

Guillaume Hupin, CNRS IJClab

Collaborators

D. Becht (CEA)

J. Herko (TRIUMF)

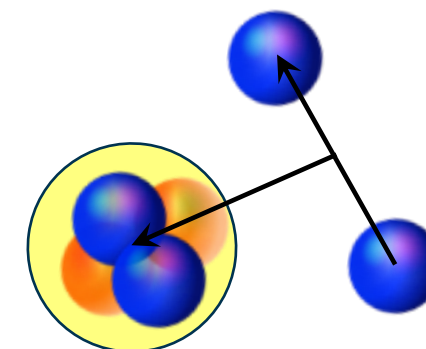
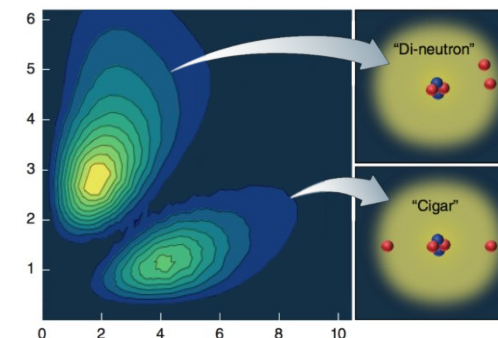
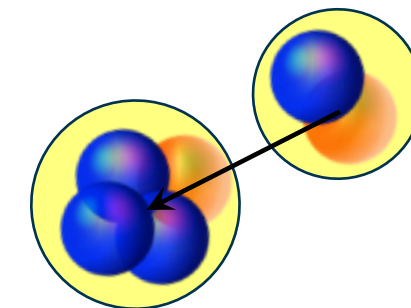
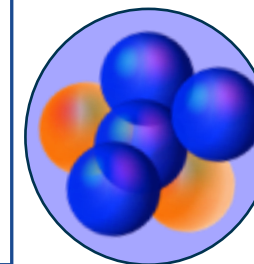
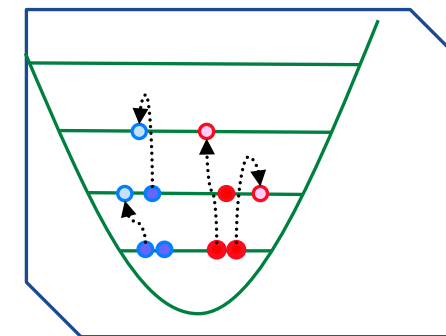
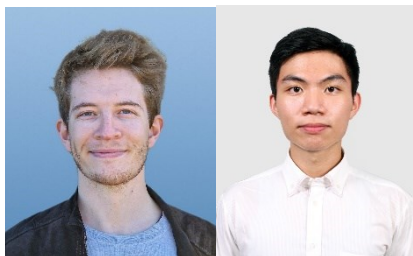
P. Navratil (TRIUMF)

K.H.D. Nguyen (CNRS/IJCLab)

N. Pillet & P. Indelicato (CEA & CNRS/LKB)

O. Yaghi (PhD, now industry)

“Nuclear theory projects”

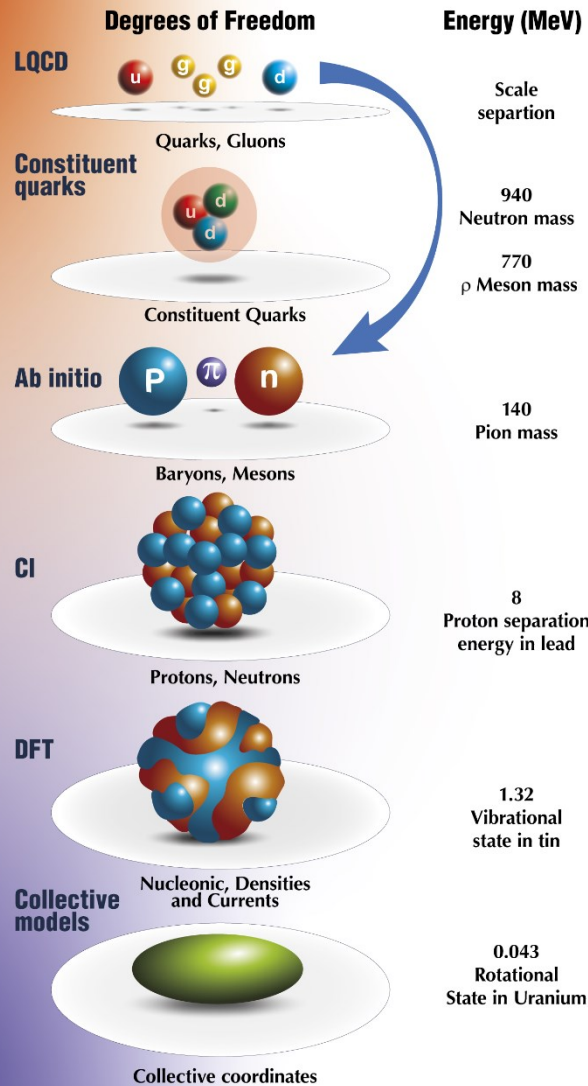




A story of multiple scale

Physics of Hadrons

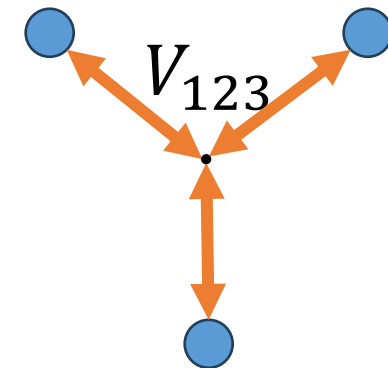
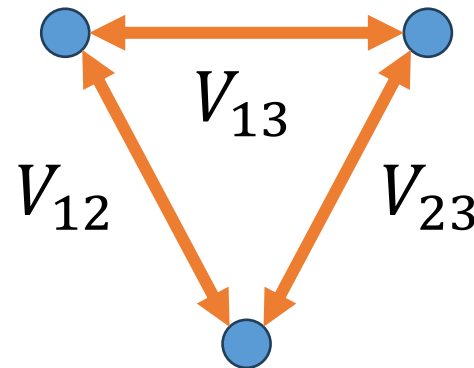
Physics of Nuclei



- Goal: Solving the Schrodinger equation (SE) for an A-body system:

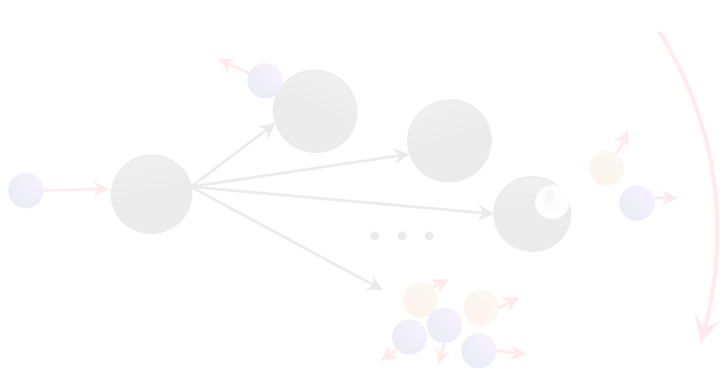
$$H|\psi^{J^{\pi T}}\rangle = E|\psi^{J^{\pi T}}\rangle$$

- Nucleons are considered as point-like particles.
- The SE is solved by considering two and many-body interactions between nucleons



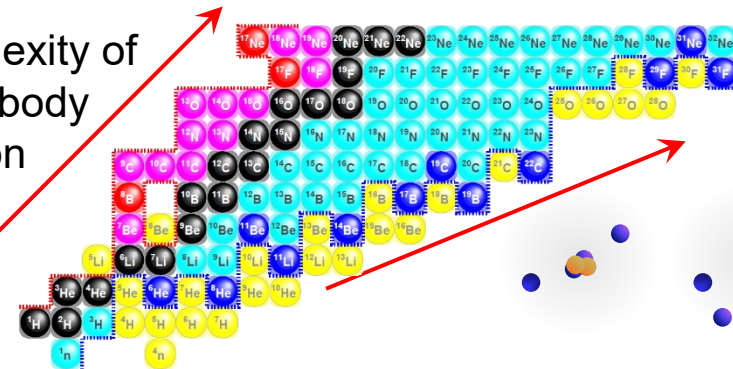


(i) Research directions



Complexity of scattering problem

Complexity of many-body solution



Both

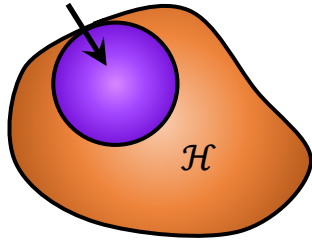
$\neq n, p$ particles interacting with strong force ($M_h \gg M_{n,p}$)

$$M_h \leq M_{n,p}$$

- ☹ Nuclear theory is **data driven**.
- ☹ The lack of accuracy of *ab initio* structure methods impedes the development of reactions modeling.

