

# On the quantum origin of a small positive cosmological constant

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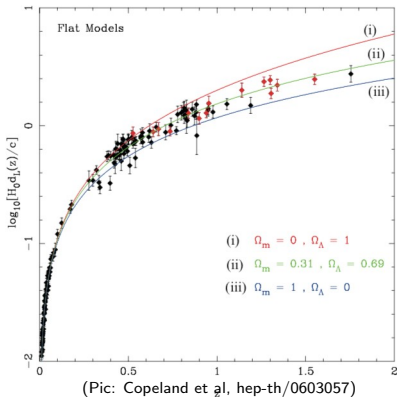
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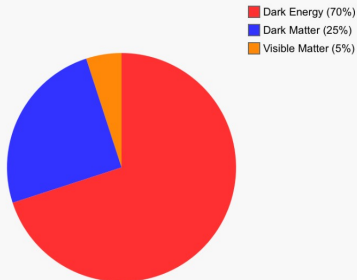
# Dark Matter, Dark Energy

$$\text{Luminosity distance } d_L(\Omega_\Lambda, \Omega_M, z) = \frac{(1+z)}{H_0} \int_0^z \frac{dz}{\sqrt{\underbrace{\Omega_\Lambda}_{0.7} + \underbrace{\Omega_M(1+z)^3}_{0.3}}}$$

$$\Omega_\Lambda = \frac{\rho_\Lambda}{\rho_{\text{crit}}} , \quad \Omega_M = \frac{\rho_M}{\rho_{\text{crit}}}$$



Matter-energy content of our Universe



- What constitutes Dark Matter?
- What constitutes Dark Energy/ $\Lambda$ ?
- Why is  $\Lambda$  positive?
- Why is  $\Lambda$  tiny, about  $10^{-123} \ell_{Pl}^{-2}$  where  $\ell_{Pl}$  is the Planck length?

$$(\rho = \int_0^{k_{max}} dk k^2 \sqrt{k^2 + m^2} \approx k_{max}^4 > 10^{50} \rho_{\Lambda})$$

- Currently  $\rho_{DM} \approx \underbrace{\rho_{\Lambda}}_{\frac{\Lambda c^2}{8\pi G}} \approx \underbrace{\rho_{crit}}_{\frac{3H_0^2}{8\pi G}} \approx 10^{-26} \text{ kg m}^{-3}$

Why? The 'coincidence problem'

# Bose-Einstein Condensate as Dark Matter

- Cold
- Dark
- Light bosons as DM  $\Rightarrow$  no small scale structure
- Macroscopic *Quantum state*
- BEC  $\Rightarrow$  DE ( $\approx$  DM) via its (repulsive) Quantum Potential
- Few assumptions and free parameters

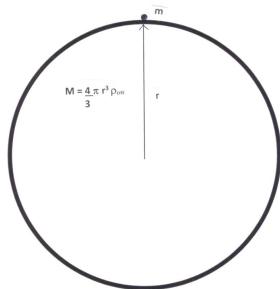
Critical temperature, universe temperature, boson mass

$$T_c(a) = \frac{\hbar c}{k_B} \left( \frac{(N/V) \pi^2}{\eta \zeta(3)} \right)^{1/3} = \frac{\hbar c}{k_B} \left( \frac{(0.25 \rho_{crit}/ma^3) \pi^2}{\eta \zeta(3)} \right)^{1/3} = \frac{4.9}{m^{1/3} a} K$$

$$T(a) = \frac{3.7}{a} K$$

$$T(a) < T_c(a) \rightarrow m < 6 \text{ eV}/c^2 \Rightarrow \text{BEC forms in the early universe}$$

# Dark Energy from BEC



Classical, Newtonian

$R_{space} = 0$ ,  $R_{spacetime} = 10^{-123} \ell_{Pl}^{-2}$ , will quantize, GR  $\rightarrow$  same results

$$m\ddot{r} = -\frac{GMm}{r^2} = -\frac{G(\frac{4}{3}\pi r^3 \rho_{crit})m}{r^2}, \quad r = r_0 a(t), \quad a = \text{scale factor}$$

