

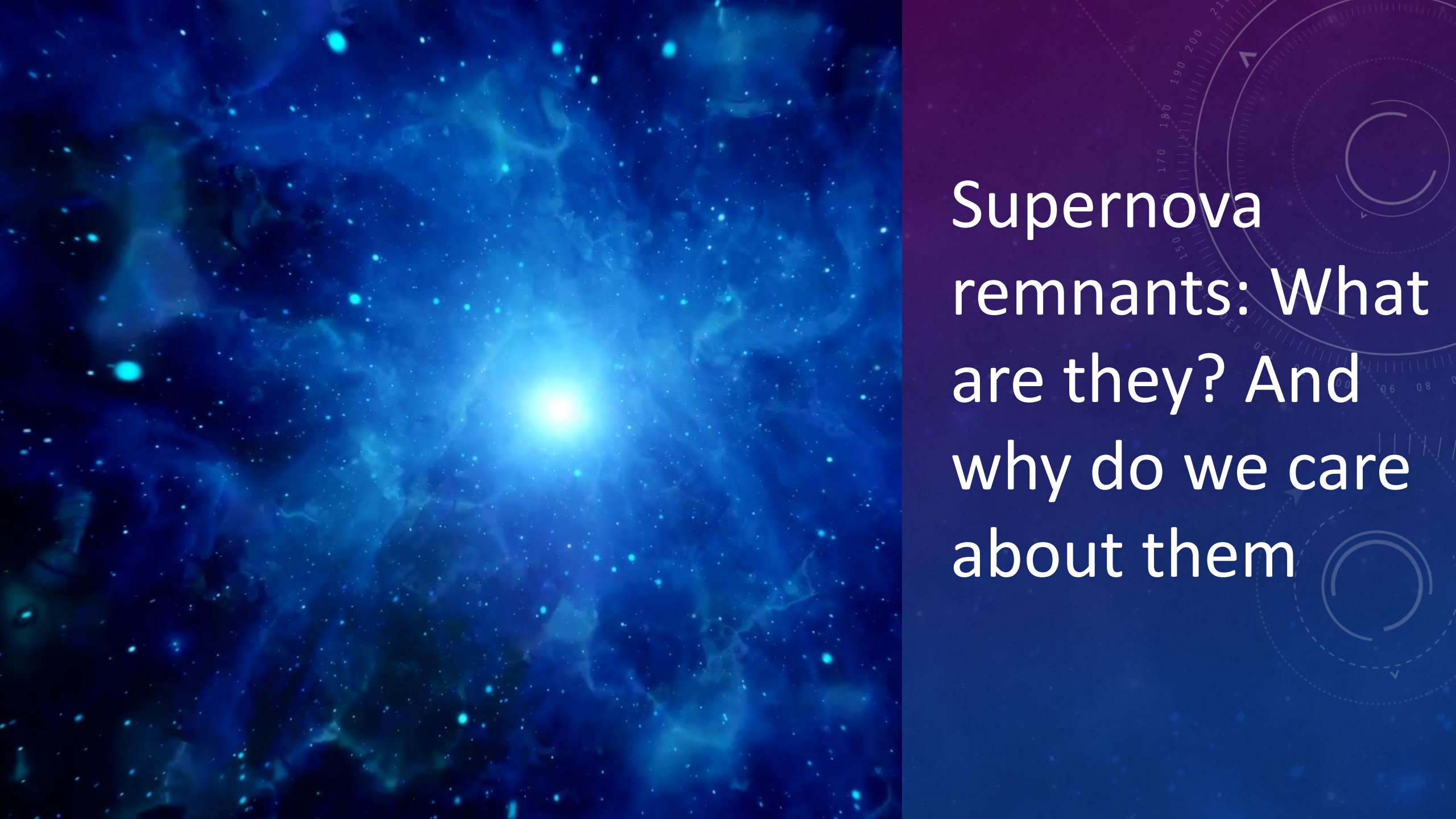
NUCLEAR/ELEMENT PRODUCTION BY SUPERNOVAE, MEASURED BY X-RAY SPECTRA

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TALKING POINTS

- Supernova remnants: What are they? And why do we care about them
- How we measured elemental abundances
- MY research.