Credits H. Lenske

active domain



- Variational;
- Orthonormal basis;
- Controllable parameters (N_{\max} , $E_{1_{\max}}$ etc...);
- UV/IR convergence.

Superposition of Slater determinants:

$$|\Psi^A\rangle = \sum_{\alpha} c_{\alpha} \Phi_{\alpha}^{\varphi}(\vec{r}_1, \dots, \vec{r}_A) = A_0 |\Phi_{0p0h}^0\rangle + \sum_{\alpha'} A_{\alpha'} |\Phi_{1p1h}^{\alpha'}\rangle + \dots$$

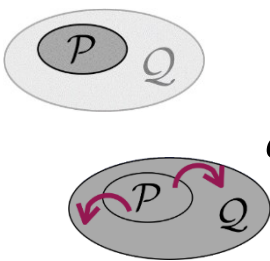
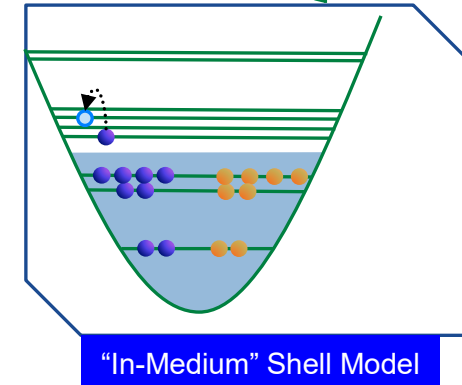
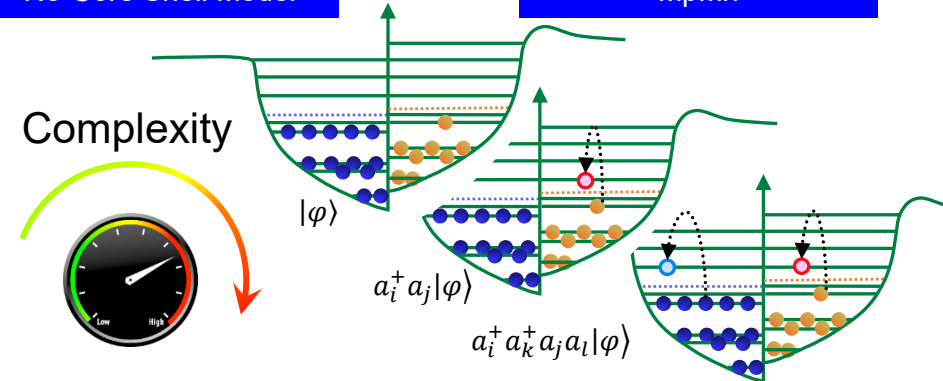
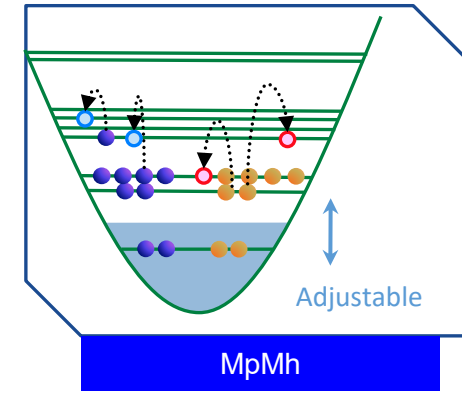
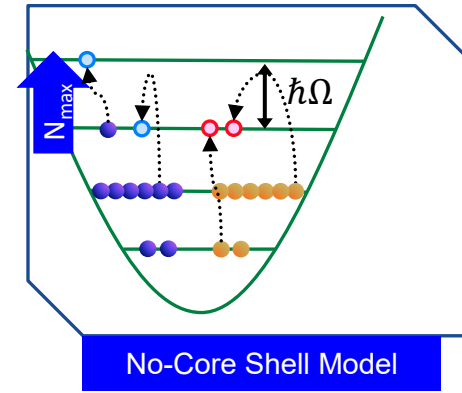
NCSM
IM-SH
MpMh

Optimization of mixing coefficients, one-body Hilbert space:

$$\delta\mathcal{E}[\Psi]_{\{A_{\alpha}^*\}} = 0 \Rightarrow \sum_{\beta} A_{\beta} \langle \Phi_{\alpha} | \hat{H} | \Phi_{\beta} \rangle = EA_{\alpha}$$

$$\delta\mathcal{E}[\Psi]_{\{\varphi_{\alpha}^*\}} = \langle \Psi | [\hat{H}, \hat{T}] | \Psi \rangle = 0 \Leftrightarrow [\hat{h}(\rho), \hat{\rho}] = \hat{G}(\sigma)$$

Generalized Brillouin (GB) equation



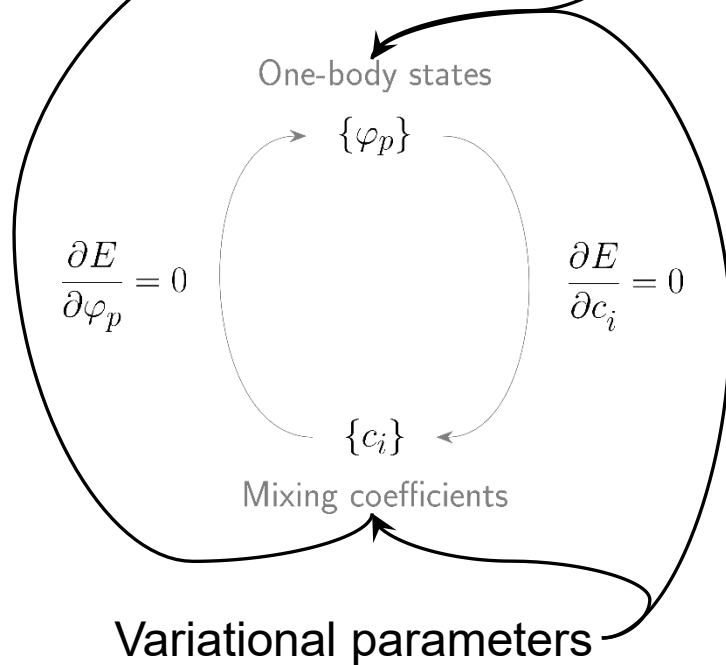


One-body Hamiltonian:

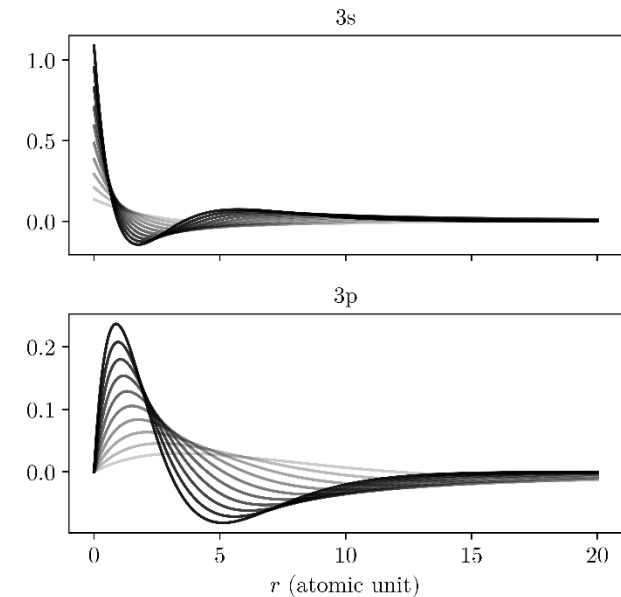
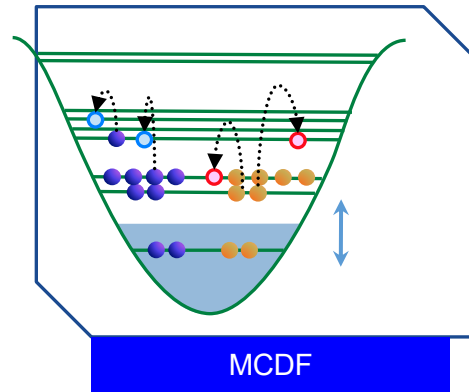
$$(c \vec{\alpha} \cdot \vec{p} + \beta c^2 + V(r)) \varphi_p(r) = 0,$$

