

Nuclear Astrophysics Opportunities with Recoil Separators

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Triangle Universities Nuclear Laboratory
Duke University

TRIUMF Science Week
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What's Coming Up?

1. Nuclear reactions in stars
2. Recoil separators around the world

DRAGON and EMMA at  **TRIUMF**, SECAR at  **FRIB**

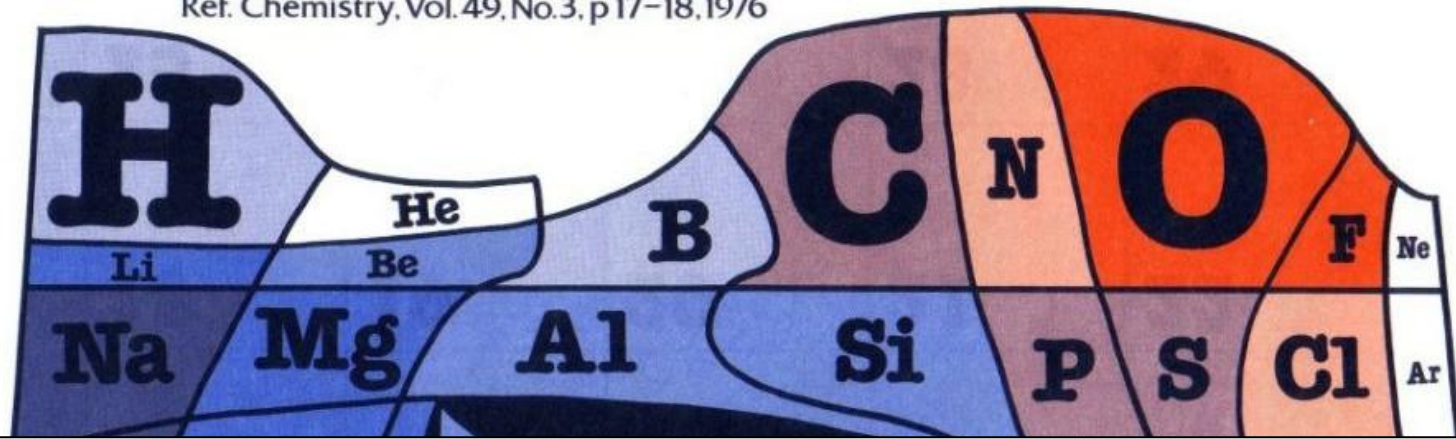
ESR at  **FAIR**
GSI

3. Summary and outlook

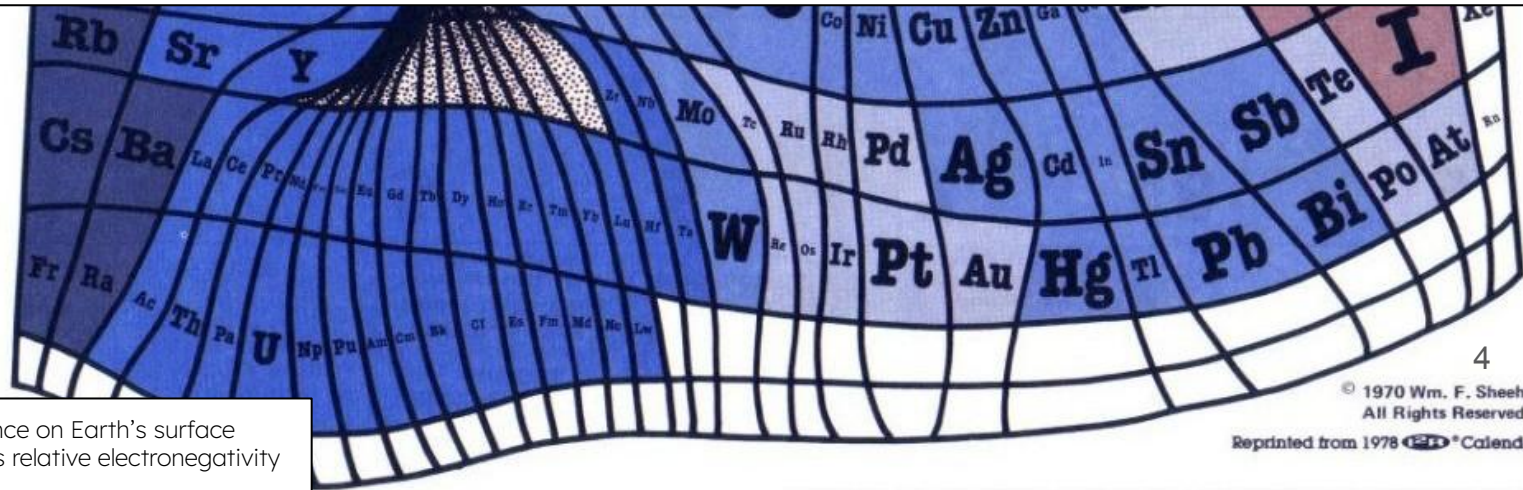


Understand how stars live and die

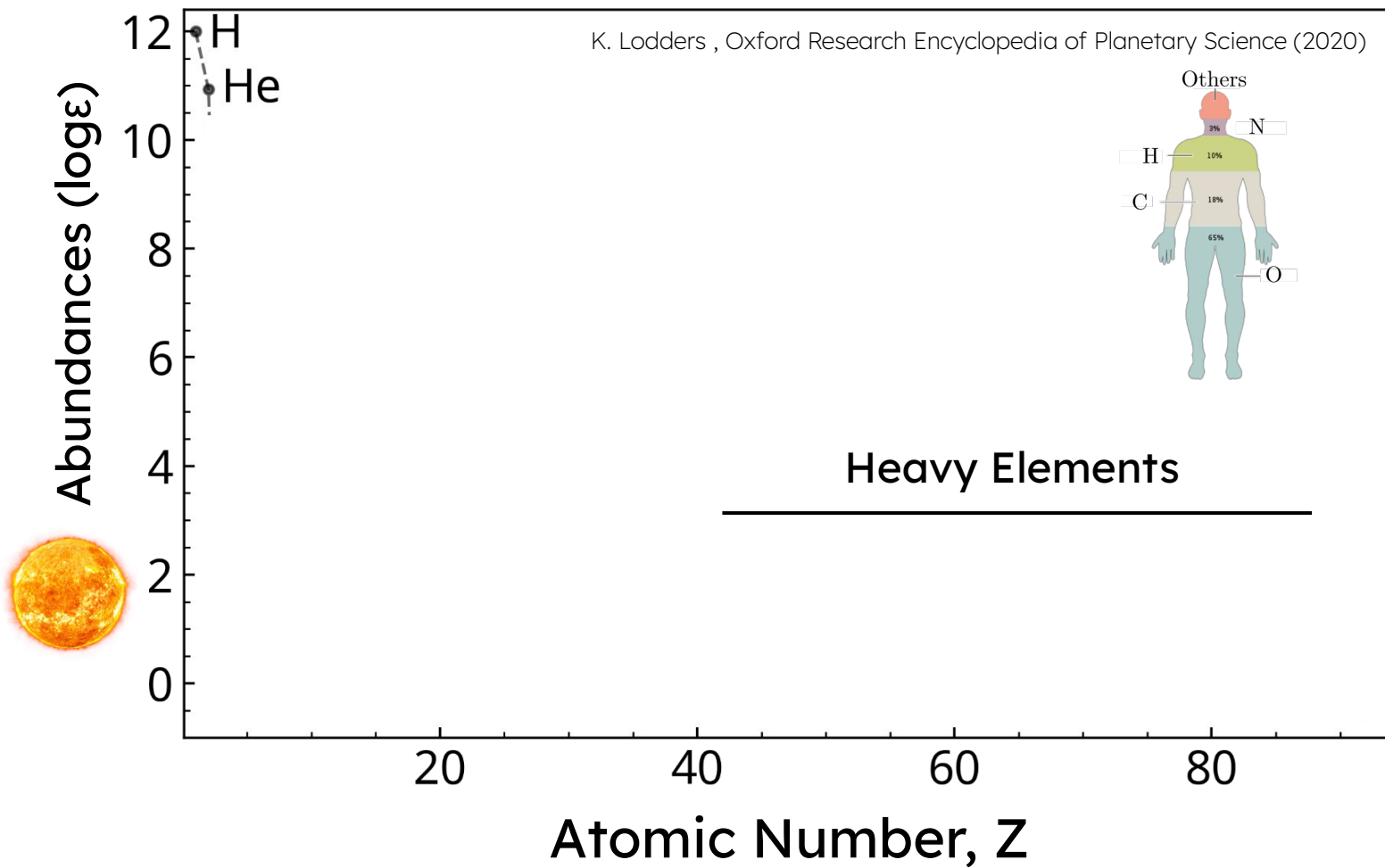
Video credit: NASA/Solar Dynamics Observatory (SDO), AIA 304 Å

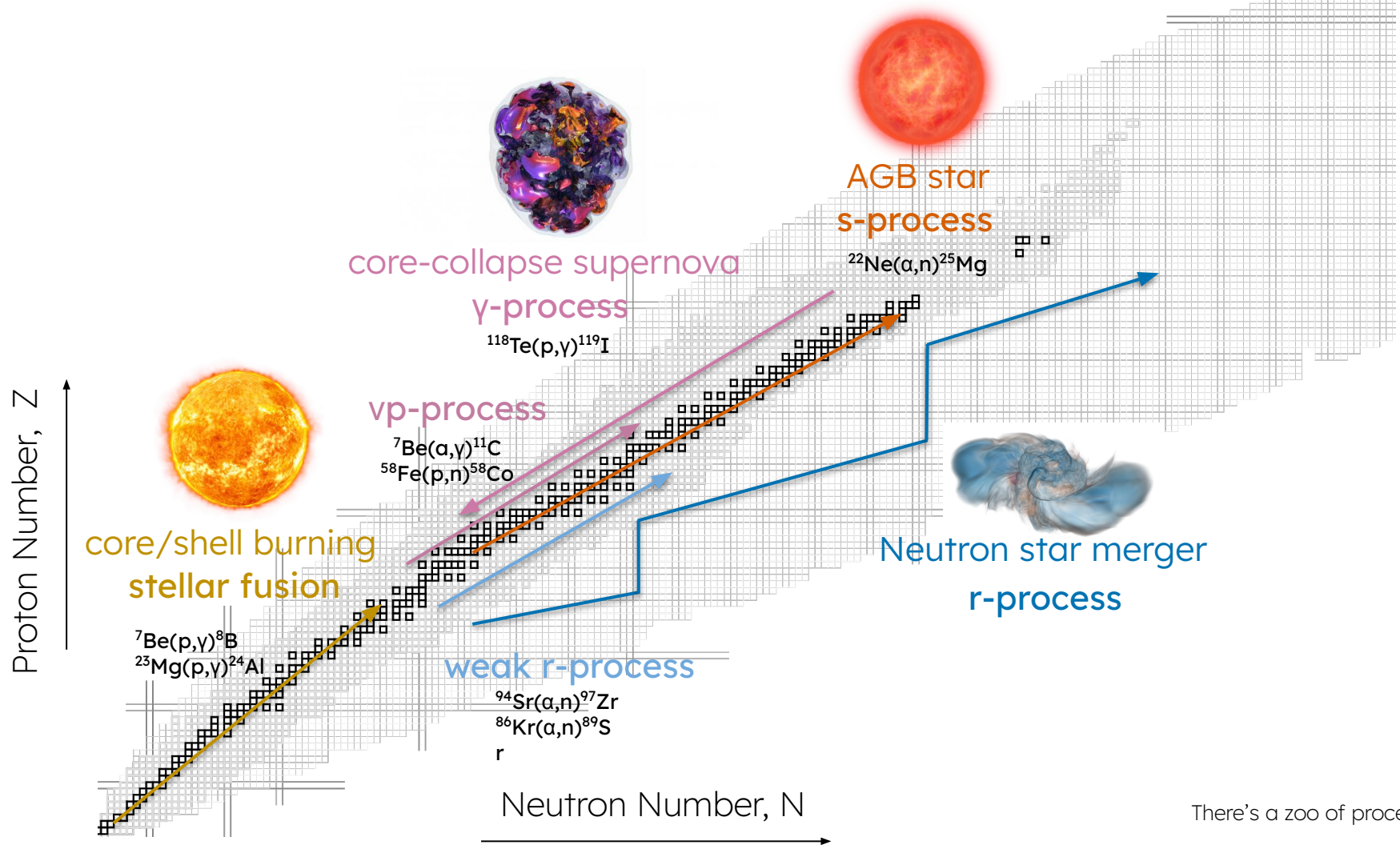


Understand the origin of the elements



Relative abundance on Earth's surface
Colour represents relative electronegativity



