

Francium

87

Uranium
&
Hydrogen

Baked into

One

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How do you end up with francium ?

- no stable isotope — shortest lifetime of first 103 elements
- only a few grams naturally on Earth
- why would we want to deal with this?

What plays the role of hydrogen in this problem?

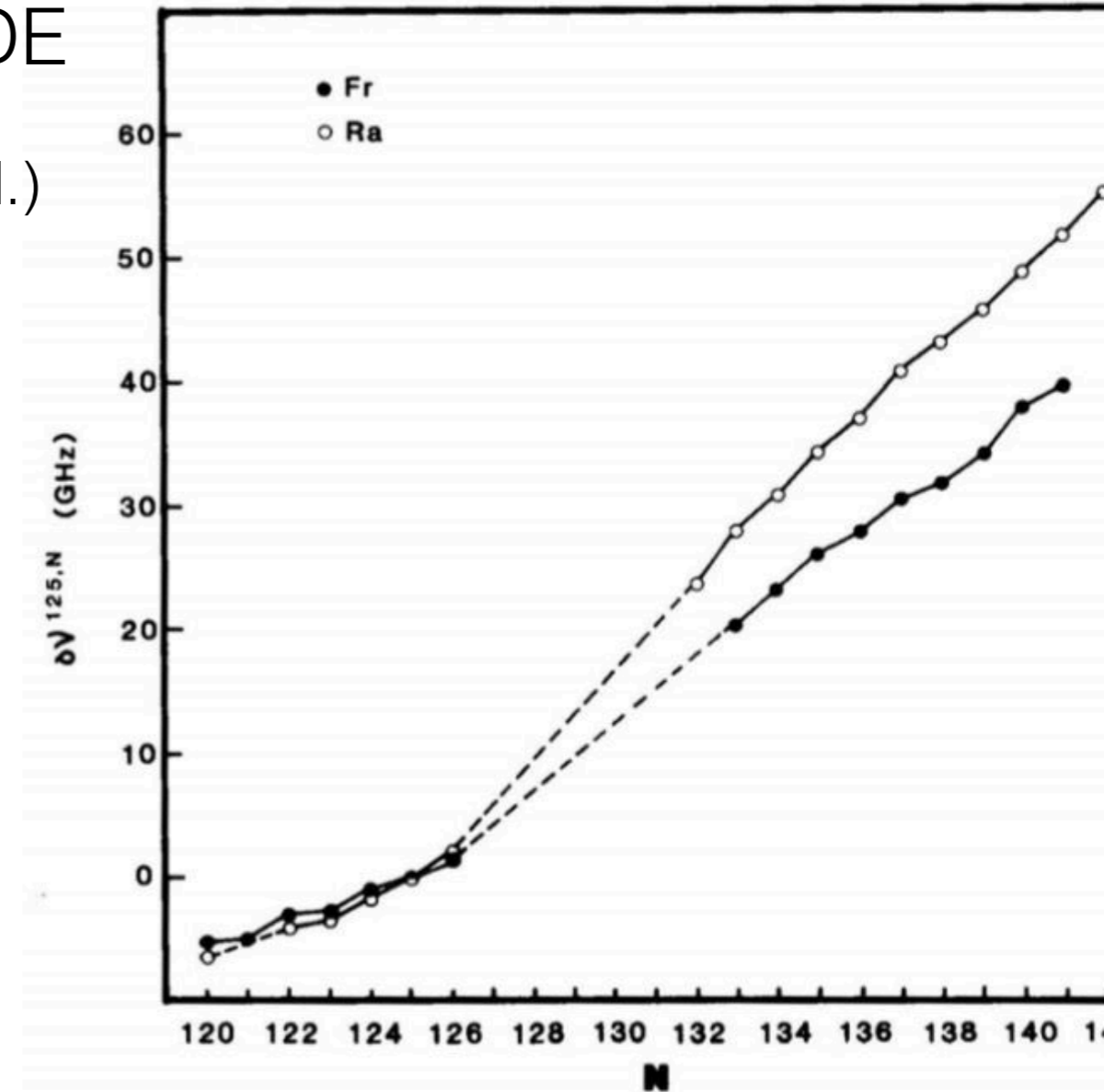
E. Fermi

(according to I.B. Khriplovich)

- many atomic tests of fundamental symmetries, need:
 - heavy nucleus (massive enhancement of effects such as parity violation or permanent EDMs)
 - simple atomic structure (aka alkali)
 - neutral atom or singly charged ion (precision spectr.)
- → Cs (historically) or Fr / Ra⁺ II (if you can afford it)

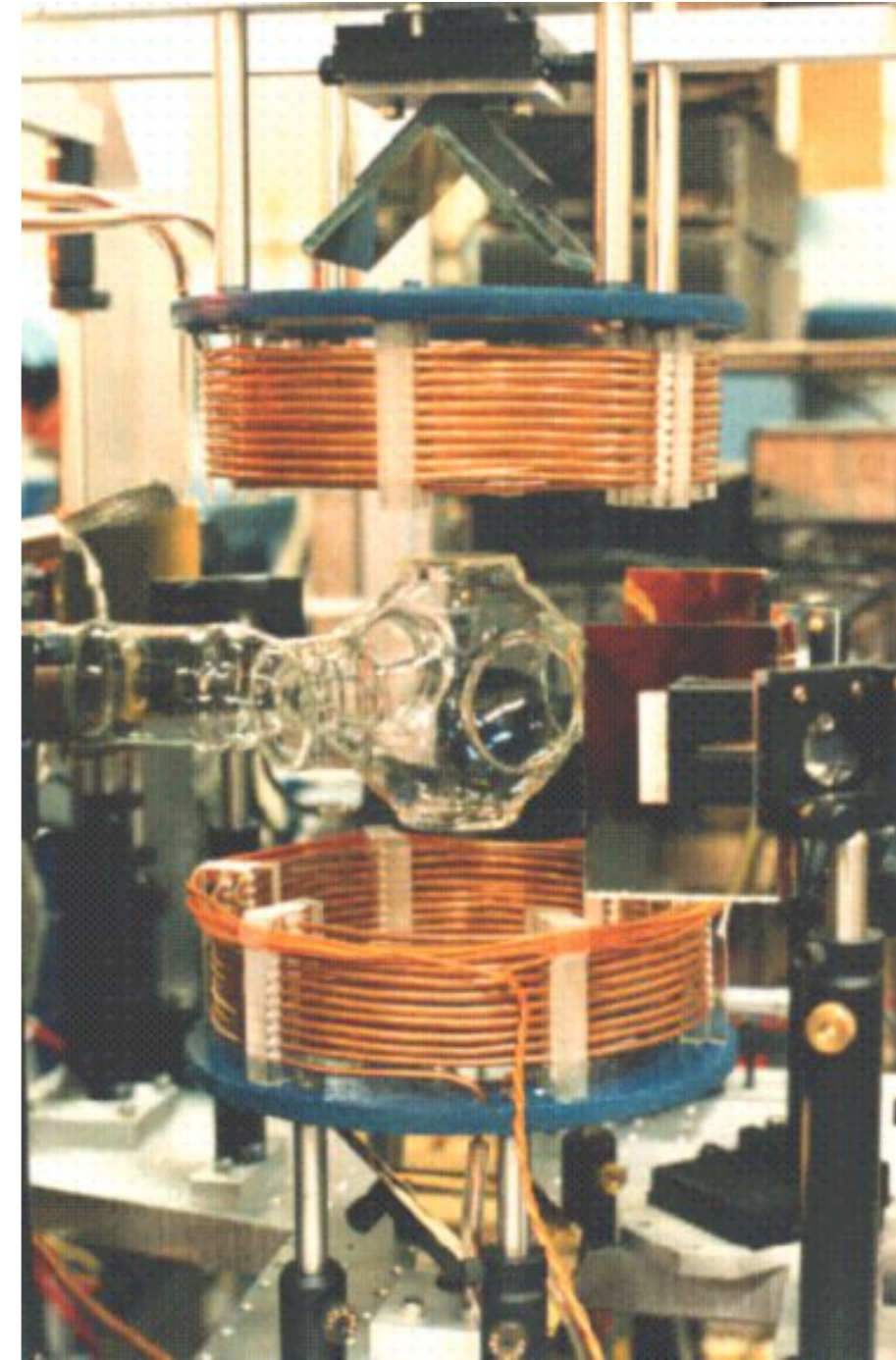
Brief history of pre-ISAC atomic Fr spectroscopy

- element discovered by Marguerite Perey (1939)
- atomic beam spectroscopy at ISOLDE
 - first observation, 1978 (Liberman et al.)
 - phase 1: 1978 - 1990
 - renewed interest: 2010's
 - mostly $7s_{1/2} - 7p_{3/2}$ isotope shifts → change in nucl. charge radius
- photo-ionization spectroscopy (Letokhov group, late 80s)
 - 1000 $^{221}\text{Fr}/\text{s}$ from ^{229}Th source
 - spectr. of high Rydberg states → ionization potential of Fr is 4.07 eV (Cs: 3.89 eV, Rb: 4.17 eV) → **relativity!**
- **Even at ISOLDE, not enough atom flux for forbidden transitions**



Brief history of pre-ISAC atomic Fr spectroscopy

- Stony Brook/Maryland group (1995-2005)
 - Gene Sprouse and Luis Orozco 1991
 - “Let’s use an atom trap”
 - atomic beam: few μ s of availability
 - magneto-optical trap: 10s of secs
 - 1995: first laser trapping of Fr
 - 8s, 9s states, lifetimes of excited states
 - $\approx 10^6$ Fr/s prod., $\approx 60,000$ trapped
- Boulder (Wieman, Lu et al.) 1997
 - laser trapped ≈ 900 ^{221}Fr from source
- Legnaro (early 2000s to present)
 - laser trap a few 1000 Fr, misc. spectr. of allowed transitions
- Tohoku, EDM; since 2010, getting ready to trap Fr



The ISAC Era

- Behr, d'Auria, Häusser, Jackson 1996
 - good Fr production at TISOL, think about ^{226}Fr trapping
 - puts TRIUMF on the Fr map!

- ISAC 2010: UC_x target at last!

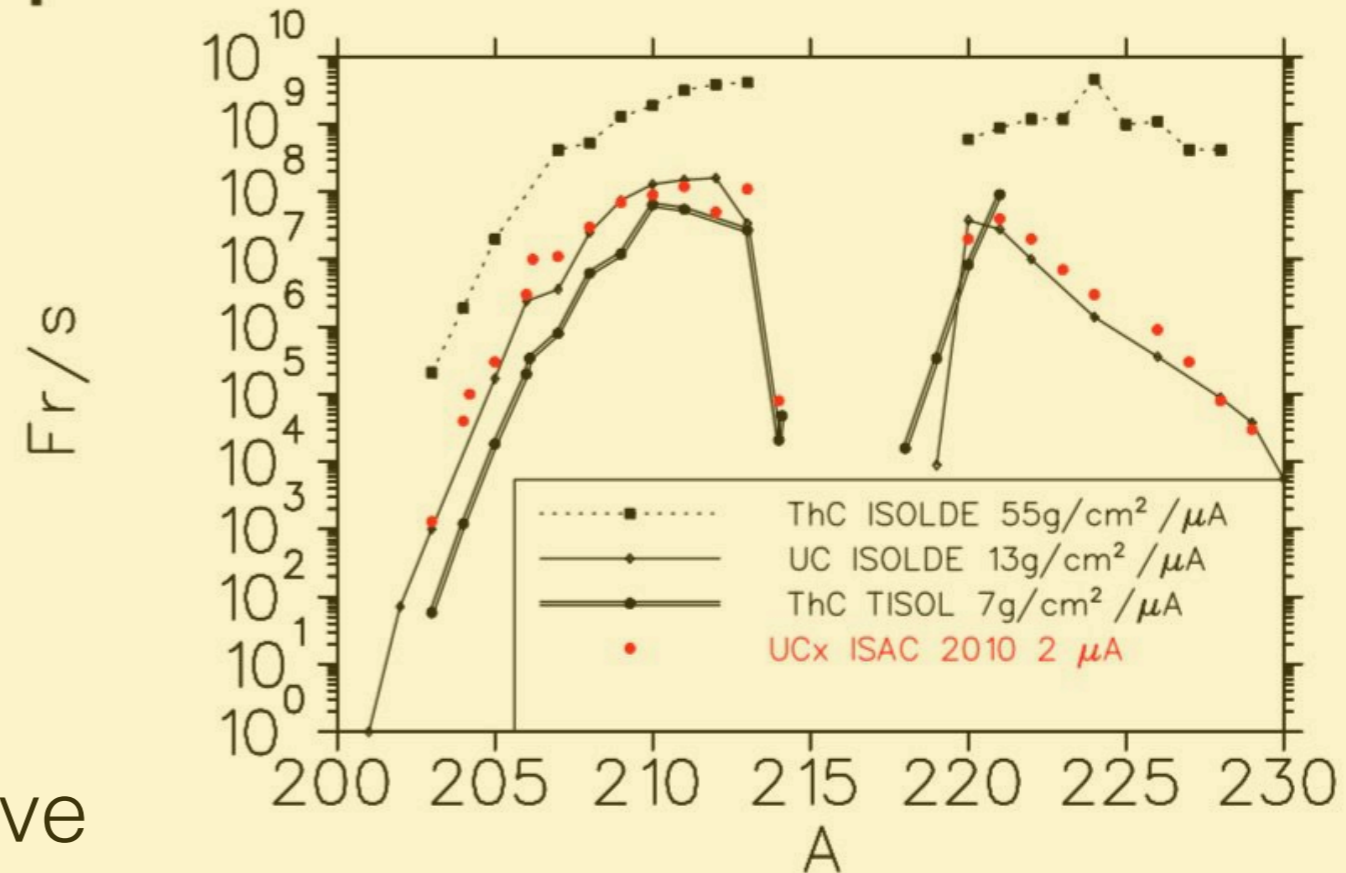
- Pearson group first out of the gate with collinear spectr.