Many body atomic wave function

$$|\Psi^n\rangle = \sum_i c_i \Phi_i^\varphi(\vec{r}_1, \dots, \vec{r}_A).$$



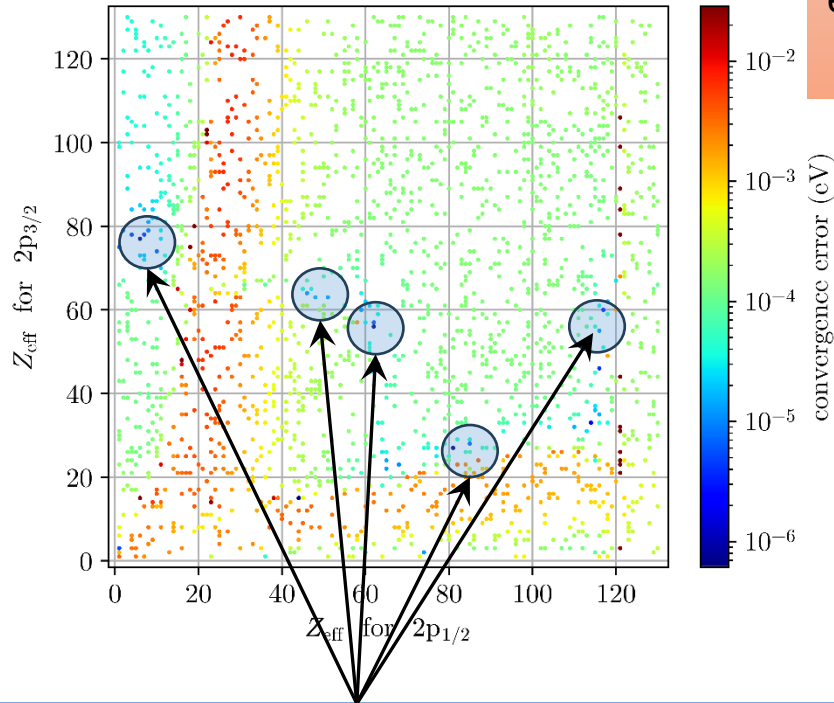
- Initial one-body states are hydrogenlike.
- To guesstimate the effect of many body correlation, orbitals are adjusted with an effective proton number Z_{eff} .
bad Z_{eff} leads to no convergence
- Finding a good combination of Z_{eff} is time consuming.





Machine Learning for Atomic Structure Codes: the proposed solution

$Z = 100$ $1s^2 2s^2$ $J = 0$



Correlation plot of the convergence error estimate as a function of the Z_{eff} of $2p_{\frac{3}{2}}$ and $2p_{\frac{1}{2}}$ of Be-like Fm.

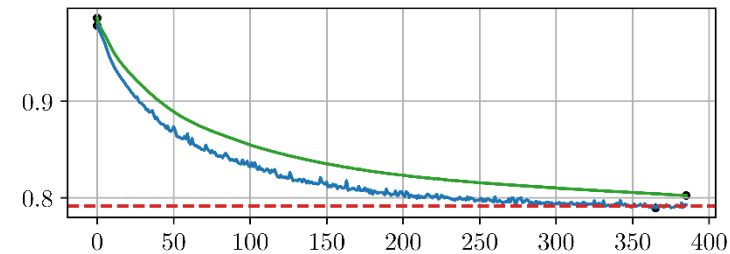
😊 This is an excellent candidate problem for **machine learning**.

🗄️ Generating data:

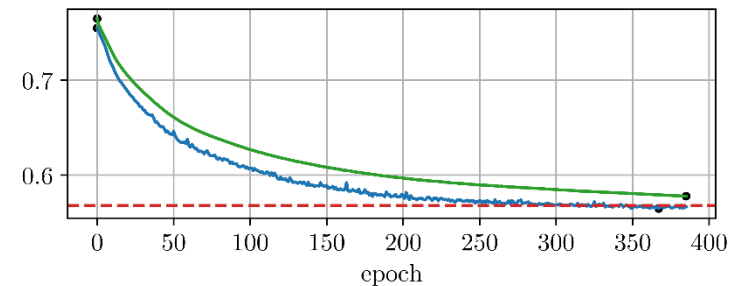
Features	Labels
⋮	⋮
state, $Z_{\text{eff}}(s)$	$\log(\text{err}_{\text{conv}})$
⋮	⋮

☹️ The physicist (or software user) is therefore required to **manually predict the hyperparameter Z_{eff}** for each successive atomic orbital of every atomic state, making the procedure very cumbersome.

Root Mean Square Error: $\frac{1}{n} \sum (\log_{10}(\text{err}_{\text{conv}}) - \text{prediction})^2$



Mean Absolute Error: $\frac{1}{n} \sum |\log_{10}(\text{err}_{\text{conv}}) - \text{prediction}|$





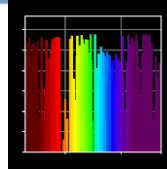
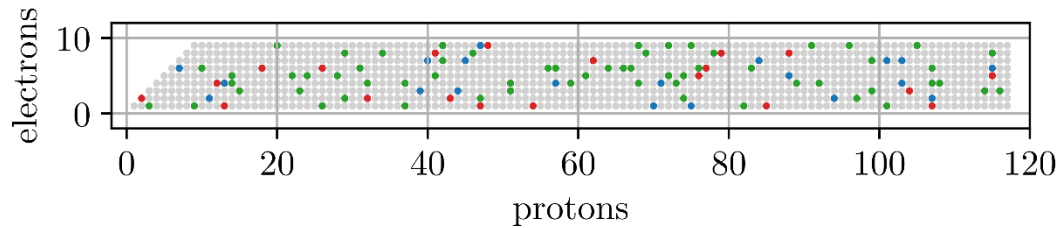
Select a target dataset:

$$1 \leq n_{\text{protons}} \leq 1181 \leq n_{\text{electrons}} \leq 10$$

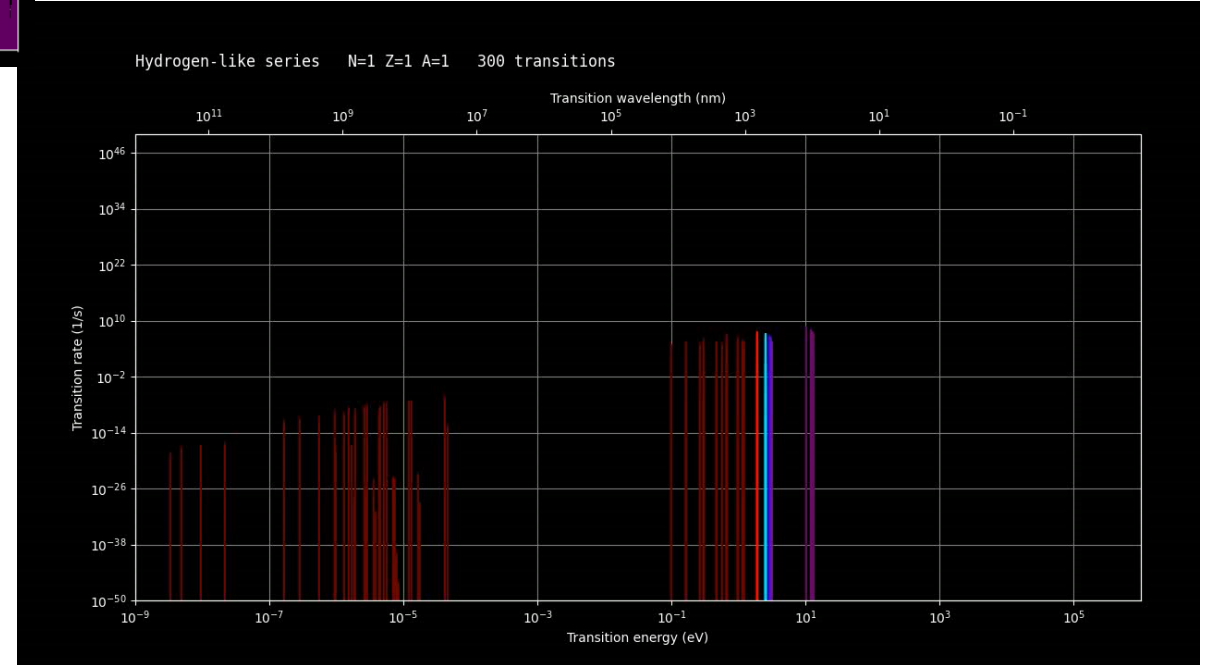
Train a neural network on 10% of it:

training + validation + test = 10% of target

Use trained model for remaining 90%



Datatom: public repository (soon)



The dataset co-generated by AI is substantially larger than the NIST repository and opens the way to applications such as:

- opacity calculations for stellar objects,

- spectroscopic identification of stellar events,
- the study of metastable states relevant to lasers and atomic clocks...