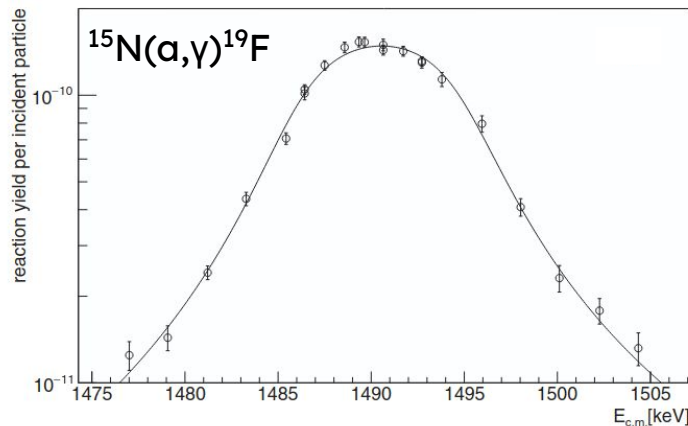


There's a zoo of processes...

observed abundances =
astrophysical conditions +
nuclear physics

From nuclear reactions to abundances

Bring the stars in the lab



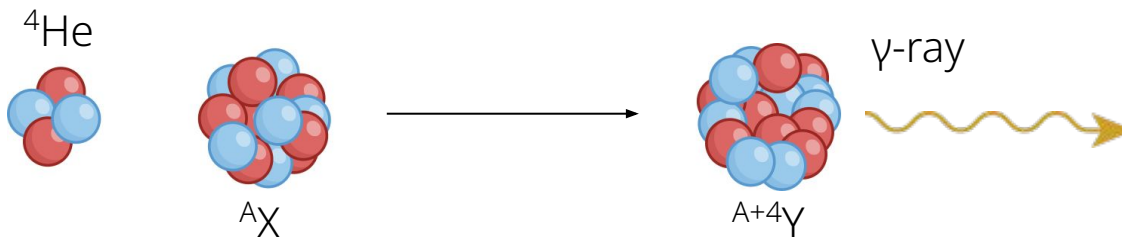
A. DiLeva *et al.*, Phys. Rev. C **95**, 045803 (2017)

$$\langle \sigma v \rangle \sim T_9^{-3/2} \sum_i (\omega \gamma)_i e^{-11.605 E_i / T_9}$$

↑
Reaction Rate

↑
Resonance strength

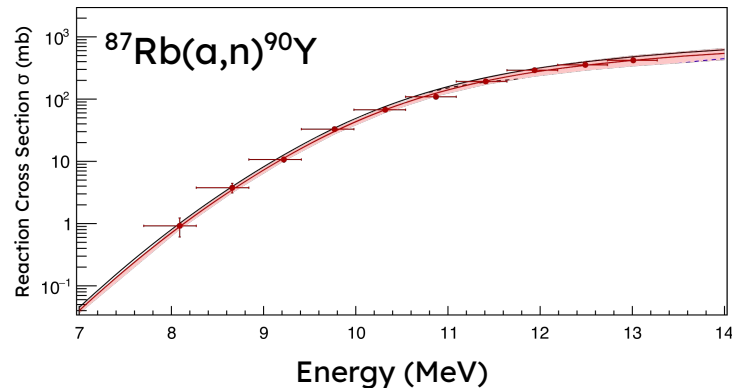
↑
Resonance energy



Detect charged-particles and (or) radiation in the lab

From nuclear reactions to abundances

Bring the stars in the lab



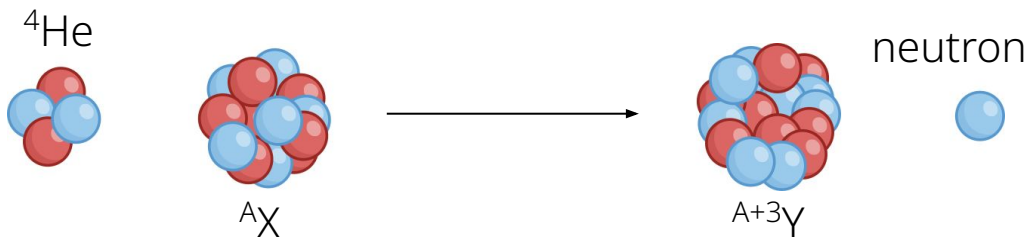
C. Fougères *et al.*, *Astrophys. J* **983**, 142 (2025)

$$\langle \sigma v \rangle \sim \int_0^\infty E \sigma(E) \exp\left(-\frac{E}{kT}\right) dE$$

Reaction Rate

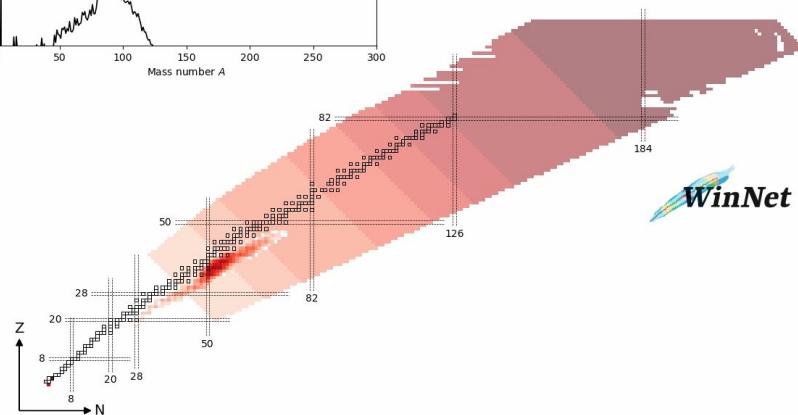
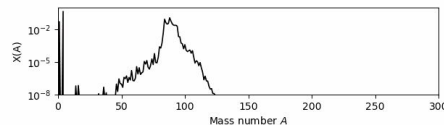
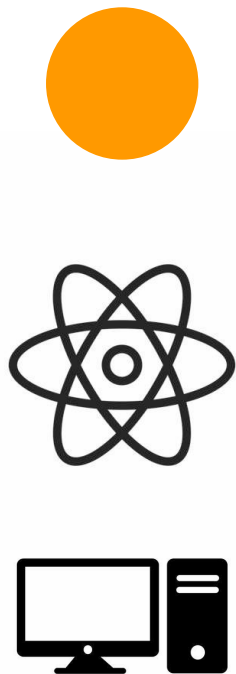
Reaction cross section

Maxwell-Boltzmann distribution



Detect charged-particles and (or) radiation in the lab

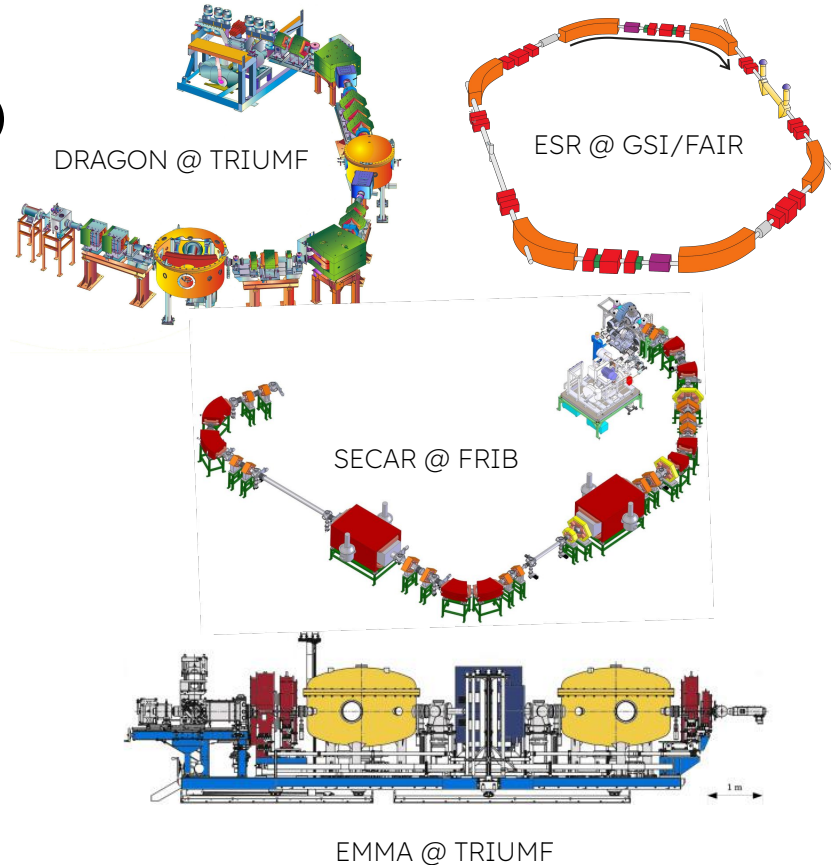
Putting everything together: nuclear reaction networks



Example of a weak r-process calculation using WinNet
<https://github.com/nuc-astro/WinNet>

Recoil separators: a versatile experimental setup

- Inverse kinematics
(suitable for short-lived isotopes)
obtain E_r and ω_y with a single measurement
- Detection of both heavy recoil and light ejectiles
- Controlled systematic uncertainties (simulations)



C. Ruiz, U. Greife and U. Hager, Eur. Phys. J. A **50**, 99 (2014)
C.R. Brune and B. Davids, Annu. Rev. Nucl. Part. Sci. **65**, 87 (2015)
C. Rolfs and C.A. Barnes, Annu. Rev. Nucl. Part. Sci. **40**, 45 (1990)

How to make a recoil separator

1. Target

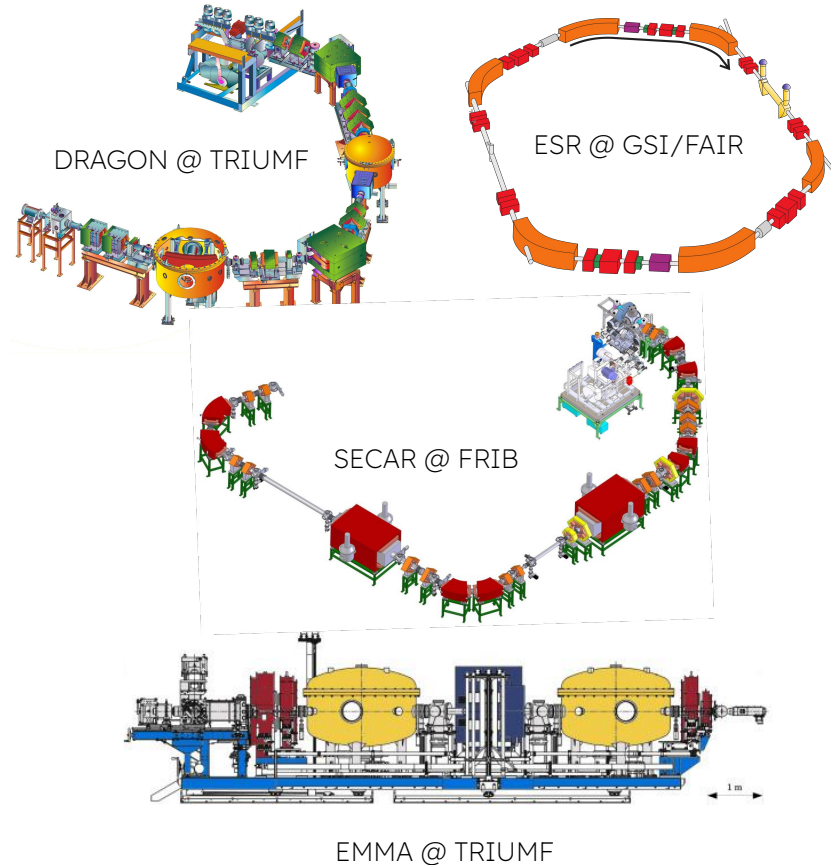
Gas (extended or jet) or solid

2. Electromagnetic separator

For beam-like particle rejection
(electric/magnetic dipoles, Wien filters, ...)

