were slightly different.



- ----- Length > 54 ns: increased from 2019 but otherwise unchanged.
- ----- Radius < 80 cm: repeat from 2019.
- ----- maxPMT < 20%: repeat from 2019.
 - Cherenkov-light > 2%: A new cut searching for potential Cherenkov light in the uncoated tubes in the first 10 ns of the event.
 - Prop20>10%: A new cut removing events with too slow of a turn on, characteristic of neutron multiple scattering.



 10^{5}

 10^{4}

 10^{3}

 10^{2}

101

 10^{0}

 10^{-1}

 10^{-2}

 10^{-3}

100

 $\cdot \mathrm{yr}^{-1}$

Differential Rate $[cts \cdot MeV^{-1}]$



Dark Sector Coupling to Meson Decays (DSCMD): 2-Mediator Models

DSCMD fits to MiniBooNE data for

 $\pi^0 \to \gamma X \ (m_X = 5 \text{ MeV})$

 10^{1}

- Target Mode: Neutrino + Antineutrino excess
- Dump Mode: null result as a constraint
- CCM tests this model using neutral pion couplings •



CCM120 2019 vs. CCM200 2021 Subtractions



- There is one major deviation from the background prediction in both data sets.
- This excess manifests in the 20-50 MeV region, and is consistent in rate even after accounting for efficiency and background differences.

CCM200 2021 Analysis: Conclusions

- The consistent anomaly in the 2 runs so far directs CCM's future plans.
- We are currently operating under the hypothesis that this anomaly is caused by very rare, very high energy neutrons.
- 3 possible sources:
 - Direct from target
 - Bounce from other flight paths
 - From bend in beamline
- 2022 (other flight paths closed) and plan to move in 2025 address these possibilities.
- If not neutrons, we plan a larger detector (CCM1000) for better PID and signal identification.



Acknowledgements







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Physics & Astronomy



Backup Deck

Outstanding Physics Puzzles The Dark Matter Puzzle "Portals" **Standard Model Dark Sector** $\mathbf{X} \wedge \wedge \wedge$ photon portal light dark matter Strong CP Problem d Down Charm S Strange g sterile neutrinos b u neutrino portal gluon photon Up Bottom н Higgs hoson VT τ W Z Tau Tau The Neutrino Mass W boson Z boson neutrino heavy neutral leptons μ Vµ Puzzle Higgs portal Muon Muon Ve neutrino е Electron Electron light mediators neutring QUARKS GAUGE BOSONS Anomalies: ALP portal MiniBooNE, Short Baseline...

The QCD Axion and Axion-Like Particles (ALPs)

• Axions are a proposed solution to the Strong-CP problem in QCD

 \rightarrow <u>Why is the neutron dipole moment so small?</u> $d_n < 10^{-26}$ e cm

- Related to *CP* violating term in QCD: $\mathcal{L} \supset \theta GG$
- $U(1)_{PQ}$ broken $\rightarrow aGG$ term to dynamically conserve CP
- But there are many other theories that predict pNGB pseudoscalars:
 - generic goldstones of broken global $U(1)_{x}$
 - String *axiverse*
 - Non-traditional QCD axions
 - Axion dark matter

CCM is sensitive to ALPs with masses up to ~100 MeV







Phenomenology: ALP Detection in CCM





 e^+

N

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DSCMD and MiniBooNE, MicroBooNE



Pion Decay Constraints versus DSCMD

- CCM and MiniBooNE much more sensitive to rare pion decays than normal searches
 - Many more pions produced.
 - Increased detector distance.
 - Appearance versus disappearance experiment.
- Neutrino search techniques highly sensitive to rare pion decay events!

PHYSICAL REVIEW LETTERS 129, 111803 (2022)

TABLE II. Relevant exotic decays of π^{\pm}/K^{\pm} and existing upper limits at 90% confidence level X stands for invisibly decaying (massive) bosons. The predicted BRs (third though last columns) are based on the following parameter choices: $[\epsilon_1, (g_1^2/4\pi)] \simeq (6.0 \times 10^{-5}, 1)$ for the single-mediator scenario, $[\epsilon_1, \epsilon_2, (g_2^2/4\pi)] \simeq (7.0 \times 10^{-5}, 1.0 \times 10^{-4}, 0.5)$ for the double-mediator scenario, $(g_{\mu}, g_n, \lambda) \simeq (5 \times 10^{-3}, 10^{-2}, 4.4 \times 10^{-4} \text{ MeV}^{-1})$ for the scalar scenario, and $(g_{\mu}, g_n, \lambda) \simeq (10^{-2}, 10^{-2}, 6.5 \times 10^{-3} \text{ MeV}^{-1})$ for the pseudoscalar scenario.