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3} \rho_{crit} = -\omega^2 \quad \text{Raychaudhuri Equation}$$

BEC in a harmonic trap for  $t \ll H_0^{-1}$  (14 Gyr)

## Quantize

$$-\frac{\hbar^2}{2m} \nabla^2 \Psi + V \Psi = i\hbar \frac{\partial \Psi}{\partial t}$$

$$\Psi(\vec{r}, t) = \mathcal{R} e^{i\frac{S}{\hbar}}, \quad \mathcal{R}, S = \text{Real}, \quad \mathcal{R}^2 = |\Psi|^2 = \rho, \quad \vec{v} = \frac{\vec{\nabla} S}{m}$$

Real and imaginary parts

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \vec{v}) = 0,$$

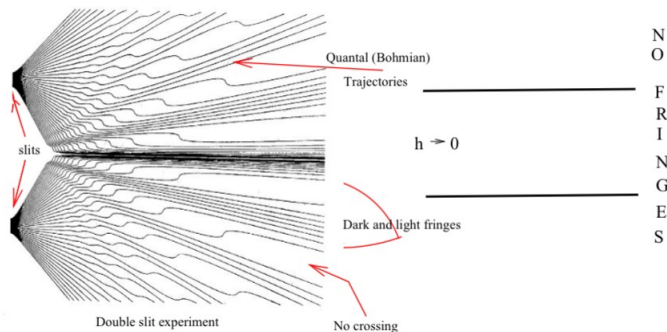
$$\begin{aligned} m \frac{d\vec{v}}{dt} &= -\vec{\nabla} V + \frac{\hbar^2}{2m} \vec{\nabla} \left( \frac{1}{\mathcal{R}} \nabla^2 \mathcal{R} \right) \\ &= -\vec{\nabla} (V + V_Q) \end{aligned}$$

$$V_Q = -\frac{\hbar^2}{2m} \frac{\nabla^2 \mathcal{R}}{\mathcal{R}}$$

$V_Q = -V + \text{const.}$  if  $\Psi =$  stationary state

# Quantum Potential - Example

## Double slit experiment



(C. Phillipides, C. Dewdney, B. J. Hiley, *Quantum Interference and the Quantum Potential*, *il nuovo cimento B*, **52**, no.1 (1979) 15-28)



# Quantum Potential $\rightarrow \Lambda$

$$m \frac{d\vec{v}}{dt} = -\vec{\nabla} (V + V_Q) \quad , \quad V_Q = -\frac{\hbar^2}{2m} \frac{\nabla^2 \mathcal{R}}{\mathcal{R}} \quad , \quad \vec{v} = \frac{d\vec{r}}{dt} \quad , \quad r = r_0 a(t)$$

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3} \rho + \frac{\hbar^2}{6m^2} \nabla^2 \left( \frac{\nabla^2 \mathcal{R}}{\mathcal{R}} \right)$$

$$\frac{\ddot{a}}{a} = -\frac{4\pi G}{3} \left( \rho + \frac{3p}{c^2} \right) + \frac{\Lambda}{3} \quad (\text{compare})$$

$$\Lambda = \Lambda_Q = \frac{\hbar^2}{2m^2 c^2} \nabla^2 \left( \frac{\nabla^2 \mathcal{R}}{\mathcal{R}} \right) = -\frac{1}{mc^2} \nabla^2 V_Q$$