3. Detection system

charged particles, γ -rays, neutrons



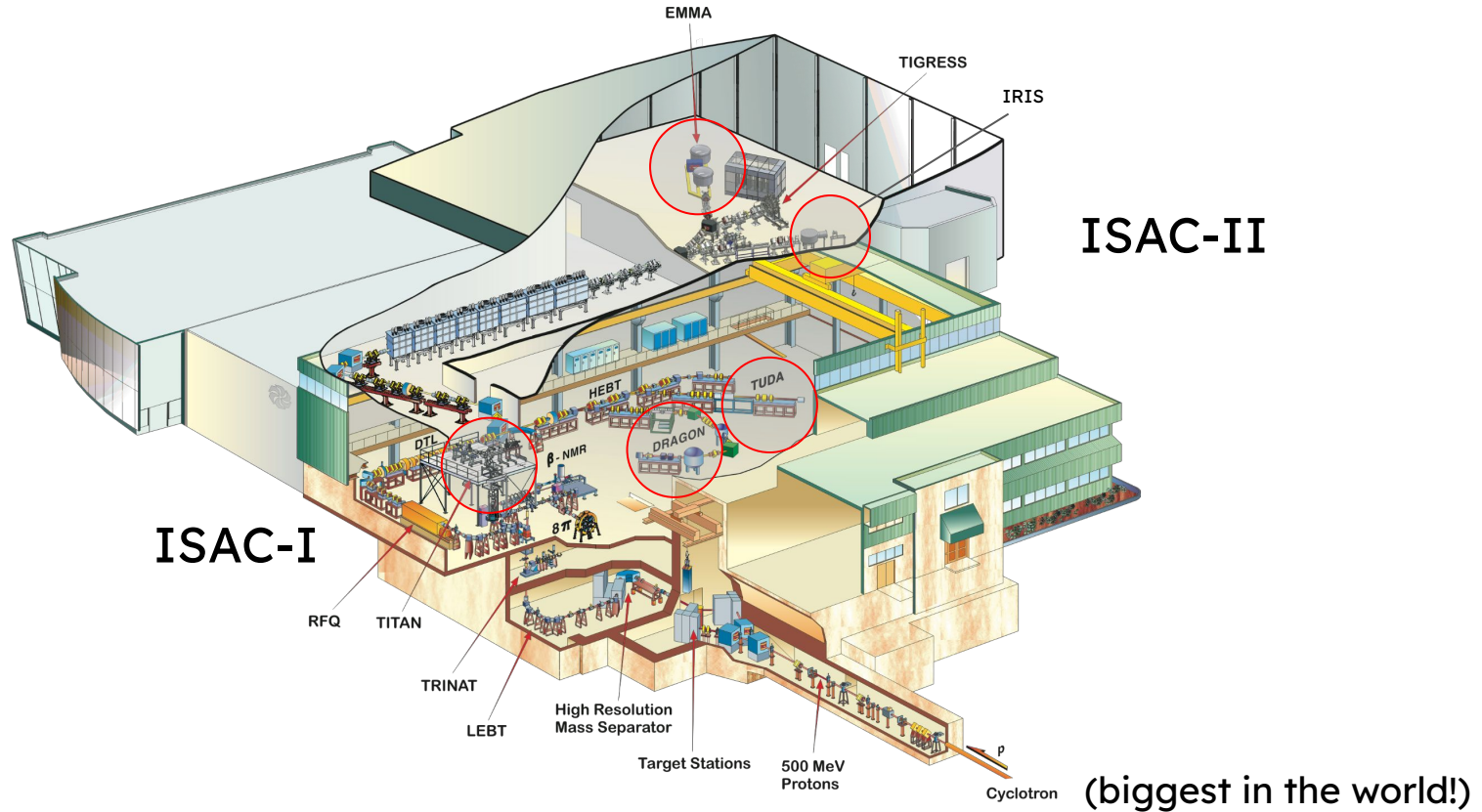


Many opportunities worldwide for nuclear astrophysics research!



Many opportunities worldwide for nuclear astrophysics research!

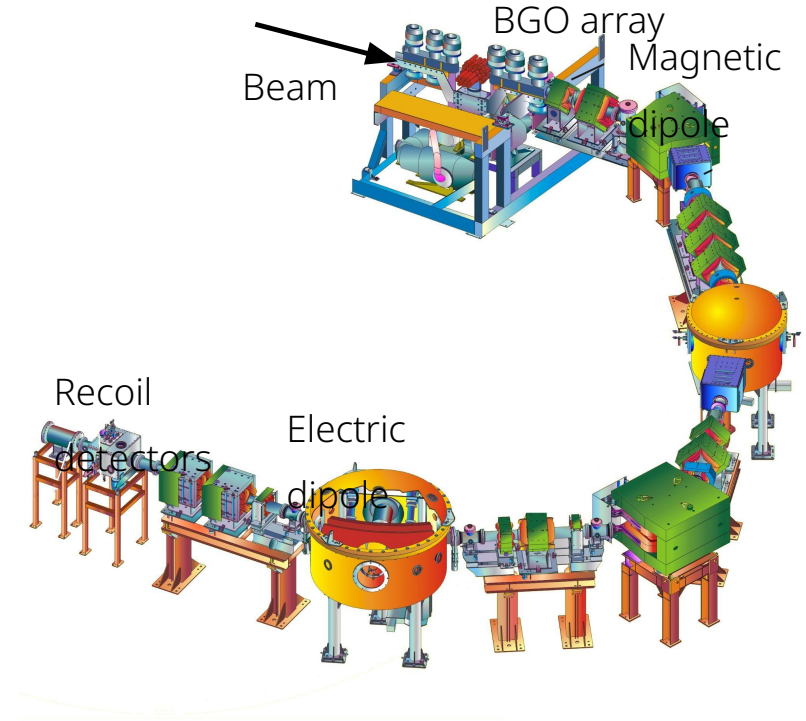
The nuclear astrophysics facilities at **TRIUMF**

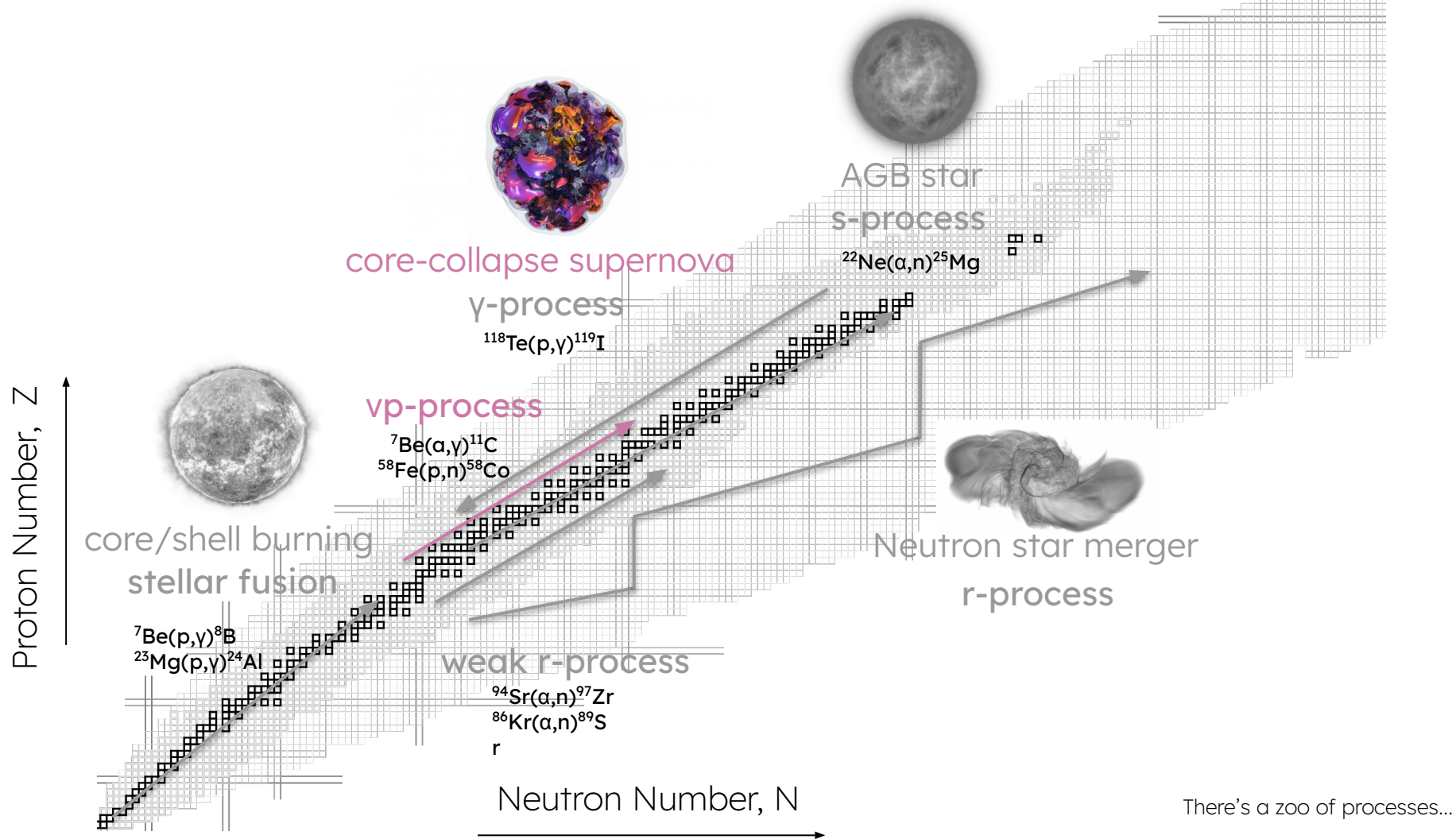


DRAGON at ISAC-I

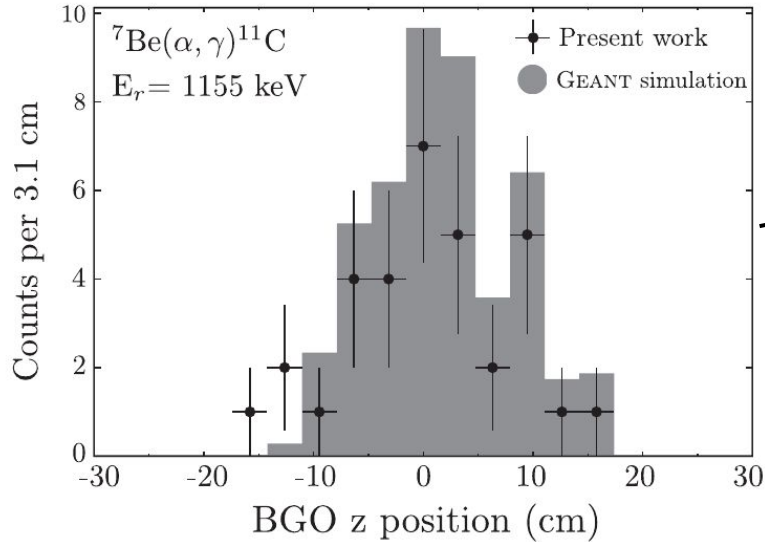
Detector of Recoils and Gammas of Nuclear reactions

- Has measured 9 out of the 14 radiative captures using radioactive ion beams
- Beam rejection $>10^{14}$ for (α, γ)
 $<10^{13}$ for (p, γ)
- Modular detector configuration in the target and focal plane



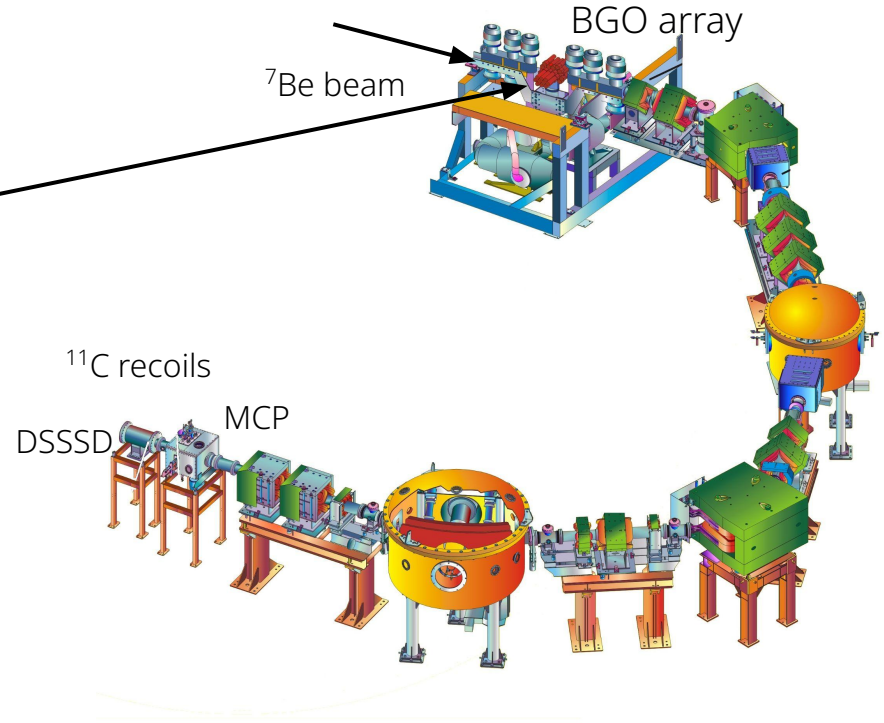


$^7\text{Be} + \text{DRAGON} = \text{❤️}$



A. Psaltis *et al.*, Phys. Rev. Lett. **129**, 162701 (2022)

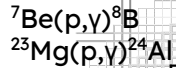
A. Psaltis *et al.*, Phys. Rev. C **106**, 045805 (2022)



$^7\text{Be}(\alpha, \gamma)^{11}\text{C}$ (S1692, Spokespersons: A. Psaltis, A.A. Chen and D. Connolly)