Channel (BR)	Limit (×10 ⁻⁸)	Model (i) (×10 ⁻¹²)		Model (ii) (×10 ⁻⁸)	
		Single	Double	ф	а
$K \rightarrow \mu \nu_{\mu} V(\phi)$ [87]	2000 (300)	500	680	230	100
$K \rightarrow e \nu_e \nu \nu$ [54]	6000	530	720		
$K \rightarrow \mu(e) \nu_{\mu(e)} ee$ [54]	7.4(2.7)	500(530)	680(720)		
$\pi \to \mu(e) \nu_{\mu(e)} X$ [88]	600(50)	0.12(25)	0.17(34)	120()	1.1(···)
$\pi \rightarrow \mu(e) \nu_{\mu(e)} ee$ [54]	-(0.37)	0.12(25)	0.17(34)		

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The CCM Experiment: The Lujan Target and Source

Extremely well understood and modeled by

Alocear Instruments and Methods in Physics Research A 594 (2008) 373–381 Nuclear Instruments and Methods in Physics Research A 632 (2011) 101–108

Complex target tuned for neutron production

Beamlines controlled by Hg window



Fig. 1. Elevation view of the Lujan Center's TMRS geometry used in our calculations. The main components are labeled: split tungsten target (W), beryllium reflector (Be), lead reflector–shield (Pb), and the steel reflector–shield (Fe).



Mam Steel Shielding - 2.5 m Hg beam windows

- Extensive shielding around target
- Simulations has confirmed hand calculated neutrino flux of ~4.74x10⁵ nu/cm²/s at 20 m
- MCNP simulation of target and ambient neutron ^{20 m Background Neutron Flux at 100 µA}



The CCM Experiment: LANSCE-Lujan Beam



 Neutrino/DM experiments require high Instantaneous Power – measure of Signal/Background: SNS (FWHM 350 nsec @ 60 Hz)= 0.060 kJ/nsec; Lujan(FWHM 138 nsec @ 20 Hz)= 0.031 kJ/nsec

Trigger by Trigger measurement of the Beam through Integrated Beam Current Monitor

- CCM includes direct measurements of the beam proton pulse through a BCM integrated directly into the DAQ and data stream.
- This BCM can be used to analyze the protons per trigger, and ensure that only the best quality triggers are included in the final analysis.



CCM120 Flow of Data **CAEN V1730** ٠ 500 MHz clock **16 Channels** Warm cable to 2¹⁶ Bit ADC Chip supply power per channel and read 2 V range signal 11 boards Decoupler box to split power and signal cables Cold cable to supply power and read Fiber Optic connection to DAQ signal computer Digitizer boards

The CCM200 Detector: Hamamatsu R5912 PMTs

- **160 PMTs** evaporatively **coated** with Tetraphenyl Butadiene (TPB)
 - Overlap between PMT detection spectrum and TPB emission spectrum
- 40 PMTs uncoated
 - Helpful for calibration purposes
- 40 Veto PMTs
 - 8 1" PMTs on the bottom looking in
 - 10 1" PMTs on the top looking in
 - 22 1" PMTs on the columns looking up and down.



- Above: TPB-coated R5912 PMT
- Top Right: Hamamatsu Efficiency Spectrum
- Bottom Right: TPB Emission Spectrum



Scintillation/WLS mechanics

- Excited argon forms a dimer, de-excites releasing a **fast singlet** and **slower triplet** excitation components
- Scintillation light spectrum peaks at **128nm**
- Emits 4*10⁴ photons/MeV of deposited energy from scintillation
- Timing of prompt and delayed light components have been measured by the DEAP collaboration (right)
- Tetraphenyl Butadiene (TPB) wavelength shifts 128nm LAr scintillation light to the visible spectrum, allows better absorption by PMTs
- Combination Coated+Uncoated tubes allows separate detection of scintillation and Cherenkov light.



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The CCM Detector: Optical LAr Detector advantages

- Optical Detectors have response times at the 2 ns level
 - LArTPCs have millisecond readouts
 - When trying to use timing-based background rejection, need FAST detector
- LAr scintillation produces 40,000 photons/MeV
 - More than any organic scintillator
 - 40 photons/keV = theoretical threshold of 10 keV at 10% photon detection level
 - LArTPCs 3-5mm wire spacing reconstructs events > MeV
- LAr does not re-absorb Cherenkov light
 - Cherenkov can separate low energy electromagnetic events from nuclear recoil
 - Cherenkov can also provide directionality for e/m like events.



