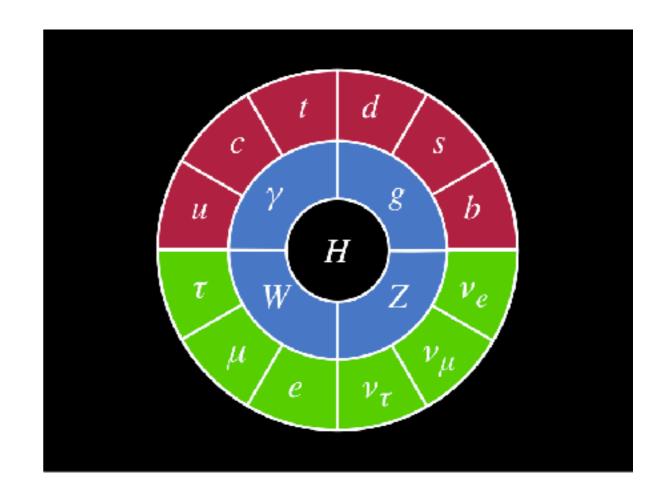
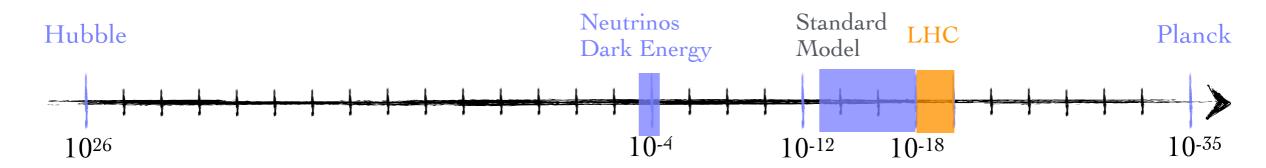
# Particle Physics with Table-Top Experiments

Asimina Arvanitaki Perimeter Institute

#### The Standard Model

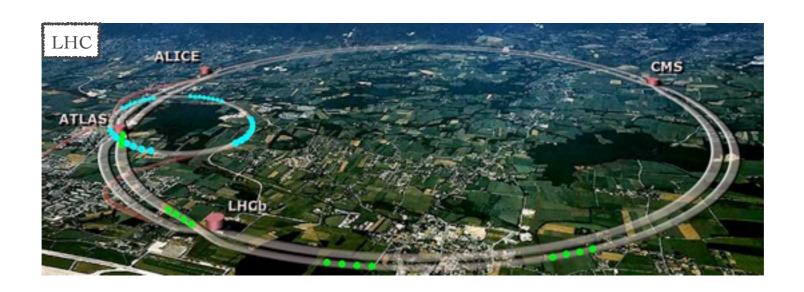


#### The Length Scales in the Universe

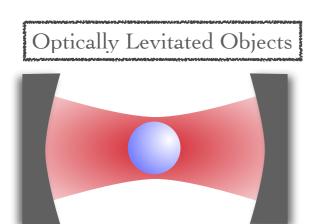


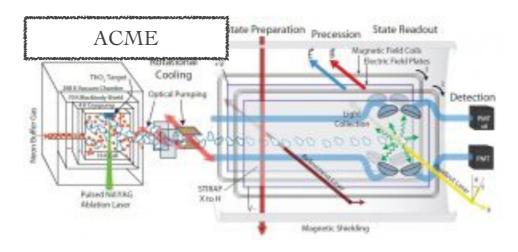
Scale in meters

## The High Energy Frontier

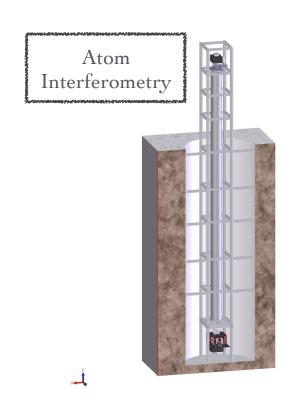


## Opportunities at the Precision Frontier

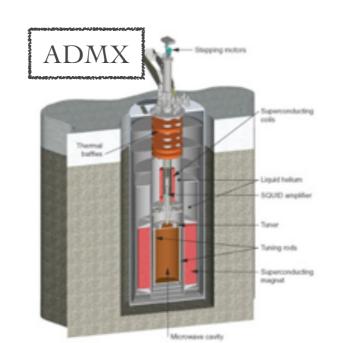


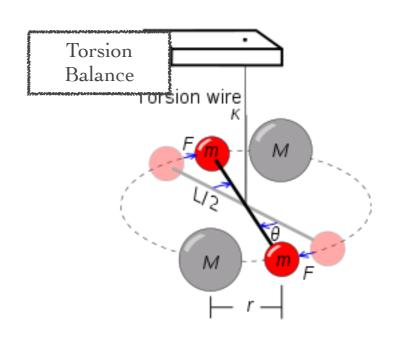




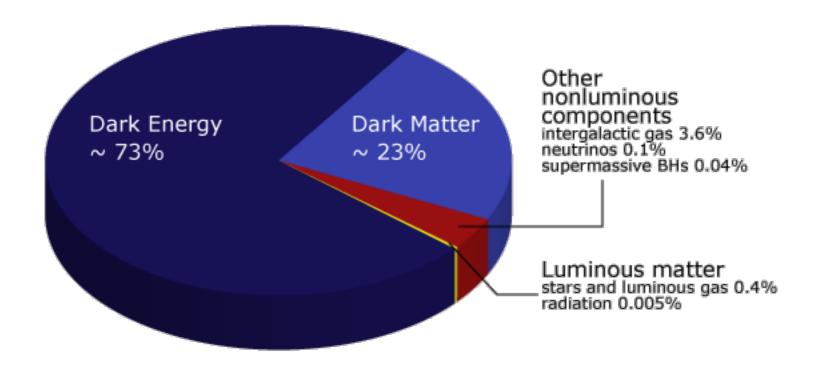








#### The Mystery of Dark Matter



#### Models of Dark Matter

• What is it made out of?

Anything from  $10^{-22}$  eV to  $10^{70}$  eV in mass

• How is it produced?

• Does it have interactions other than gravitational?

#### Outline

Light Bosonic Dark Matter

• Light Bosonic Dark Matter

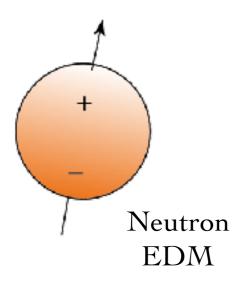
• Atomic Clocks

Molecules

• Black Holes

#### Why is the Electric Dipole Moment of the Neutron Small?

The Strong CP Problem and the QCD axion



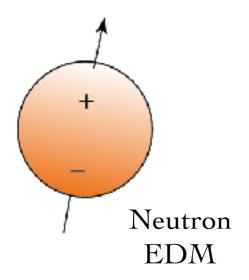
$$\frac{g_s^2}{32\pi^2}\theta_s\vec{E}_s\cdot\vec{B}_s$$

 $EDM \sim e~fm~\theta_s$ 

Experimental bound:  $\theta_s$  < 10-10

## Why is the Electric Dipole Moment of the Neutron Small?

The Strong CP Problem and the QCD axion



$$\frac{g_s^2}{32\pi^2}\theta_s\vec{E}_s\cdot\vec{B}_s$$

EDM ~ e fm  $\theta_s$ 

Experimental bound:  $\theta_s$  < 10-10

#### Solution:

 $\theta_s$  ~a(x,t) is a dynamical field, an axion

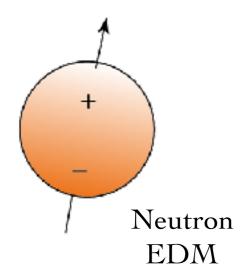
Axion mass from QCD:

$$\mu_a \sim 6 \times 10^{-11} \text{ eV} \frac{10^{17} \text{ GeV}}{f_a} \sim (3 \text{ km})^{-1} \frac{10^{17} \text{ GeV}}{f_a}$$

f<sub>a</sub>: axion decay constant

## Why is the Electric Dipole Moment of the Neutron Small?

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f<sub>a</sub>: axion decay constant