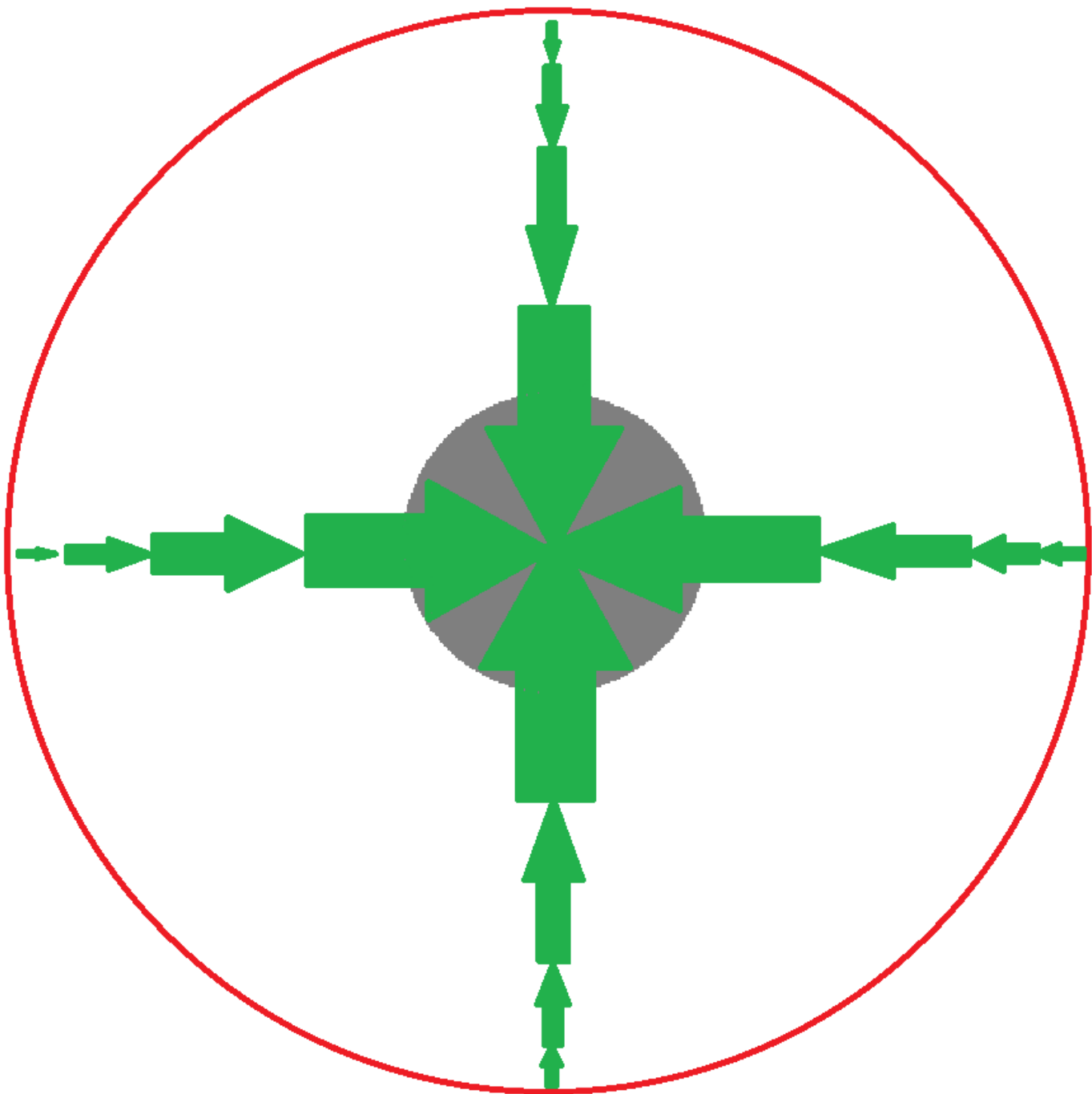
The image features a cosmic scene on the left, dominated by a bright, glowing blue and cyan nebula. A central star is the primary light source, surrounded by intricate, filamentary structures of gas and dust. The right side of the image is a dark blue vertical panel with a purple-to-blue gradient. It contains faint, semi-transparent technical graphics, including circular gauges with numerical scales (150, 170, 180, 190, 200, 210) and concentric circles, suggesting a scientific or data-driven theme.