Simulated hits from Cherenkov light in CCM200 detector. Visible Cherenkov light precedes scintillation by up to 8 ns.

Data Selection: CCM120 Energy Scale from Radioactive Sources

- Impurities from not recirculating or filtering the argon led to low light levels.
 ~2 ppm O₂ reduced the 128 nm light attenuation length from ~20 m to ~50 cm
- According to simulations the 4.7 PE peak for Co57 is an artifact of the event cuts, the real peak is ~1.8 PE
- Na22 33.2 ± 8.9 PE for 2.2 MeV
- Both Co57 and Na22 rates are within 25% of simulation prediction
- Nuclear Recoil (neutron, LDM, CEvNS) energy scale approx. ½ of electron event (Na22, ALP) scale.



A. A. Aguilar-Arevalo et al. "First dark matter search results from Coherent CAPTAIN-Mills". In: *Physical Review D* 106.1 (July 2022).

Contamination



tion electron charge and scintillation light". In: Journal of Instrumentation 5.05 (2010), P05003-P05003.

- CCM observed LAr absorption lengths at 128 nm of 55 cm, . along with an additional quenching factor of 55%.
- Both Values are consistent with oxygen contamination just • below the 2ppm level.

Antígona Segura et al. "Biosignatures from Earth-Like Planets Around M Dwarfs". In: Astrobiology 5 (Jan. 2006), pp. 706-25.



CCM200 Shielding Layout at Lujan (23 m from target)

More steel and concrete shielding added. Roof shielding added in October 2021.

Filter skid (MicroBooNE design): 4A molecular sieve material to remove water contamination and Cu Alumina to remove oxygen contamination. Ten ton LAr recirculation turn over time of ~three hours.

ALP Event Characteristics

- In CCM, we expect ALPs to produce high energy photon or electron scatterings.
- Electromagnetic-like signature
 - Triplet light 75% of total scintillation
 output
 - Quenching Factor of 1 with pure argon
- => Long in length and high in energy
- Due to long length running into beam neutrons, prompt window set to 40ns (from 90) to avoid beam energy in long events.
 - The redone energy calibration for 40ns prompt window is shown to the right.
 - 2.2 MeV = 22.1 PE => 10 PE/MeV



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Background Measurements

 CCM made several direct measurements of various background types.



- Thermal neutron detector measures steady state neutrons
- Germanium detector measures gamma rays at various energies.



Figure 3.22: An EJ-301 detector (left) and the placement around CCM of 5 such detectors (right).



Figure 3.24: The gamma spectrum as determined by the germanium detector comparing beam on versus beam off

CCM120 Analysis: Measuring Steady State Backgrounds

- Prompt light only analysis
- Dynamic event lengths allow a poor-persons PID
 - Maximize dark matter over Ar39
 puts the length cut at 44 ns
- Pre-beam is flat in time (no bias) allowing a good prediction of what to expect in the prompt speed of light window (ROI)
- ROI is a beam-related background free region, so the prediction on the number of events is statistical only (systematics will be on DM signal)
- Ideal for Machine Learning techniques



Background Prediction

- CCM uses direct measurements of the steady state background to predict the ROI background.
- However, the 'steady state' is not steady

 checks revealed a background that
 was slightly decaying in rate across all
 energies, significant on the order of
 microseconds.
 - Beam activated radioactivity (known to exist from Germanium background measurements) was a possible candidate.
- Used linear fits of 6us of background to determine that expected within the ROI.



Figure 6.1: The 19PE bin's start time distribution. In blue are the prebeam ROI-length (150ns) time windows, black is the actual data value, and red is the background prediction from the linear fit (also in red).

distribution shown in blue on the lower right.



CCM120 Electromagnetic Search: Data and Background Prediction



(Left) The data and background spectra from the background prediction and the measured data in the beam ROI. (Right) The subtraction between data and background prediction about 0.

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- Tripathi, Mayank. University of Florida ProQuest Dissertations & Theses, 2024.31144965.
- Added Concrete and Roof shielding between CCM120 (2019) and CCM200 (2021) runs leads to reduced background
 - rates Nearly to level of eliminating beam-related background/activation, from comparison of beam-on vs. beam-off background.
- Factor of ~7 reduction in LDM search backgrounds.
- Factor of \sim 5 reduction in E/M search • backgrounds. Reduction greatest at low energy, reducing as energy increases.