Mediates new forces and can be the dark matter

Extra dimensions

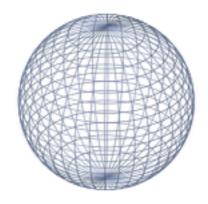
• Extra dimensions

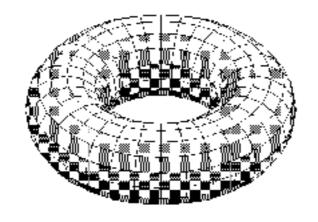
• Gauge fields

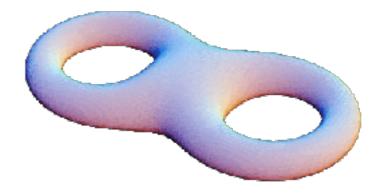
Extra dimensions

Gauge fields

● Topology



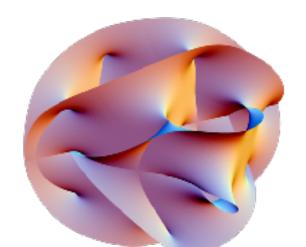




• Extra dimensions

Gauge fields

● Topology



• Extra dimensions

• Gauge fields

● Topology



Give rise to a plenitude of Universes

Extra dimensions

• Gauge fields

● Topology



Give rise to a plenitude of massless particles in our Universe

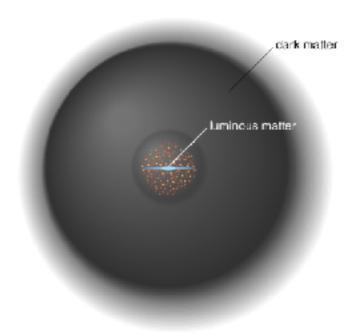
#### A Plenitude of Massless Particles

• Spin-0 non-trivial gauge field configurations: String Axiverse

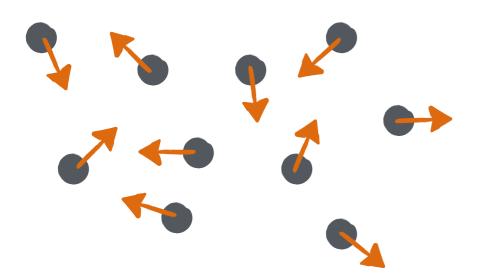
• Spin-1 non-trivial gauge field configurations: String Photiverse

• Fields that determine the shape and size of extra dimensions as well as values of fundamental constants: Dilatons, Moduli, Radion

Dark Matter Particles in the Galaxy

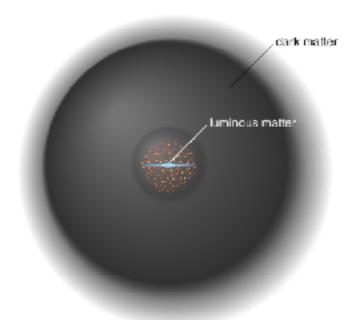


Usually we think of ...

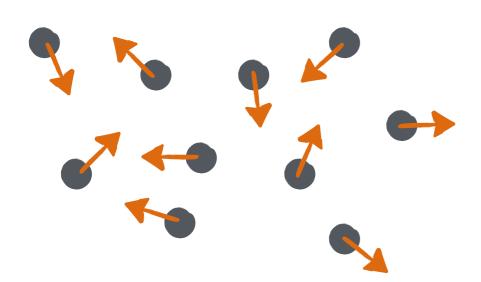


like a WIMP

Dark Matter Particles in the Galaxy

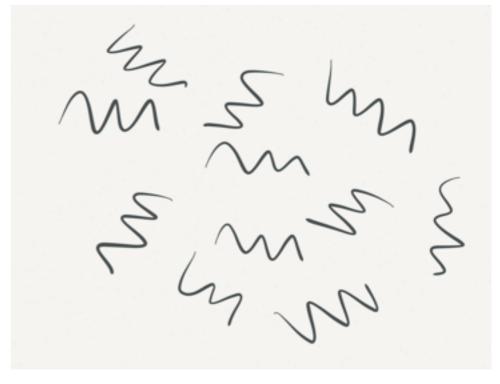


Usually we think of ...



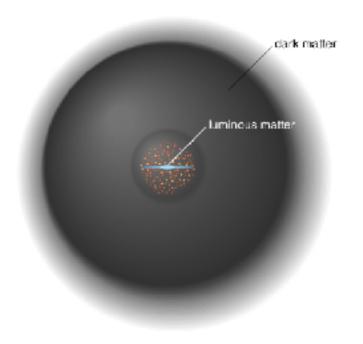
like a WIMP

instead of...



$$\lambda_{DM} = \frac{\hbar}{m_{DM}v}$$

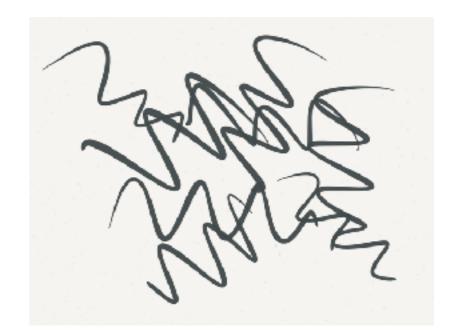
Dark Matter Particles in the Galaxy



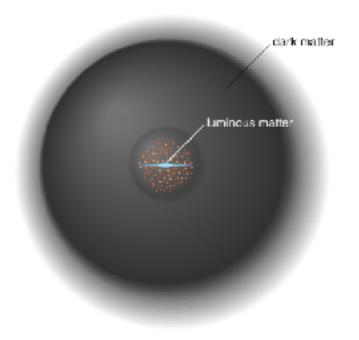
Decreasing DM Mass



$$\lambda_{DM} = \frac{\hbar}{m_{DM}v}$$



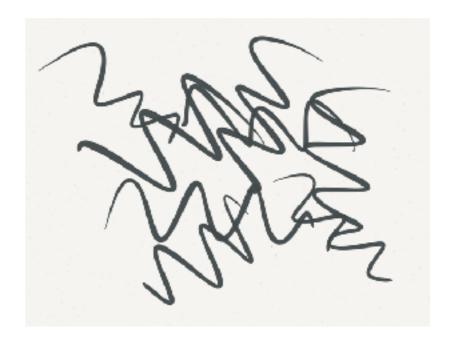
Dark Matter Particles in the Galaxy



Decreasing DM Mass



$$\lambda_{DM} = \frac{\hbar}{m_{DM}v}$$



Equivalent to a Scalar wave

#### Going from DM particles to a DM "wave"



When 
$$n_{DM} > \frac{1}{\lambda_{DM}^3}$$

In our galaxy this happens when  $m_{DM} < 1 \text{ eV/c}^2$ 

we can talk about DM  $\phi(x,t)$  and locally

$$\phi(t) \approx \phi_0 \cos \omega_{DM} t$$

with amplitude

$$\phi_0 \propto \frac{\sqrt{\rm DM \ density}}{\rm DM \ mass}$$

with frequency

$$\omega_{DM} pprox \frac{m_{DM}c^2}{\hbar}$$

and finite coherence

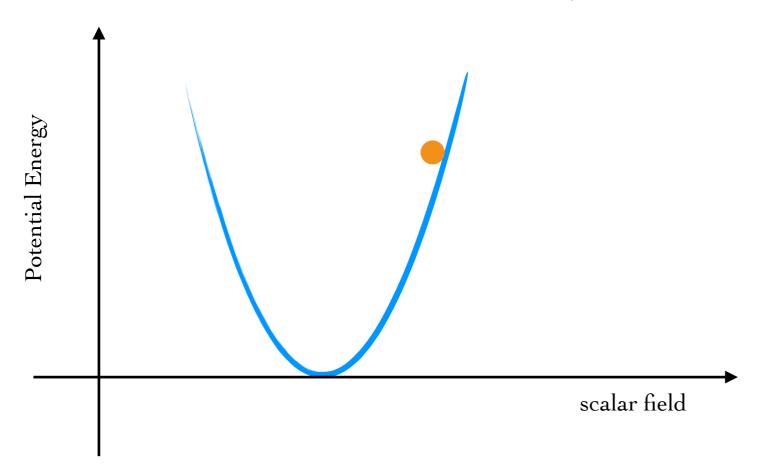
$$\delta\omega_{DM} \approx \frac{m_{DM}v^2}{\hbar} = 10^{-6}\omega_{DM}$$