Machine Learning for Nuclear Structure Codes: the problem

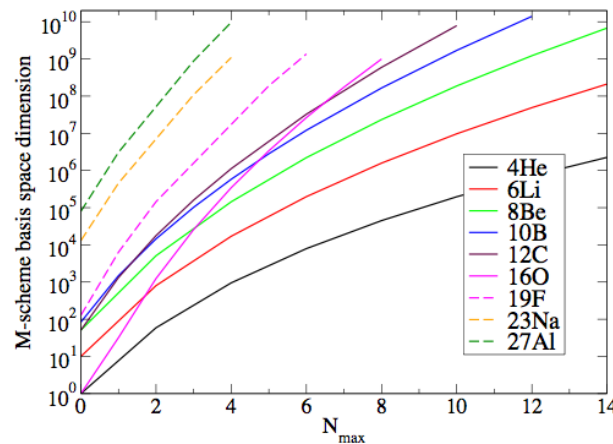
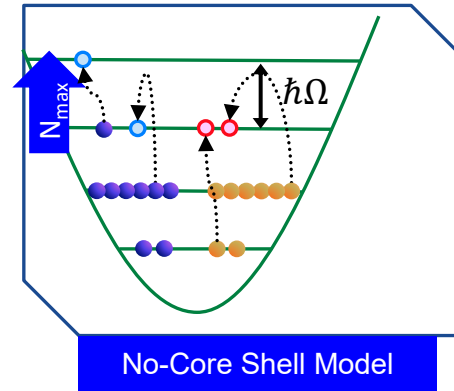
One-body basis states:

$$\varphi_{nl\frac{1}{2}j}(\vec{x}, \vec{\sigma}) = [R_{nl}(x)Y_l(\hat{x})\chi_{\frac{1}{2}}(\vec{\sigma})]^J,$$

Many body nuclear wave function

$$|\Psi^n\rangle = \sum_i c_i \Phi_i^\varphi(\vec{r}_1, \dots, \vec{r}_A).$$

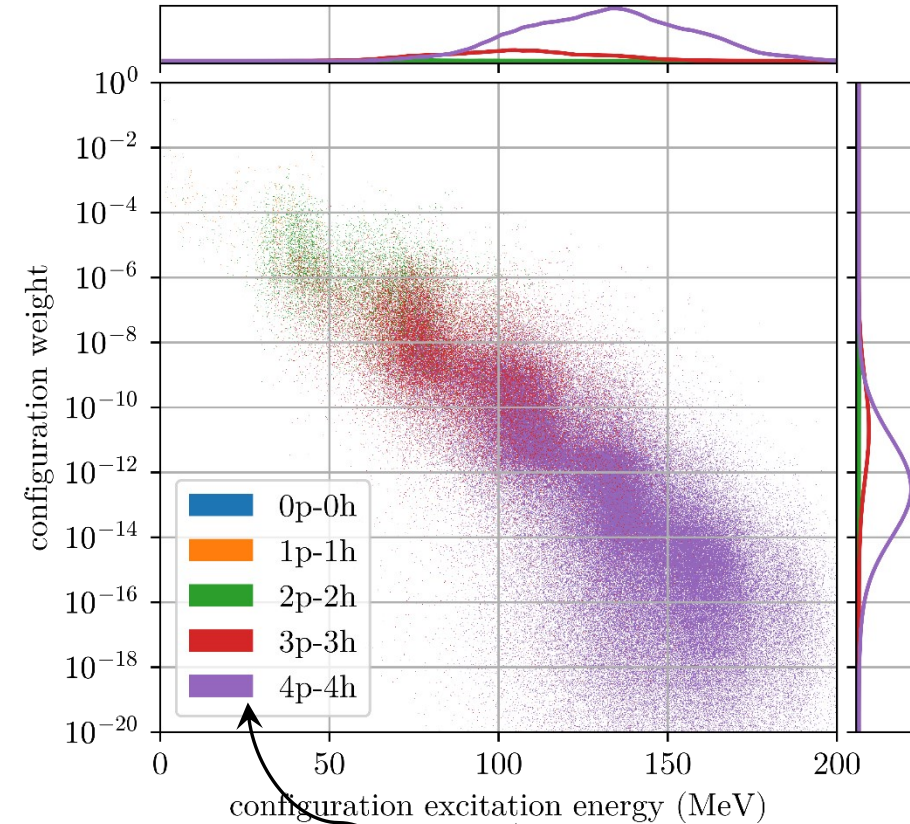
$$\phi_n^A = \frac{1}{\sqrt{A!}} \begin{vmatrix} \phi_i(\vec{r}_1) & \dots & \phi_i(\vec{r}_A) \\ \vdots & \ddots & \vdots \\ \phi_l(\vec{r}_1) & \dots & \phi_l(\vec{r}_A) \end{vmatrix} = a_l^\dagger \dots a_i^\dagger |0\rangle$$



Usual truncation schemes:

- Excitation order
- Excitation energy
- Importance truncation

$Z=13 \quad A=22 \quad \pi=1 \quad \text{eigenstate 3}$



Usual schemes:

- Excitation order
- Excitation energy

Importance truncation

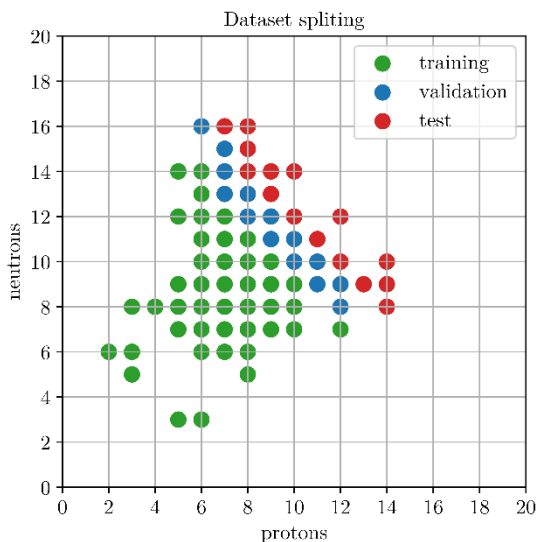


Select a target dataset (exact solution):

$$1 \leq n_{\text{protons}}, n_{\text{neutrons}} \leq 14; A \leq 18$$

Train a neural network on 10% of it:

training + validation + test



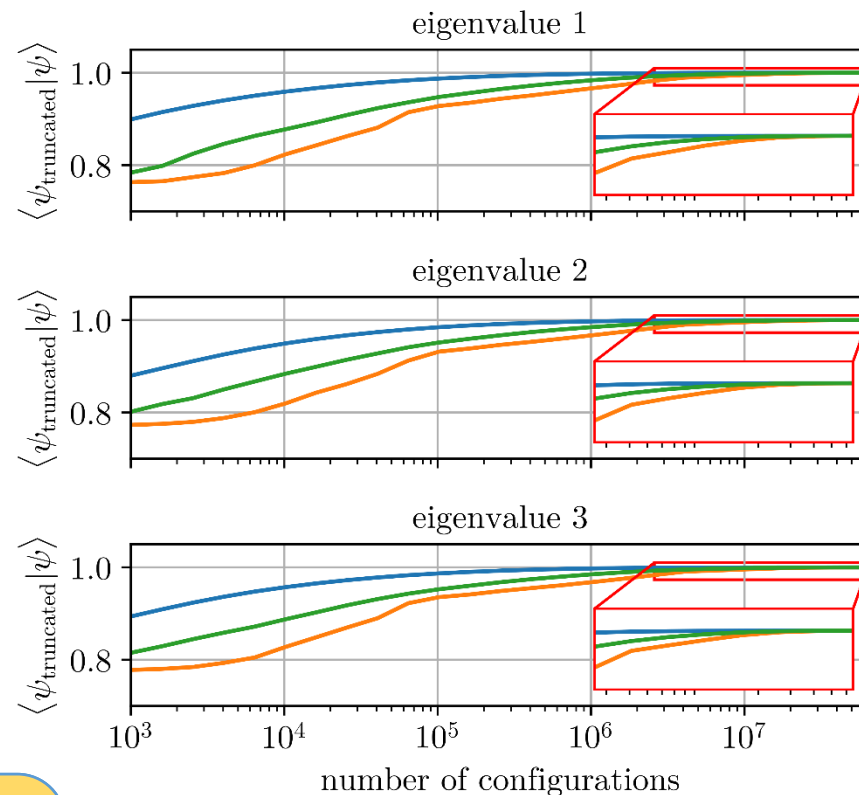
Generating data:

Features	Labels
⋮	⋮
state, $ \text{config}_i\rangle$	$\log_{10}(c_i^2)$
⋮	⋮

Each Slater determinants configuration are encoded via a transformer layer

The AI augments, **rather than replaces**, the MB solver by automatically selecting the relevant subspace for diagonalization, with state-dependent variations.