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DM cuts

2021: Prebeam Modeling

- In the 2021 data set, the prebeam used for background prediction was not simple. Effects included:
 - An oscillation
 - Sharp dips
 - An overall trend
 - Issues near the window start
 - And overall higher error
- We compensated with a full fit function on a limited prebeam:

$$y = a0 * sin(a1 * (x - a2)) + exp(a3 + a4 * x) - exp(a5 + a6 * x) + a7$$





CCM200 2021 Analysis: Data Subtraction



The background prediction and measured data in the ROI (left) and the data - background subtraction (right) for the 2021 electromagnetic search of CCM.

BSM Signal Prediction Pipeline

CCM120 Engineering 6-week Run

- 1.79 × 10²¹ Protons on Target (**POT**)
- Electromagnetic / ALP Search
 - 294590 observed EM events in ROI
 - **Backgrounds**: 294614.3 ± 241.7 (syst) ± 542.8 (stat)
- <u>Nuclear Recoil / LDM Search</u>
 - 115005 **observed** in ROI
 - 16.5±338.4 subtraction events
- Separation of energy spectra from keV to o(100) MeV across different models

CCM200 Projection

- 3 year run (2023-2025)
- 2.25 × 10²² POT





Lujan Target Physics: Electromagnetic and Hadronic Particle Showers

- GEANT4 simulated fluxes from 800 MeV protons on W target geometry
- Physics list: QGSP_BIC_HP
- Hadronic cascades + meson production

 \circ π^{0} , π^{+} , π^{-}

- Production of secondaries: e⁺, e⁻, γ's:
 - \circ e⁺e⁻ Pair production
 - Bremsstrahlung
 - \circ π^0 decays
 - Neutron/proton scattering
 - etc...
- Each of these processes serves as a potential dark sector flux source







2019 Search for ALPs: Electron Couplings

A. A. Aguilar-Arevalo et al. "Axion-Like Particles at Coherent CAPTAIN-Mills". In: (Dec. 2021). arXiv: 2112.09979 [hep-ph].



Pure ALP-electron coupling

With electron loop-induced photon coupling and $a \rightarrow \gamma \gamma$

Neutrino Physics at CCM

- CEvNS measurements
- Charged-current quasi-elastic (CCQE) cross section measurement
- Relevant for physics
 @DUNE:
 - Supernovae neutrino physics
 - sub-GeV atmospheric neutrinos



More measurements needed here!

Heavy Neutral Lepton Search

 CCM200 3 year run sensitivity to HNL production from neutrino upscattering in shielding and detector materials only



 Lujan facility is spacious enough to allow for increased shielding in future runs



Lujan Improved S/B with Upgrade 30 nsec Beam

lized Plight Path 3 Start Time 0.01 - v from u-decay am Pulse from *π*-decay Prompt neutrino/DM signal window ~190 nsec. 0.008 0.006 TOF technique unique and Neutron powerful for isolating Measured Wave prompt signal and 0.004 Beam TO Gamma-Flash measuring backgrounds and errors from pre-beam. Key 0.002 is to shorten the beam width. 10 Derivative 10 10 10 10

-800

Shorter beam pulse reduces random backgrounds from Ar39 decay and neutron activation

If we shorten PSR pulse from 300 nsec to 30 nsec (Blue), would increase signal efficiency and reduce random backgrounds, estimate increase S/B (30 nsec) > 100.

-200

-400

• Factor ~10 reduction in random backgrounds from Ar39 and neutron activation

-600

• Factor ~10 reduction of EM events relative to nuclear scattering using Singlet/Triplet light PID.

Future Upgrades at CCM

Upgrades can greatly reduce

backgrounds to improve sensitivity:

- Short Pulse from LANSCE AOT
 - Working toward 120ns pulse
- Cherenkov light
 - Direction reconstruction
- Underground Argon
- CCI (coherent Cesium Iodide second detector)
 - \circ tonne-scale

Lujan/CCM makes up for less power with large, sensitive, and fast 10-ton LAr detector!



Proton Improvement Plan Beam Dump (PIP-II BD) [arXiv:2203.08079]

New Dedicated Facility: FNAL Proton Beam Dump @ 1 GeV

- Single-phase, 100 ton scintillation-only LAr detector o(20 m) proximity to target
 - Same tech as CCM200
 - 800 MeV 2 GeV beam possibilities
- Projected world-leading sensitivity;
 - Dark matter
 - ALPs
 - Sterile Neutrinos
 - v cross section measurements, etc.