## Light Scalar Dark Matter

• Just like a harmonic oscillator

$$\ddot{\phi} + 3 H \dot{\phi} + m_{\phi}^2 \phi = 0$$

$$\ddot{x} + \gamma \ \dot{x} + \omega^2 x = 0$$



Frozen when: Hubble >  $m_{\phi}$ 

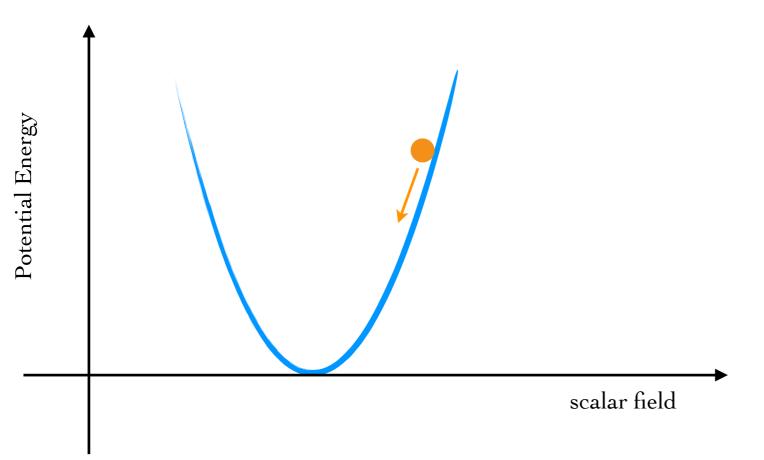
Initial conditions set by inflation

#### Light Scalar Dark Matter

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$$\ddot{x} + \gamma \ \dot{x} + \omega^2 x = 0$$



Frozen when: Hubble >  $m_{\phi}$ 

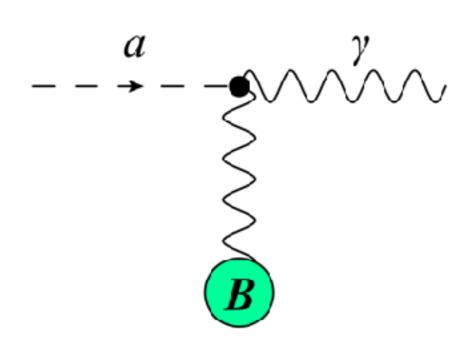
Oscillates when: Hubble  $< m_{\phi}$ 

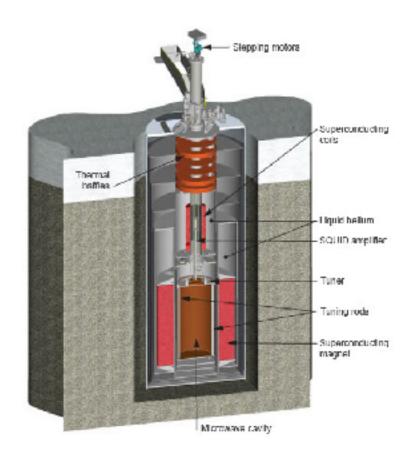
 $\rho_{\phi}$  scales as a-3 just like Dark Matter

#### Axion Dark Matter

Some examples

Axion-to-photon conversion (ex. ADMX)





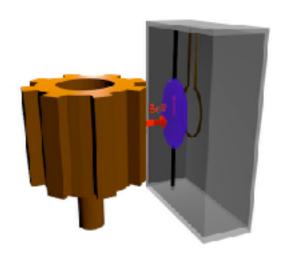
Cavity size = Axion size

#### Axion Dark Matter

Some examples



• Axion Force experiments and DM experiments



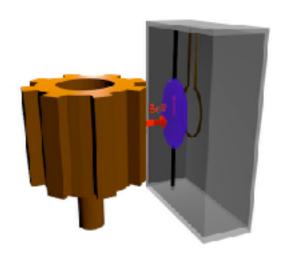
ARIADNE experiment

#### Axion Dark Matter

Some examples



• Axion Force experiments and DM experiments



ARIADNE experiment

#### Dark Photon Dark Matter

Some examples

• Detected if kinetically mixed with the photon

$$\mathcal{L} \supset \epsilon F_{EM} F_{DM}$$

• Detected like a photon (ADMX)

DM electric field  $\sim \sqrt{\rho_{DM}} \sim 50 \text{ V/cm}$ 

#### Moduli Dark Matter

• Couple non-derivatively to the Standard Model (as well as axions with CP violation)

Examples of couplings

$$\mathcal{L} = \mathcal{L}_{SM} + \sqrt{\hbar c} \frac{\phi}{\Lambda} \mathcal{O}_{SM}$$

$$\mathcal{O}_{SM} \equiv m_e e \bar{e}, \ m_q q \bar{q}, \ G_s^2, \ F_{EM}^2, \dots$$

#### Fundamental constants are not really constants

## Keeping the DM time with Atomic Clocks

with Junwu Huang and Ken Van Tilburg (2014)

## Oscillating Atomic and Nuclear Energy Splittings due to Dark Matter

Optical Splittings

$$\Delta E_{
m optical} \propto lpha_{EM}^2 m_e \sim {
m eV}$$

Hyperfine Splittings

$$\Delta E_{
m hyperfine} \propto \Delta E_{
m optical} lpha_{EM}^2 \left(rac{m_e}{m_p}
ight) \sim 10^{-6} \, {
m eV}$$

Nuclear Splittings

$$\Delta E$$
 (m<sub>p</sub>,  $\alpha_s$ ,  $\alpha_{EM}$ )~ 1 MeV

#### Atomic Clocks

Kept tuned to an atomic energy level splitting

#### Current definition of a second:

the duration of 9192631770 periods of the radiation corresponding to the transition between the two hyperfine levels of the ground state of the caesium 133 atom

• Have shown stability of 1 part in 10<sup>18</sup>

Compared to 1 part in 10<sup>13</sup> expected by DM

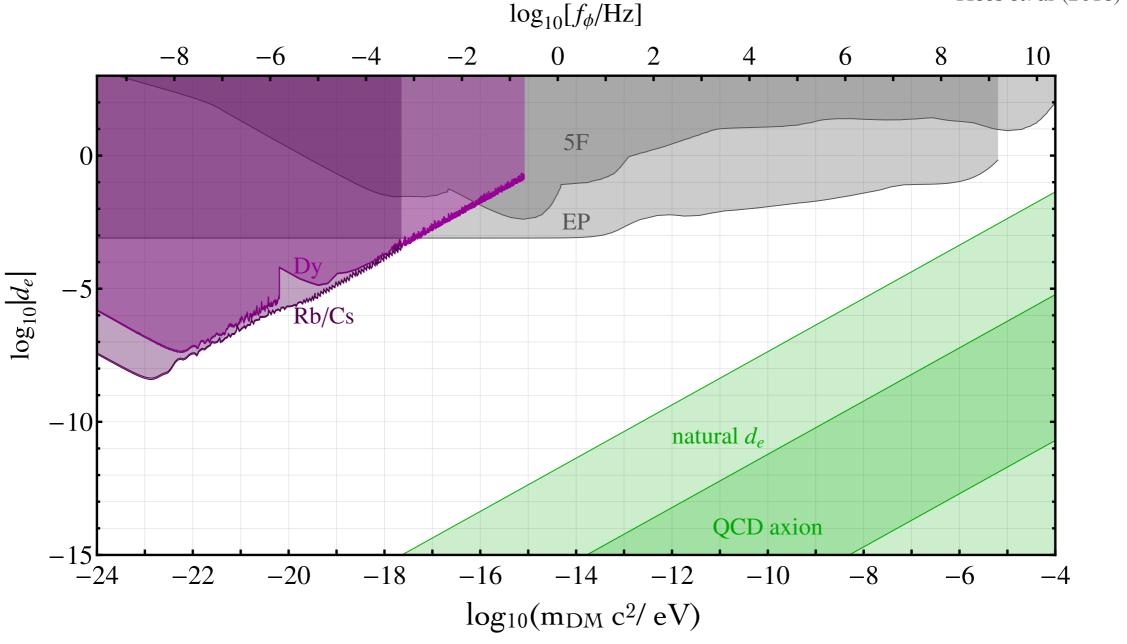
• Have won several Nobel prizes in the past 20 years

#### The Dy isotope and Rb/Cs Clock Comparison

sensitivity to  $\alpha_{EM}$  variations

Ken Van Tilburg and the Budker group (2015)

Hees et. al (2016)



## The Sound of Dark Matter

with Ken Van Tilburg and Savas Dimopoulos (2015)

#### Oscillating interatomic distances

• The Bohr radius changes with DM

• 
$$r_B \sim (\alpha m_e)^{-1}$$

$$\frac{\delta r_B}{r_B} = -\left(\frac{\delta \alpha_{EM}}{\alpha_{EM}} + \frac{\delta m_e}{m_e}\right)$$