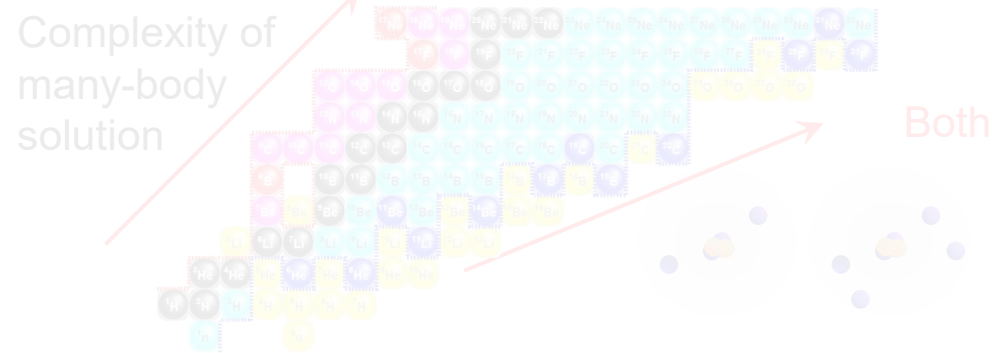
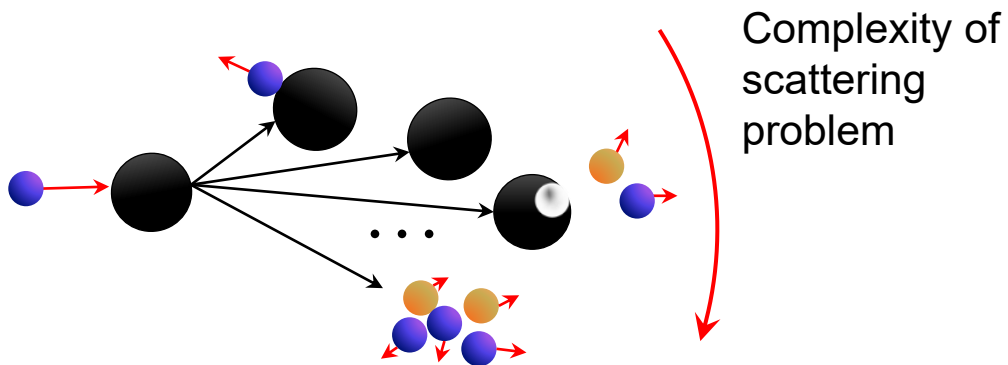
$$Z = 13 \quad A = 22 \quad \pi = 1$$



Using real weights: $c_i^2 \geq cste$;
 Using model;
 Using excitation energies: $E_i \leq cste$



(ii) Research directions



$\neq n, p$ particles interacting with strong force ($M_h \gg M_{n,p}$)

$$M_h \leq M_{n,p}$$

- ☹ Nuclear theory is **data driven**.
- ☹ Few-body techniques scale **very bad** with the number of constituents in the continuum.

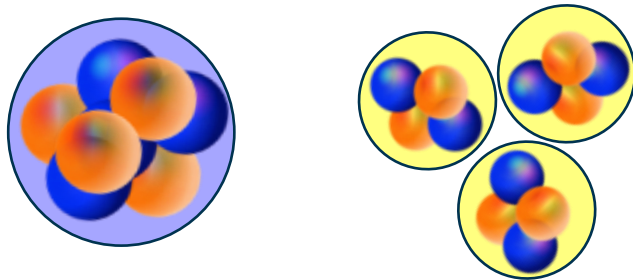
Credits H. Lenske



We want to develop a unique tool applicable to both nuclear structure and reactions, to enhance our understanding of the strong force at low energy.

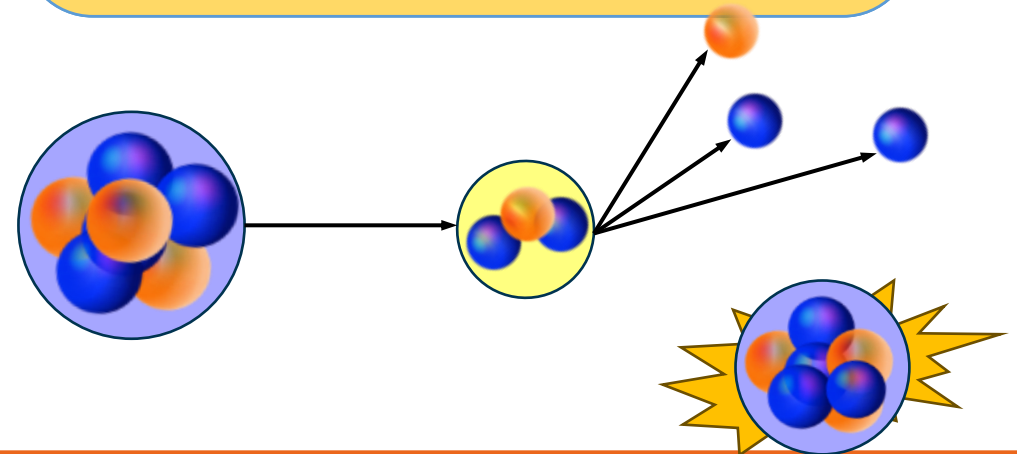
Structure

- ☺ Extraction of resonance properties for fine tuning interactions and applications to astrophysics.



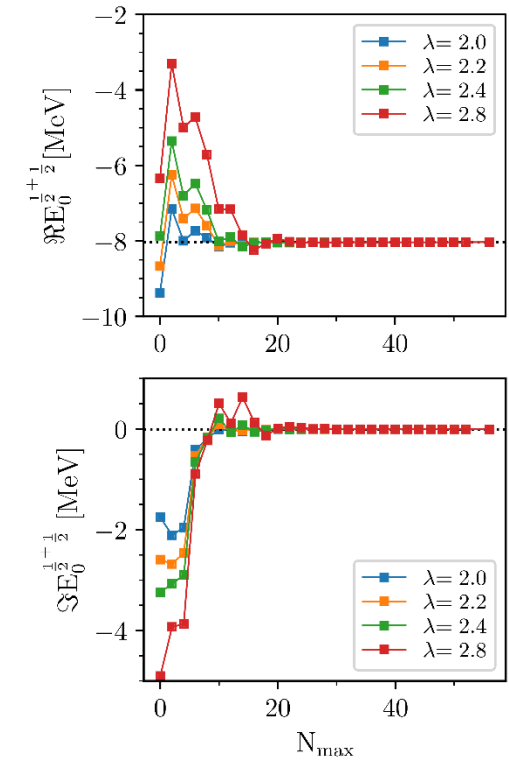
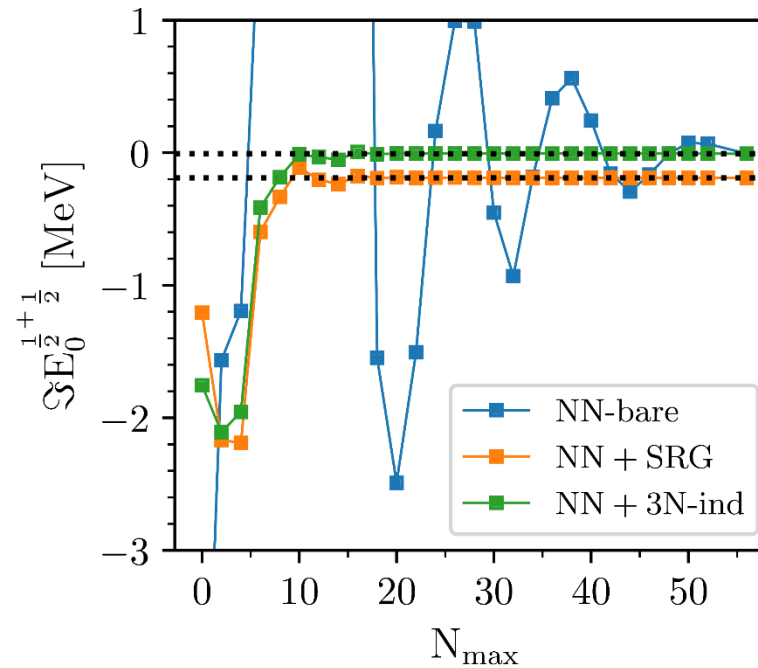
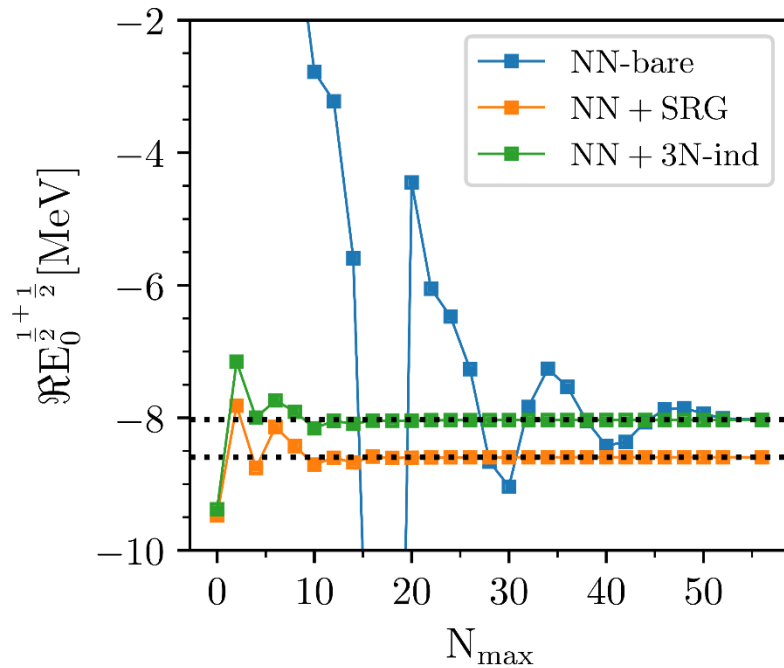
Reactions

- ☺ Enabling us to study nuclear breakup reactions including Final State Interaction (FSI)
- ☺ Calculation of complex charged (and multi-neutron) nuclear decay.