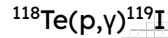
Proton Number, Z

core/shell burning
stellar fusion

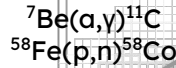


core-collapse supernova

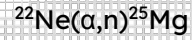
γ -process



vp-process

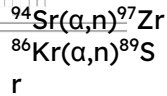


AGB star
s-process



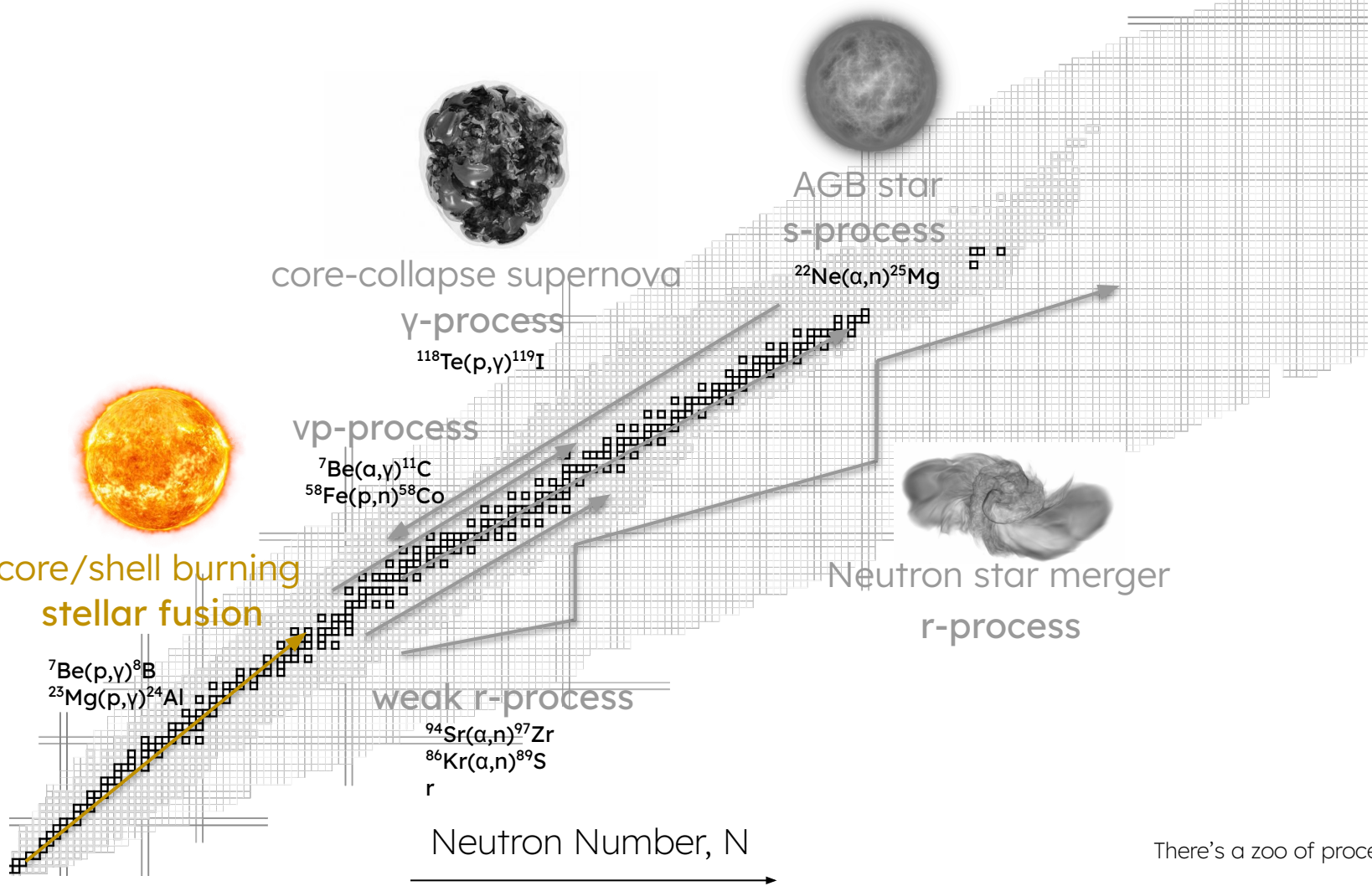
Neutron star merger
r-process

weak r-process

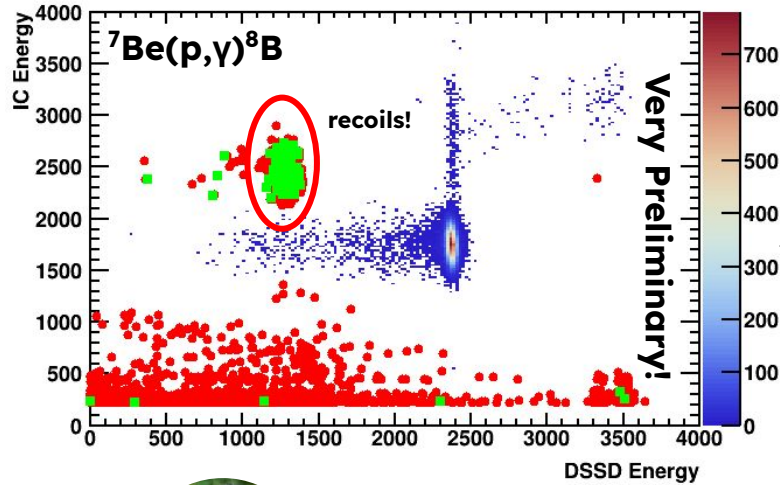


Neutron Number, N

There's a zoo of processes...

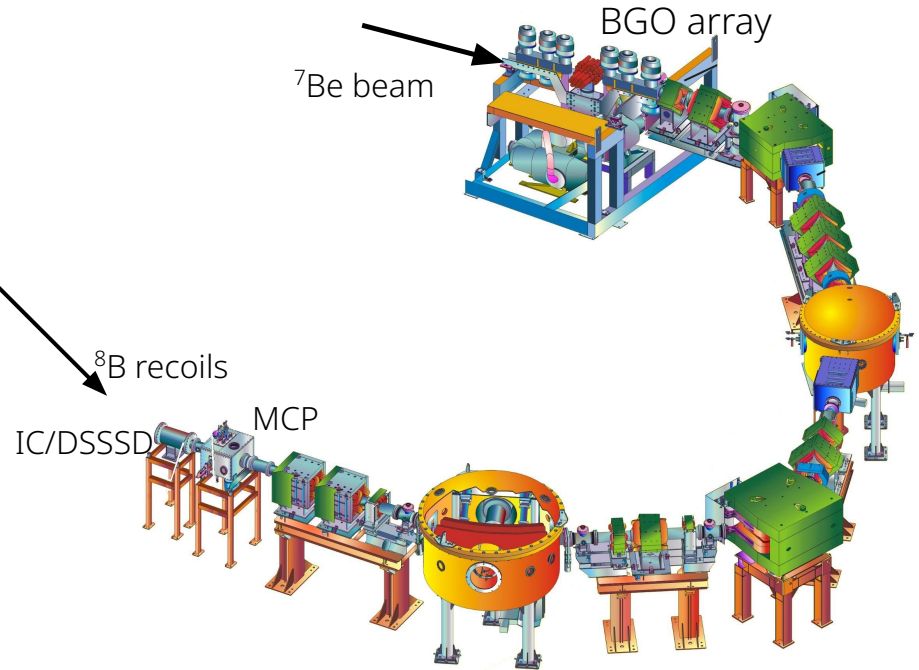


$${}^7\text{Be} + \text{DRAGON} = \text{♥}$$

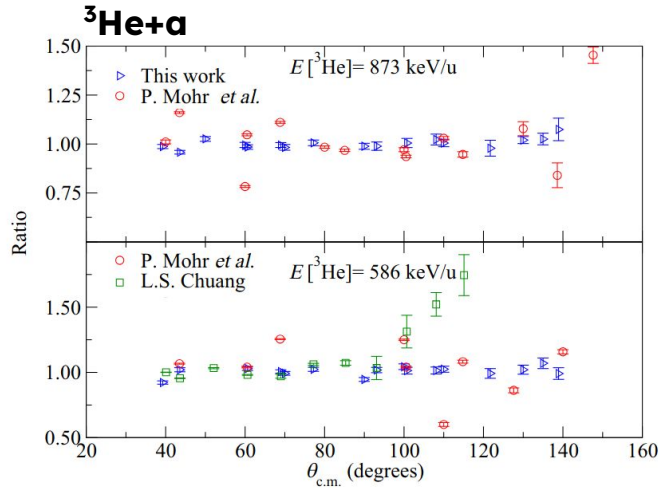


Mallory Loria
(UVic/TRIUMF graduate student)

${}^7\text{Be}(p,\gamma){}^8\text{B}$ (S2013, Spokespersons: M. Williams, B. Davids and U. Greife)



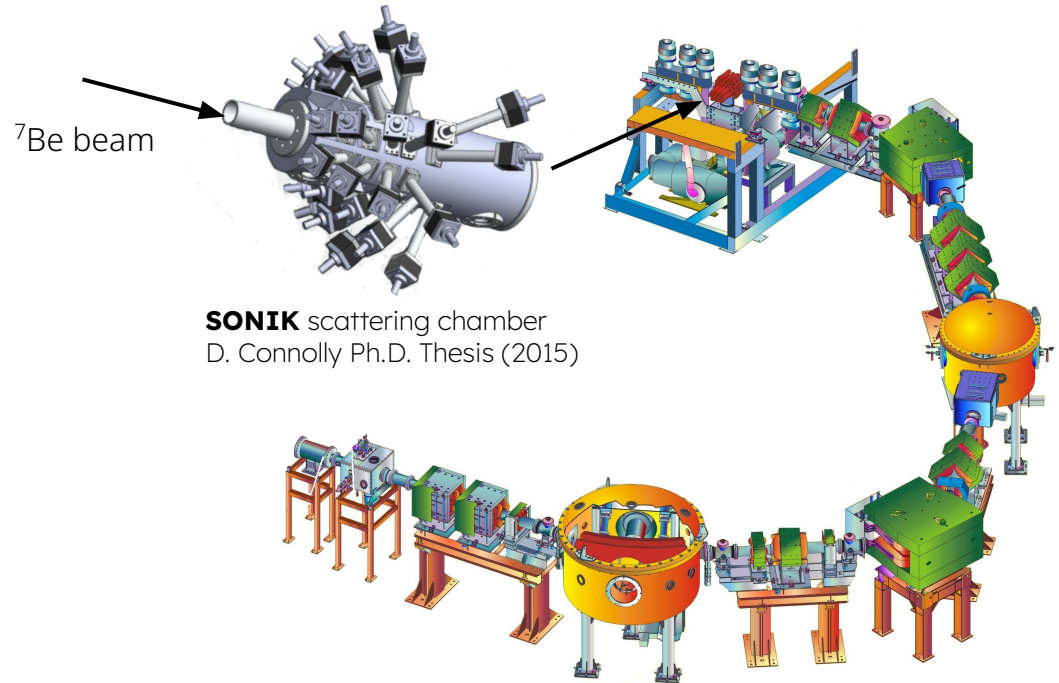
$^7\text{Be} + \text{DRAGON} = \text{❤️}$