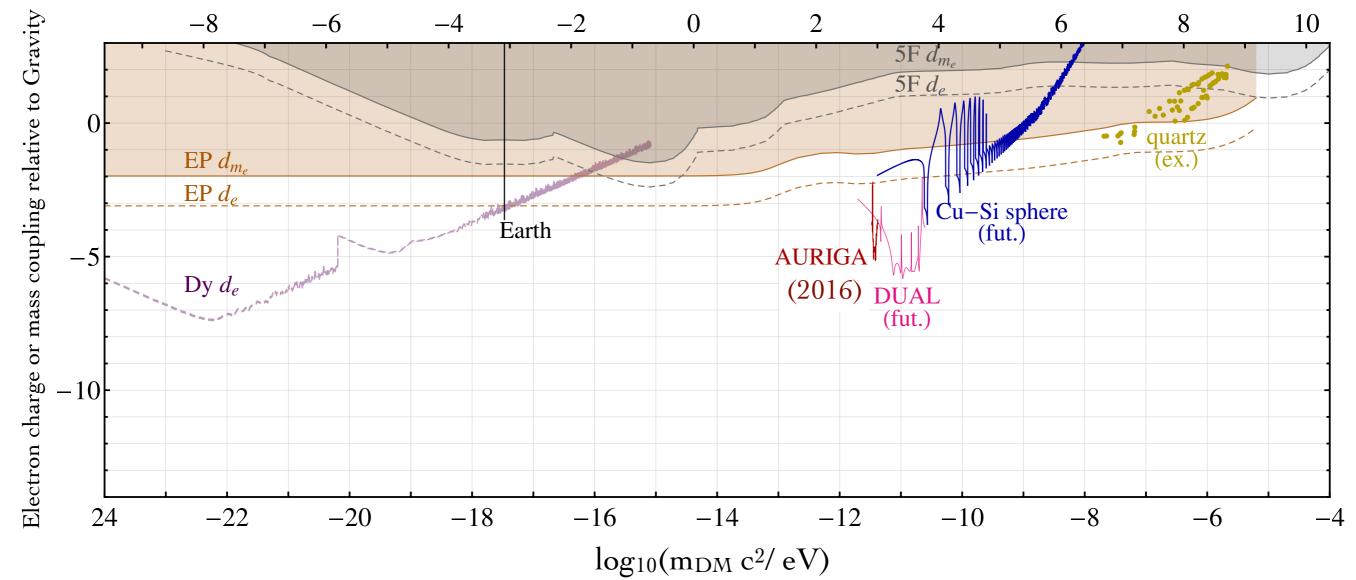
The size of solids changes with DM

• 
$$L \sim N \ (\alpha \ m_e)^{-1}$$
 
$$\frac{\delta L}{L} = -\left(\frac{\delta \alpha_{EM}}{\alpha_{EM}} + \frac{\delta m_e}{m_e}\right)$$

For a single atom  $\delta r_{B} \sim 10^{-30}$  m Need macroscopic objects to get a detectable signal

#### What can be done in the future?

 $\log_{10}[f_{\phi}/\mathrm{Hz}]$ 



## Resonant Absorption of Dark Matter

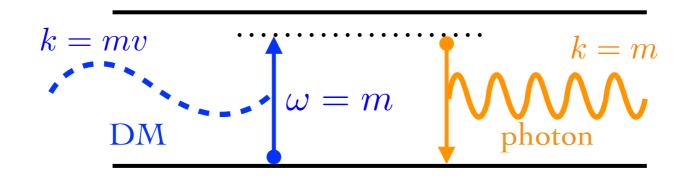
with Savas Dimopoulos and Ken Van Tilburg (2017)

# Resonant Absorption of Dark Matter in Molecules





## Resonant Absorption of Dark Matter in Molecules



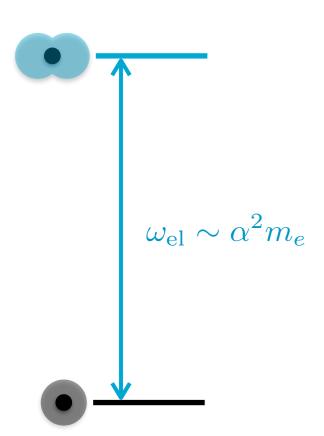
$$\Gamma_{abs} = N \Omega^2 \tau$$

N: number of molecules

 $\Omega$ : Interaction energy between DM and molecule (ie Rabi frequency)

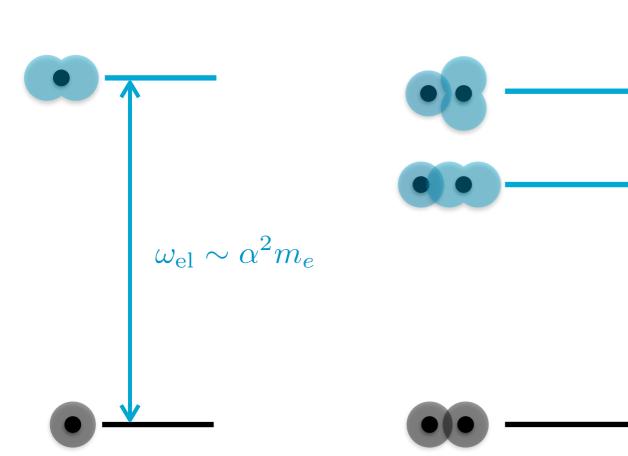
τ: Smallest of excitation lifetime, inter-molecule collisions, DM coherence time...