Supernova
remnants: What
are they? And
why do we care
about them

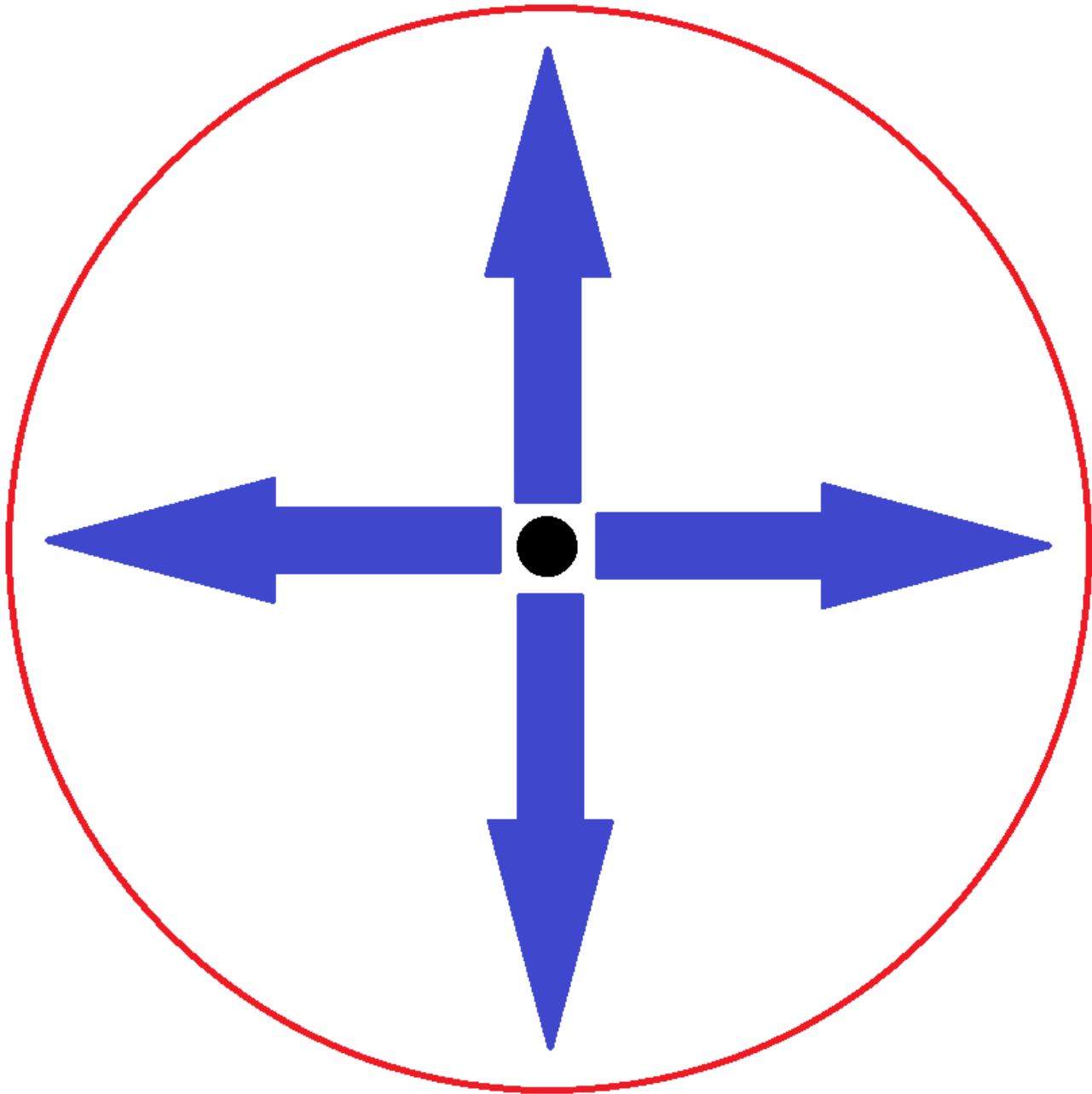


SUPERNOVA

- Death of large stars or white dwarfs
- Massive explosion
- Creation of heavy elements



CONTRACTION



REACTION





SUPERNOVA