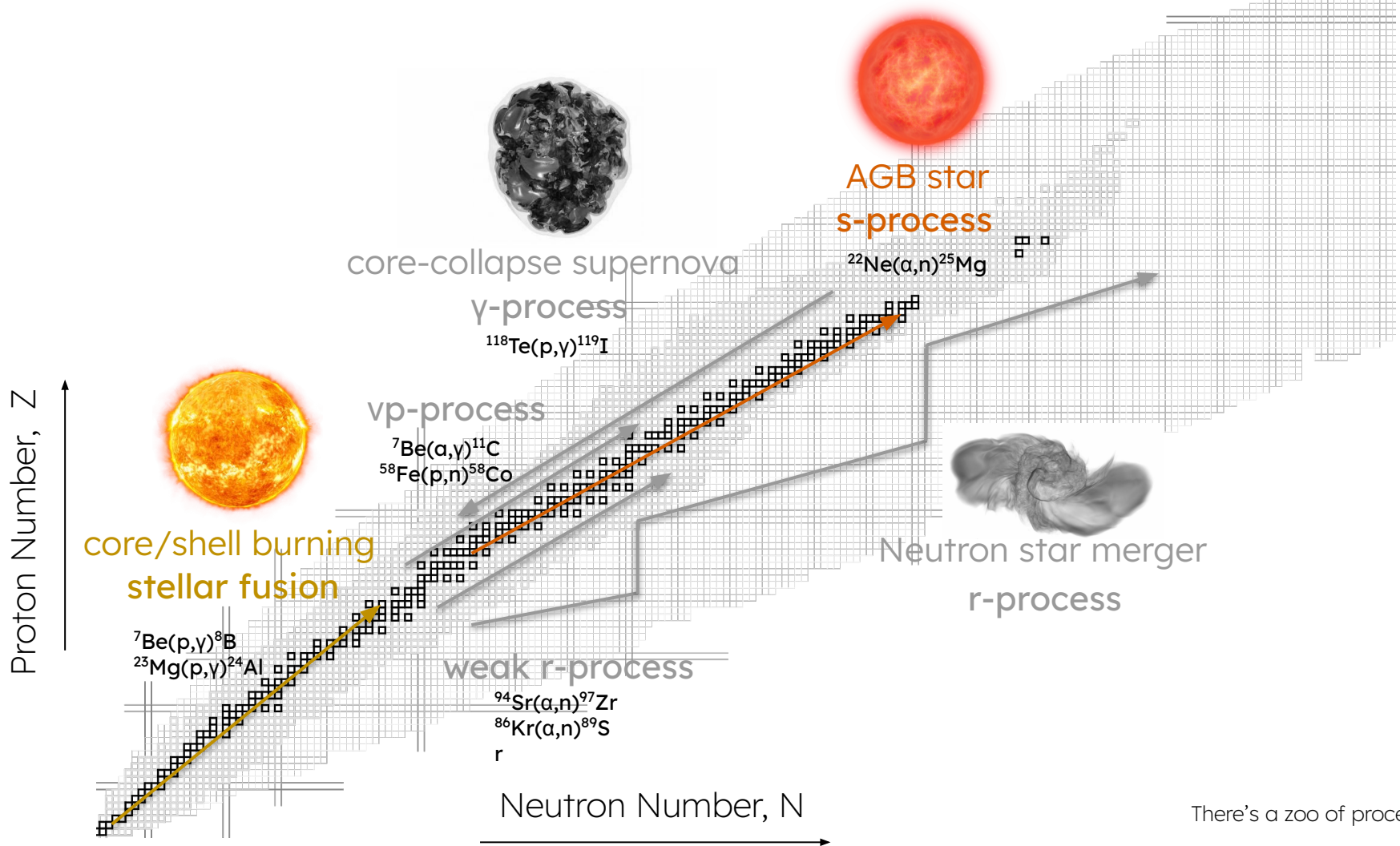


S.N. Paneru *et al.*, Phys. Rev. C **109**, 015802 (2024)

$^7\text{Be} + \alpha$ (S2018, Spokespersons: R.J. deBoer and C. Ruiz)

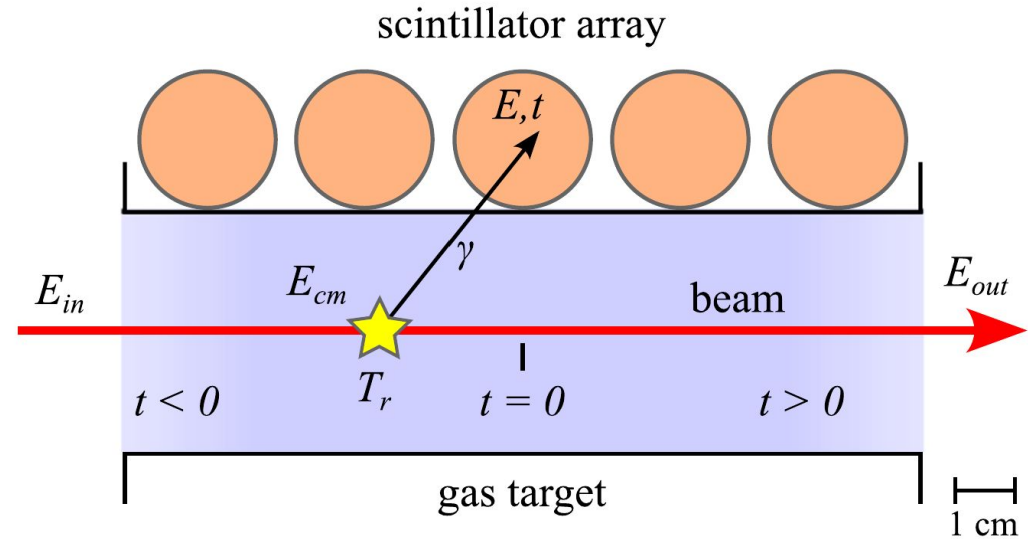
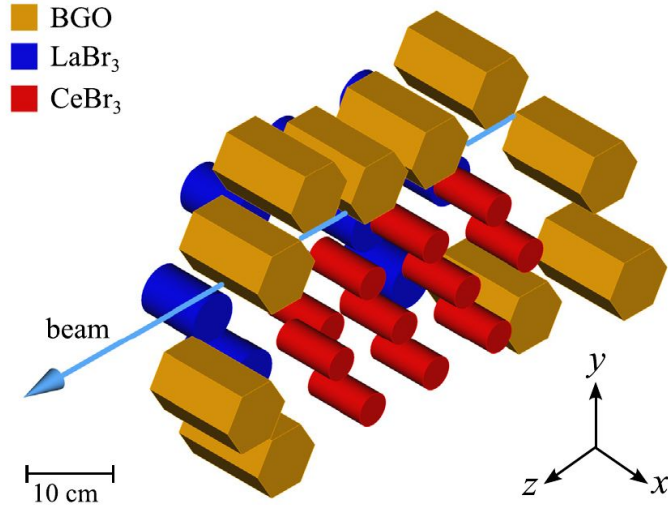
$^7\text{Be} + p$ (S2034, Spokespersons: C. R. Brune and C. Ruiz)





There's a zoo of processes...

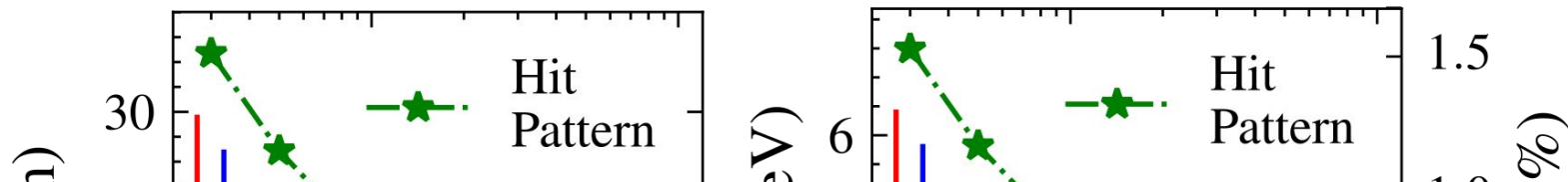
Resonance timing method at DRAGON



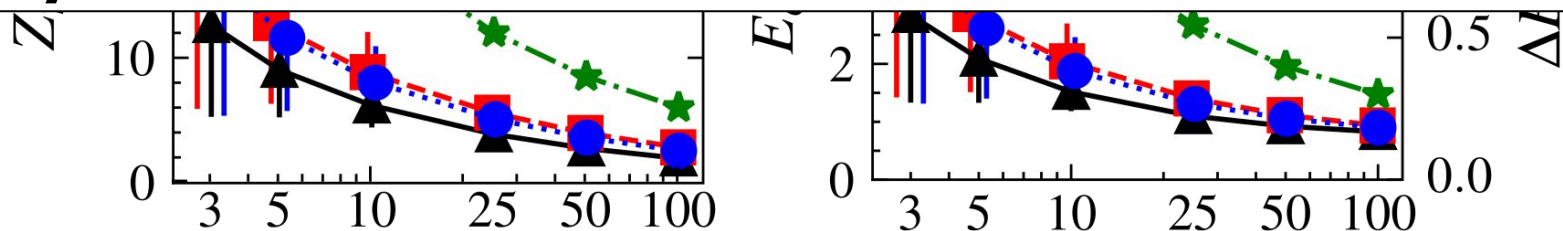
$$\langle \sigma v \rangle \sim T_9^{-3/2} \sum_i (\omega \gamma)_i e^{-11.605 E_i / T_9}$$

Resonance timing method at DRAGON

—▲— $\text{LaBr}_3/\text{CeBr}_3$ -■- BGO ···●··· All



Only a small number of recoils needed!



Sample Size

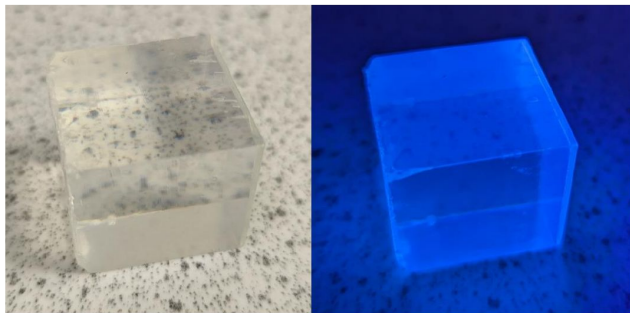
D.A. Hutcheon *et al.*, Nucl. Instrum. Methods Phys. Res A **689**, 70 (2012)

G. Christian *et al.*, Eur. Phys. J. A **50**, 75 (2014)

G. Christian *et al.*, Nucl. Instrum. Methods Phys. Res A **1072**, 170199 (2025)



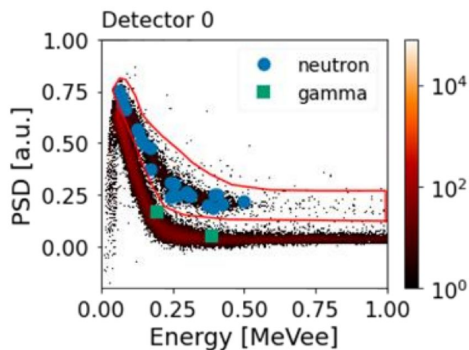
Upgrade DRAGON to measure (α, n) reactions



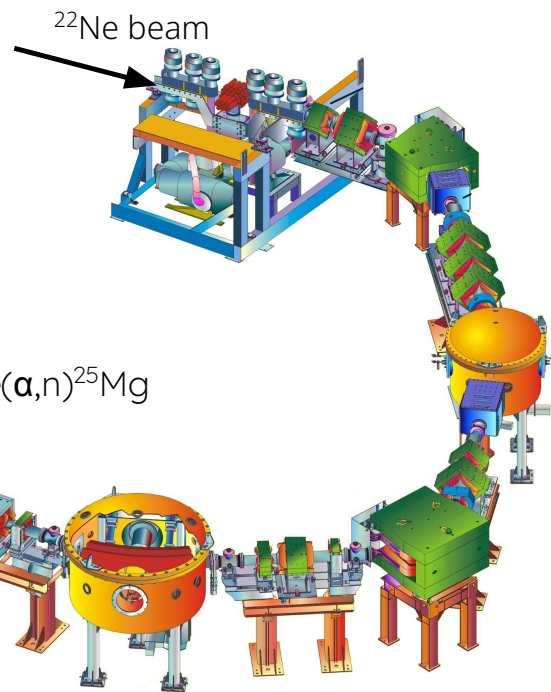
Organic Glass Scintillator (OGS) produced by BlueShift Optics



Ben Reed
(SMU/TRIUMF postdoc)



Proof-of-principle experiment $^{22}\text{Ne}(\alpha, n)^{25}\text{Mg}$
neutron source for the s-process



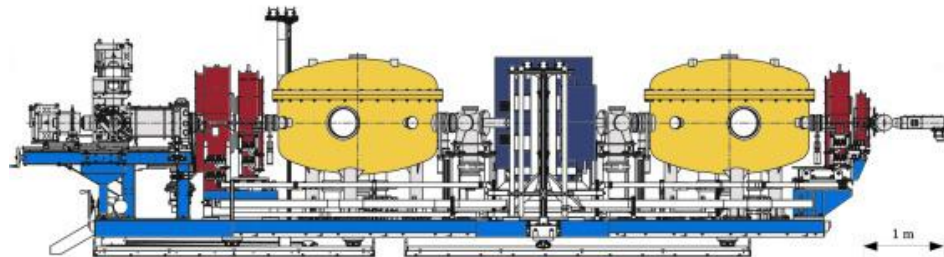
DEMAND

Direct Experimental Measurements of Astrophysical
reactions using Neutron Detectors

EMMA at ISAC-II

ElectroMagnetic MAss Analyzer

- Solid Si:He targets for α -induced reactions ($n \sim 10^{19}$ atoms/cm²)
Polyethylene (CH₂)_n for (p, γ)
- γ -ray detection from TIGRESS
12 clover detectors (HPGe)
- Coincidences with detection of heavy recoils at EMMA's focal plane
(PGAC/ionization chamber/silicon detector)