atom



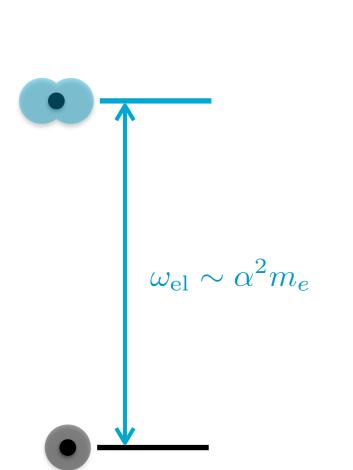
atom

diatomic molecule



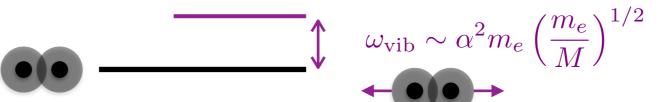


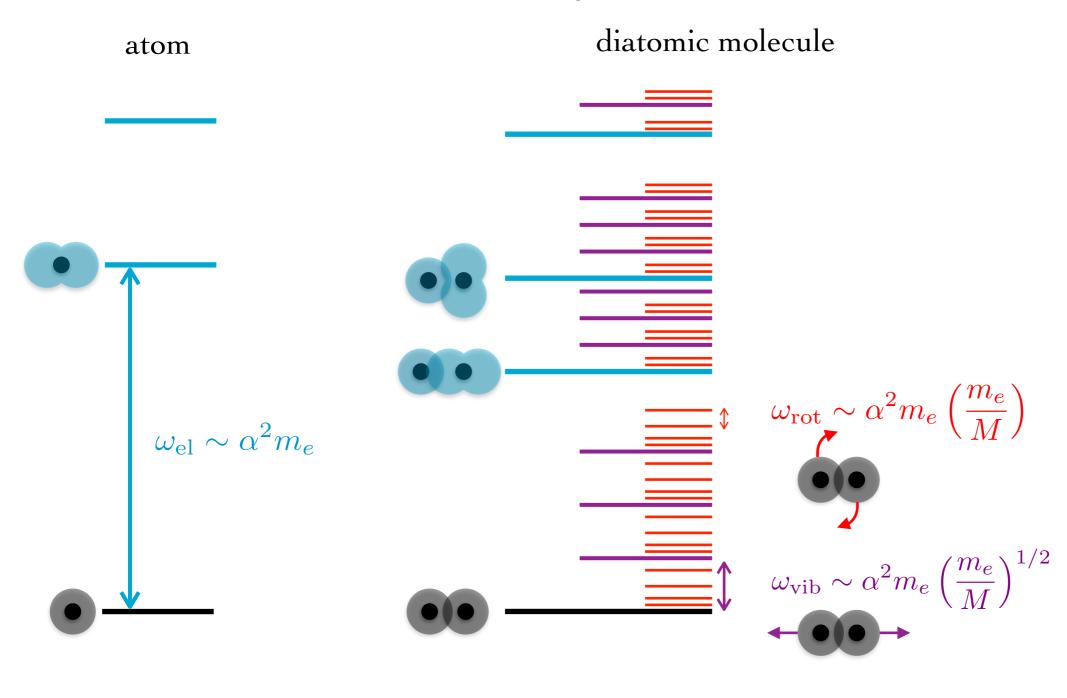
diatomic molecule

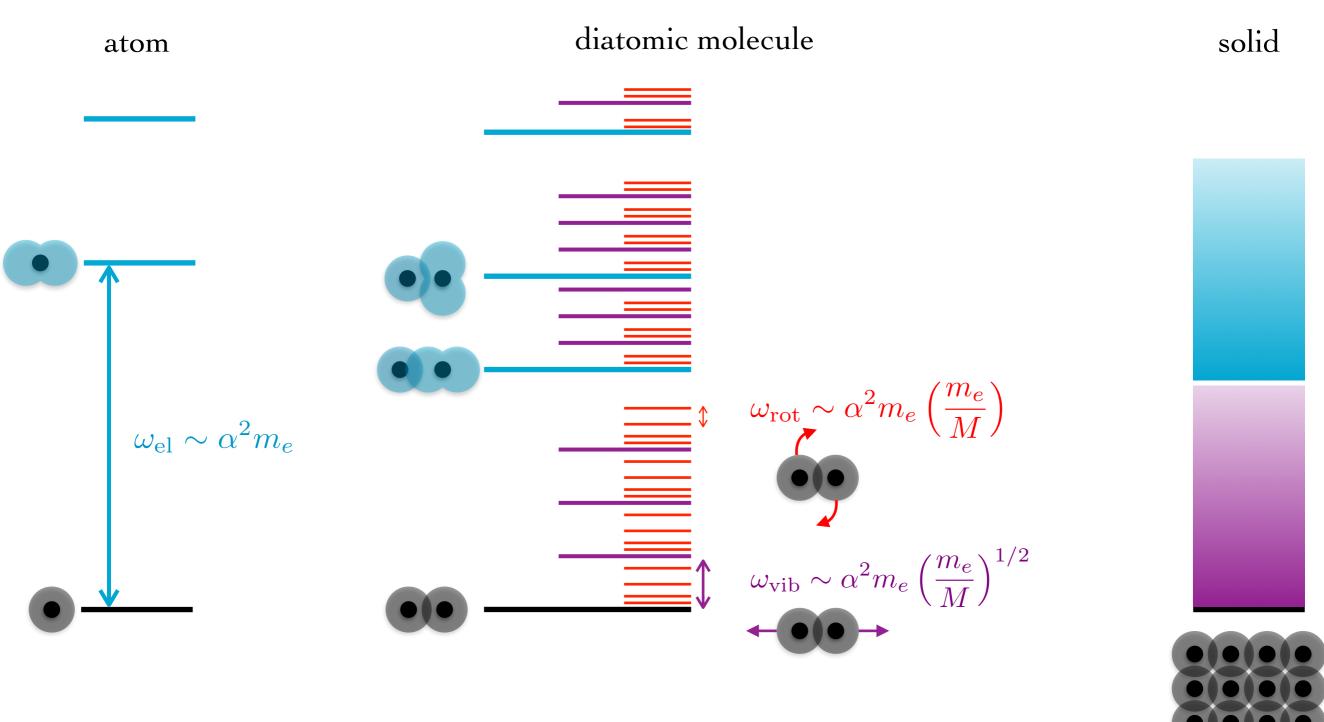






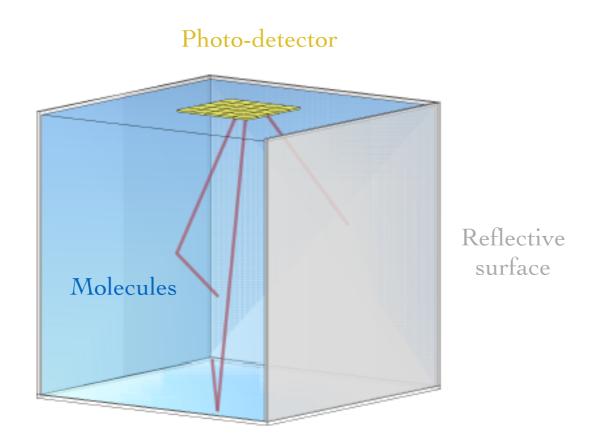






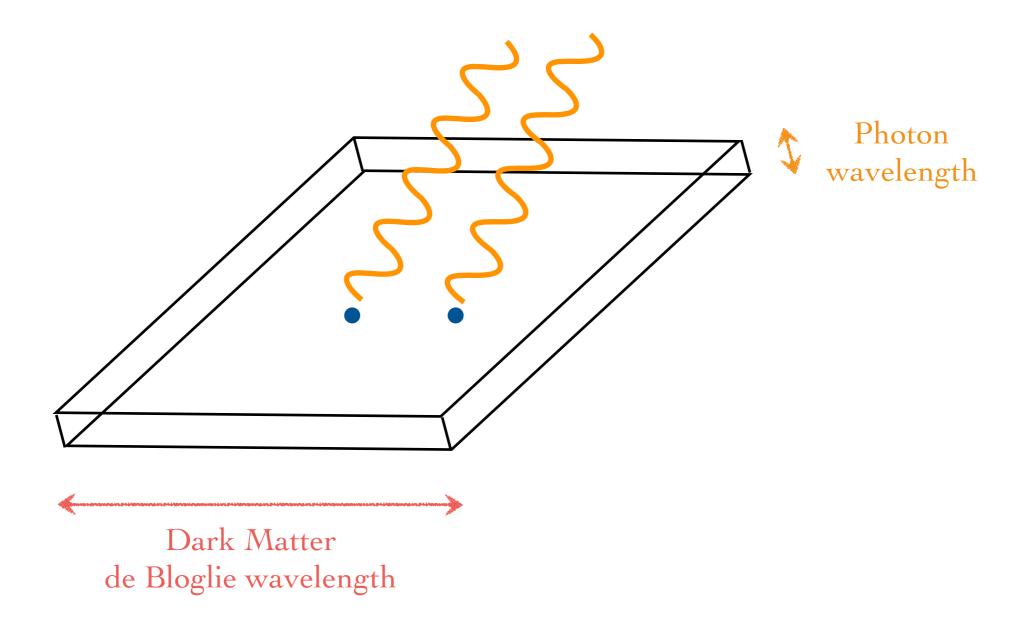
Molecules are a multi-mode resonator with well-understood spectrum

## Proposed Set-ups



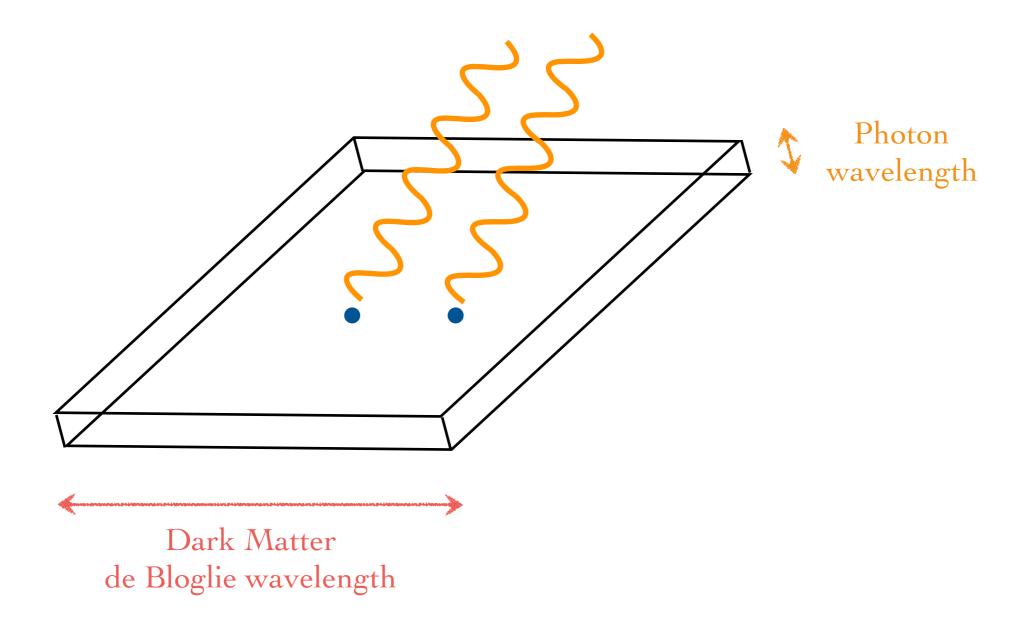
Bulk detector

## Cooperative emission effects



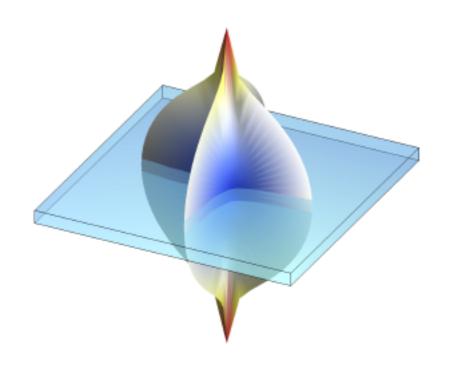
Fluorecent photon interference produces interesting emission patterns