- 2011 FrPNC beamline (thanks, Gordon!)

- 2012: FrPNC Fr laser trap goes live

- order of 10^6 trappable, in particular isotopes 207-213, which are 'nice' in terms of nuclear structure

TRINAT at TISOL: A good match for surface ion source for alkali production

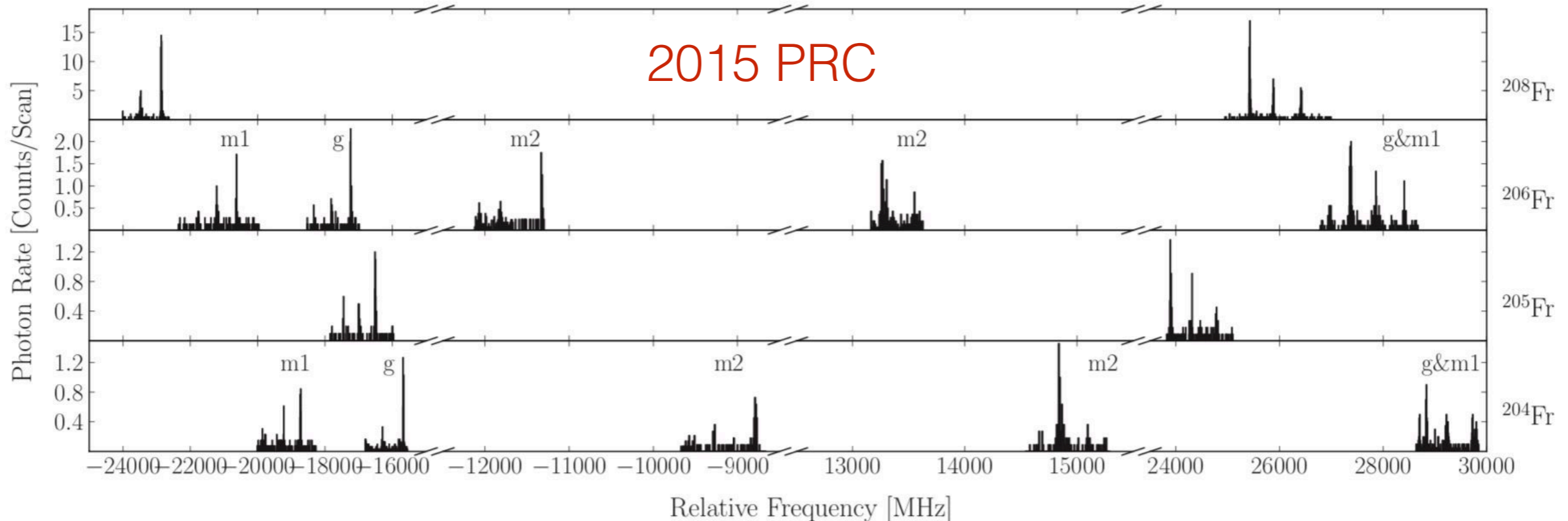
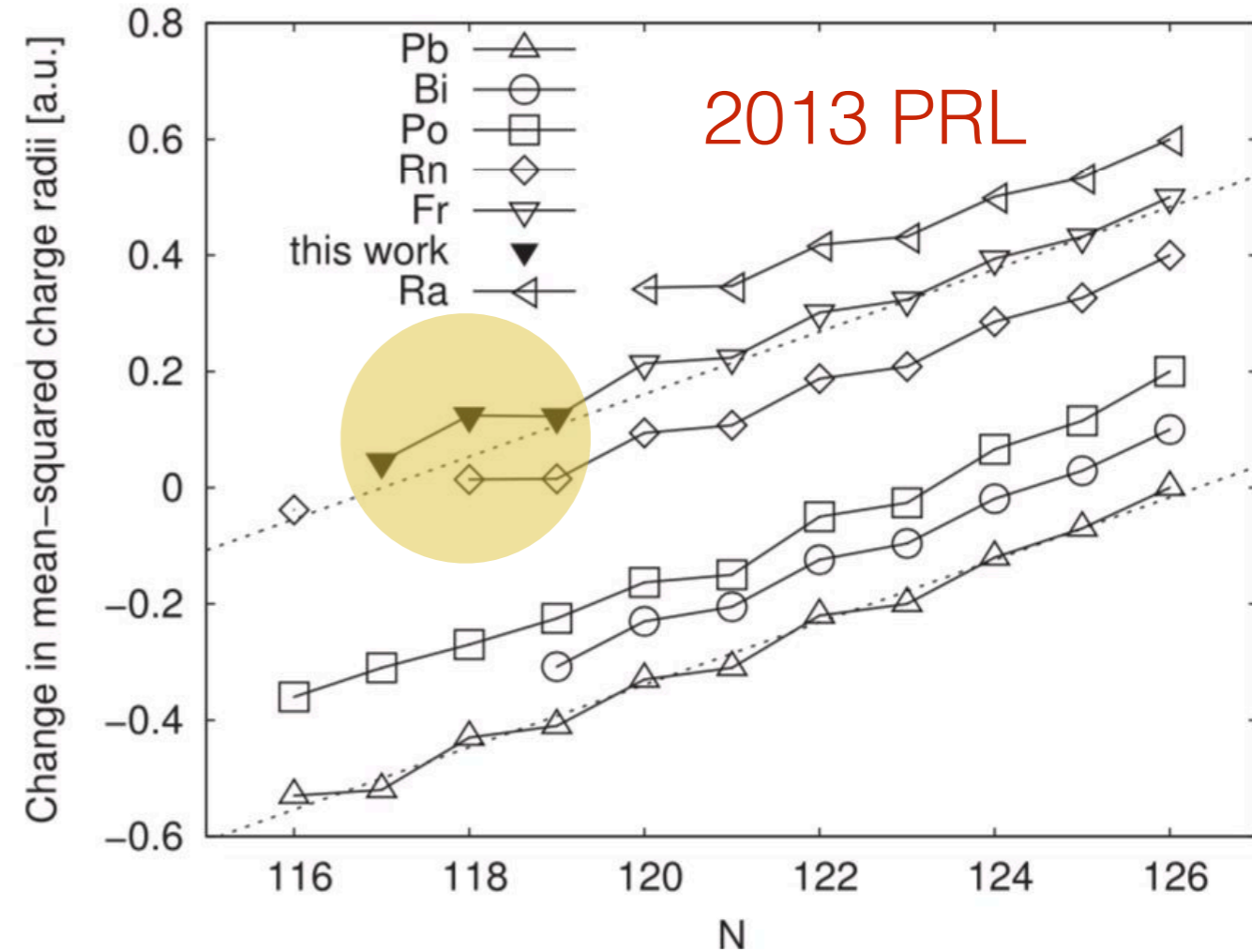


Collinear Fr spectroscopy at ISAC

- M. Pearson, A. Voss et al.
 - high-frequency intensity mod.
→ improved sensitivity

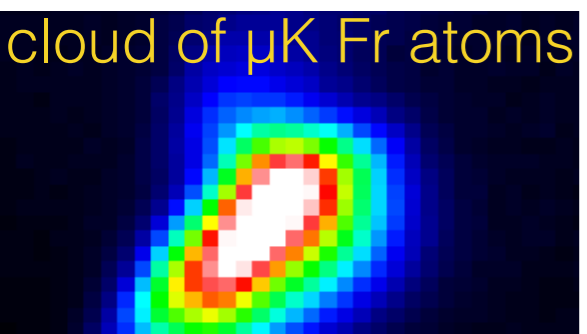
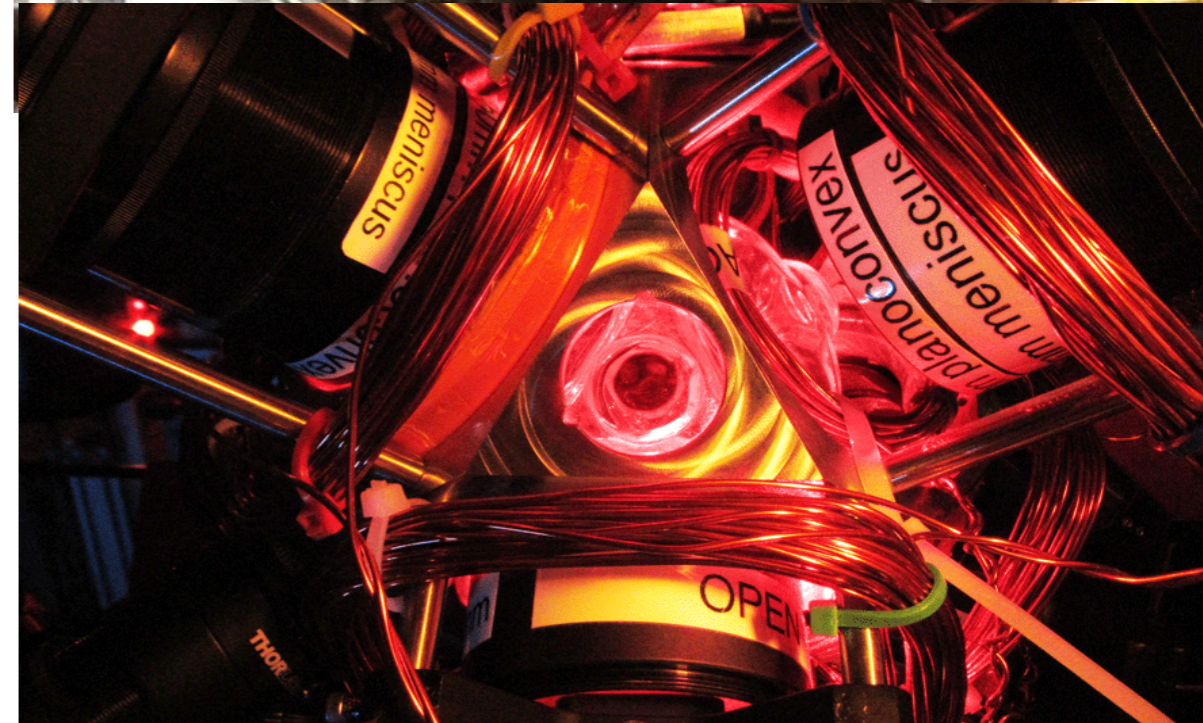
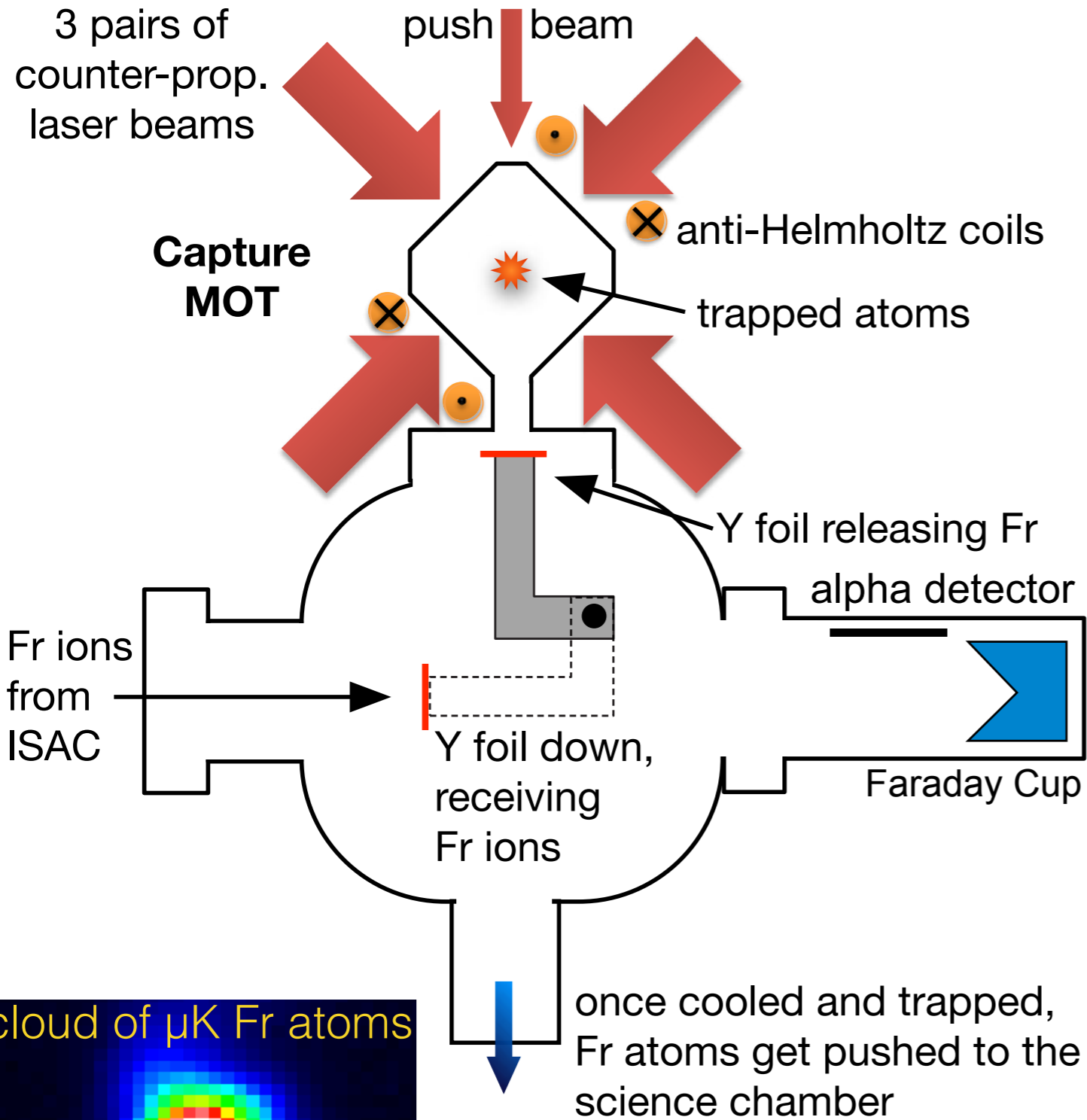
Voss et al. PRL 111, 122501 (2013)