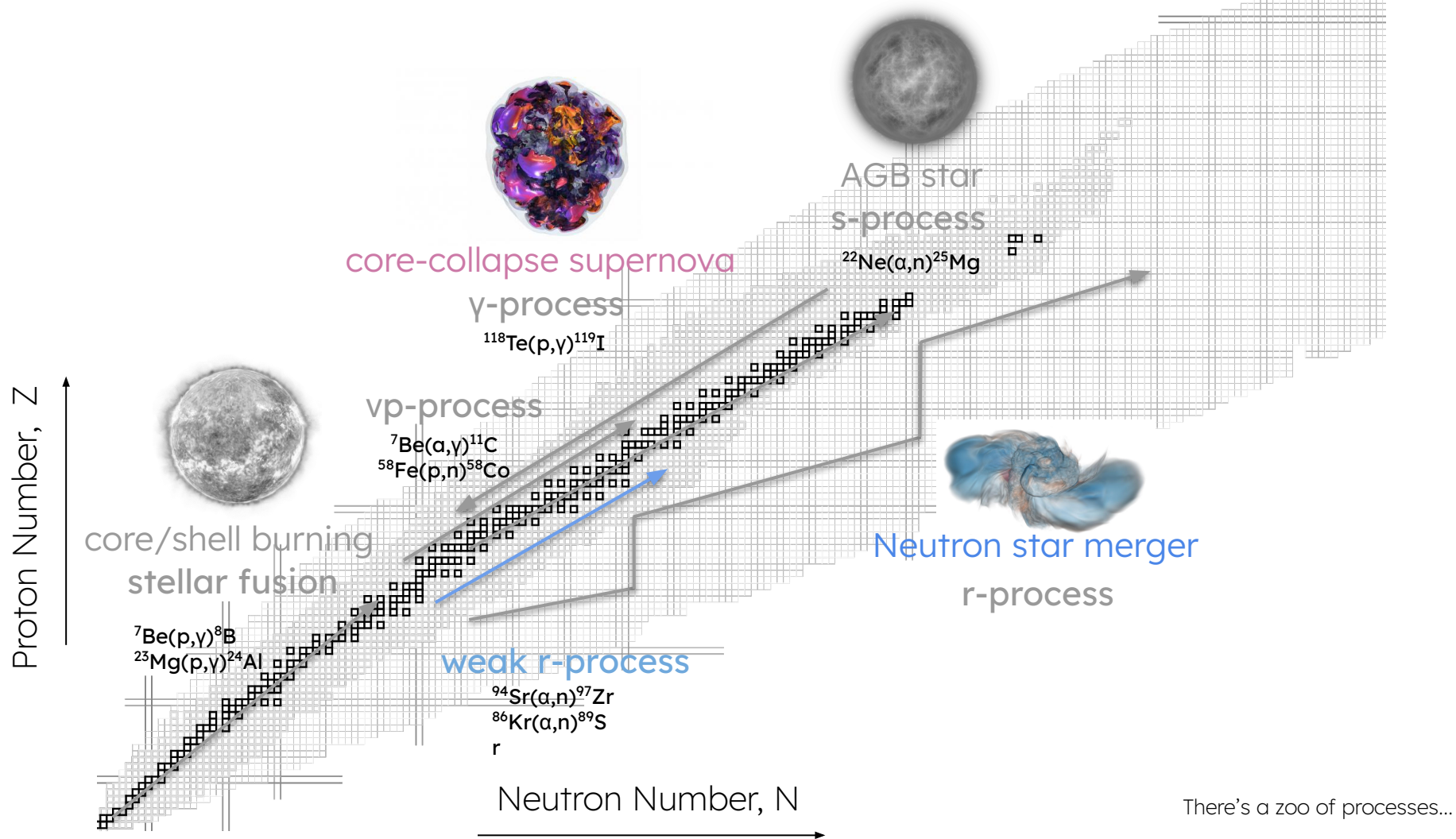


B. Davids & C.N. Davids, Nucl. Instrum. Methods Phys. Res. A **544**, 565 (2005)

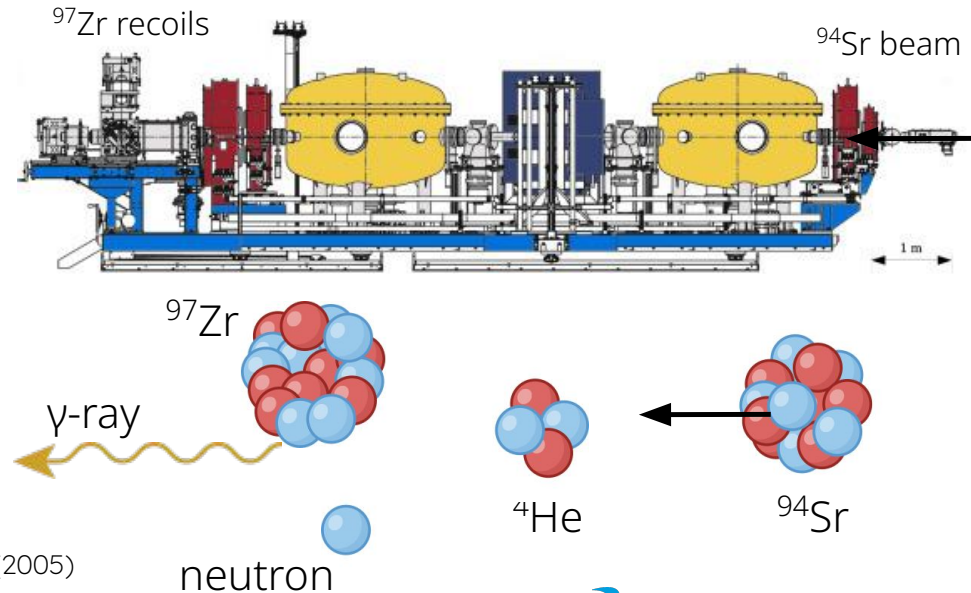
B. Davids et al., Nucl. Instrum. Methods Phys. Res A **930**, 191 (2019)

C.E. Svensson *et al.*, J. Phys. G **31**, S1663 (2005)

V. Godinho *et al.*, ACS Omega **1**, 1229 (2016)



First direct weak r-process measurement using a radioactive ion beam

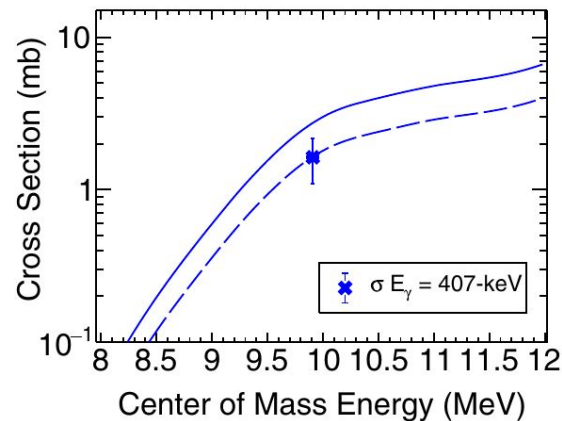
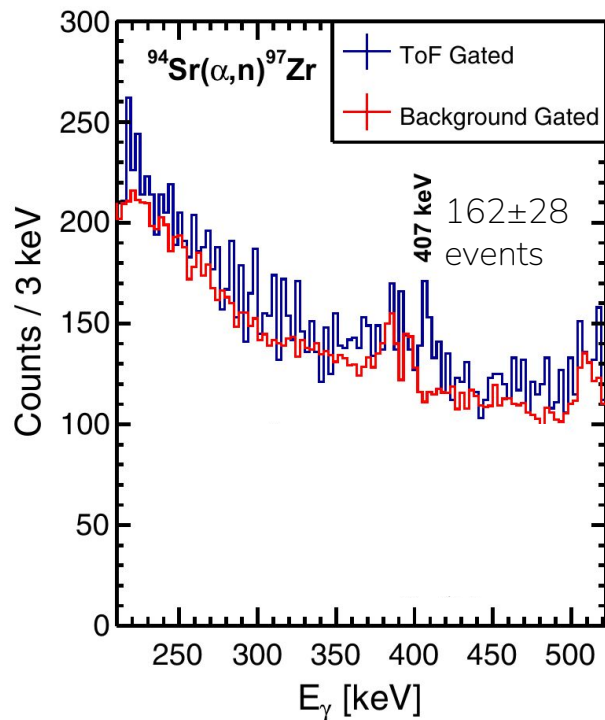
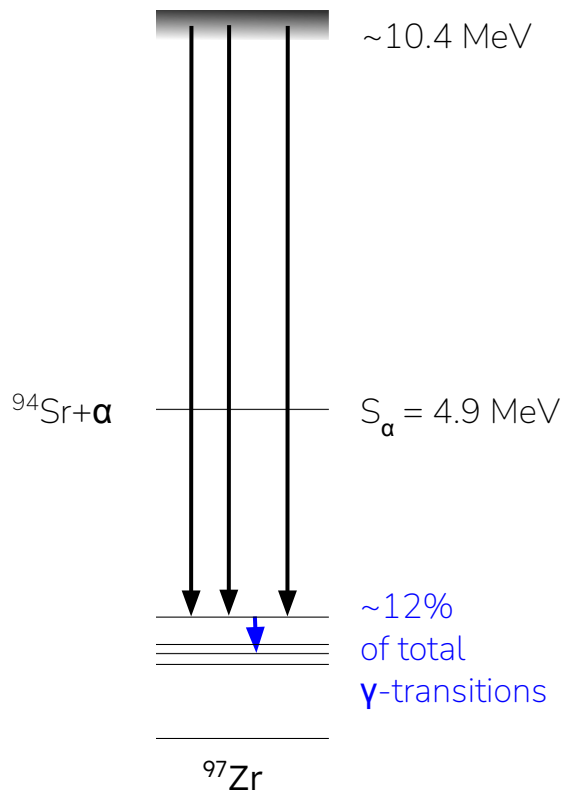


C.E. Svensson *et al.*, J. Phys. G **31**, S1663 (2005)

B. Davids & C.N. Davids, Nucl. Instrum. Methods Phys. Res. A **544**, 565 (2005)

B. Davids *et al.*, Nucl. Instrum. Methods Phys. Res. A **930**, 191 (2019)

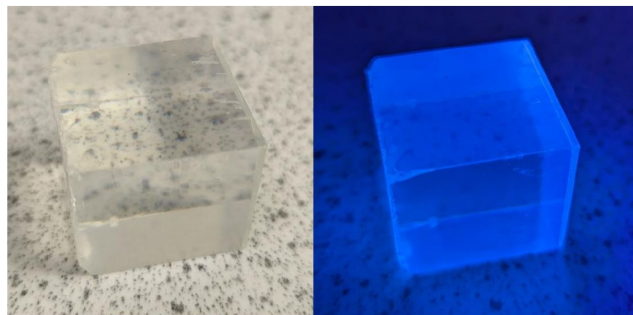
First direct weak r-process measurement using a radioactive ion beam



Matt Williams
(Rutherford Fellow at U Surrey)

DEMAND+EMMA+TIGRESS to measure (α, n) and (d, n) reactions

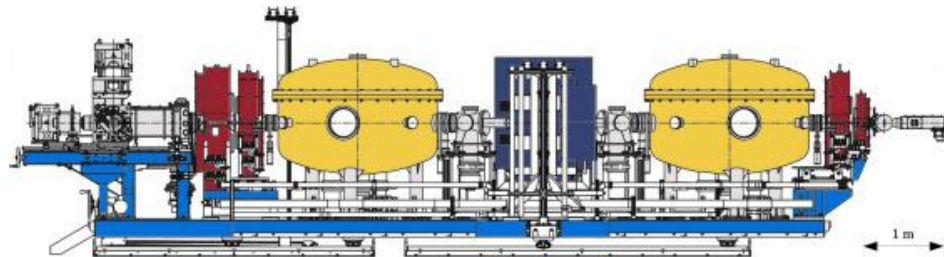
triple coincidence – heavy recoil (EMMA) + neutron (DEMAND) + γ -ray (TIGRESS)



Organic Glass Scintillator (OGS) produced by BlueShift Optics



Ben Reed
(SMU/TRIUMF postdoc)



DEMAND

Direct Experimental Measurements of Astrophysical reactions using Neutron Detectors

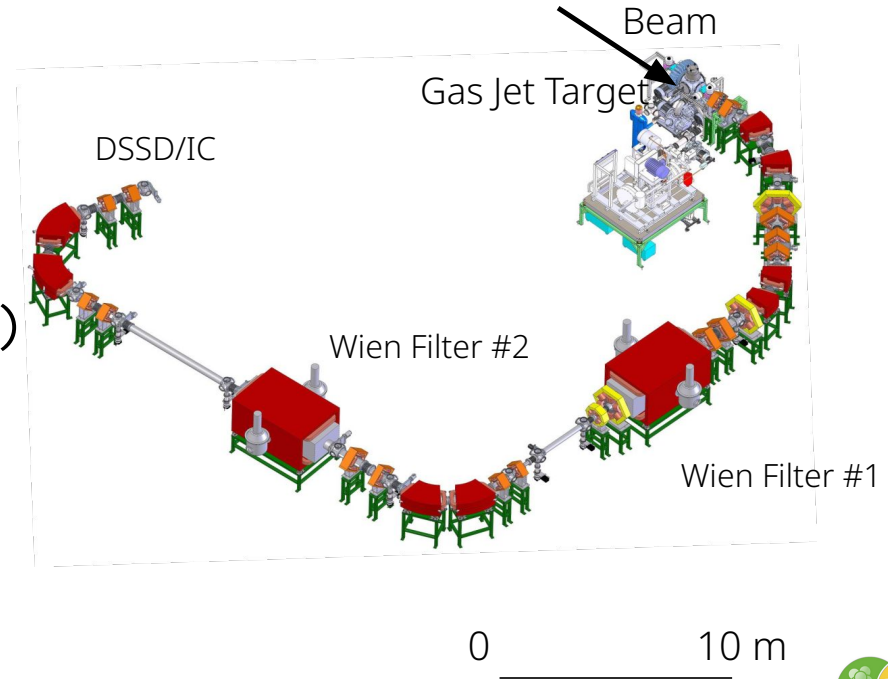


Many opportunities worldwide for nuclear astrophysics research!