## Cooperative emission effects

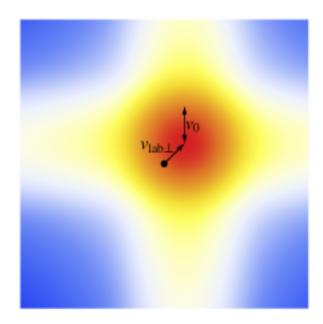


Fluorecent photon interference produces interesting emission patterns

## Cooperative emission effects

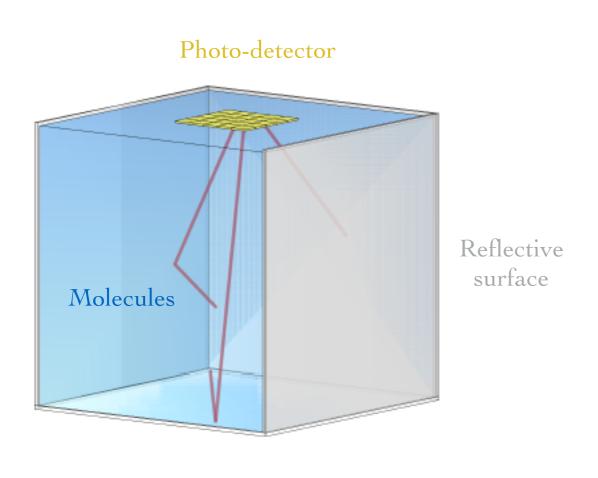


Fluorescent power emitted from a slab

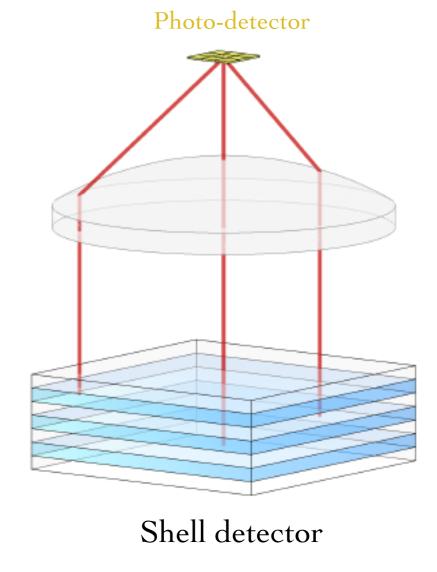


Angular distribution of power close to the slab center

## Proposed Set-ups

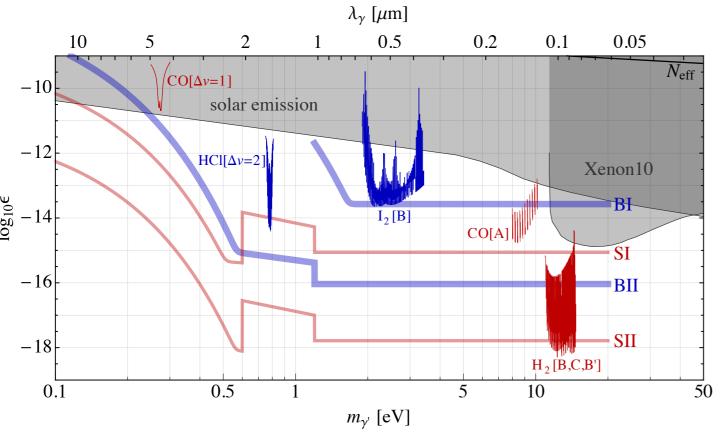


Bulk detector



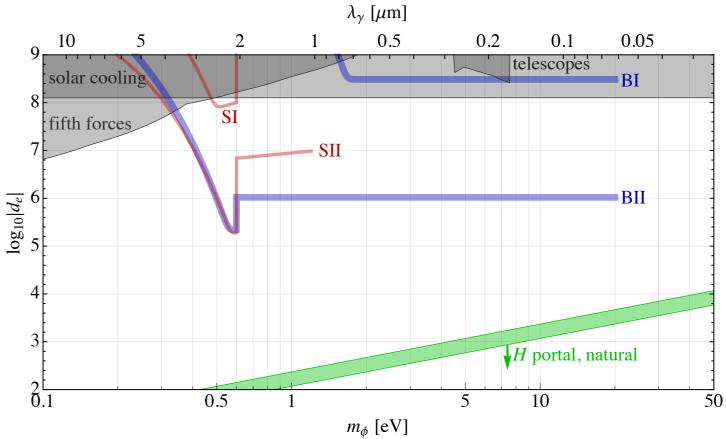
Alternating shells of gas filled and empty containers of photon wavelength width