Voss et al. PRC 91, 044307 (2015)



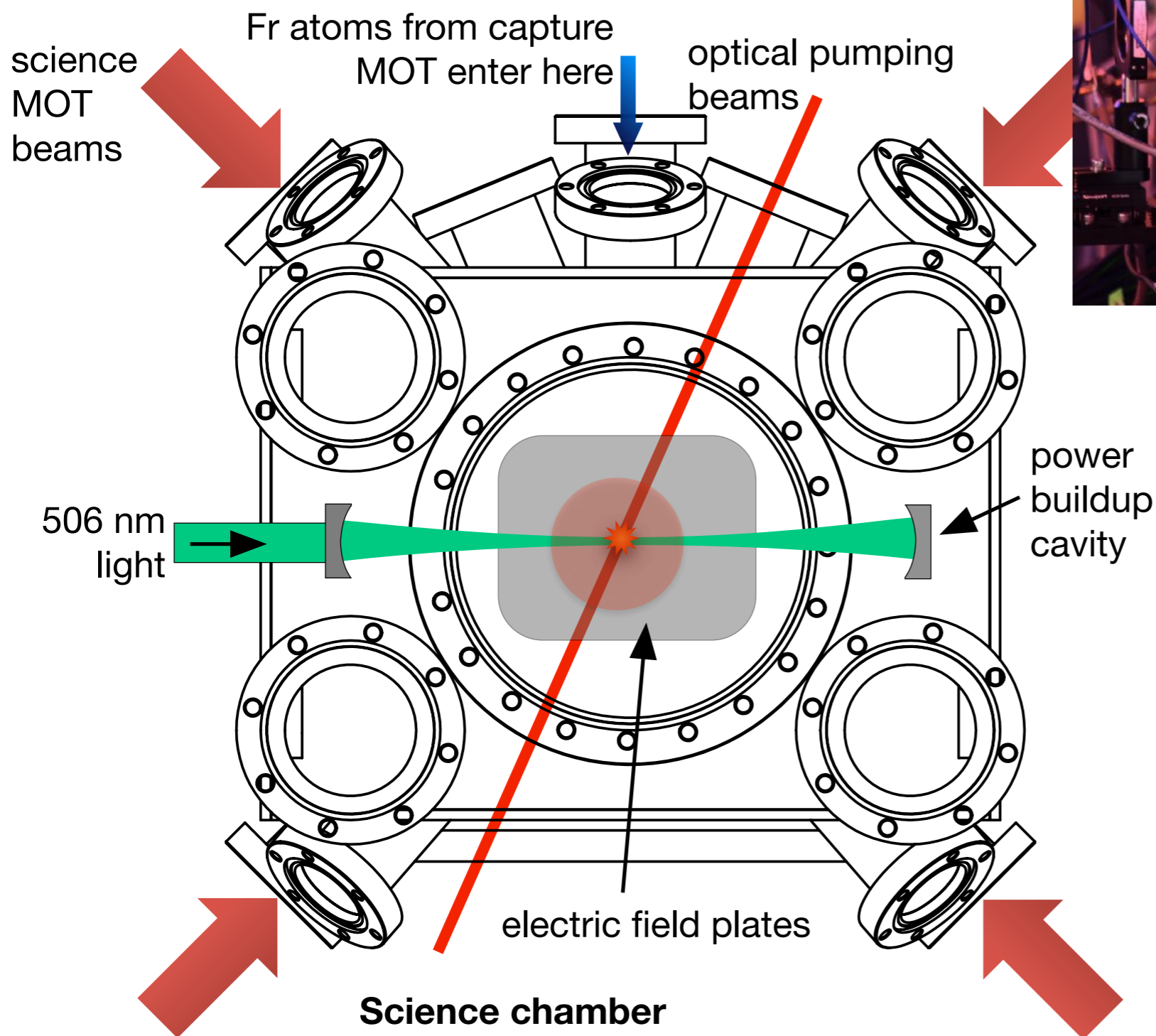
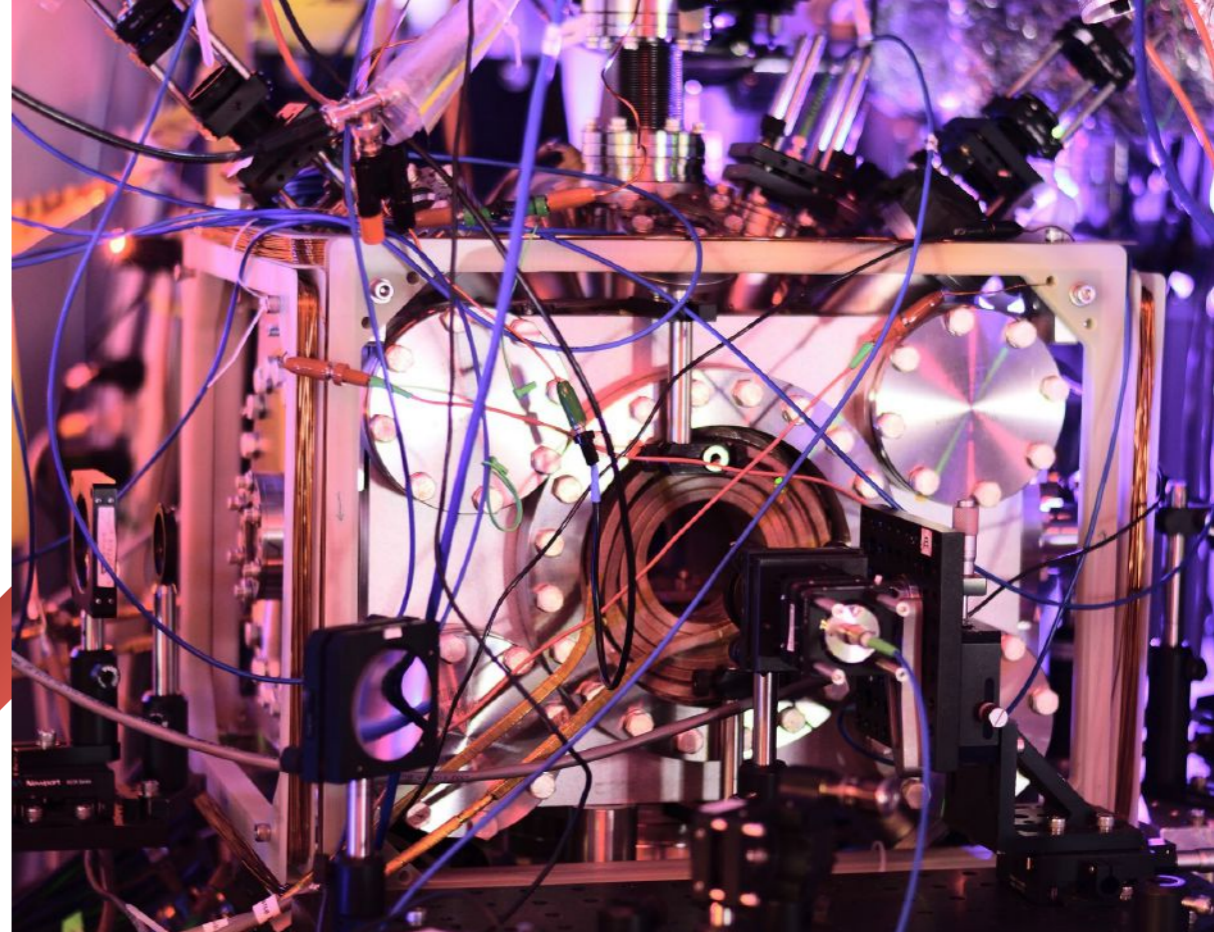
The Francium Trapping Facility at TRIUMF/ISAC

part 1: online capture trap



2017: sed -i 's/yttrium/zirconium/g' {} \;