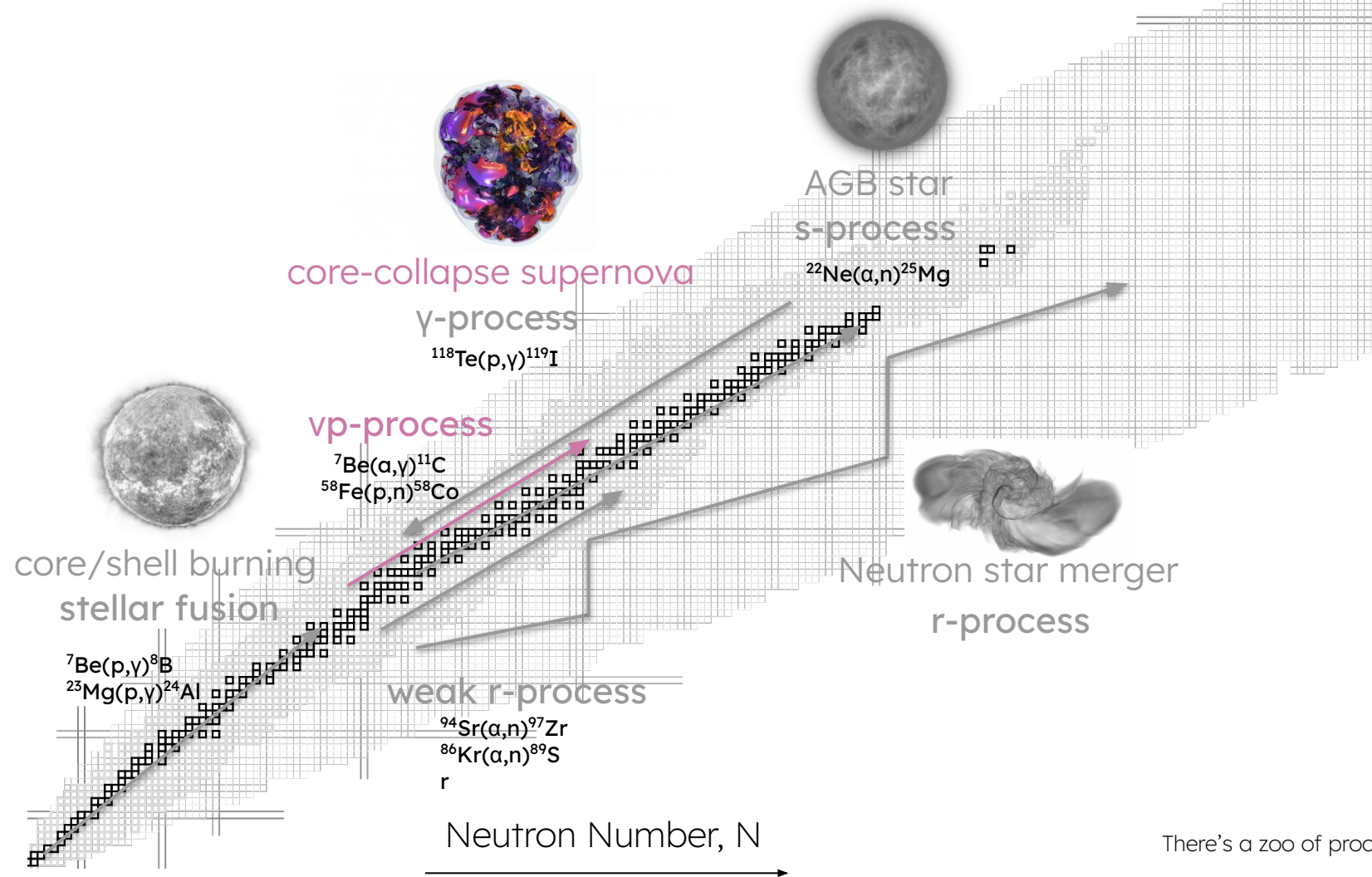
SECAR at FRIB

SEparator for CApture Reactions

- JENSA gas jet ($n \sim 10^{19}$ atoms/cm²) or extended windowless target
- BGO array (γ) or LENDA bars (n) for ejectile detection
- Two velocity (Wien) filters
- DSSD/IC at the last focal plane



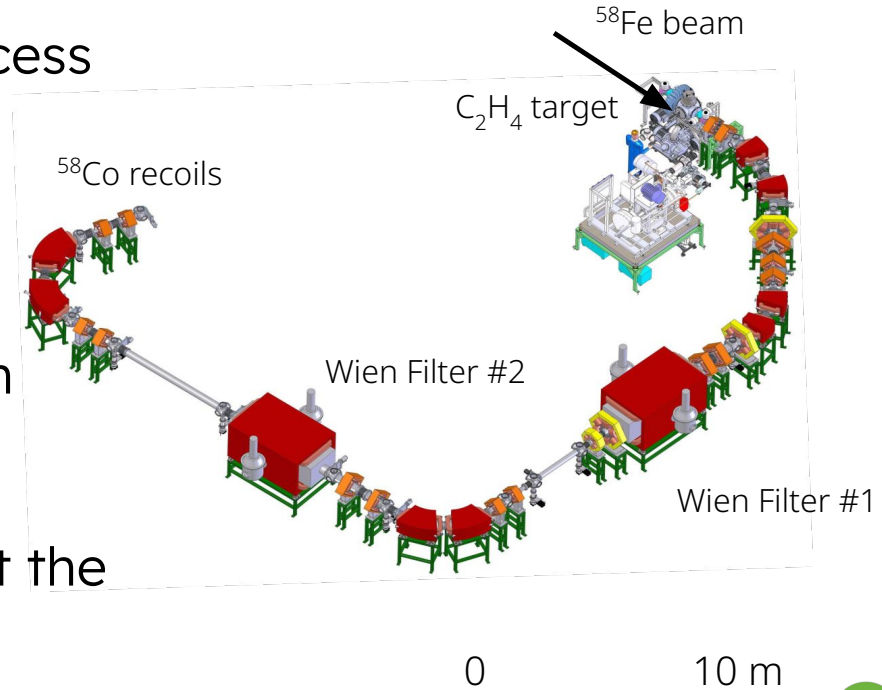
Proton Number, Z



There's a zoo of processes...

First results from SECAR: $^{58}\text{Fe}(p,n)^{58}\text{Co}$

- Study of the inverse of (n,p) reactions, relevant to the vp-process
- C_2H_4 solid target ($n \sim 8 \times 10^{18}$ atoms/cm 2)
- 4 organic scintillators for neutron detection
- Ionization chamber and DSSD at the focal plane

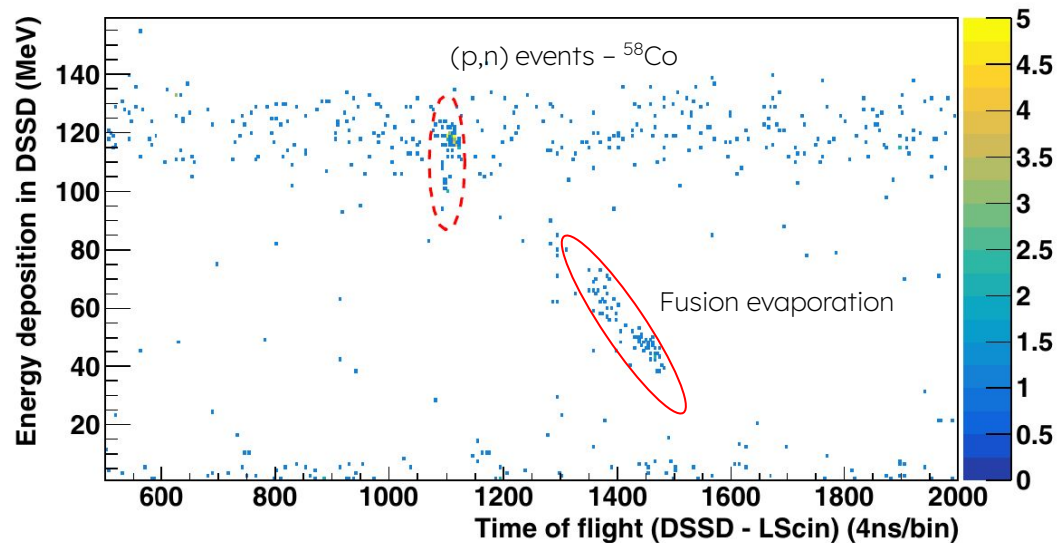


G. Berg et al., Nucl. Instrum. Methods Phys. Res. A **877**, 87 (2018)

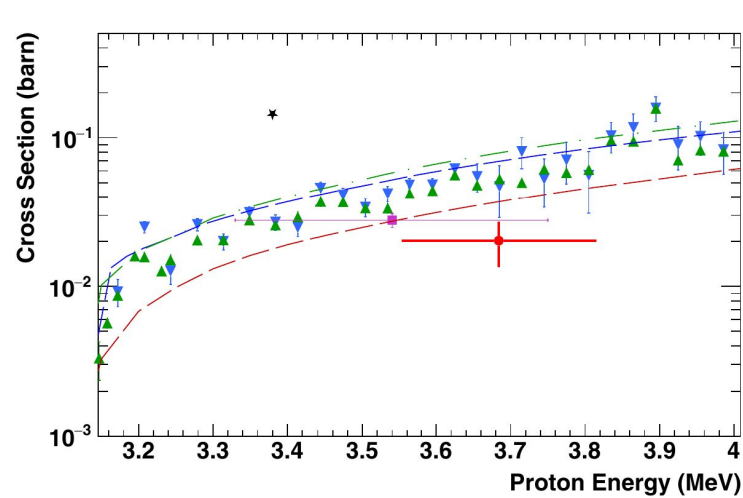
P. Tsintari et al., Phys. Rev. Res. **7**, 013074 (2025)

P. Gastis et al., Nucl. Instrum. Methods Phys. Res. A **985**, 164603 (2021)

First results from SECAR: $^{58}\text{Fe}(p,n)^{58}\text{Co}$



Focal Plane data



Final Cross Section

G. Berg et al., Nucl. Instrum. Methods Phys. Res A **877**, 87 (2018)
P. Tsintari et al., Phys. Rev. Res. **7**, 013074 (2025)
P. Gastis et al., Nucl. Instrum. Methods Phys. Res. A **985**, 164603 (2021)



Pely Tsintari
(Postdoc at FRIB)



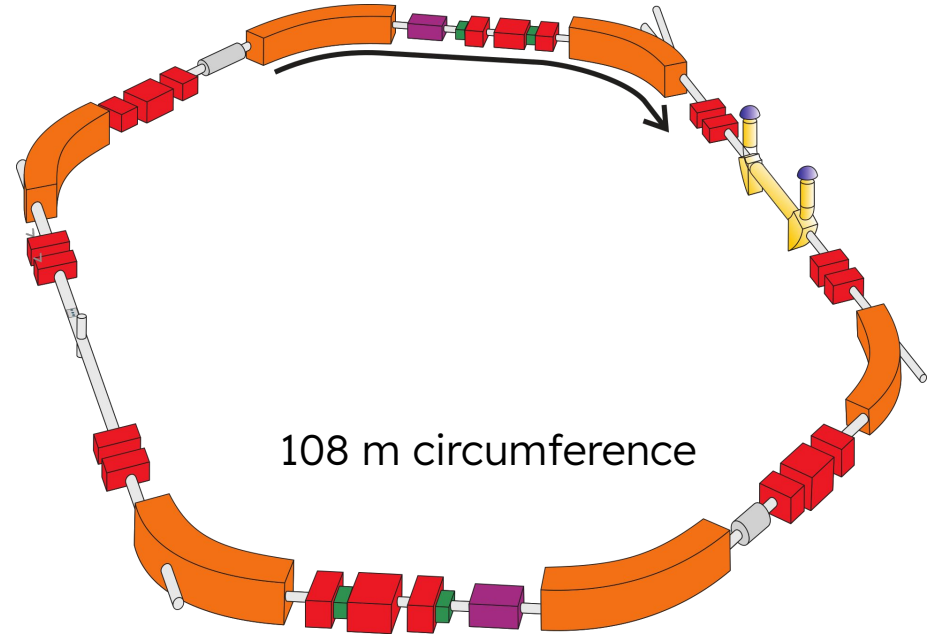


Many opportunities worldwide for nuclear astrophysics research!