## Sensitivity Plots Some examples



Dark Photon Dark Matter

Electron charge modulus Dark Matter



## Advantages of Molecule based detector

• Sensitivity to many Dark Matter candidates

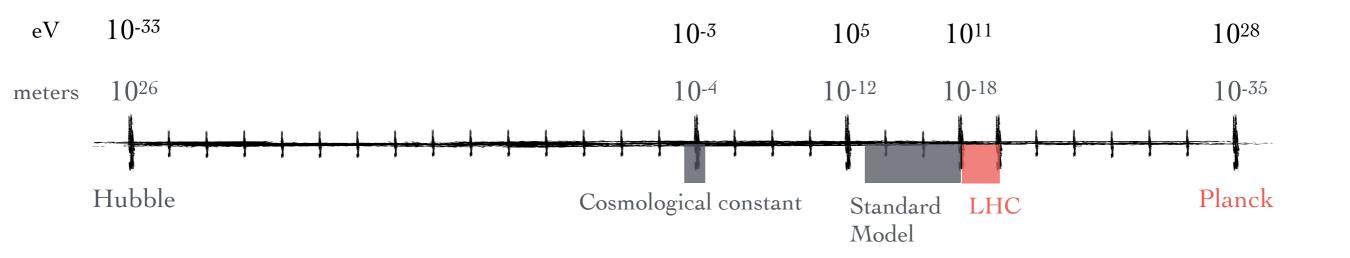
• Excellent Energy Resolution

Tunable

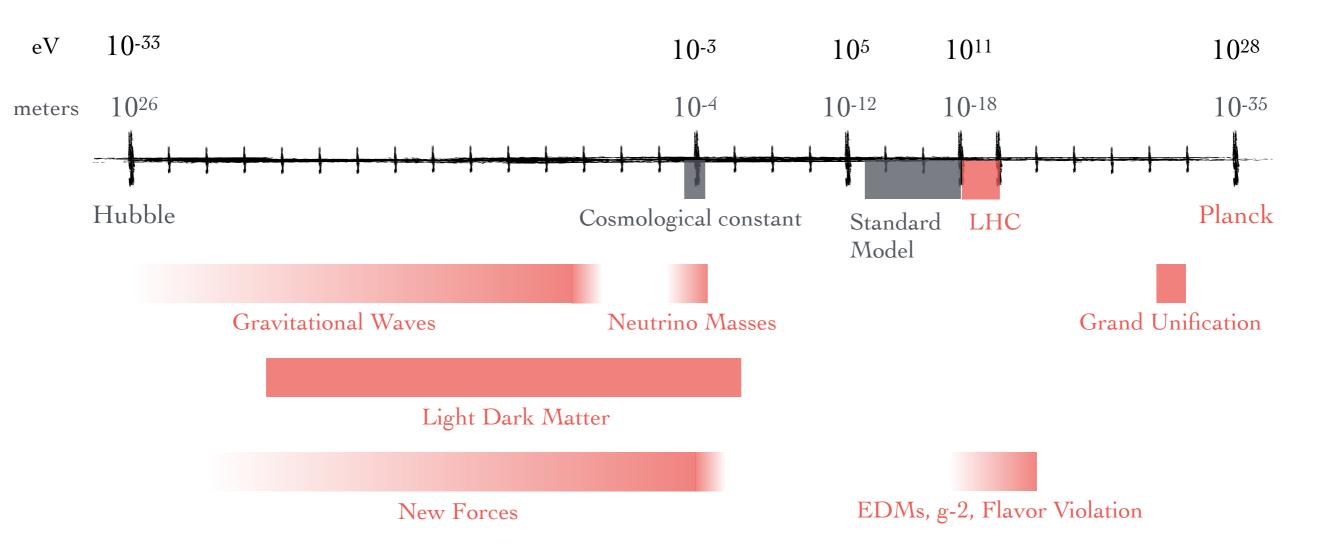
• Background rejection

• Directional information for Dark Matter

#### The Scales in Our Universe



#### The Scales in Our Universe

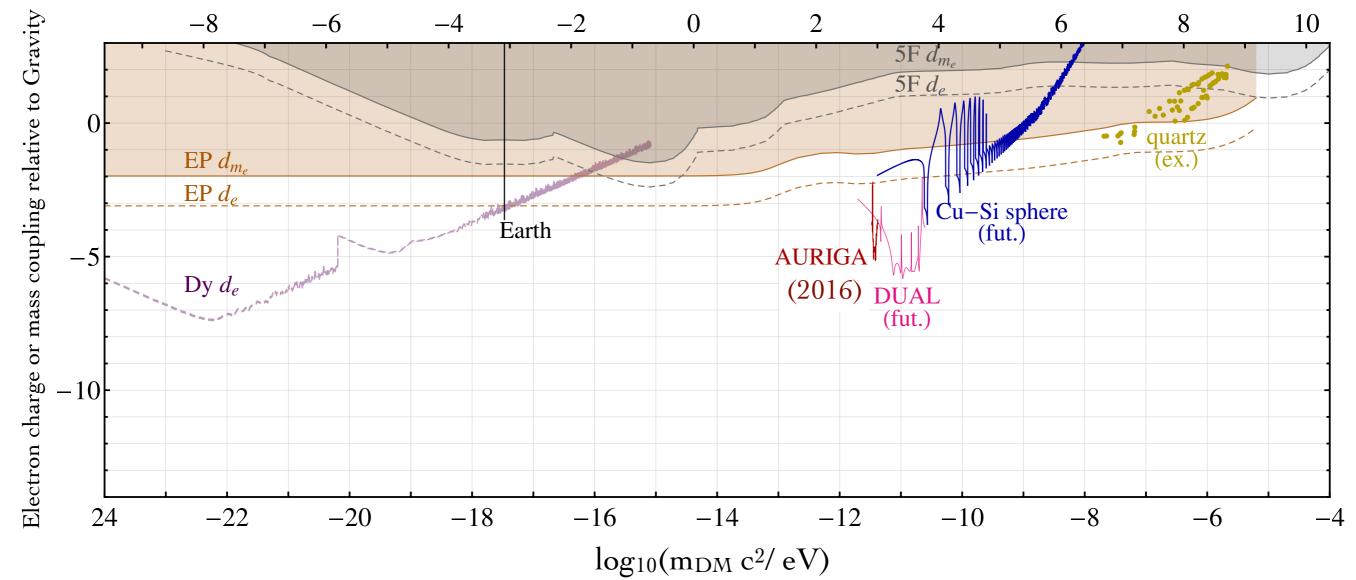


There are more things in heaven and earth, Horatio, Than are dreamt of in your philosophy.

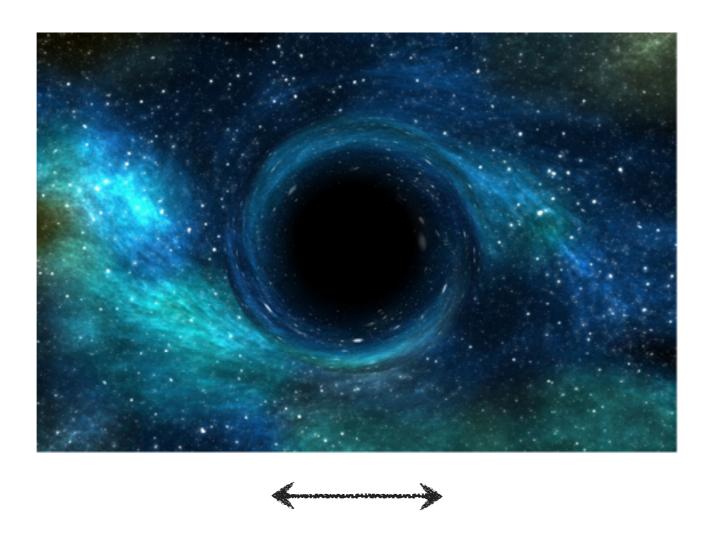
- Hamlet

#### What can be done in the future?

 $\log_{10}[f_{\phi}/\mathrm{Hz}]$ 

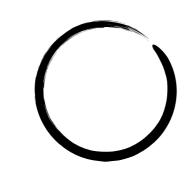


#### Black Holes as Nature's Detectors



1 km -10 billion km

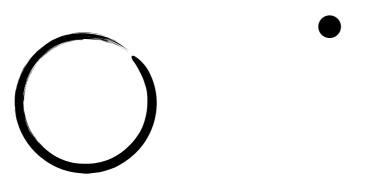
They can detect bosons of similar in size