Part2: Science chamber

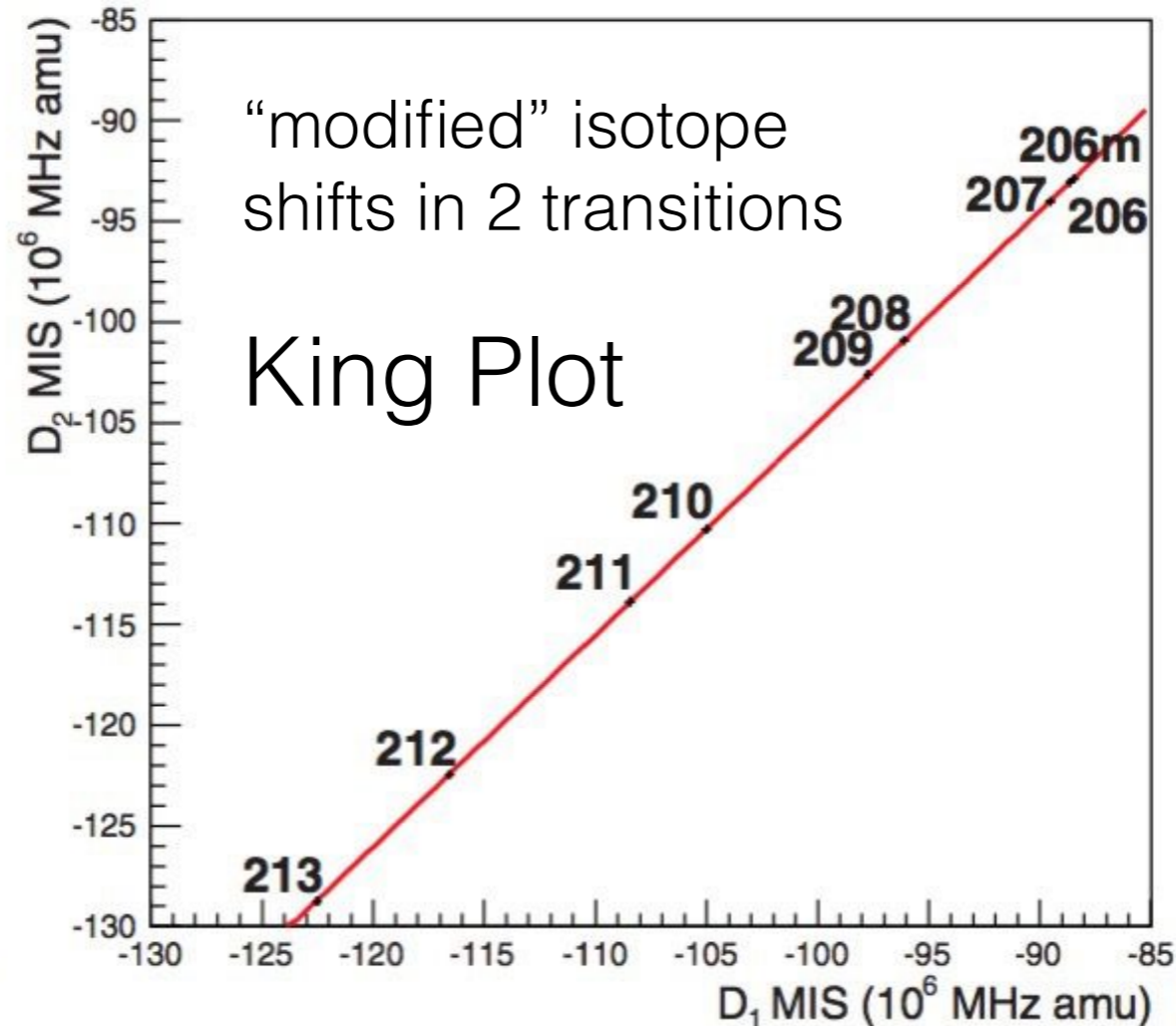
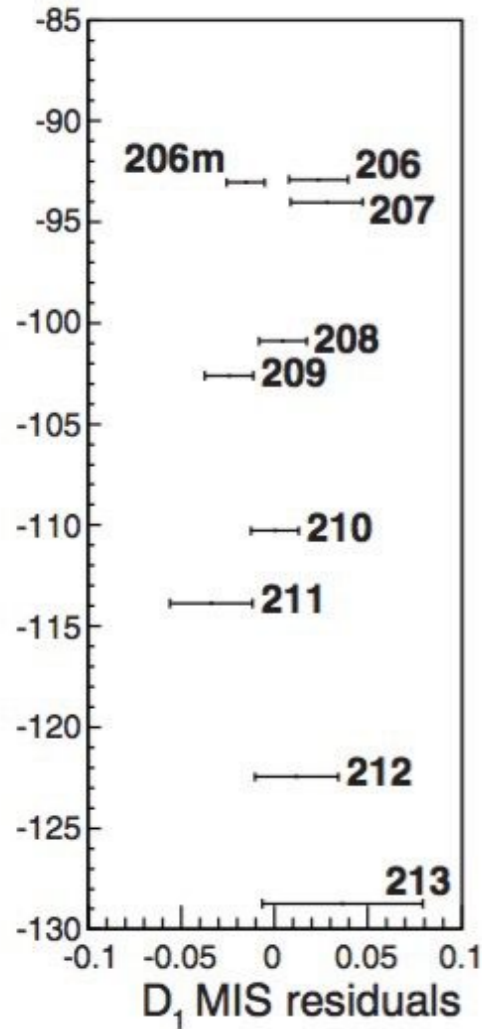


transparent field plates

operate APV experiment
inside MOT
temporal APV sequence

UHV-compatible
power buildup cavity
(1000x)

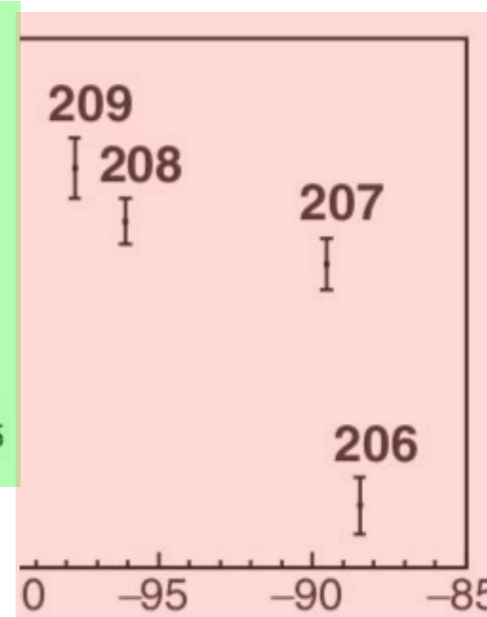
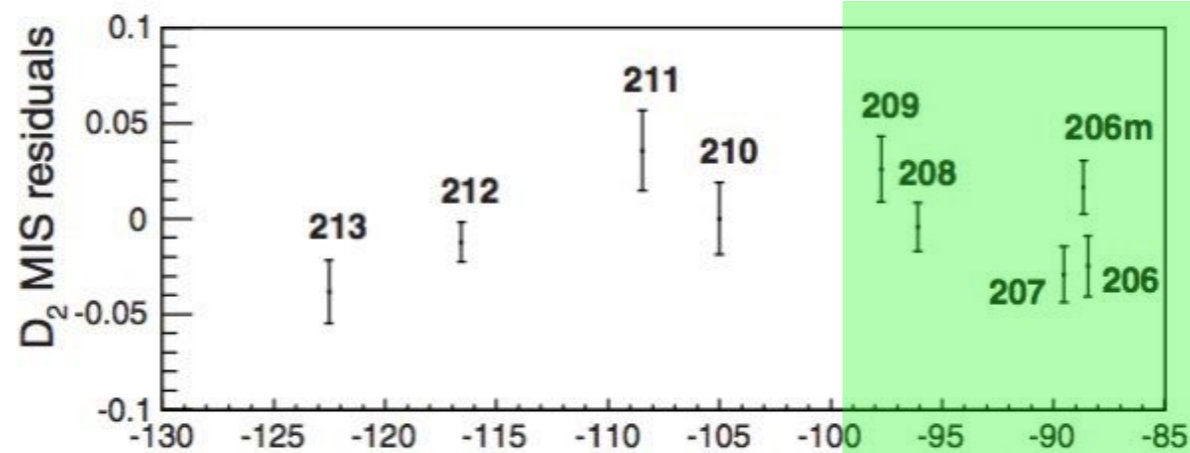
ISAC Fr trap: tune up with allowed transitions



D1 isotope shifts in a string of light francium isotopes

Benchmarks state-of-the-art atomic theory in Fr by Safranov and others.

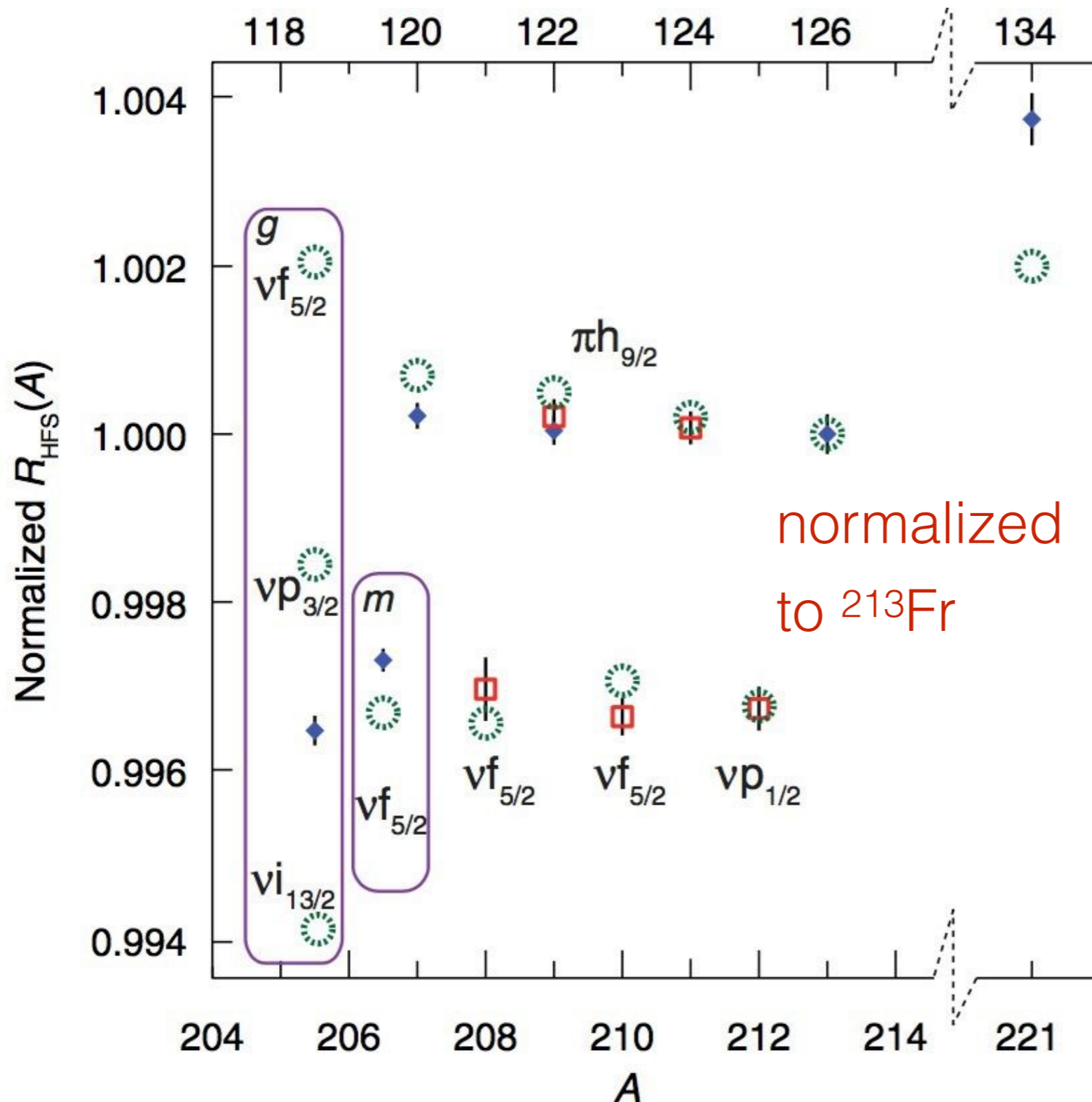
Fit Results
$\chi^2 / \text{ndf} = 7.00094 / 7$
slope = 1.0521 ± 0.0008
int = 194 ± 78 GHz amu



Fresh D2 data for 206 by Voss et al.