Triton g.s. convergence with CS-Hamiltonian ($\theta = 0.3$ rad)

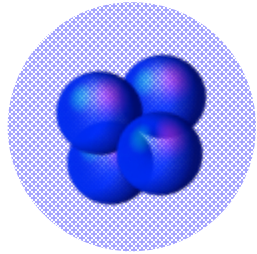


Maximum model space achievable $N_{\max} \sim 60$ (30 nodes a box in excess of 15 fm)

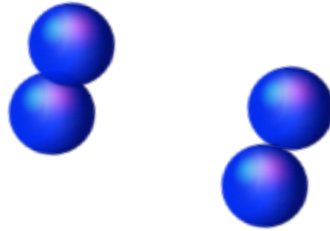
- SRG leads to faster convergence;
- Induced 3N-interaction needs to be accounted for (as expected from literature).



4-neutron: a resonance ?



or



Claimed by experiment

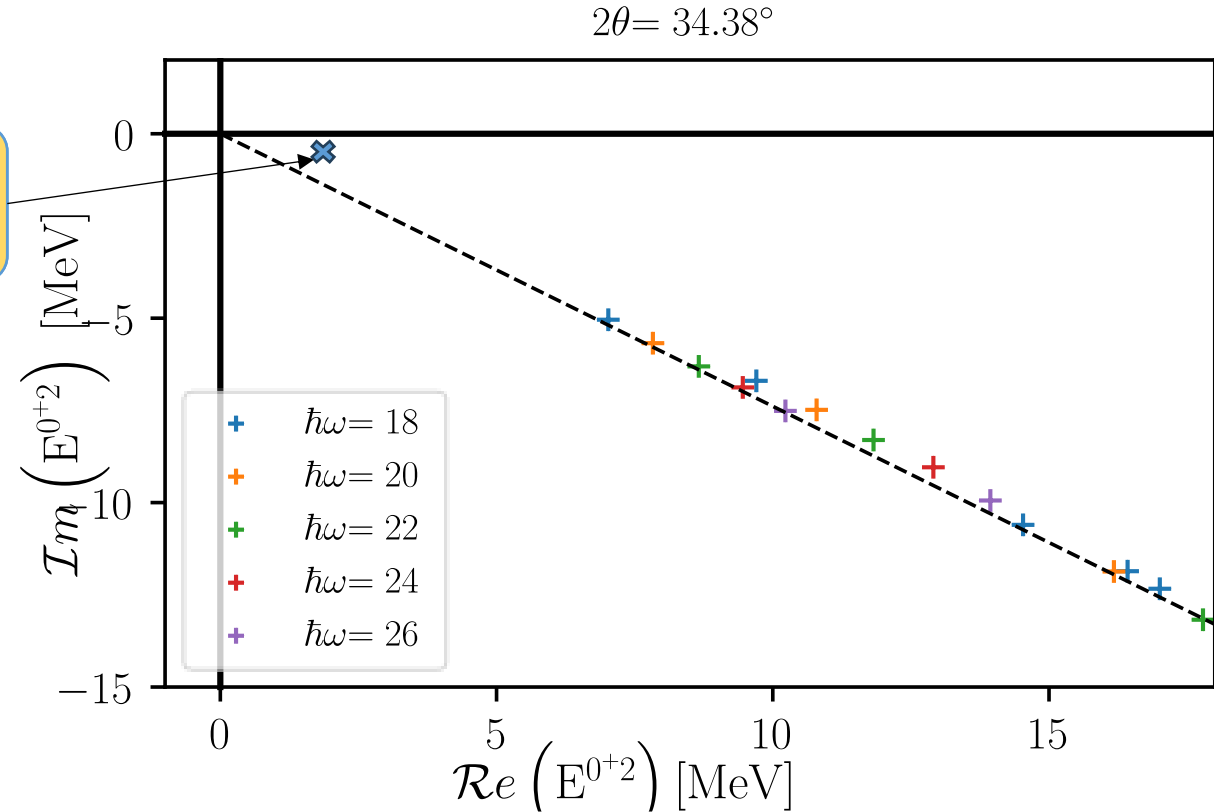
CS shows no indication of such a resonance in 1^+ , 1^- , 0^+ or 0^- .

Lower bound of :

$$\Gamma_r = 1.9 E_r$$

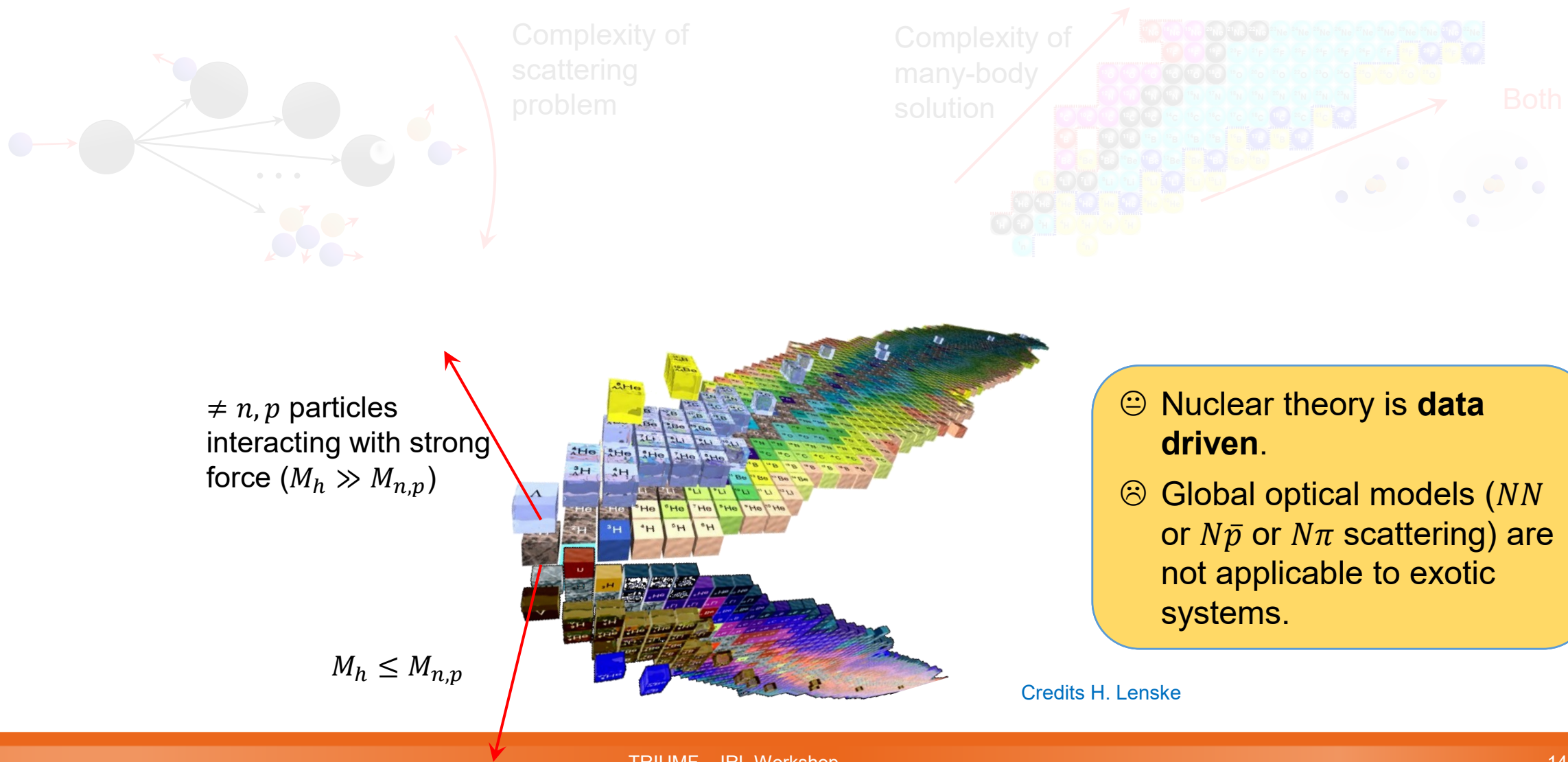
or :

$$\Gamma_r = 4.5 \text{ MeV}$$





(iii) Research directions

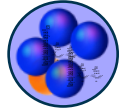




- One way to solve the many-body problem when two scales appear

$$\Psi_{NCSM}^{(A)} = |A\lambda J^\pi T\rangle = \sum_{\alpha} c_{\alpha} |A\alpha j_z^\pi t_z\rangle$$