Super-radiant scattering of a massive object

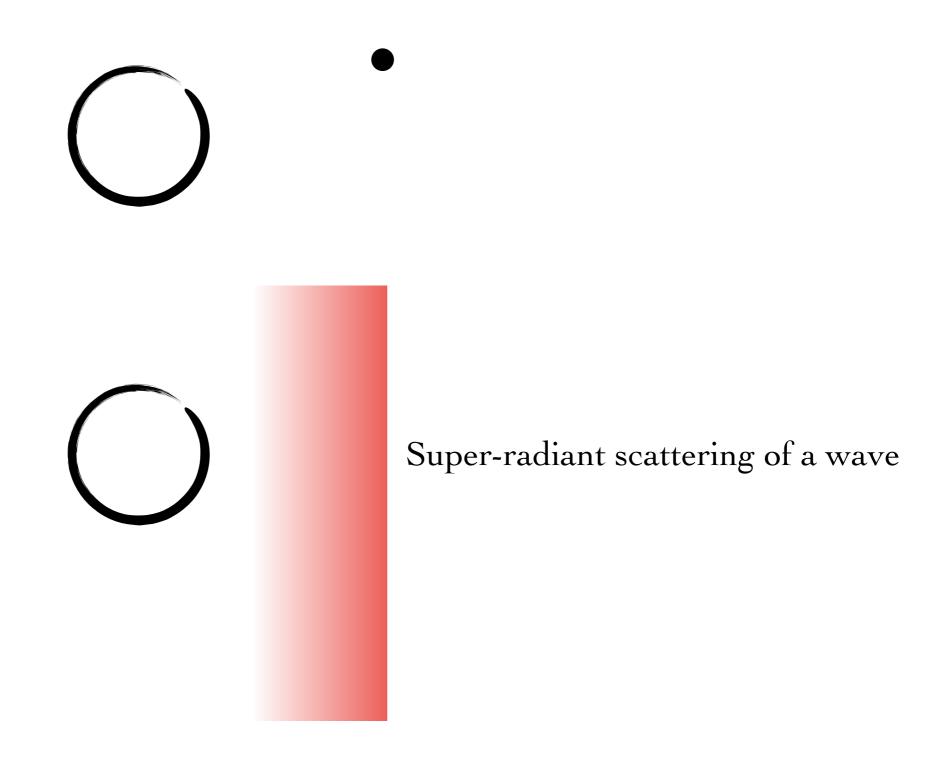


Super-radiant scattering of a massive object



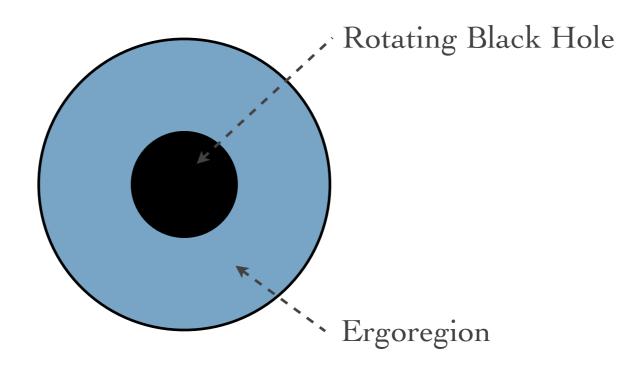


Super-radiant scattering of a wave



## Black Hole Superradiance

Penrose Process

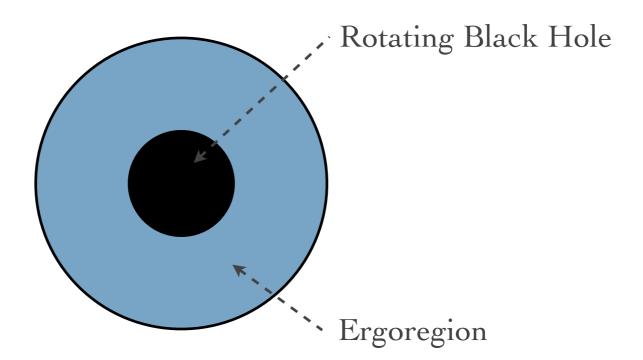


Ergoregion: Region where even light has to be rotating

## Black Hole Superradiance

Penrose Process

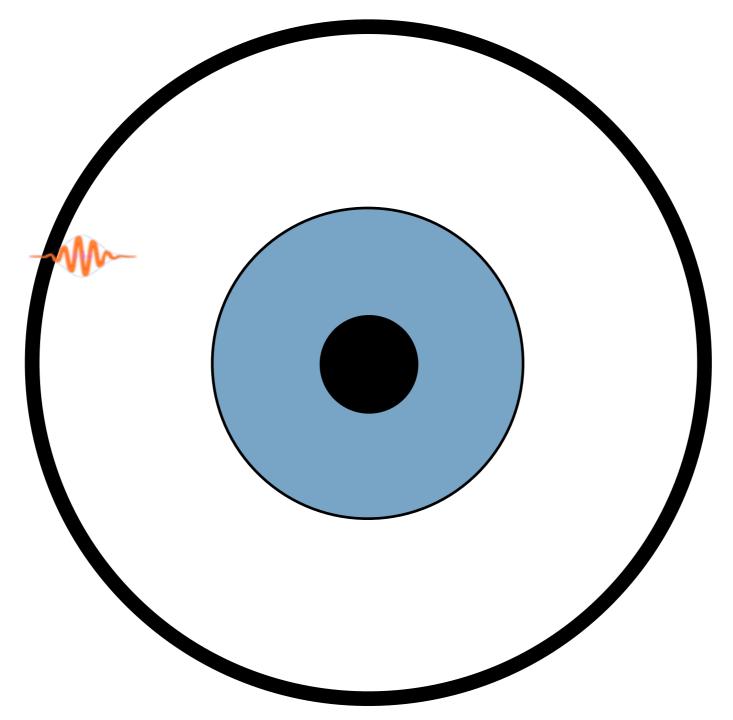




Extracts angular momentum and mass from a spinning black hole

#### Black Hole Bomb

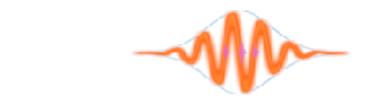
Press & Teukolsky 1972

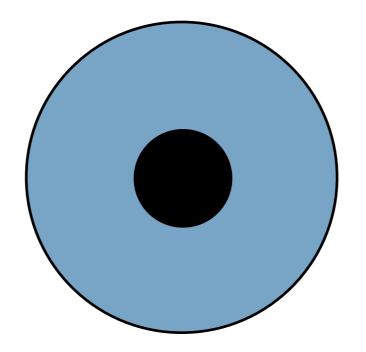


Photons reflected back and forth from the black hole and through the ergoregion

#### Black Hole Bomb

Press & Teukolsky 1972

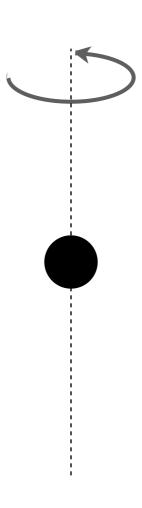




Photons reflected back and forth from the black hole and through the ergoregion

## Superradiance for a massive boson

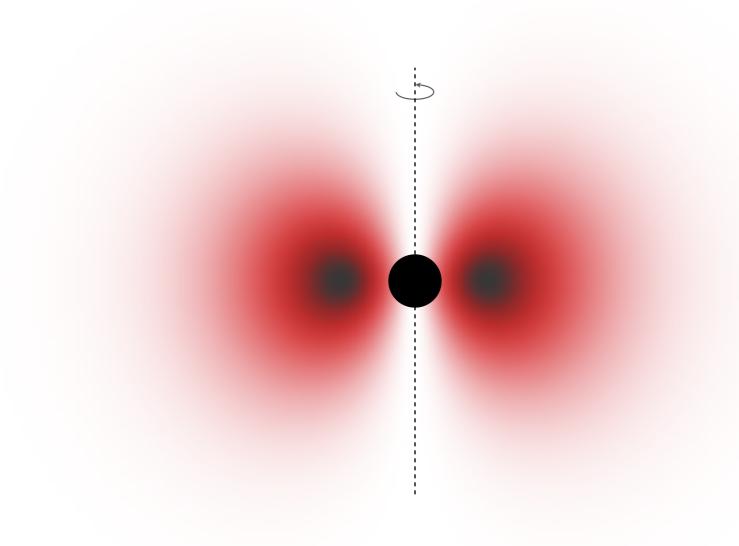
Damour et al; Zouros & Eardley; Detweiler; Gaina (1970s)



Particle Compton Wavelength comparable to the size of the Black Hole

## Superradiance for a massive boson

Damour et al; Zouros & Eardley; Detweiler; Gaina (1970s)



Particle Compton Wavelength comparable to the size of the Black Hole

## Gravitational Atom in the Sky



Fine-structure constant:

$$\alpha = G_{\rm N} M_{\rm BH} \mu_a = R_g \mu_a$$

Principal (n), orbital (l), and magnetic (m) quantum number for each level

$$E_{\text{binding}} = -\frac{\alpha^2 \mu_a}{2n^2}$$

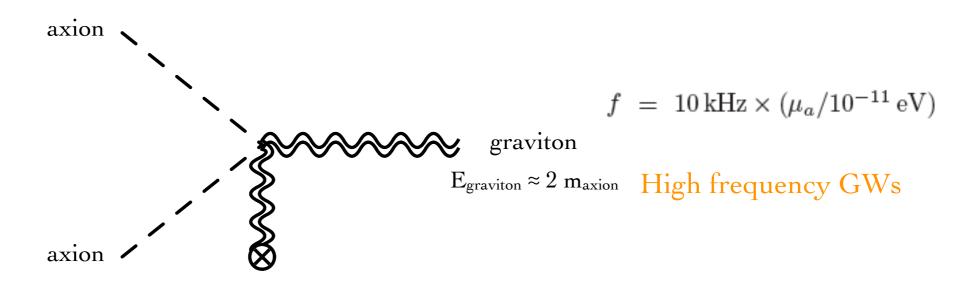
Main differences from hydrogen atom:

Levels occupied by bosons - occupation number >10<sup>77</sup>

In-going Boundary Condition at Horizon

## Super-Radiance Signatures

GW annihilations



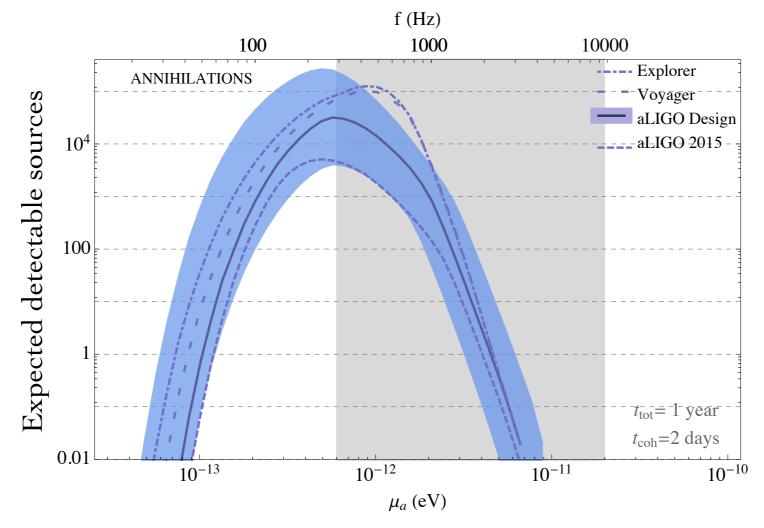
• Signal enhanced by the square of the occupation number of the state

$$h_{\rm peak} \simeq 10^{-22} \left(\frac{1 \,\mathrm{kpc}}{r}\right) \left(\frac{\alpha/\ell}{0.5}\right)^{\frac{p}{2}} \frac{\alpha^{-\frac{1}{2}}}{\ell} \left(\frac{M}{10 M_{\odot}}\right)$$

• Signal duration determined by the annihilation rate (can last thousands of years)

#### Expected Events from Annihilations

• Large uncertainties coming from tails of BH mass distribution



Pessimistic: flat spin distribution and 0.1 BH/century

Realistic: 30% above spin of 0.8 and 0.4 BH/century

Optimistic: 90% above spin of 0.9 and 0.9 BH/century