Hyperfine anomaly (Bohr-Weisskopf effect)

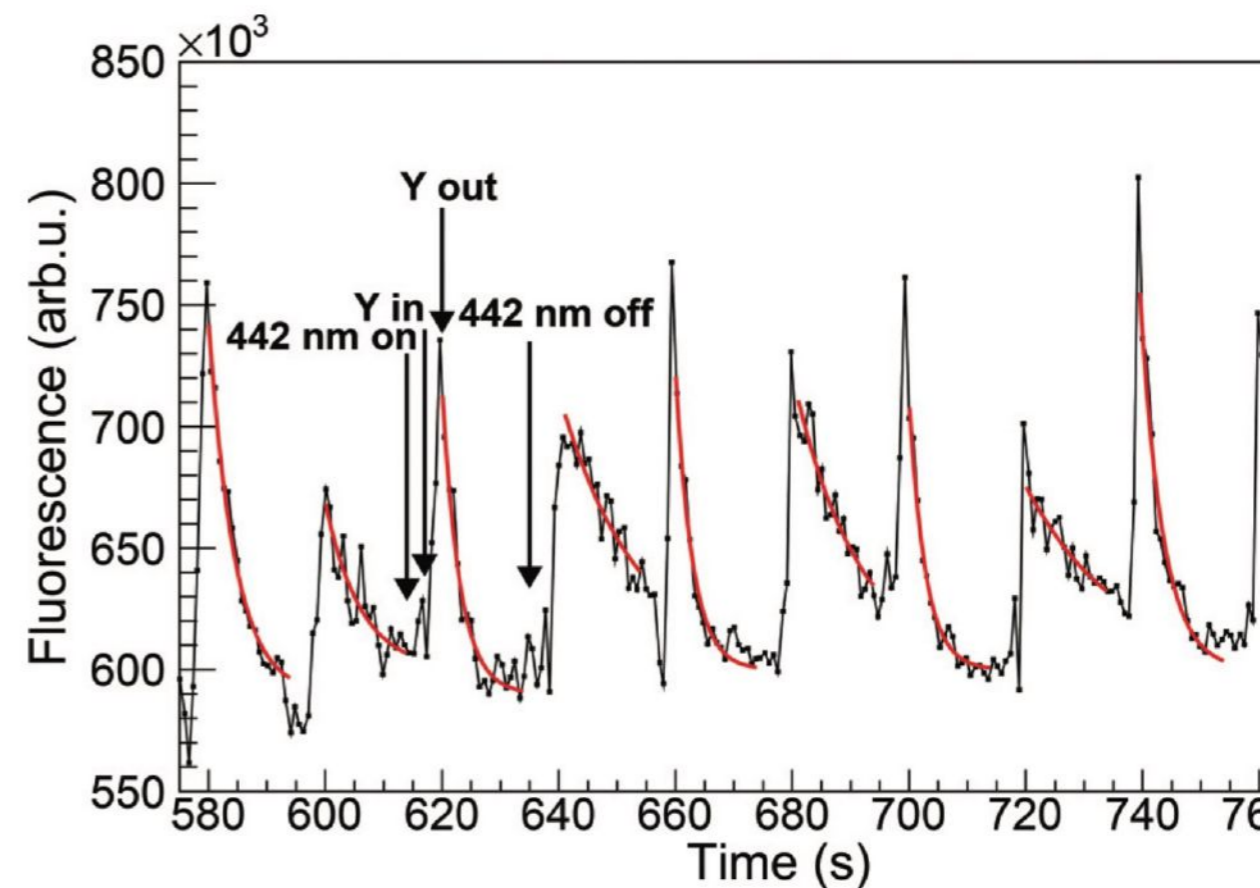
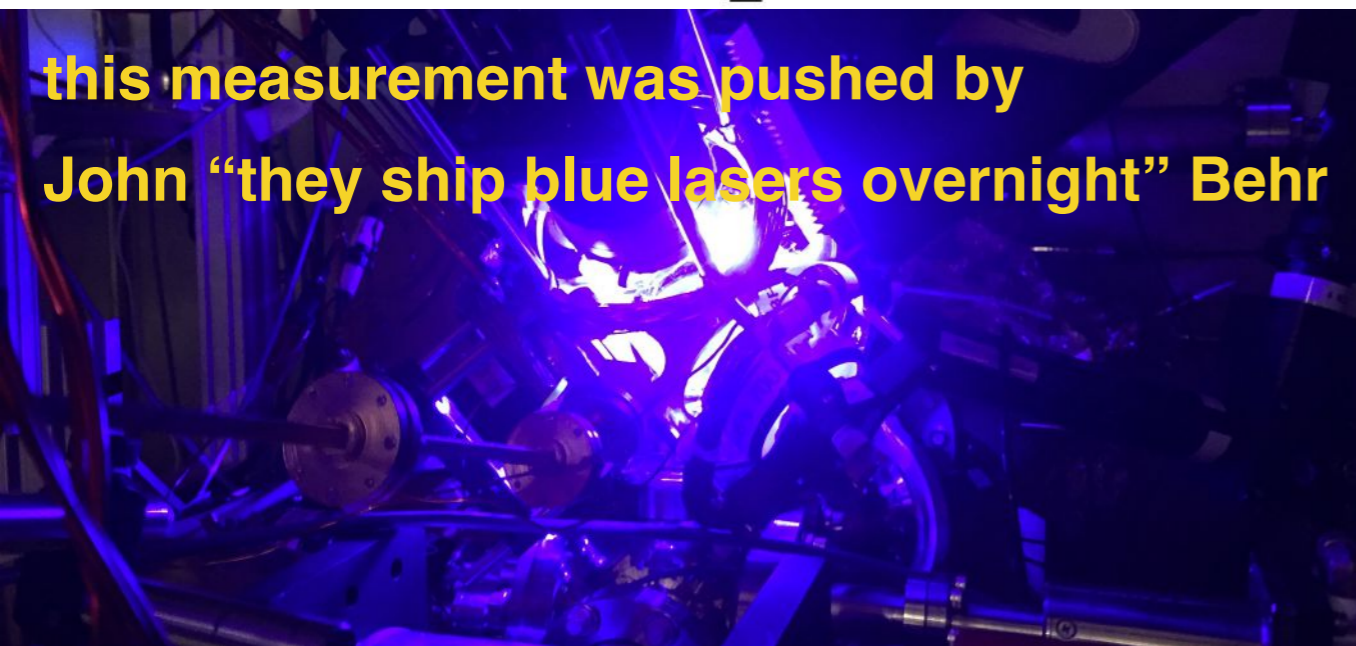
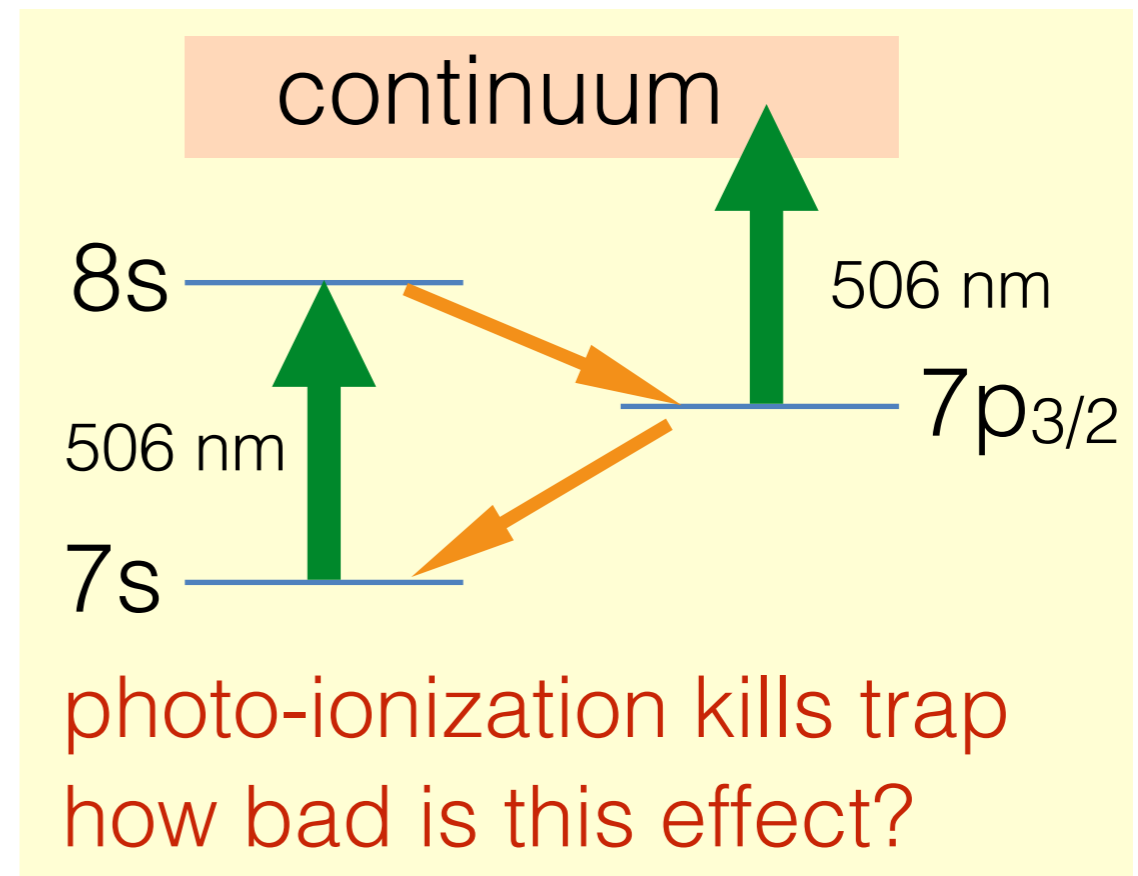
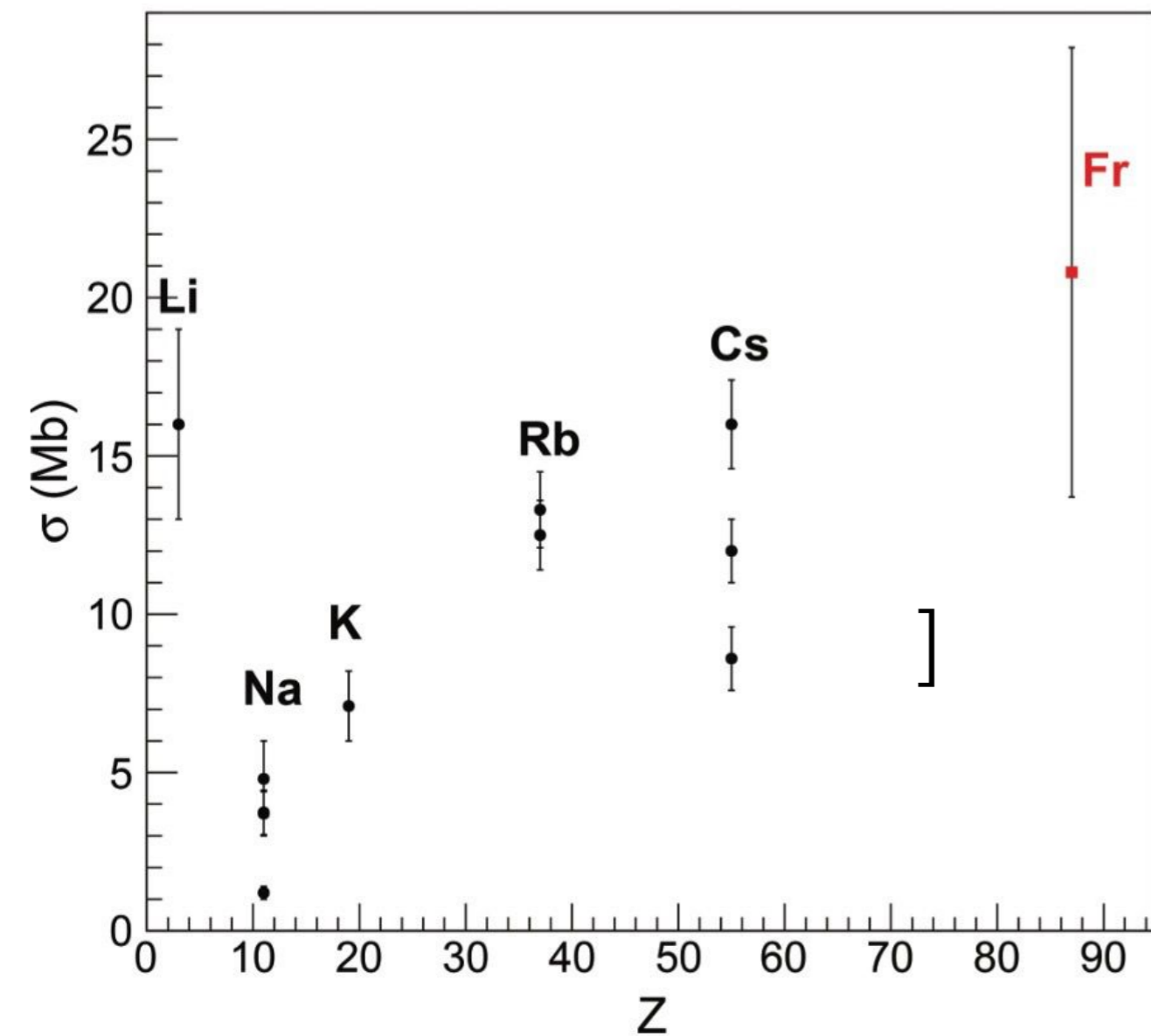
Zhang et al., Phys Rev Lett 115, 042501 (2015)



- nuclear magnetization distrib. differs from isotope to isotope
- \rightarrow ratio of $7p_{1/2}$ and $7s_{1/2}$ hyperfine splittings isotope dependent