# Quantum Potential of BEC

## Quantize

BEC wavefunction

$$\Psi = \underbrace{R(a)}_{\text{DM decay}} \times \underbrace{e^{-r^2/\sigma^2}}_{\text{HO GS}},$$

$$R(a) = \frac{R_0}{a^{3/2}}, \quad \rho_{DM} = |\Psi|^2 \propto \frac{1}{a^3}, \quad \int dV |\Psi|^2 = N, \quad \sigma^2 = \frac{2\hbar}{m\omega}$$

$$V_Q = -\frac{\hbar^2}{2m} \frac{\nabla^2 \Psi}{\Psi} = \frac{3}{2} \hbar\omega - \frac{1}{2} m\omega^2 r^2 \quad (\text{Inverted HO!})$$

$$\frac{\ddot{a}}{a} = -\frac{4\pi G \rho_{crit}}{3} + \frac{\Lambda_Q}{3}$$

$$\Lambda_Q = -\frac{1}{mc^2} \nabla^2 V_Q = \frac{3\hbar^2}{2} = 4\pi G \rho_{crit} = \text{constant!}$$

$$\rho_\Lambda = \frac{\Lambda}{8\pi G} = \frac{\rho_{crit}}{2} \approx \rho_{DM}, \quad p_\Lambda = -\frac{\Lambda c^2}{8\pi G} = -\rho_\Lambda$$

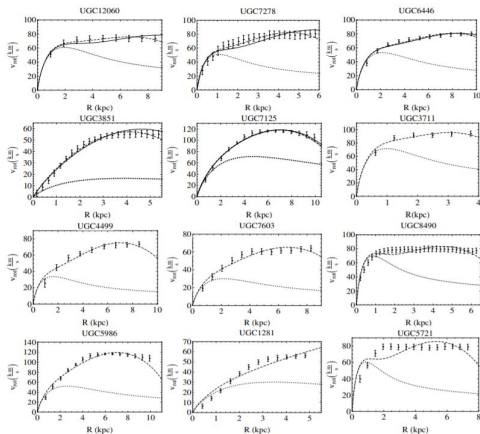
$$= -\frac{4\pi G}{3} (\rho_{DM} + \rho_\Lambda + 3p_\Lambda) = 0 \quad \text{Einstein's static Universe!}$$

S.Das, R. K. Bhaduri, Class. Quant. Grav. 32 105003 (2015) and arXiv:1812.07647

# Results for $m = 10^{-32} \text{ eV}/c^2$

$$i\hbar \frac{\partial}{\partial t} \psi(\mathbf{r}, t) = \left[ -\frac{\hbar^2}{2m} \Delta + mV_{\text{ext}}(\mathbf{r}) + \lambda\rho(\mathbf{r}, t) \right] \psi(\mathbf{r}, t).$$

$$V_{\text{ext}} = U_Y(\mathbf{r}) = - \int \frac{G\rho_{\text{BEC}}(\mathbf{r}')}{|\mathbf{r} - \mathbf{r}'|} e^{-\frac{|\mathbf{r} - \mathbf{r}'|}{R_S}} d^3\mathbf{r}'.$$



E. Kun, Z. Keresztes, S. Das, L.A. Gergely, *Symmetry*, **10**, 520 (2018), arXiv:1905.04336

## Summary

- What constitutes DM? *BEC*
- What constitutes DE? *Quantum potential of the BEC*
- Why is  $\Lambda$  positive? *Because  $V_Q = -V > 0$*
- Why is  $\rho_{DM} \approx \rho_\Lambda \approx \rho_{crit}$ ? *Because  $|V_Q| = |V|$*

## Remarks

- We get  $\rho_\Lambda = \not\approx \rho_{DM}$ , because  $\rho_{DM} \propto 1/a^3$ ,  $\rho_\Lambda \propto \text{constant}$  and all bosons  $\neq$  GS
- Newtonian  $\rightarrow$  Relativistic (Straightforward)
- *i.e.* Raychaudhuri Equation  $\rightarrow$  Raychaudhuri Equation +  $V_Q$ . Results don't change
- Prediction: ultralight bosons of  $m < 6\text{eV}/c^2$  including the 'preferred' value of  $m = 10^{-22} \text{eV}/c^2$  (prevents small scale structure)
- What are these bosons? gravitons, axions, ...
- Better fitting with galaxy data
- Prediction:  $\Lambda$  has changed in the far past and will change in the future