- Death of large stars or white dwarfs
- Massive explosion
- Creation of heavy elements

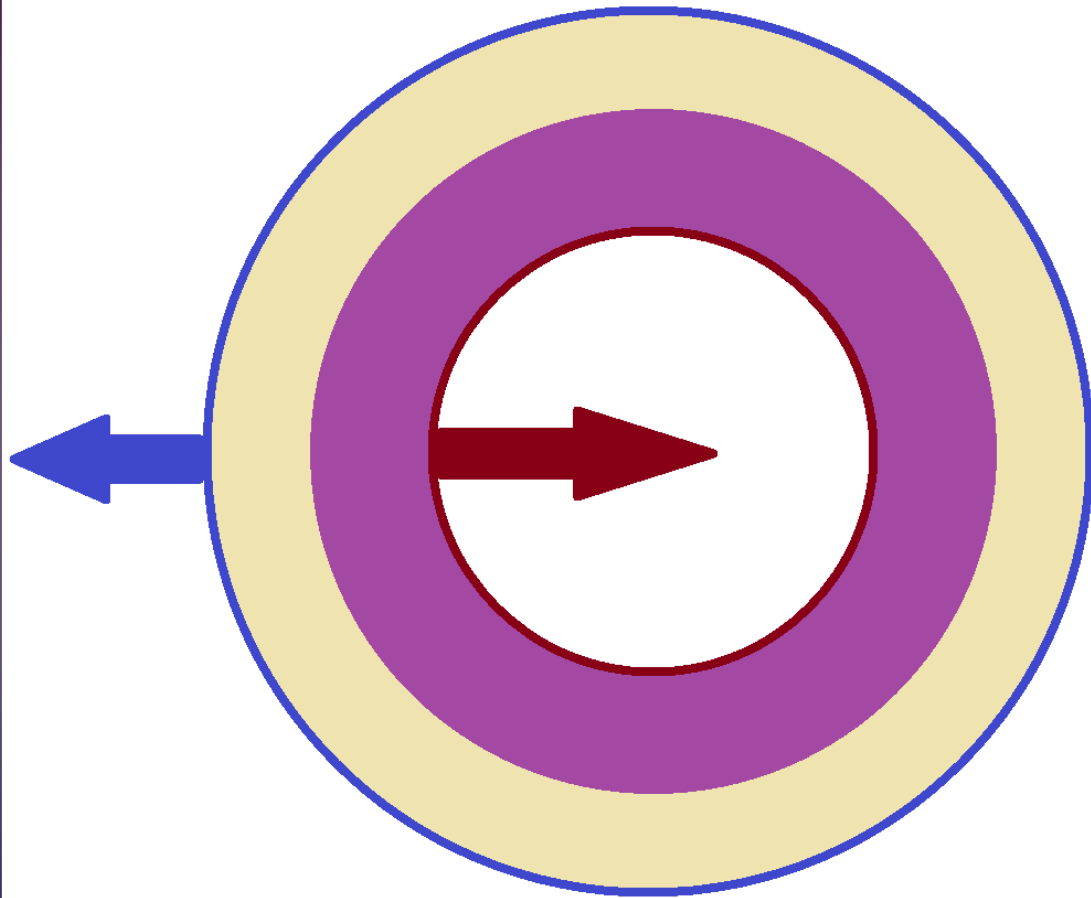
SUPERNOVA REMNANTS

Expanding balls of
hot gas and plasma

Interactions with the
surrounding dust and
gas (Interstellar
medium)