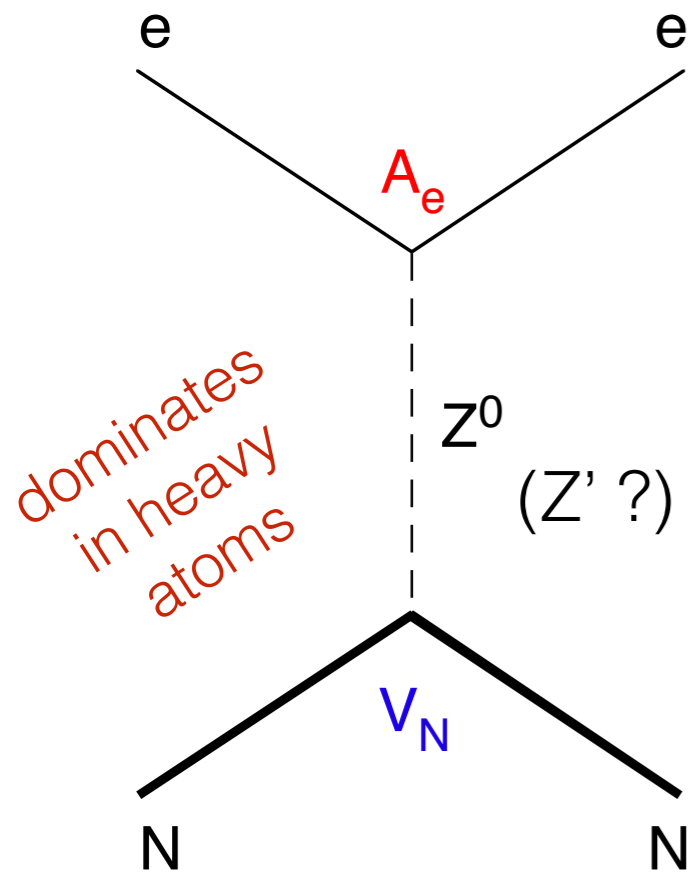
Reconfirms that in terms of nuclear structure, 208-213 are “good” nuclei for APNC/anapoles

Francium $7p_{3/2}$ photo-ionization — Collister et al. 2017, Can J Phys



Atomic parity violation

- Z-boson exchange between atomic electrons and the quarks in the nucleus



nuclear spin indep. interaction:
coherent over all nucleons

H_{PNC} mixes atomic s & p states

$\langle n's' | H_{\text{PNC}} | np \rangle \propto Z^3$ times a relativistic factor
Drive s \rightarrow s E1 transition!

 Bouchiat & Bouchiat 1974

Z^3 law easy enough to motivate

The NSI APV Hamiltonian for a point-like nucleus

$$H_{\text{PNC}}^{\text{nsi}} = \frac{G}{\sqrt{2}} \frac{Q_W}{2} \gamma_5 \delta(\mathbf{r}) \quad Q_W = 2(\kappa_{1p}Z + \kappa_{1n}N)$$

$$\kappa_{1p} = \frac{1}{2}(1 - 4 \sin^2 \theta_W), \kappa_{1n} = -\frac{1}{2}$$

The "nuclear weak charge" Q_w
contains the weak interaction physics

≈ 0.02

$\rightarrow \text{APV} \approx Q_w(n)$

$$\langle n'L' | H_{\text{PNC}}^{\text{nsi}} | nL \rangle = \frac{G}{\sqrt{2}} \frac{Q_w}{2} \langle n'L' | \delta(r) \vec{\sigma} \cdot \vec{p} | nL \rangle$$

$$\propto \langle n'L' | \frac{d}{dr} | nL \rangle |_{r=0} \quad R_{nL} \approx r^L Z^{L+1/2}$$

\Rightarrow at $r = 0$ only $R_{ns}, \frac{d}{dr} R_{np}$ are finite

H_{PNC} mixes s and p states $\langle ns | H_{\text{PNC}}^{\text{nsi}} | n'p \rangle \propto Z^3$
Bouchiat, 1974

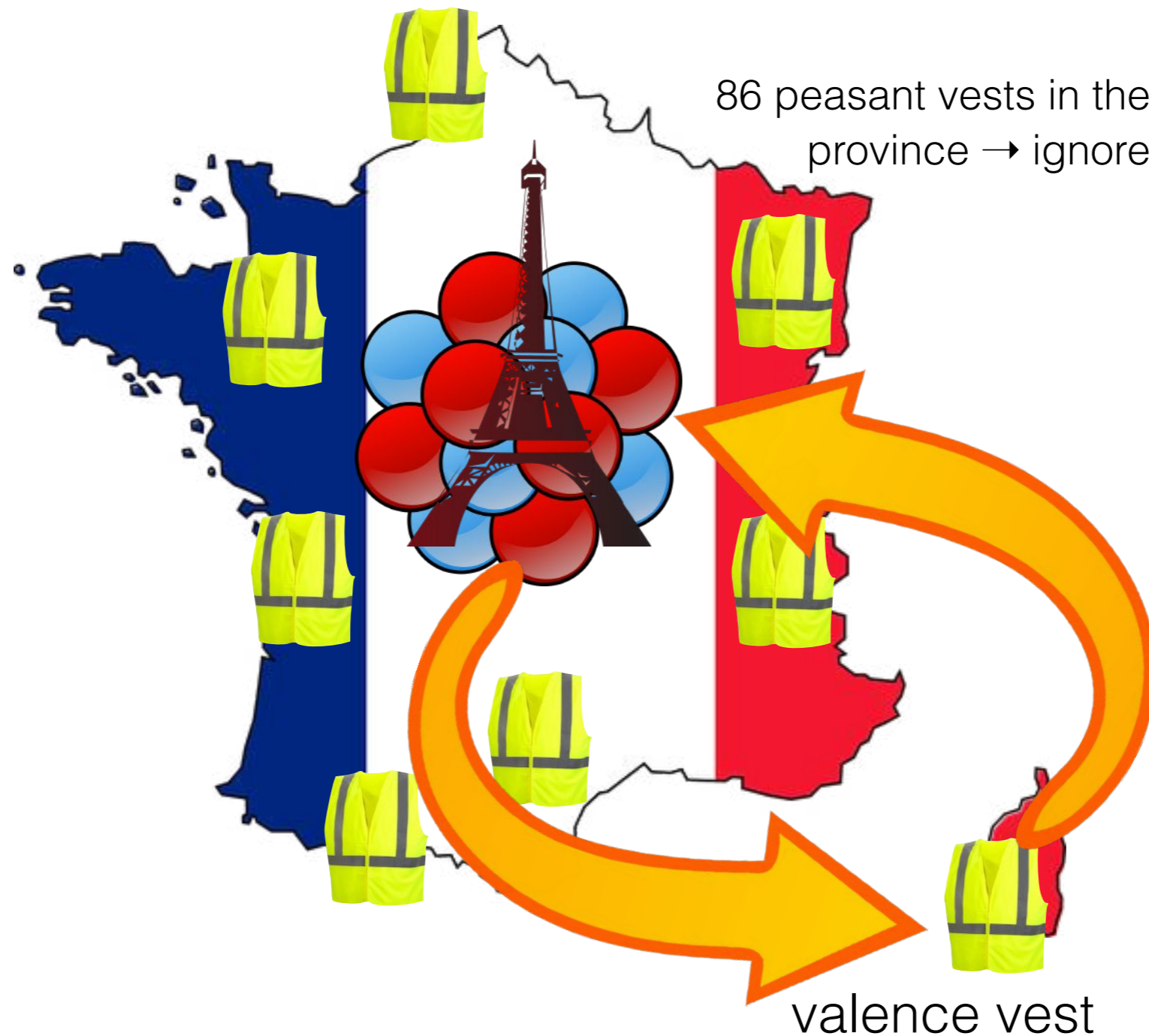
but let's try some intuition

Atomic parity violation in Fr

France and the Rutherford atom are very similar:

Gigantic thing in the middle and nothing worth mentioning around it

(according to the inhabitants of the gigantic thing — don't shoot the messenger)

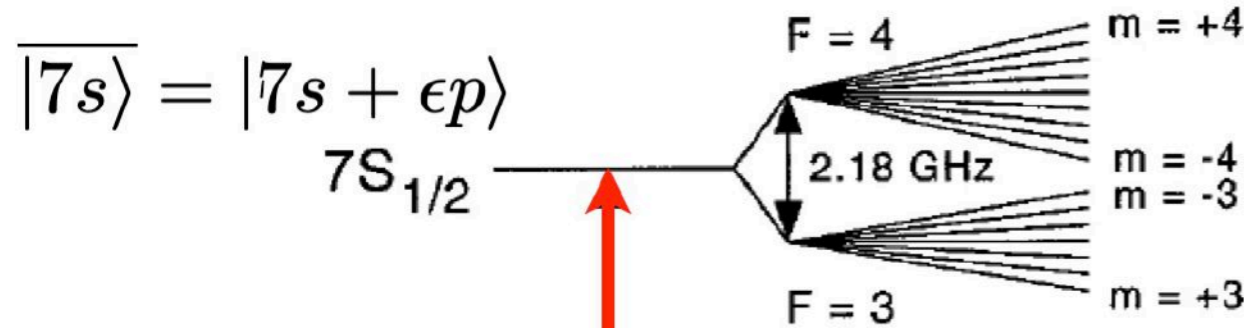


How is this measured ?

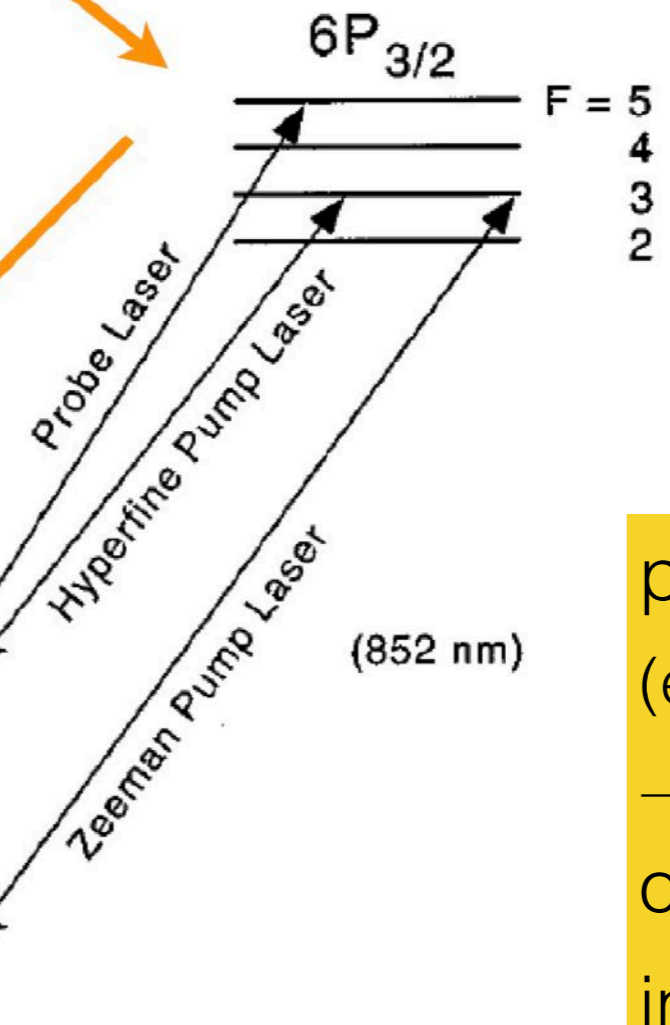
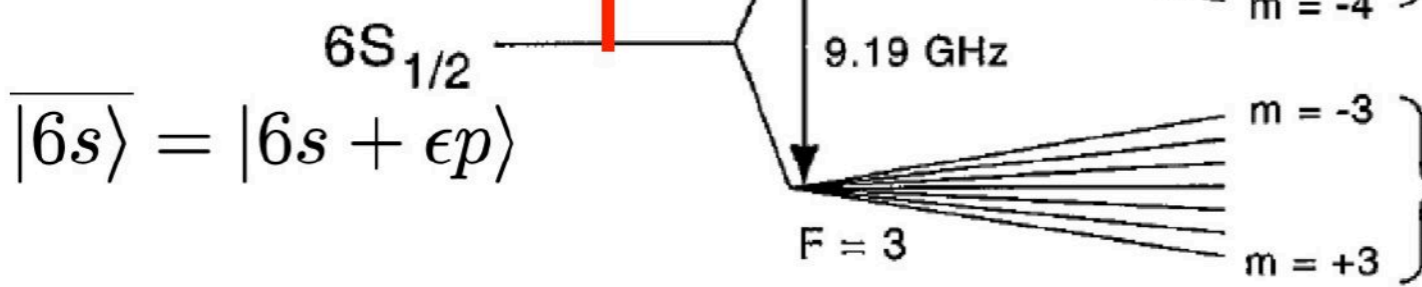
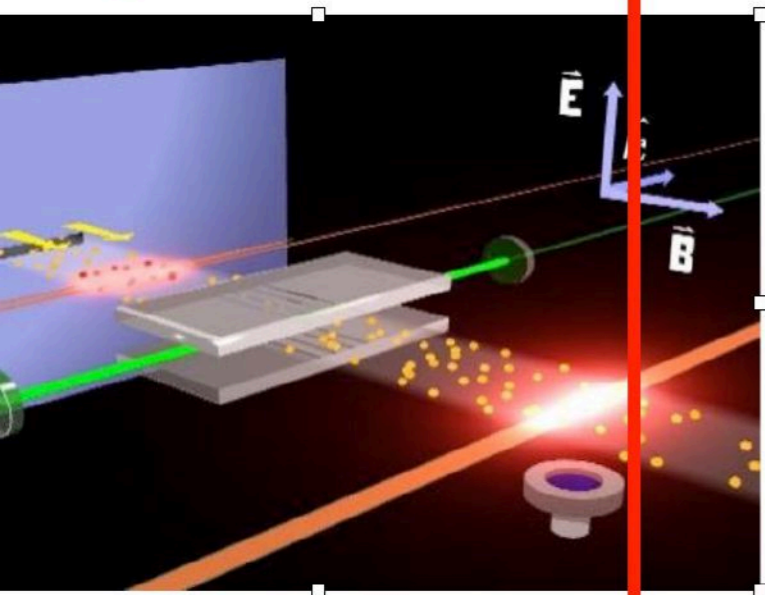
Plan to follow Wieman's Boulder Cs experiment