Mixing coefficients (unknown) A-body harmonic oscillator states

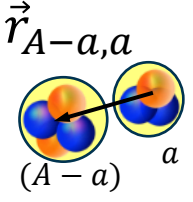


$$|A\lambda J^\pi T\rangle_{SD} \phi_{00}(\vec{R}_{c.m.}^A)$$

Second quantization

$$\Psi_{RGM}^{(A)} = \sum_v \int d\vec{r} g_v(\vec{r}) \hat{A}_v |\Phi_{v\vec{r}}^{(A-a,a)}\rangle$$

Relative wave function (unknown) Antisymmetrizer Channel basis



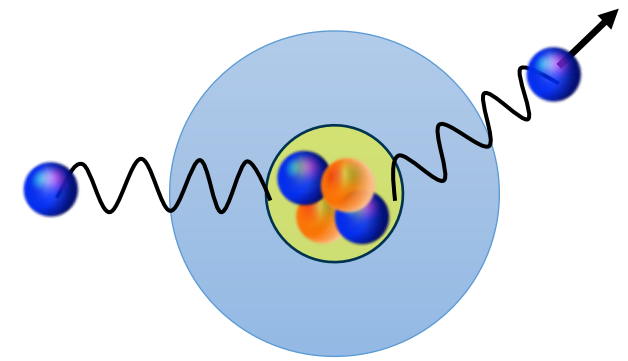
$$\psi_{\alpha_1}^{(A-a)} \psi_{\alpha_2}^{(a)} \delta(\vec{r} - \vec{r}_{A-a,a})$$

Cluster expansion technique

Many-body basis is twice as large as Ψ_{NCSM}

- $\psi_{\alpha_1}^{(A-a)} \in \mathcal{H}^{N_{\max}}$
- $\psi_{\alpha_2}^{(a)} \in \mathcal{H}^{N_{\max}}$

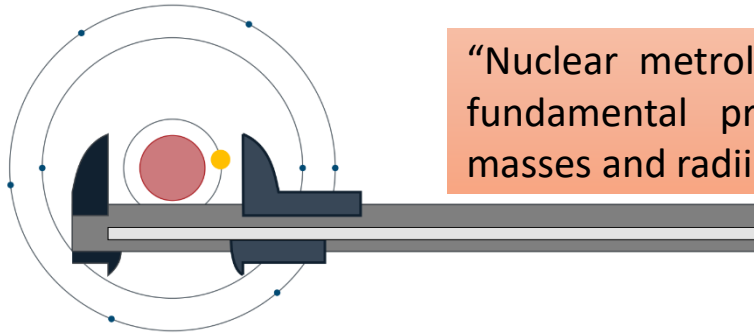
Can address bound and low-lying resonances (short range correlations)



NCSM/RGM
Cluster formalism for elastic/inelastic

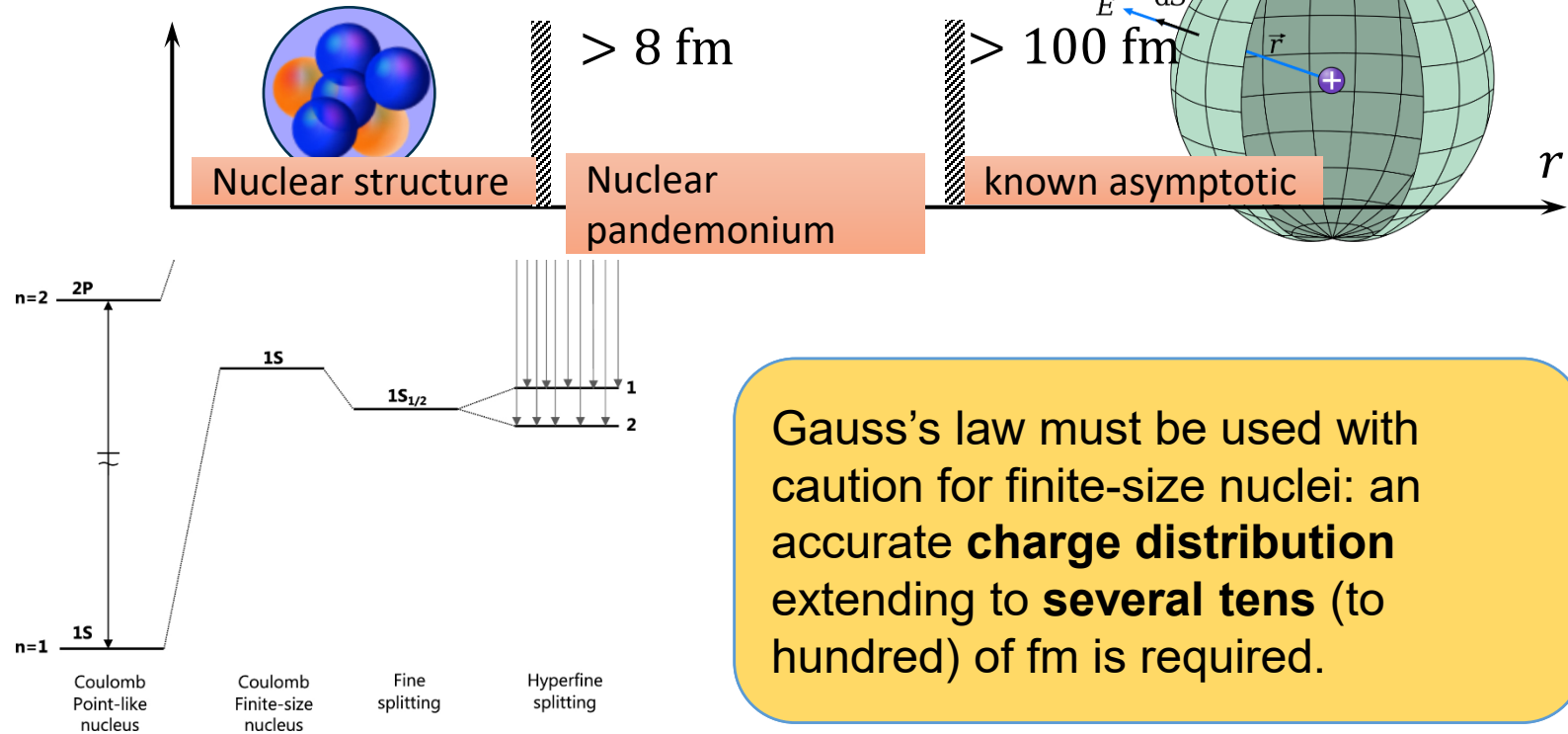
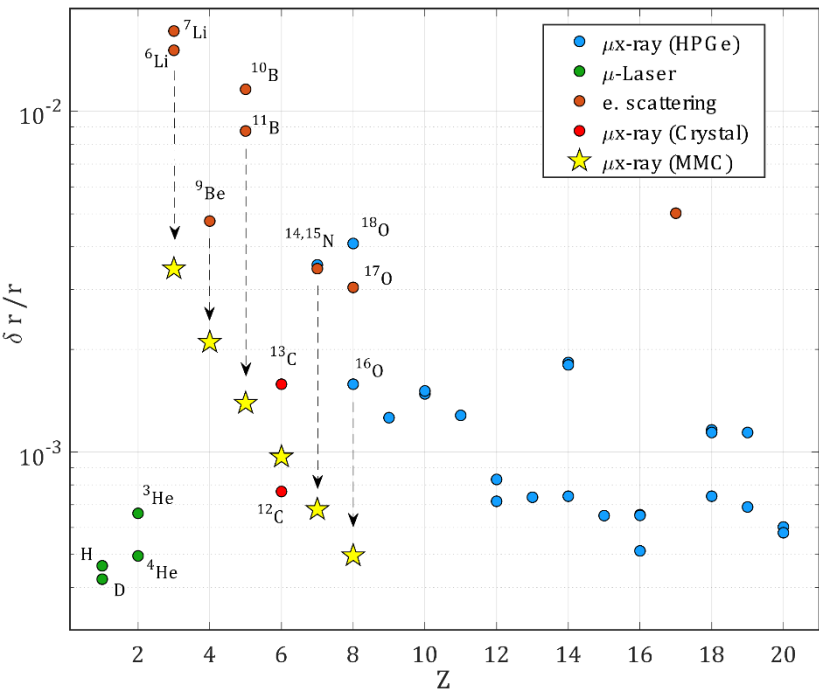


Quartet experiment: connecting *ab initio* atomic theory to nuclear *ab initio* structure



“Nuclear metrology”: determining fundamental properties such as masses and radii..