Spreads heavy
elements

FORWARD AND REVERSE SHOCKS



- Two shocks
- One into ISM, one into material from star
- Emission between shocks
- Little to no hydrogen in purple region

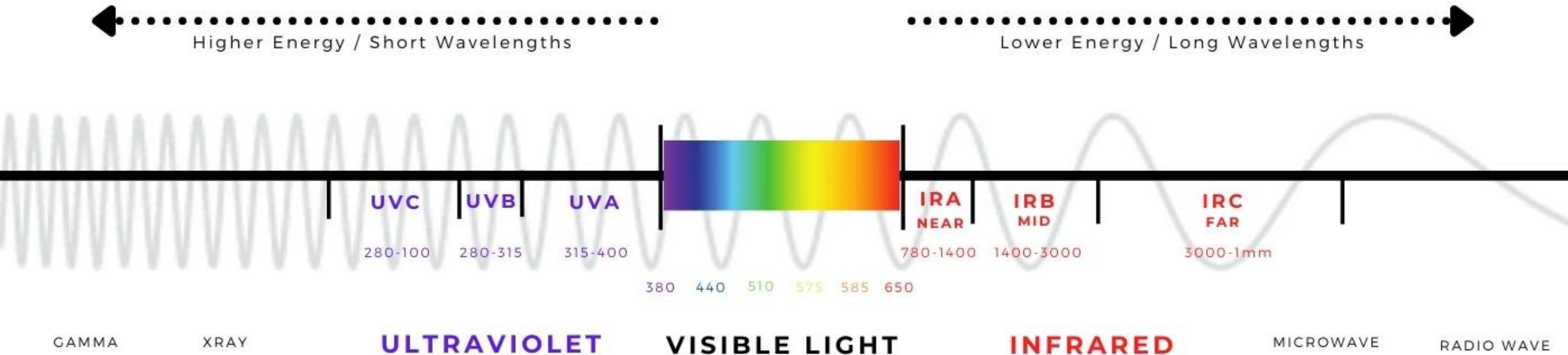
HOW WE MEASURED ELEMENTAL ABUNDANCES

X-ray Spectroscopy



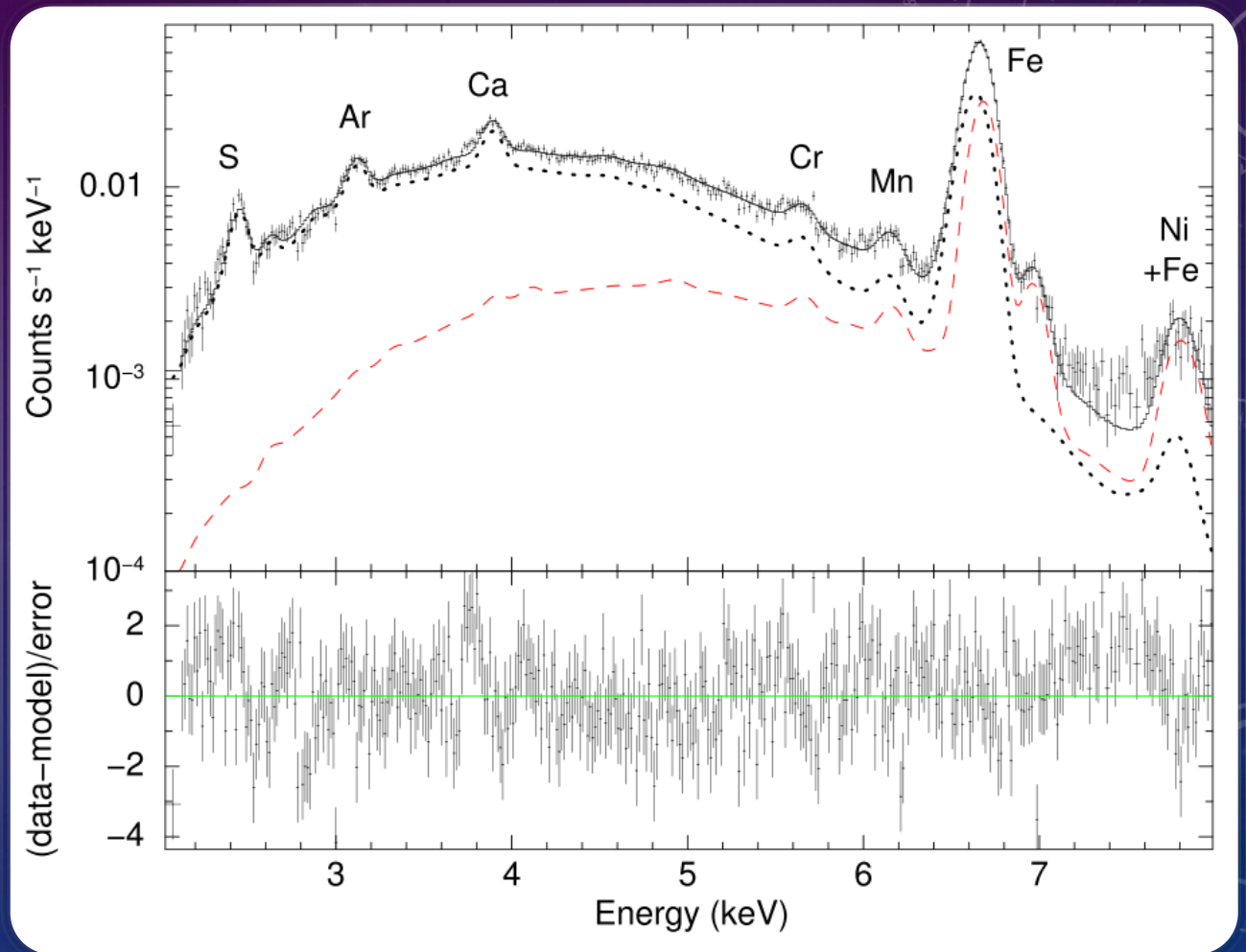
X-RAY SPECTRUM

- Photon between 50 to 50,000 time more energetic than visible light
- Typical studies focus of around 1-10keV (1.2-0.12 nm)



X-SPEC

- Calculates model spectrum and fits to observations
- Uses solar amounts of hydrogen
- Gives abundances from the model spectrum



Spectrum from: Zhou, S.-C. Leung, Z. Li, K. Nomoto, J. Vink, and Y. Chen. Chemical abundances in sgr a east: Evidence for a type iax supernova remnant. *The Astrophysical Journal*, 908(1):31, Feb. 2021.

MY RESEARCH



ELECTRON/HYDROGEN DENSITY DENSITY

Element
measurements
based off electron
to hydrogen ratio

Heavy elemental
abundances are
significantly
underestimated

THE PROCESS

01

Find the elemental abundance of supernova remnants previously studied

02

Correct the abundances while assuming little to no hydrogen

03

Calculate the masses of each element (assuming we have distance)

04

End goal of creation catalog of abundances and masses

KEY POINTS

What is a Supernova remnant

- Super nova remnants are spheres of gas and plasma expanding into the surrounded interstellar medium. Supernova create heavy elements.

How are the element abundances measured

- Elemental abundances are measured using the x-ray spectrum fitting software x-spec. This process assumes solar hydrogen abundances.

My contributions

- I will correct the elemental abundance measurements in regions of supernova remnants that are hydrogen deficient, where the assumption of x-spec are no longer accurate.