ESR at GSI/FAIR

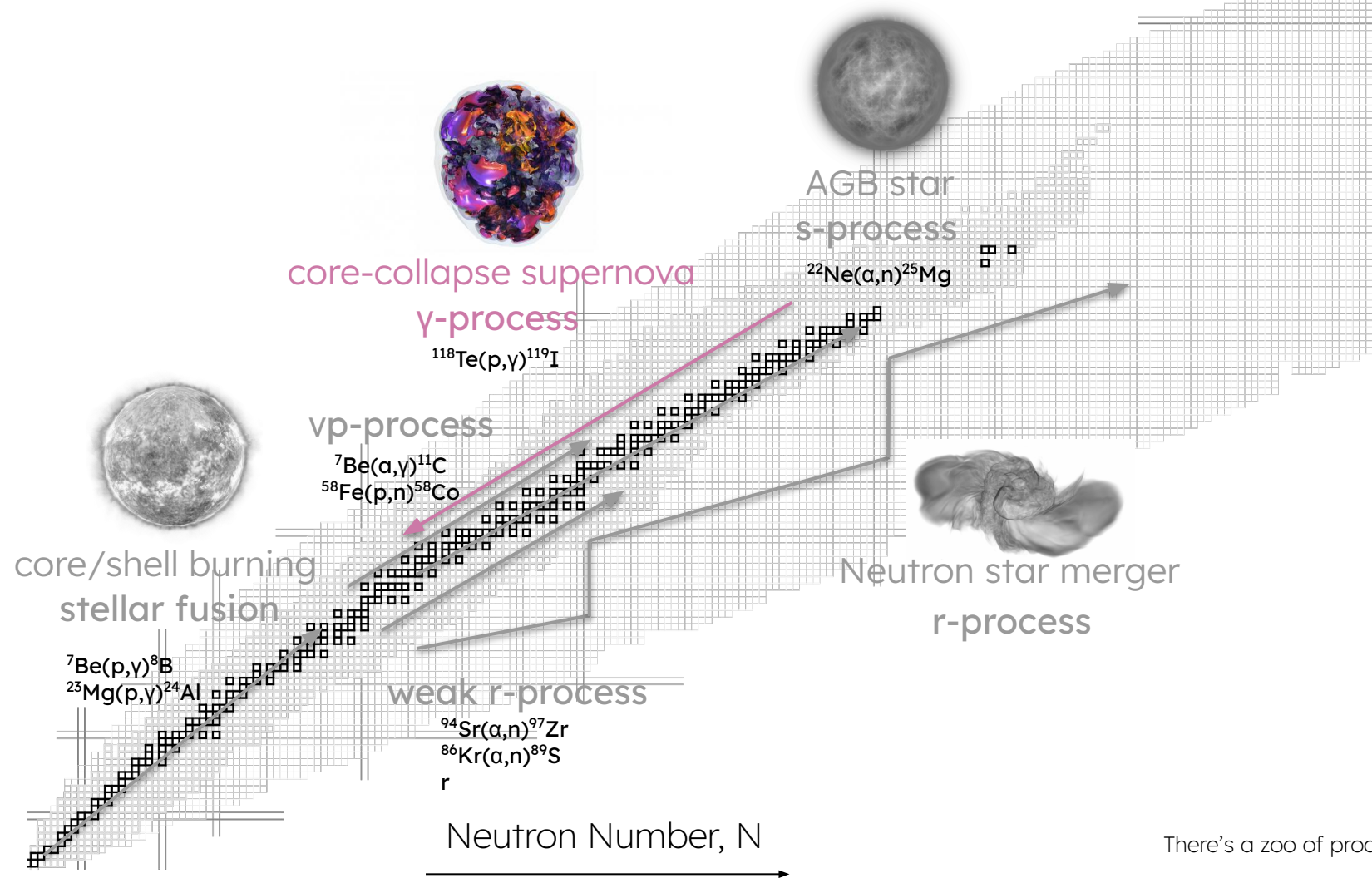
Experimental Storage Ring

- Thin hydrogen jet target
 $n = 10^{14}$ atoms/cm²
- Revolution frequency
 $f = 300\text{--}400$ kHz
- Beam is electron-cooled
to $\Delta p/p \sim 10^{-5}$
- Beam lifetime ~ 1.5 s (vacuum)



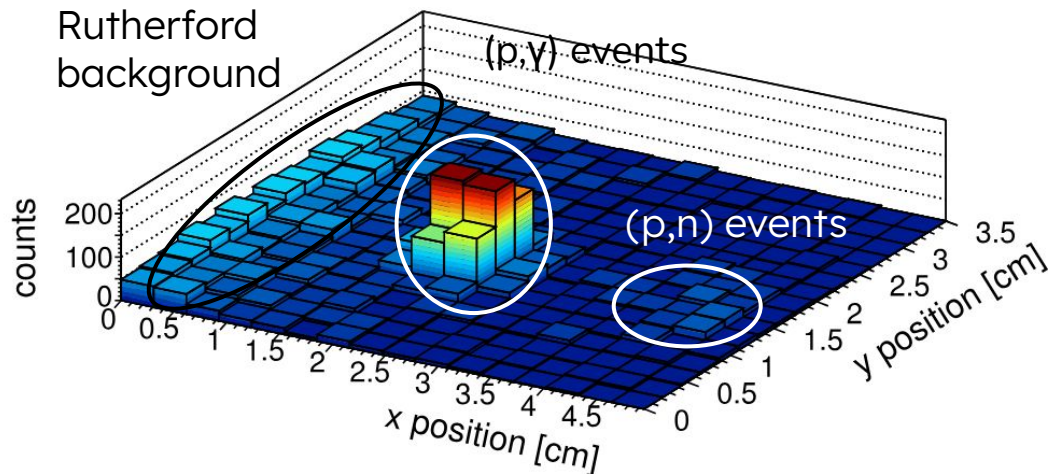
Primarily used for mass measurements and decay studies

Proton Number, Z

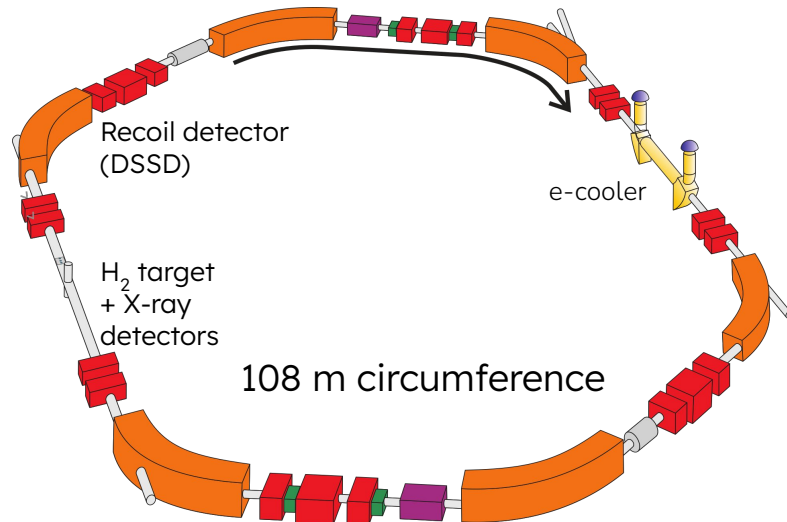


There's a zoo of processes...

First direct γ -process measurement using a radioactive ion beam and a storage ring



Events from both the (p, γ) and (p,n) reactions were detected!



J. Glorius *et al.*, Phys. Rev. Lett. **122**, 092701 (2019)
L. Varga *et al.*, Phys. Rev. Lett. **134**, 082701 (2025)
S.F. Dellmann *et al.*, Phys. Rev. Lett. **134**, 142701 (2025)



László Varga
(Postdoc at TU Munich)



Sophia Dellmann
(Postdoc at LANL)





Many opportunities worldwide for nuclear astrophysics research!

Come work with us at SMU!

Support available for undergraduate and graduate students. Lots of nuclear astrophysics projects to work on! 😊



Halifax, NS 🦞 🌊 ⚓

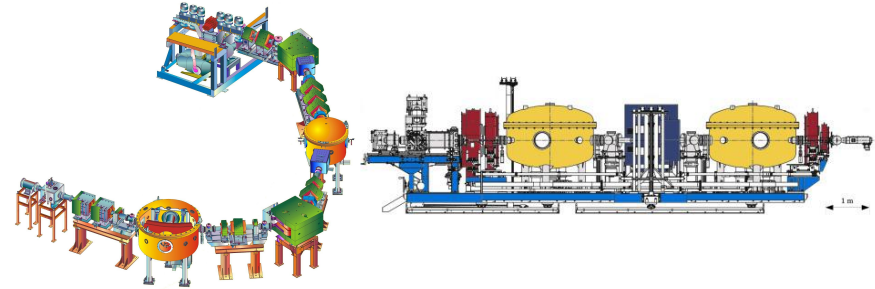
 <https://psaltisa.github.io>



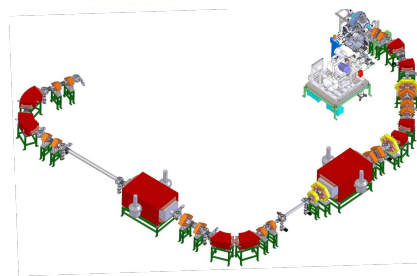
Acknowledgements



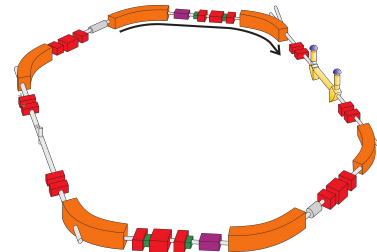
C. Ruiz, B. Davids, D. Hutcheon, A. Laird,
A. Lennarz, D. Connolly, **M. Williams**,
G. Lotay, **G. Christian**, **B. Reed**, A. Chen,
C. Brune, J. Fallis, **M. Loria**, M. Alcorta,
U. Greife, P. Machule, and the
DRAGON and EMMA collaborations



H. Schatz, F. Montes, **P. Tsintari**,
L. Wagner, K. Chipps, M. Couder, C. Deibel,
G. Perdikakis, J. Pereira, M. Smith, and the
SECAR collaboration



Y. Litvinov, J. Glorius, R. Reifarth,
S.F. Dellmann, **L. Varga**, I. Dillmann,
and the
ESR collaboration

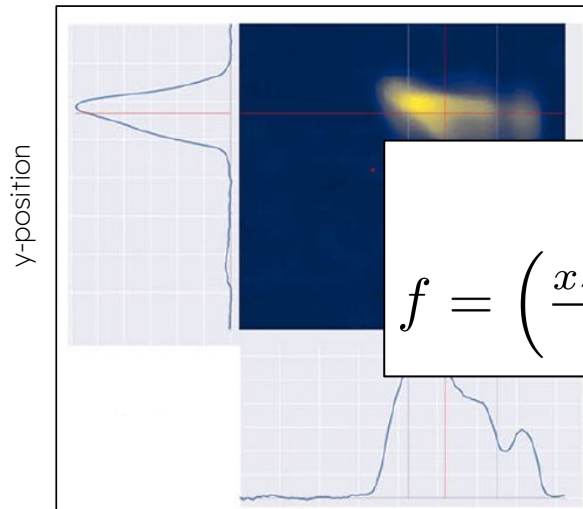


Summary & Outlook

1. Nuclear reactions affect stellar evolution in various environments.
2. Recoil separators are key instruments to study important astrophysical reactions.
3. Exciting opportunities arise in radioactive ion beam facilities around the world (ARIEL/TRIUMF, FRIB, FAIR and others).

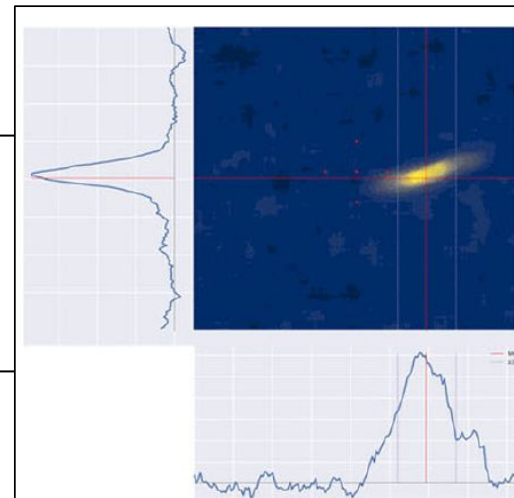
Use of ML for beam optimization at SECAR

Before optimization



x-position

After optimization



x-position

$$\text{minimize } f = \left(\frac{x_i + w_i^x - x_{\text{init}}}{20} \right)^4 +$$