The aim is to improve the measurement of absolute charge radii of light nuclei by a factor of 10 (**on light systems!**).



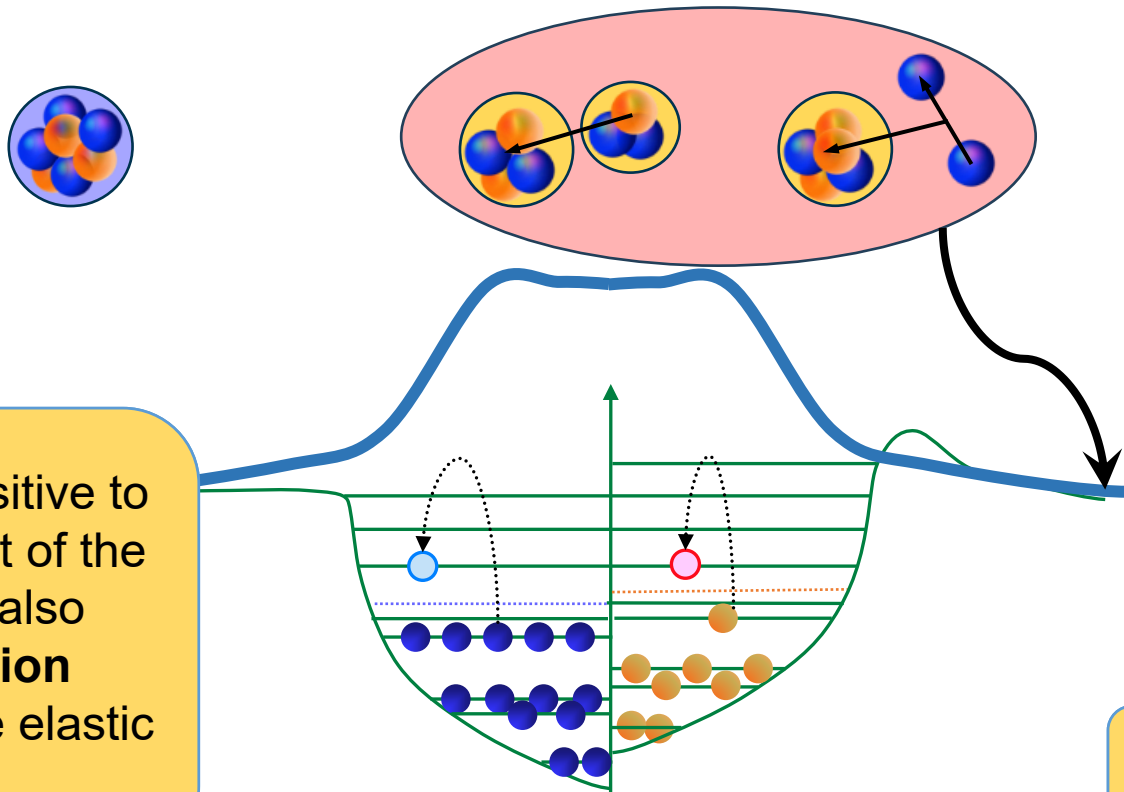
Gauss's law must be used with caution for finite-size nuclei: an accurate **charge distribution** extending to **several tens** (to hundred) of fm is required.



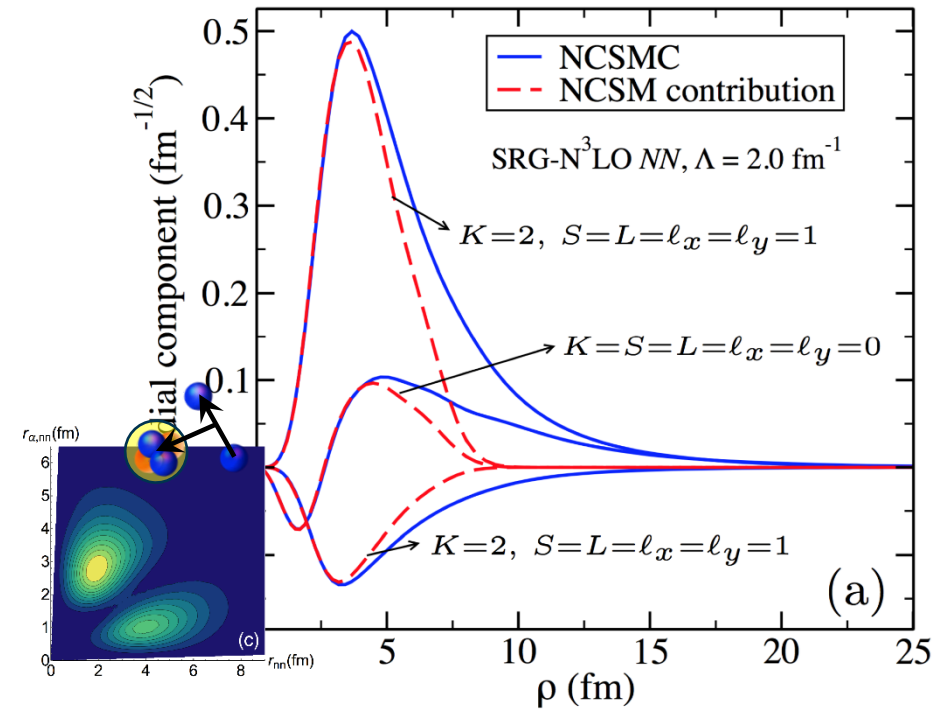
Tail of the nuclear wave function may contains many long-range structure

- The nuclear bound state wave function can be decomposed into all the neighboring reaction channels

$$\Psi_{\text{NCSMC}}^{(A)} = \Psi_{\text{NCSM}}^{(A)} + \Psi_{\text{RGM}}^{(A-1,1)} + \dots + \Psi_{\text{RGM}}^{(A-3,3)} + \dots + \Psi_{\text{RGM}}^{(A-2,1,1)} + \dots$$



Observables sensitive to the long-range part of the wave function are also **sensitive to reaction channels** near the elastic threshold.



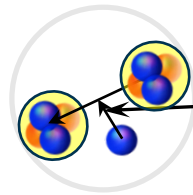
$N_{\text{max}} = 12$ model space i.e. $r \approx 8$ fm HO spatial box

Long-distance components are mandatory to model ${}^6\text{He}$.



Doable with continuum

- Nuclear size 10^{-5} a.u.
- Muonic 1s Bohr radius 10^{-3} a.u.



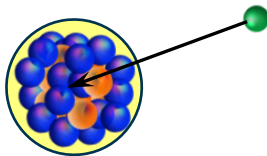
not scaled

$$r_n = \frac{\hbar c}{Z\alpha} \frac{m_{\mu^-}}{2\mu} n^2$$



200 times closer than the electron, 10 times more sensitive to nuclear density

Maybe in 2326



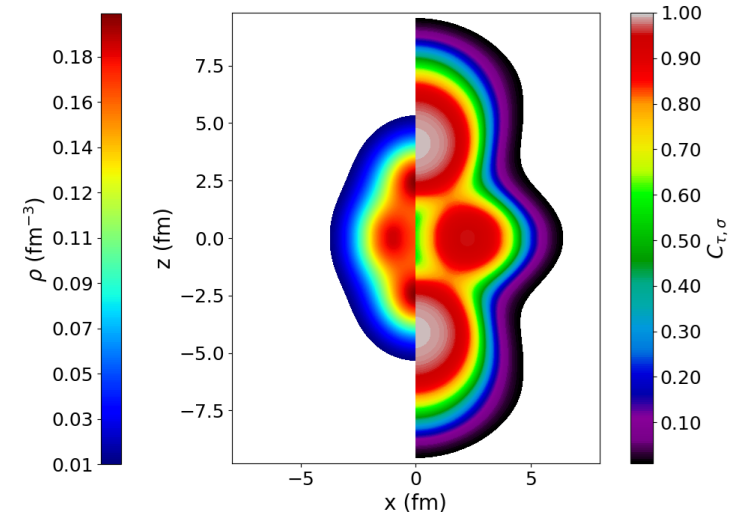
- Bound state properties with continuum

$$\langle \Psi_{\text{NCSMC}}^{(A)} | \hat{O} | \Psi_{\text{NCSMC}}^{(A)} \rangle$$

to obtain the charge density $\hat{\rho}(r)$ at large r and the magnetic distribution.

- We will also explore the interest of computing localization probabilities:

$$D_{\tau\sigma}(r) = \left(\tau_{\tau\sigma} - \frac{1}{4} \frac{(\nabla \rho_{\tau\sigma})^2}{\rho_{\tau\sigma}} \right)$$



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Link with DFT (new constraints)



Thank you !