Q_w (Cs 133) to $\approx 0.5\%$

Oscillator strength	Cs	Fr
$E1_{\text{Stark}}$ @ 500 V/cm	10^{-11}	10^{-10}
M1	10^{-13}	10^{-11}
$E1_{\text{pnc}}$	10^{-22}	10^{-21}
$E1_{\text{stark}} - E1_{\text{pnc}}$ IF	10^{-16}	10^{-15}



$$|E1_{\text{Stark}} + M1 + E1_{\text{PNC}}|^2$$

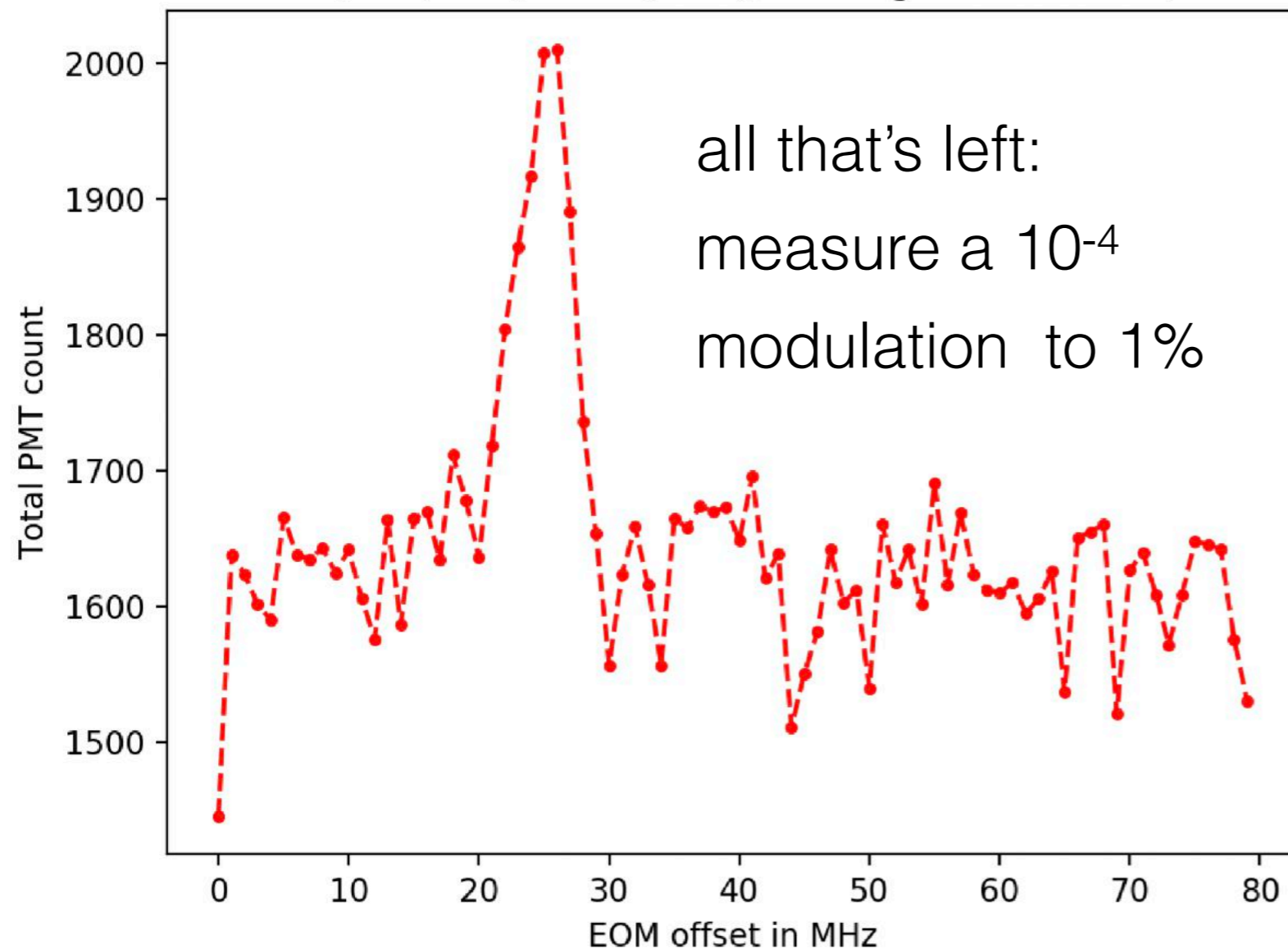


parity flip
(e.g. E-field reversal)
→ 10^{-4} modulation of decay fluoresc.
in Fr

The latest results: 7s-8s at last

- September 2018: Observed the β -type Stark-induced 7s-8s transition (basis of the planned APV measurement)

Fr211, 7s (F=5) \rightarrow 8s(F=4), beta signal at 6124 V/cm



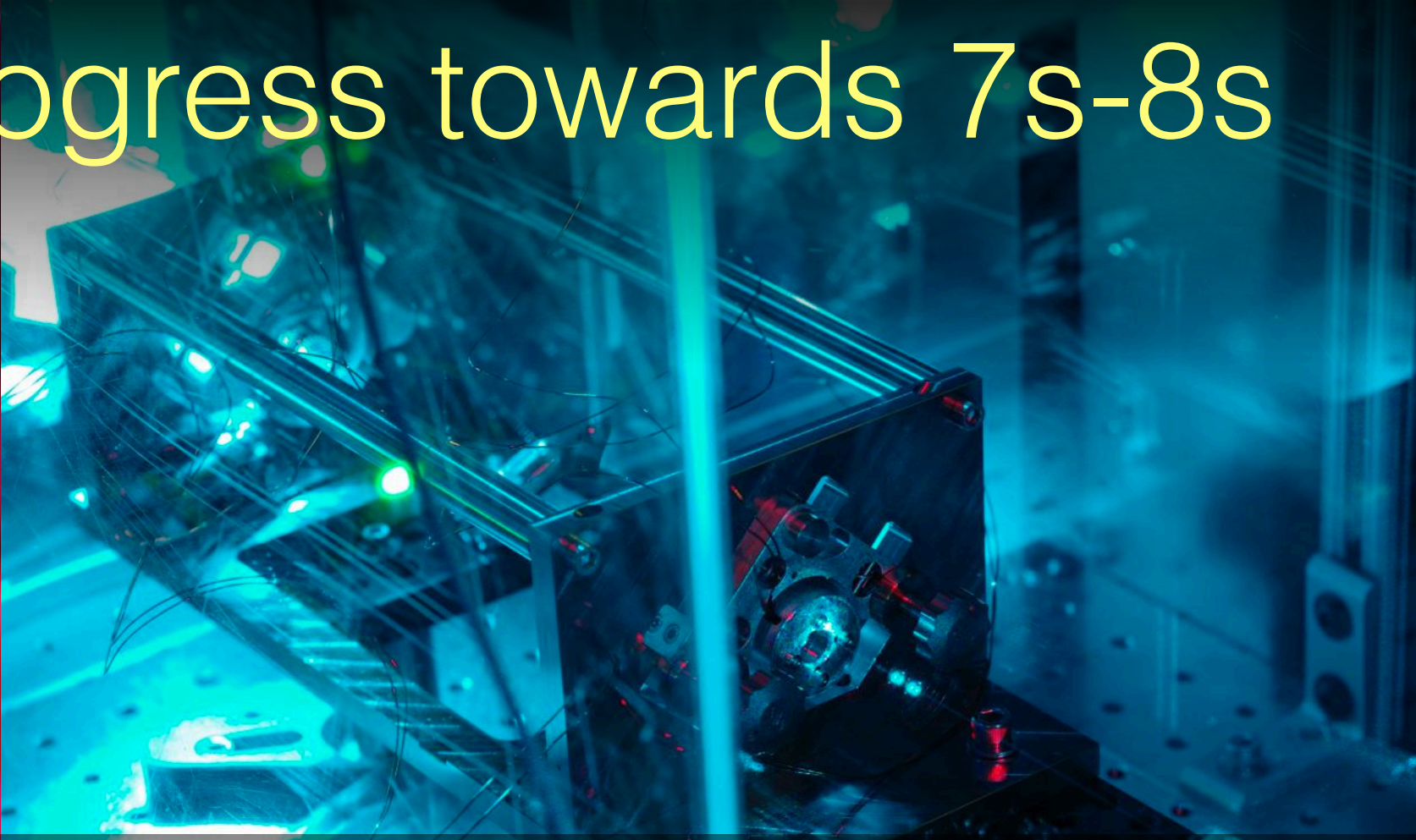
- $10^8 - 10^9$ x weaker than typ. E1 trans.
- in least stable of first 103 elements
- true milestone
- the bad:
 - for APNC need to use 10x less E-field \rightarrow 100x less signal
- the good:
 - some aces up our sleeves

Technical progress towards 7s-8s

Photos: M. Kossin



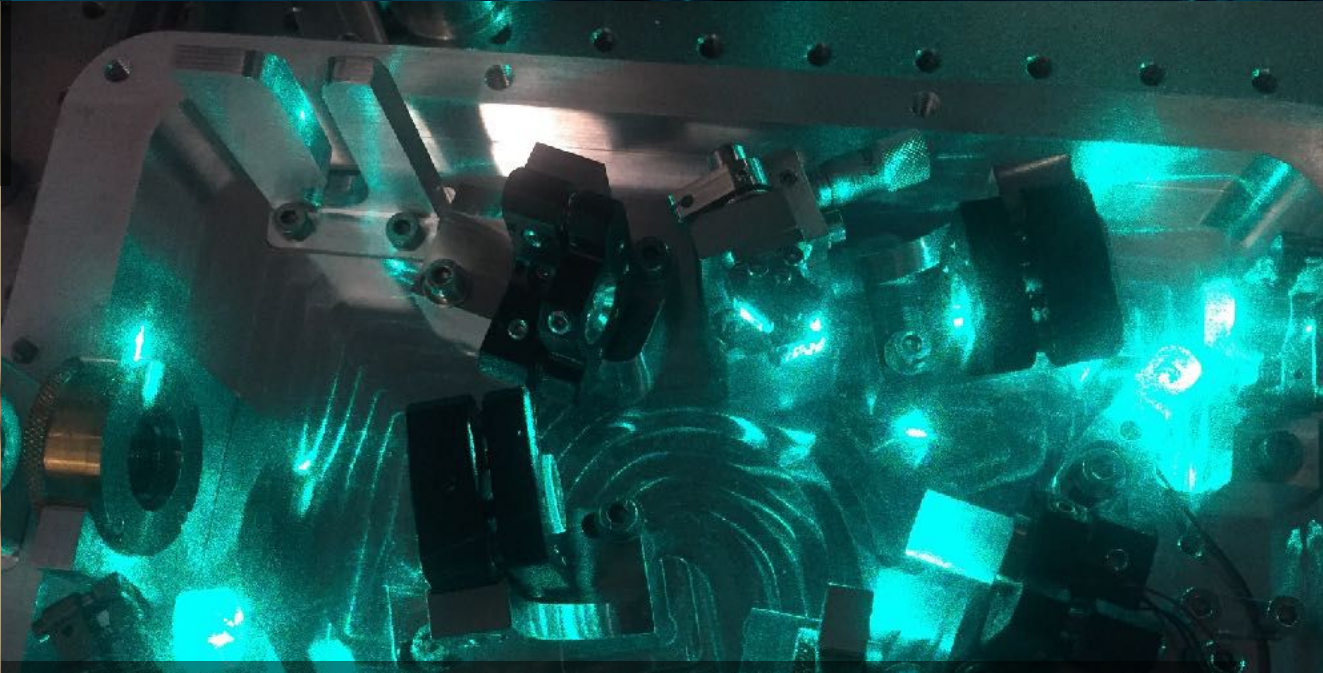
done: ultra-stable (10^{-10} level)
ref. cavity for 7s-8s spectr.



UHV-compatible 506 nm power buildup
cavity (1000-2000 x) in development

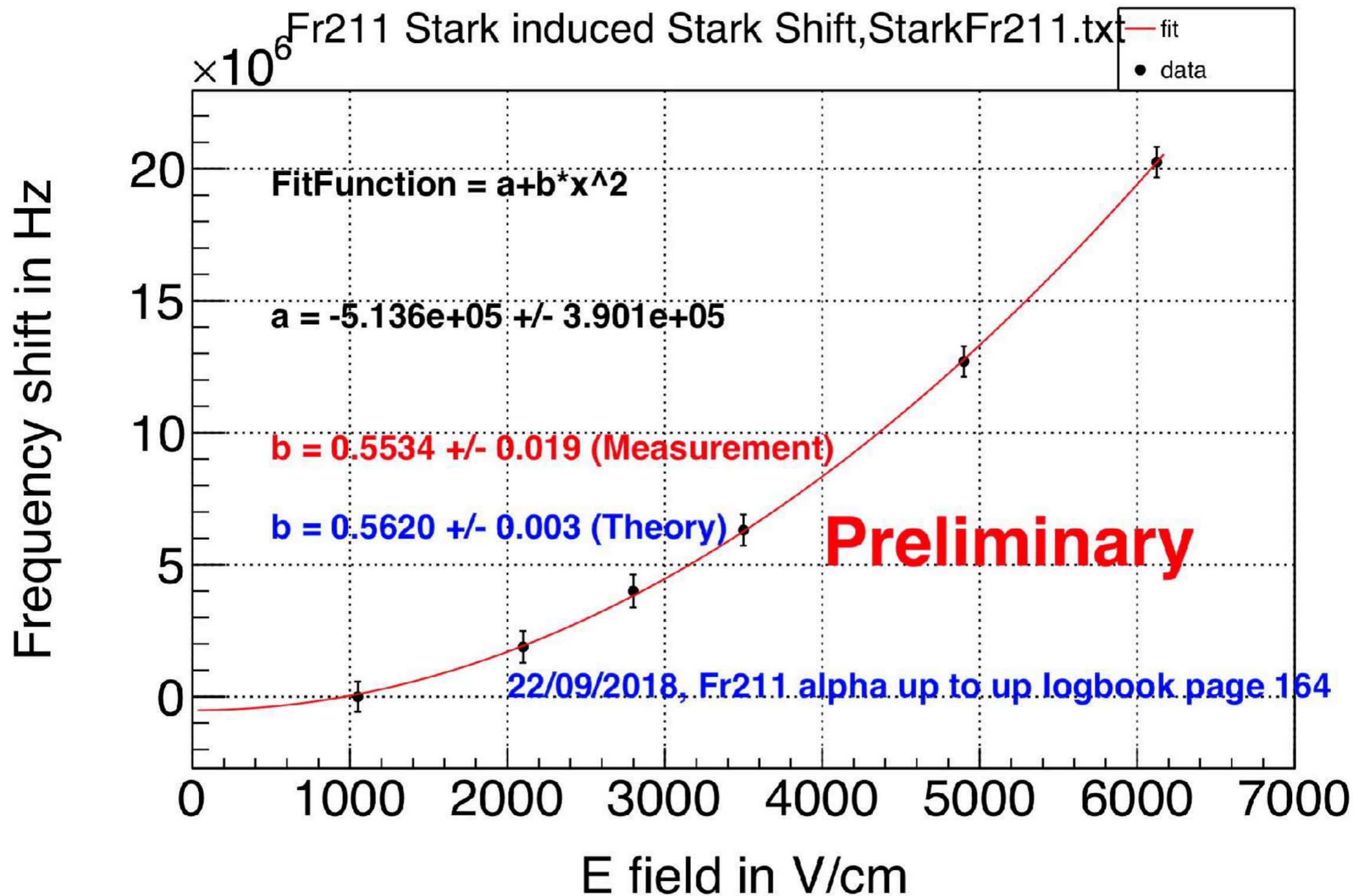


done: transparent field plates for operating
laser trap internally



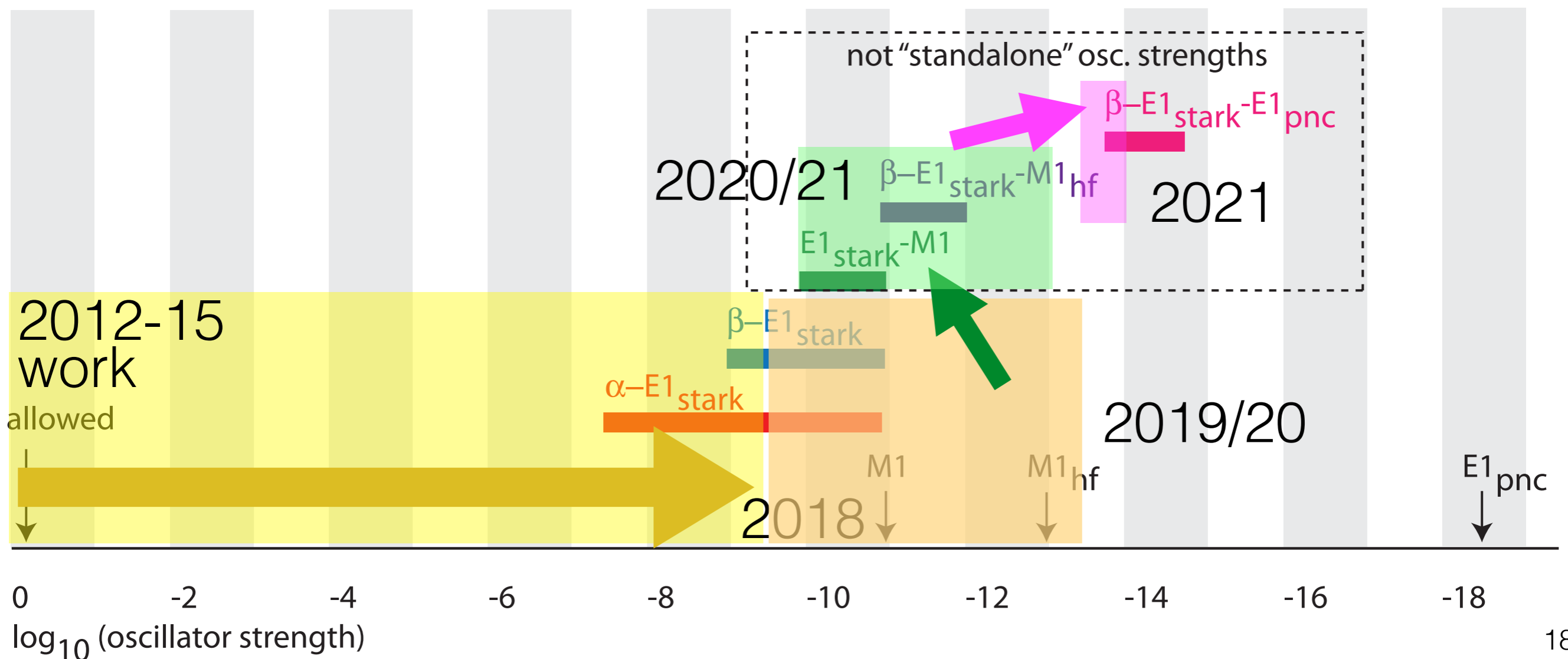
506 nm 7s-8s light (... until the great
power outage of 2018)

Fr 7s - 8s DC Stark effect

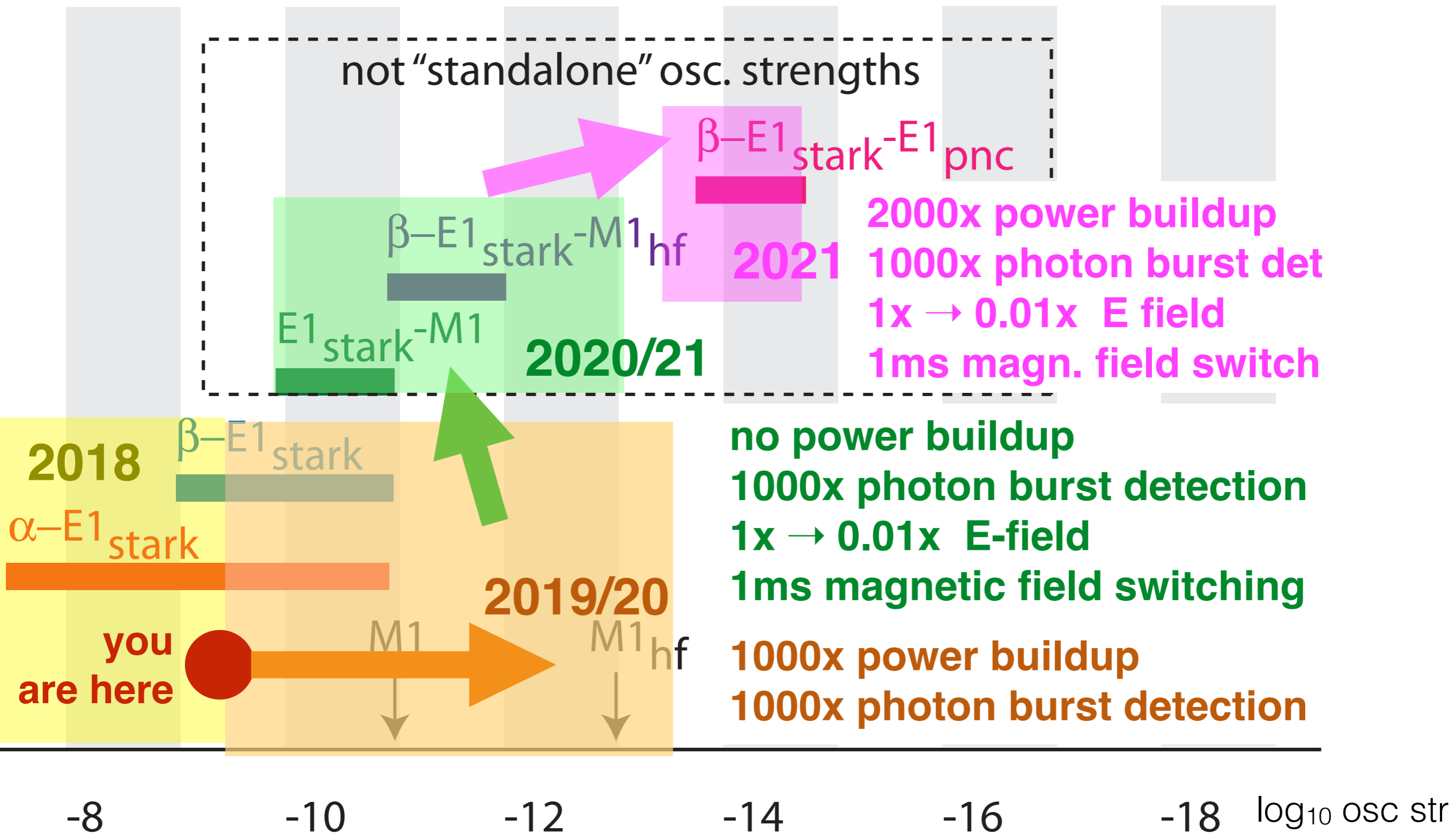


The road ahead

- Now in a position to move forward with the 7s - 8s spectroscopic program towards APV
- Over ≈ 3 years, expect to work our way down towards detecting the APV amplitude



The road ahead



The FrPNC team

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E. Gomez — San Luis Potosi

S. Aubin — William & Mary

Alumni:

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J. Zhang (PhD 2015, U Maryland)

R. Collister (Phd 2015, U Manitoba)

A. DeHart (MSc 2018, U Manitoba)

A. Senchuk (PhD 2018, U Manitoba)

grad student

postdoc / res. assoc.

Funding:

NSERC, NRC/TRIUMF, U Maryland, U Manitoba

DOE: Fr trapping facility infrastructure

NSF: previously



Sept 2018 beamtime